



Long-lived charged particles at FCC-hh and FCC-he

[Sho IWAMOTO](#) (岩本 祥)

2 Nov. 2017

Seminar @ Tohoku University

Based on

hh: Jonathan L. Feng, SI, Yael Shadmi, Shlomit Tarem [[1505.02996](#)]

(collected in FCC-hh report [[1606.00947](#)])

he: Kechen Wang, SI, Monica D'Onofrio, Georges Azuelos [17???.?????]

(subgroup in BSM@ep collaboration)

0. Collider Physics basic

1. LLCPs

2. Future colliders (FCC-hh and FCC-he)

3. LLCP searches at FCC-hh

4. LLCP searches at FCC-he

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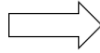
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c-tagging for new physics at the LHC

[Sho IWAMOTO](#) (岩本 祥)

25 Oct. 2017

Workshop on beyond standard model and the early universe
@ Tohoku University

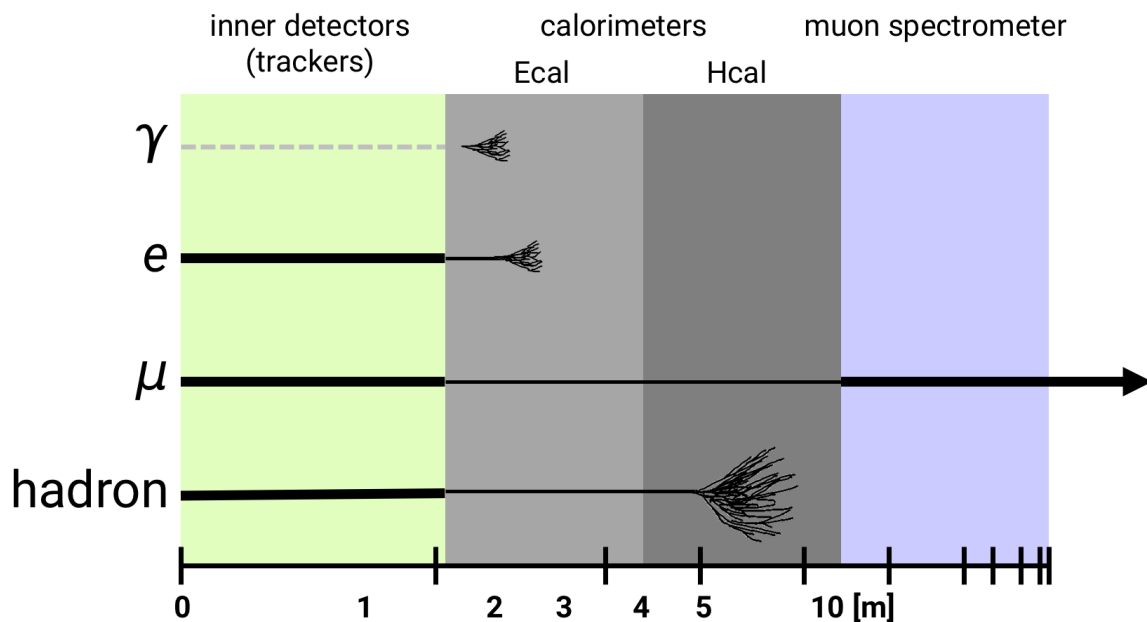
Sho Iwamoto, Gabriel Lee, Yael Shadmi, Yaniv Weiss [[1703.05748](#)]

(and discussion with Jonathan Shlomi @ Weizmann/ATLAS)

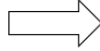


c-tagging for new physics at the LHC

Standard technique of Collider detectors

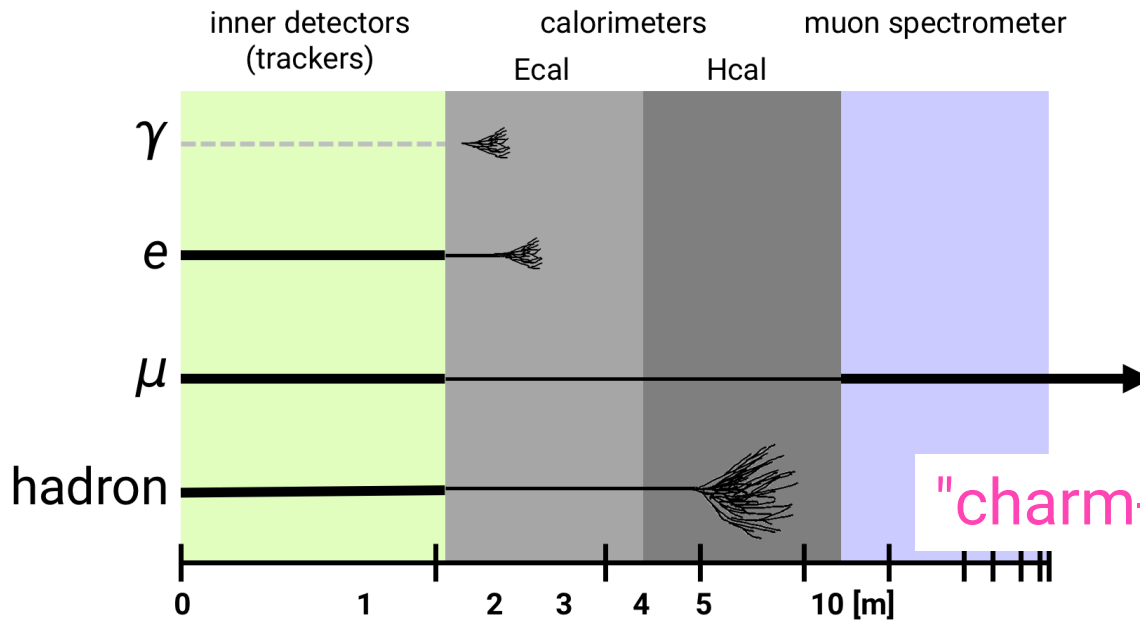


- τ -lepton? \longrightarrow τ -tagging
- which quark? or gluon? \longrightarrow quark-flavor tagging



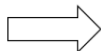
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Standard technique of Collider detectors



"charm-tagging"

- τ -lepton? \longrightarrow τ -tagging
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c-tagging for new physics at the LHC

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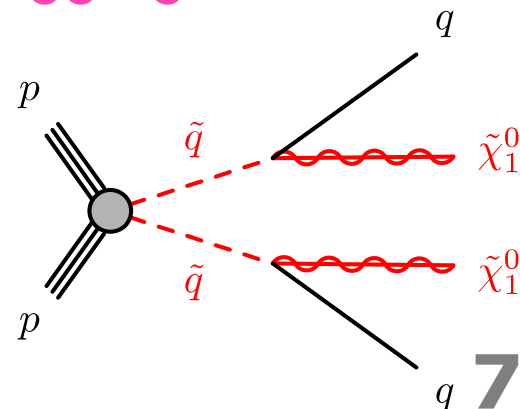
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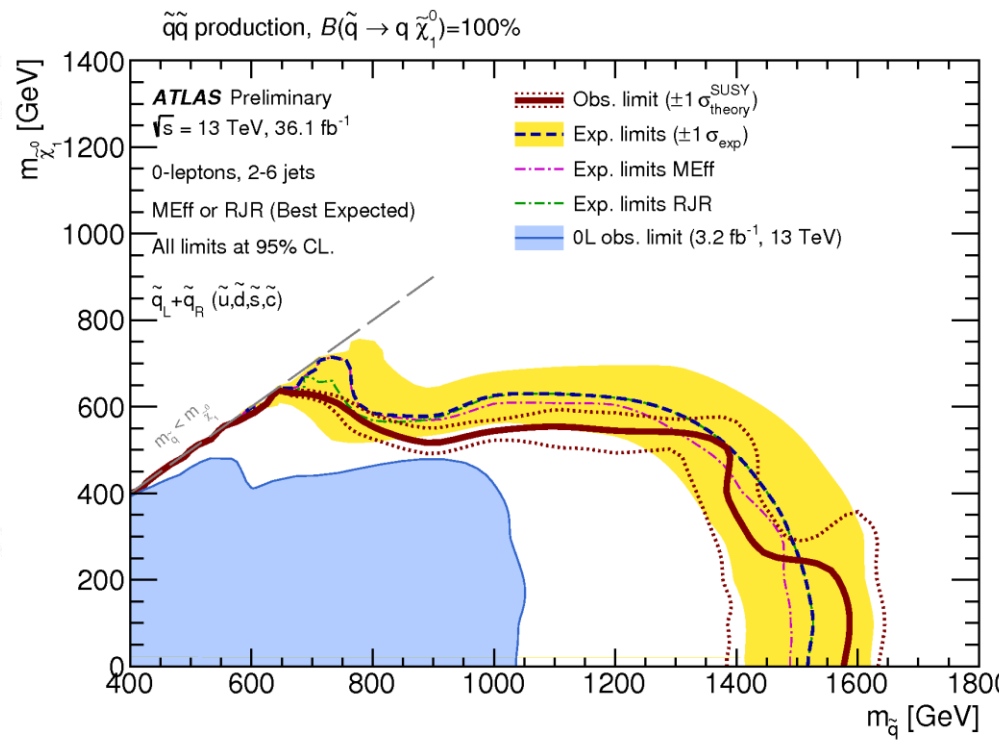
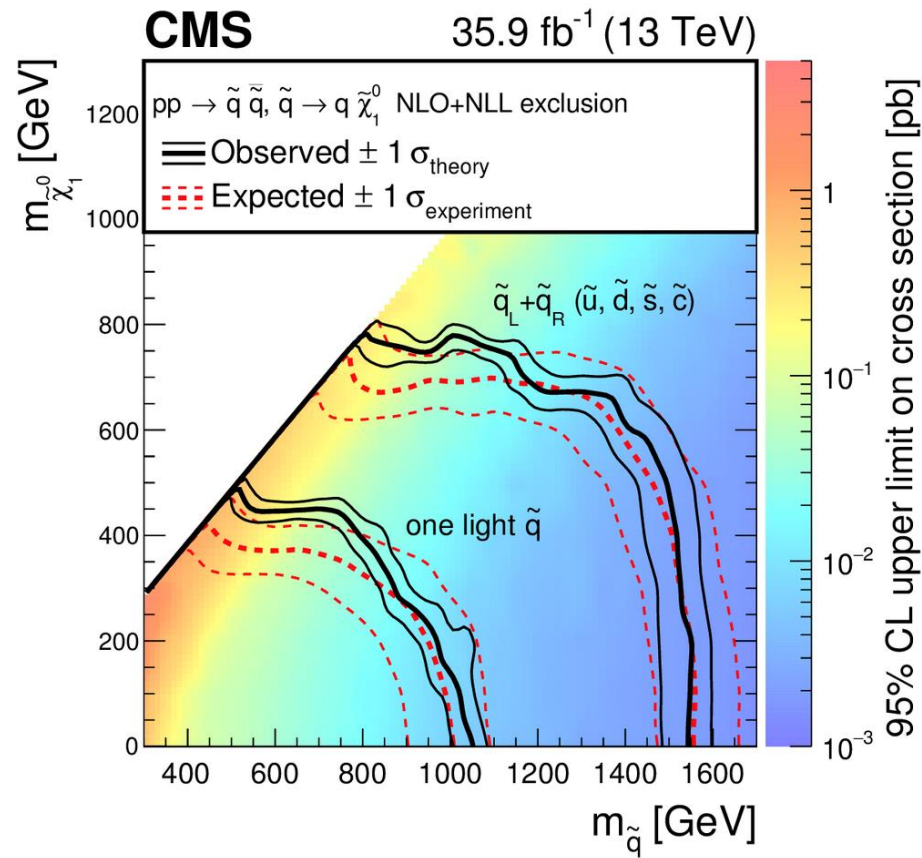
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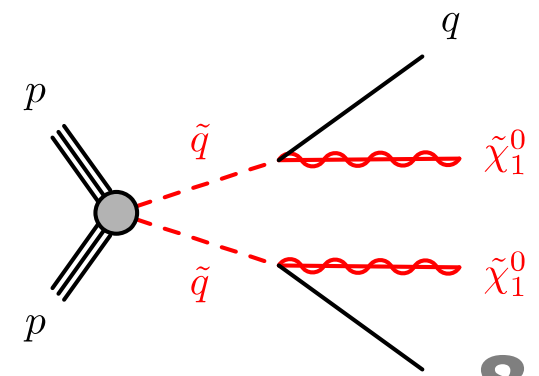
"charm-tagging" in



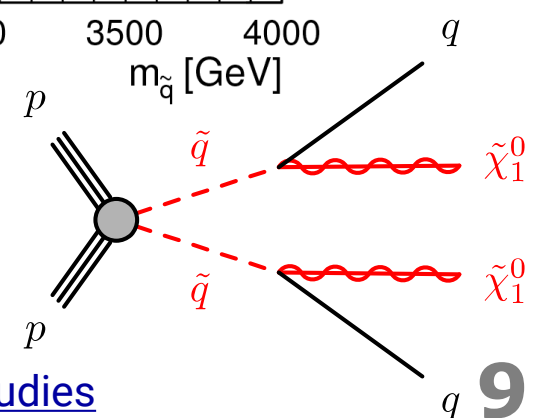
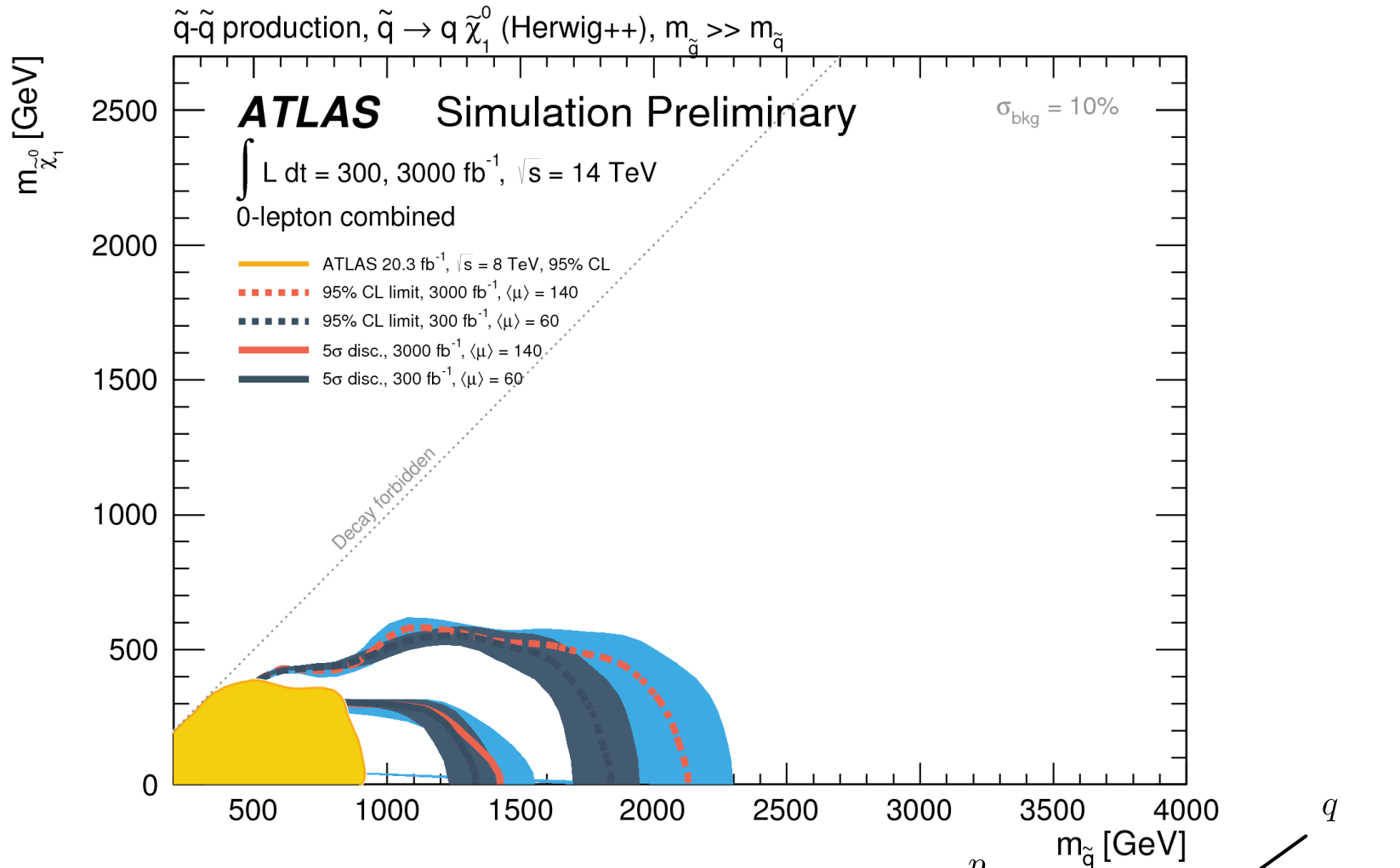
SUSY searches at the LHC : Nothing found



[1705.04650](https://arxiv.org/abs/1705.04650) / [ATLAS-CONF-2017-022](https://arxiv.org/abs/ATLAS-CONF-2017-022)



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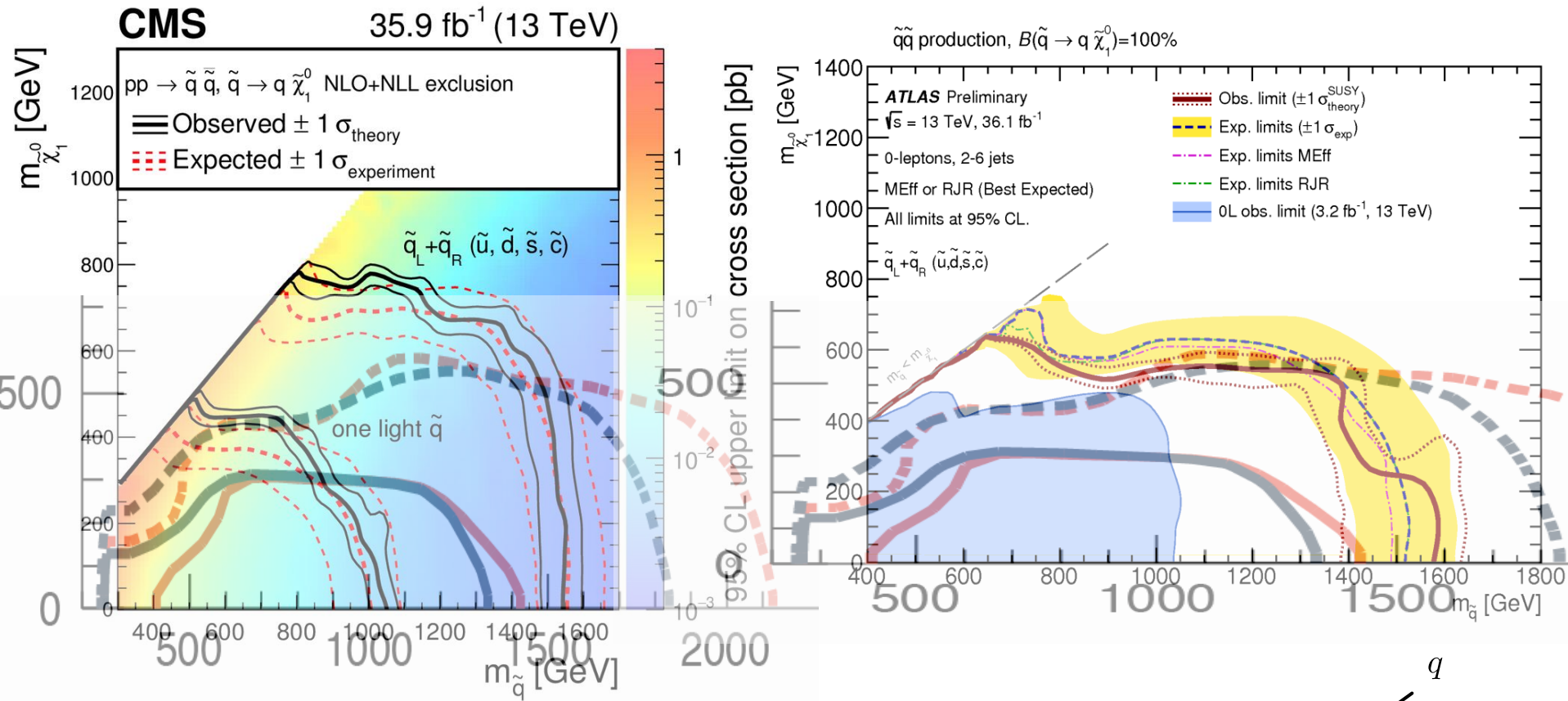


[ATL-PHYS-PUB-2014-010](https://twiki.cern.ch/twiki/bin/view/ATLASPublic/ATL-PHYS-PUB-2014-010)

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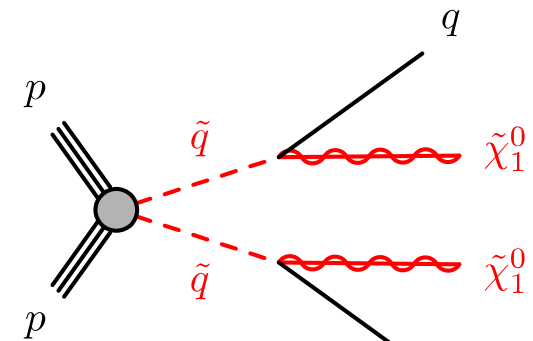


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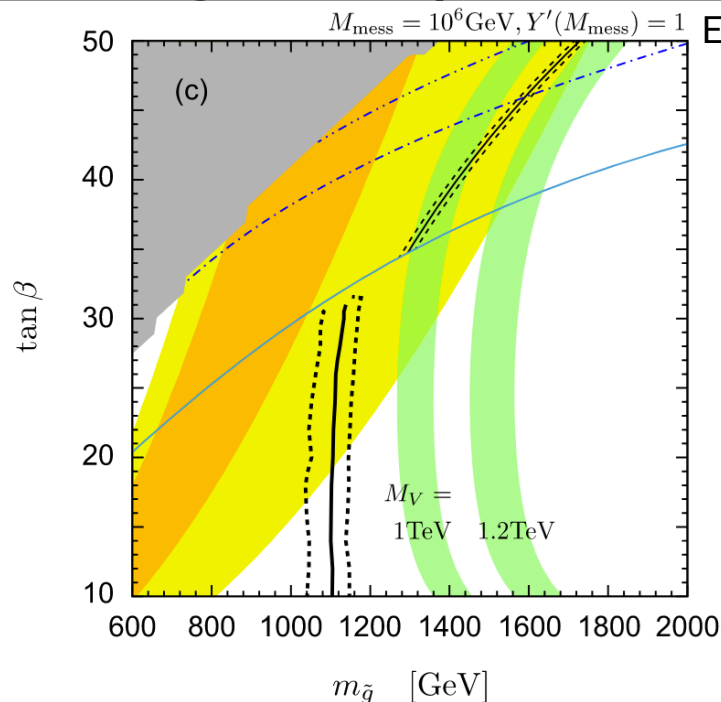
[ATL-PHYS-PUB-2014-010](https://arxiv.org/abs/1408.0244)

<https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResultsFP>

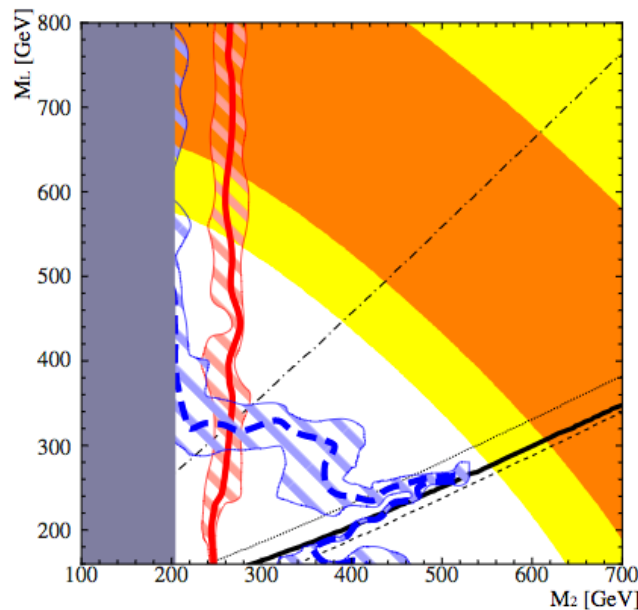
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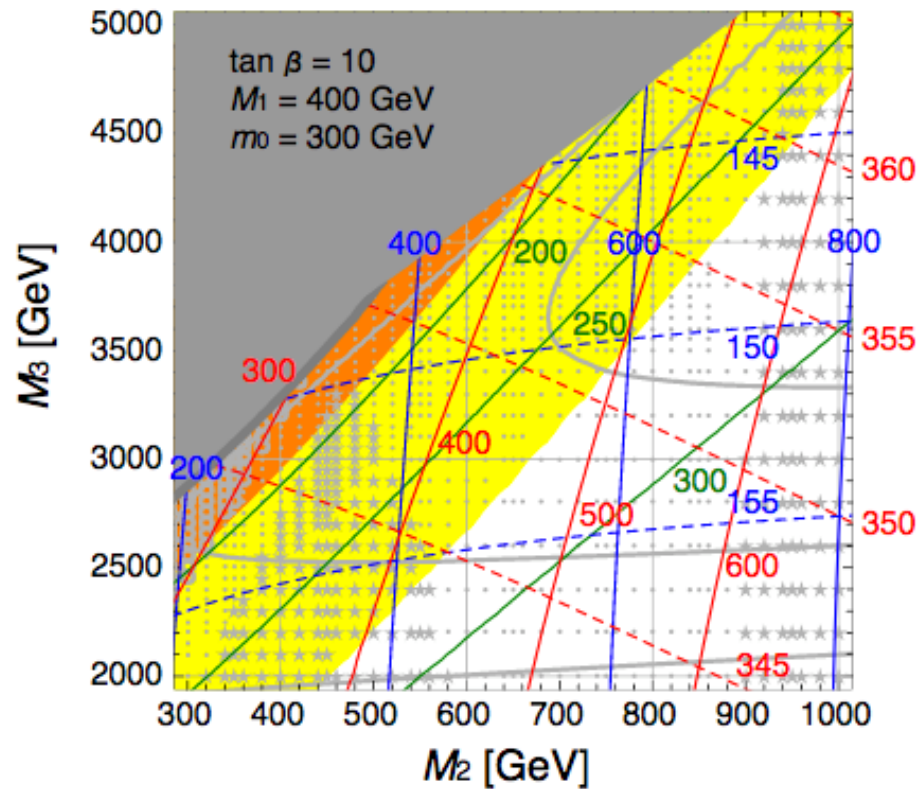
Muon $g-2$ anomaly?



Endo, Hamaguchi, Ishikawa, Si, Yokozaki [1212.3935]



(c) $\mu = M_2/2, m_R = 3 \text{ TeV}$



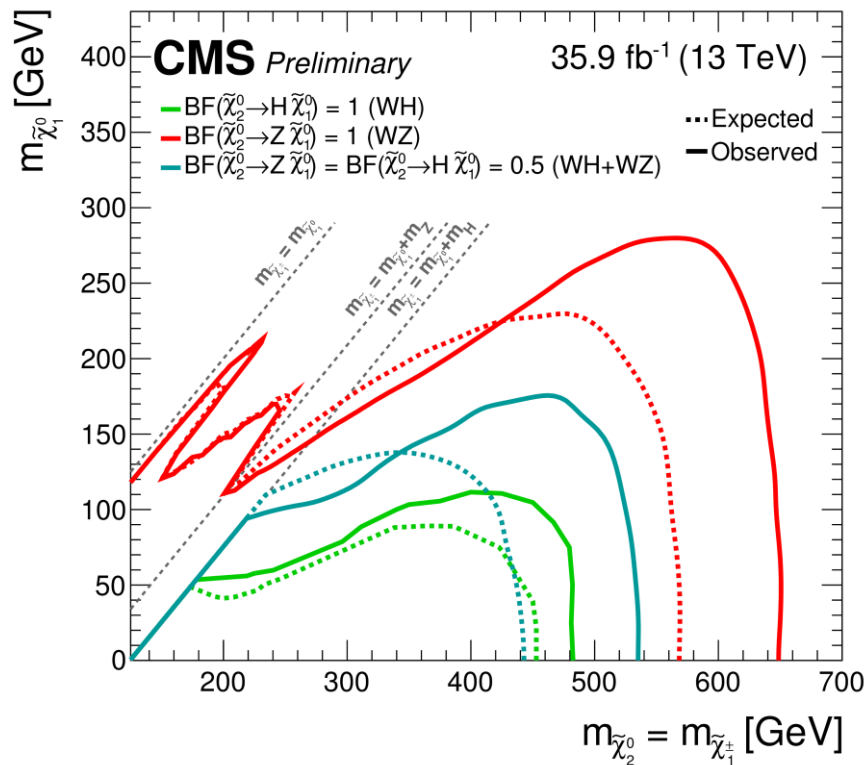
Si, Yanagida, Yokozaki [1407.4226]

Bino(LSP), **Wino**, **L/R-slepton**, and **stau** masses are shown.

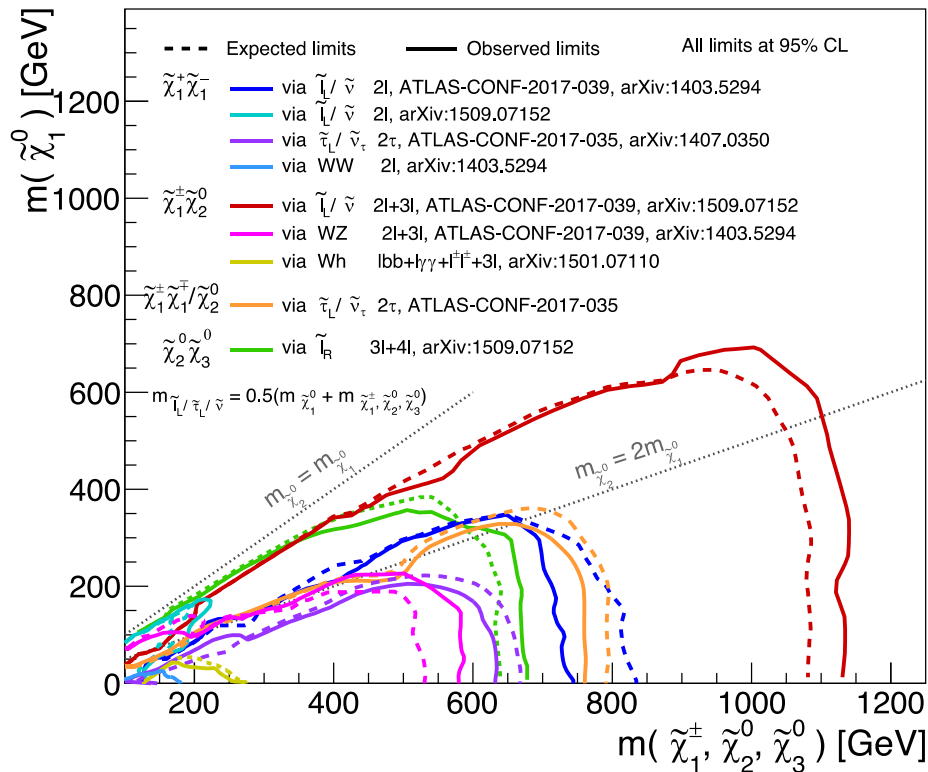
Endo, Hamaguchi, Si, Yoshinaga [1303.4256]

Searches for electroweakino

$$pp \rightarrow \tilde{\chi}_2^0 \tilde{\chi}_1^\pm$$



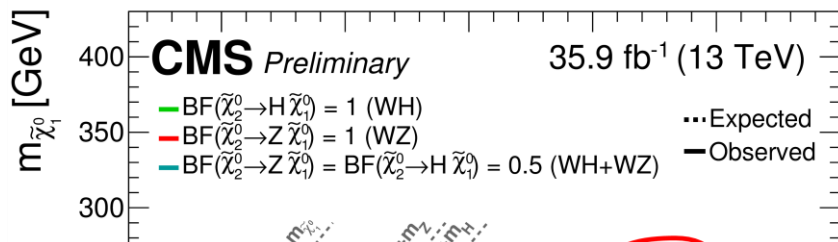
May 2017 **ATLAS Preliminary** $\sqrt{s}=8,13$ TeV, 20.3-36.1 fb⁻¹



CMS Summary plot <https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResultsSUS>
[ATLAS-CONF-2017-039](https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResultsSUS)

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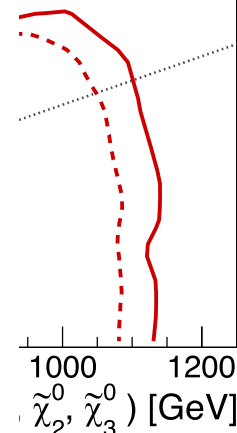
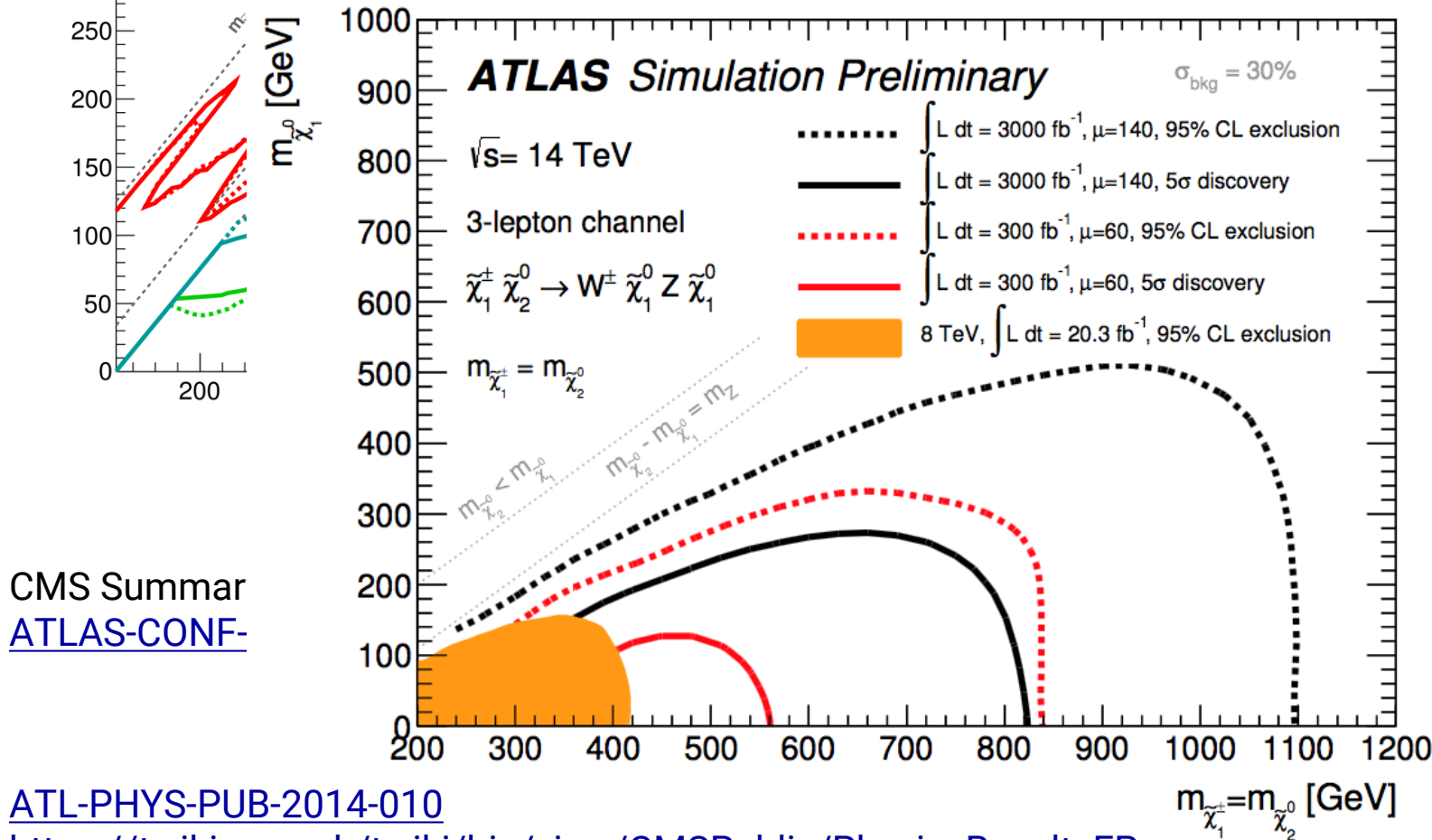
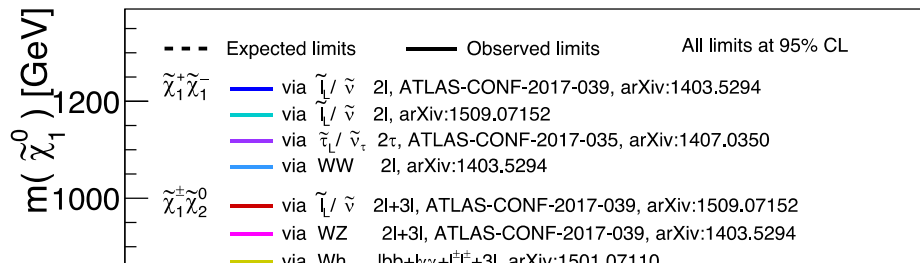
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May 2017

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CMS Summer
ATLAS-CONF-

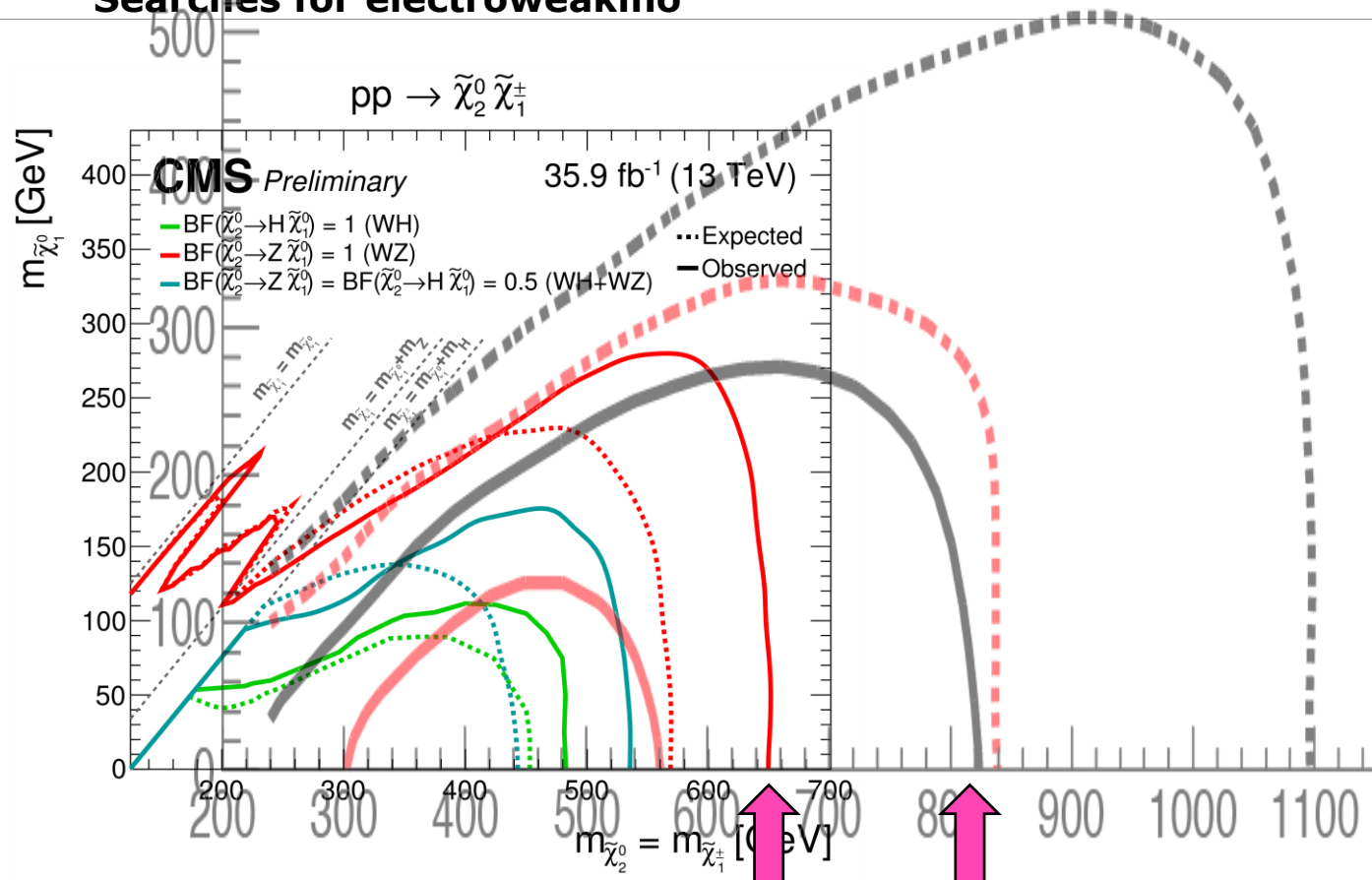
US

ATL-PHYS-PUB-2014-010

<https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResultsFP>

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Searches for electroweakino



discovery@3000/fb

corresponding exclusion@36/fb

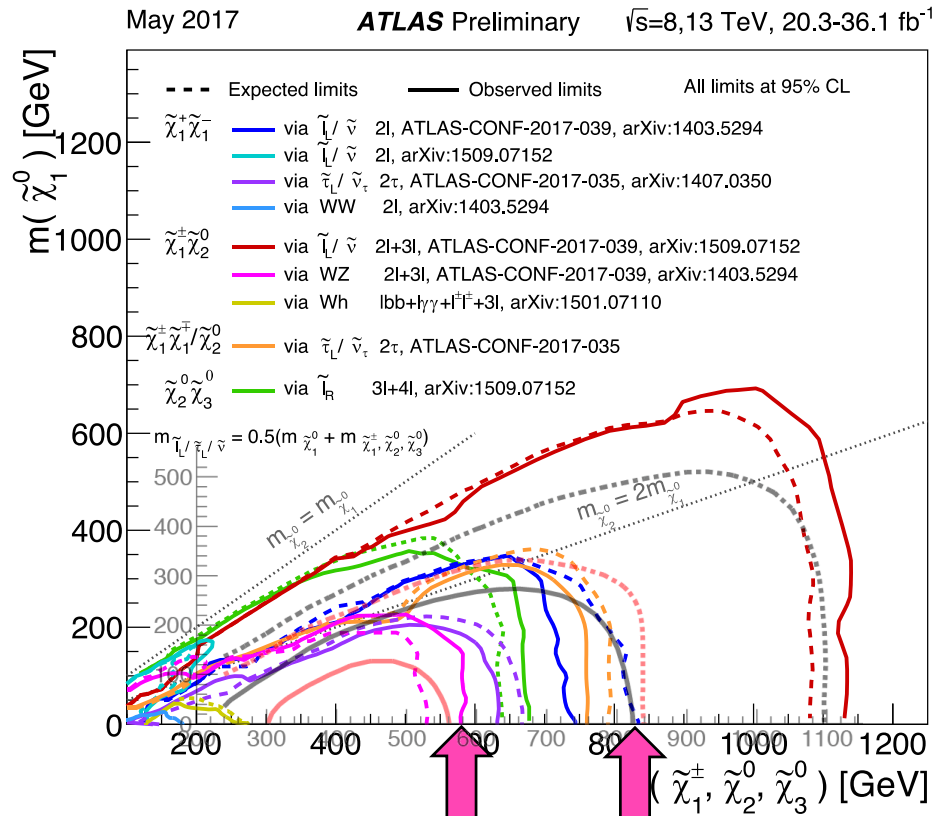
CMS Summary plot <https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResultsSUS>

[ATL-PHYS-PUB-2014-010](https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResultsFP)

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[ATLAS-CONF-2017-039](https://arxiv.org/abs/1705.02362)

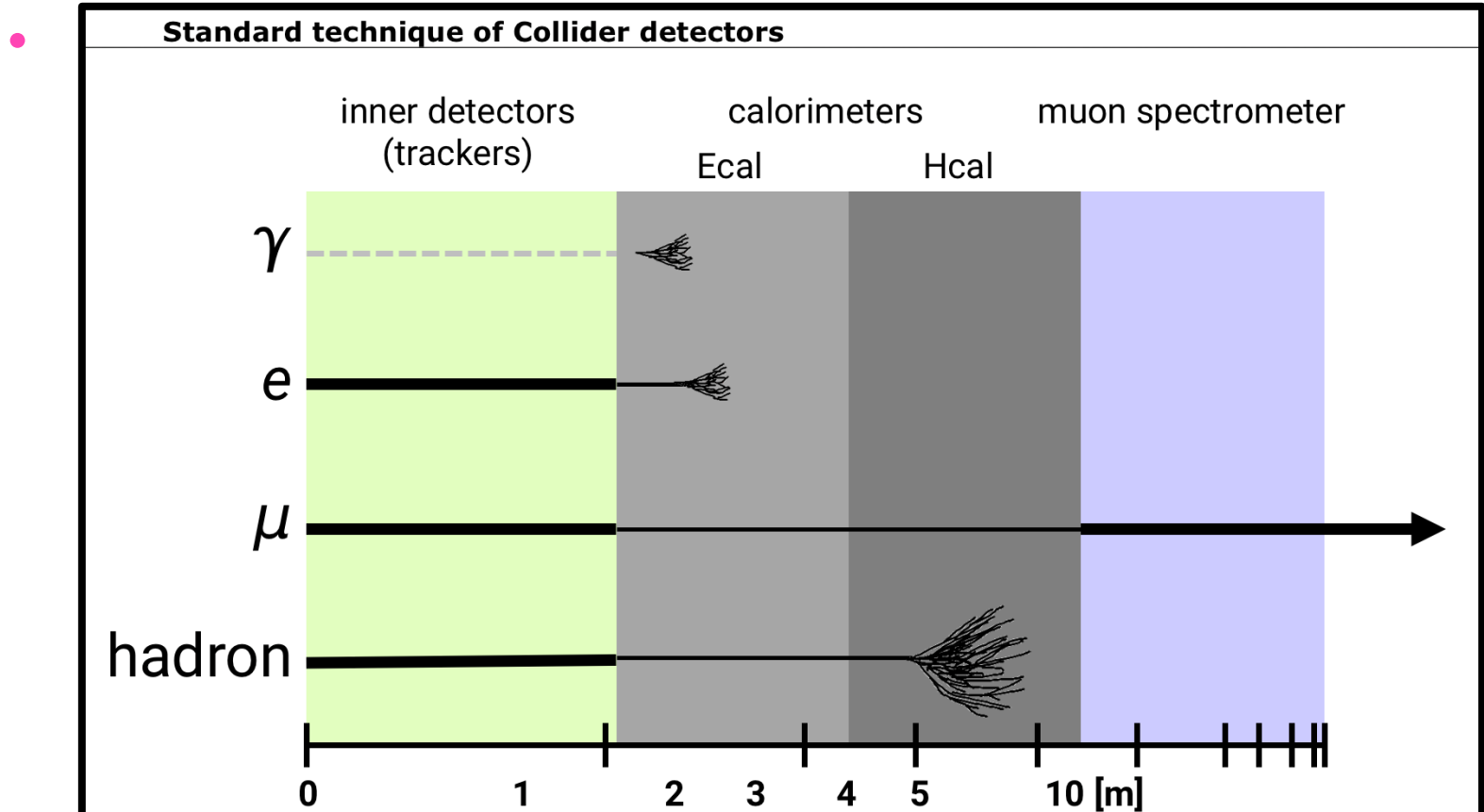
[ATL-PHYS-PUB-2014-010](https://arxiv.org/abs/1403.5294)

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0. Collider Physics basic

- What do "discovery" and "exclusion" mean?



- τ -lepton? \longrightarrow τ -tagging
- which quark? or gluon? \longrightarrow quark-flavor tagging

0. Collider Physics basic

- What do "discovery" and "exclusion" mean?
- How do they distinguish "e" and " μ "?

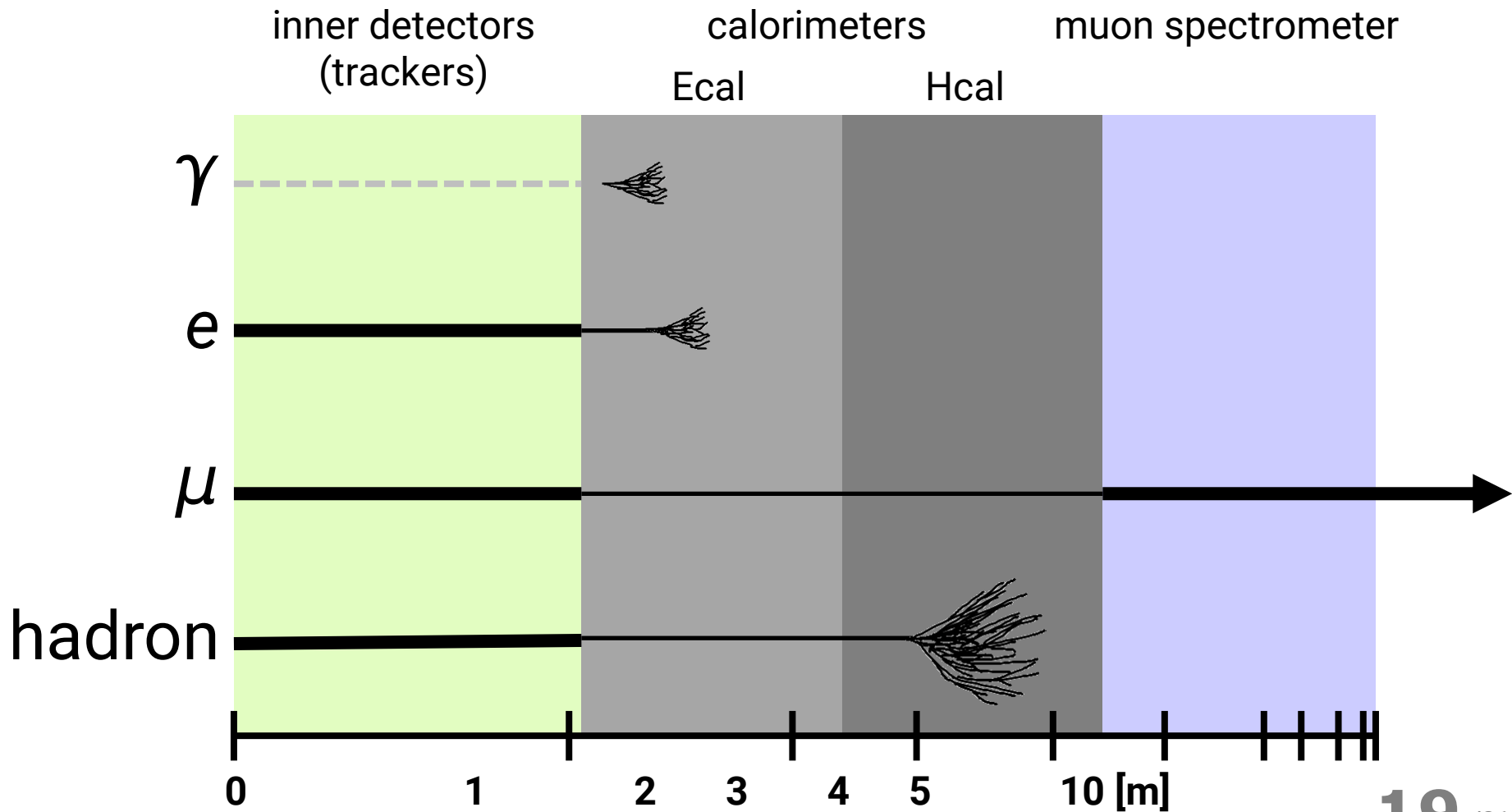
0. Collider Physics basic

- What do "discovery" and "exclusion" mean?

→ **blackboard**

0. Collider Physics basic

- How do they distinguish "e" and " μ "?



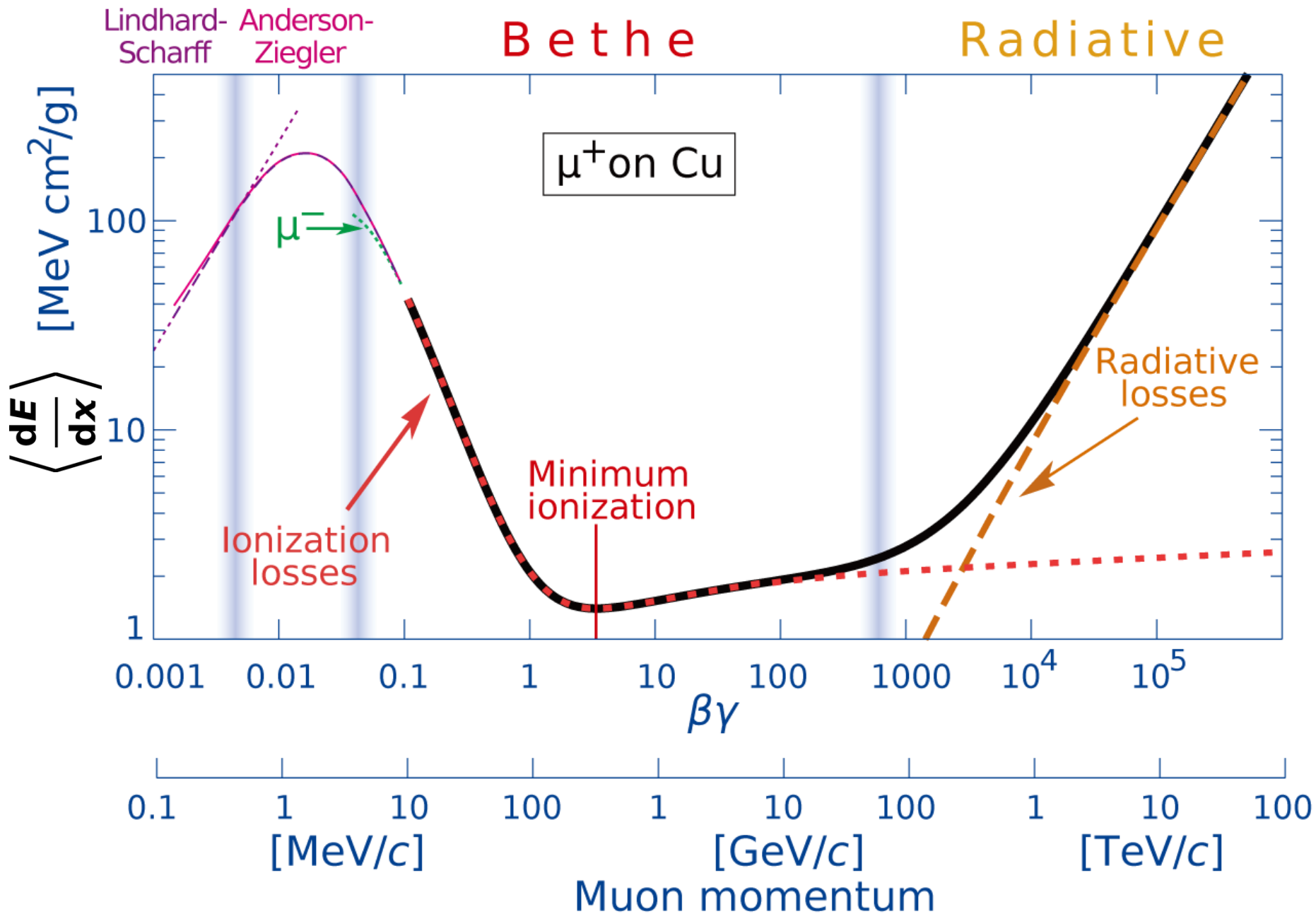


Figure from Groom, Mokhov, Striganov, Atom. Nucl. Data Tab. **78** (2001) 183-356
 [also in PDG Review "Passage of particles through matter"]

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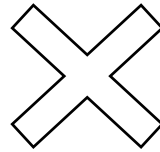
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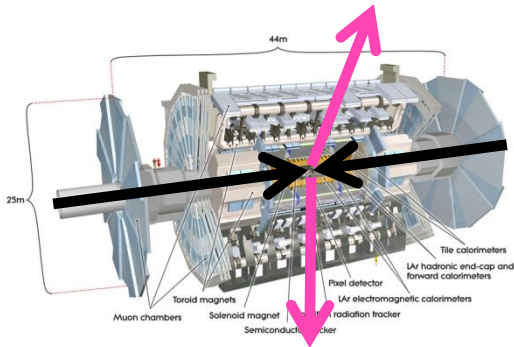
(subgroup in BSM@ep collaboration)

- Particle property $m > O(100)\text{GeV}$
 - Heavy colored \rightarrow hadronize \rightarrow $\left\{ \begin{array}{l} \text{R-hadron /} \\ \text{stopping particles} \end{array} \right.$
 - **Heavy non-colored**
 - Light non-colored, milli-charged \rightarrow dedicated searches



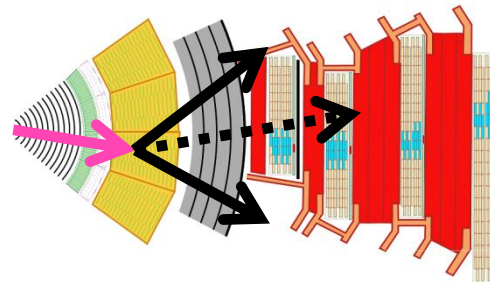
■ Lifetime

➤ “stable”



$$c\tau \gtrsim 1 \text{ m}$$

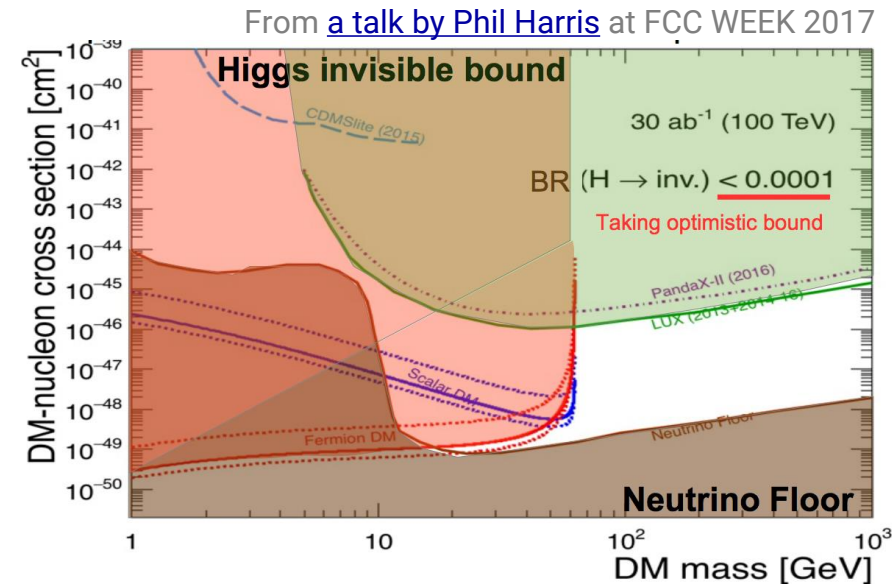
➤ “in-flight decay”



$$c\tau \sim 1 \text{ mm} - 1 \text{ m}$$

Why should we search for LLCP? – three viewpoints

- experimental
 - ...why not?
 - relevant for detector design
- phenomenology



long lifetime
→ an actor in early Universe

FCC-hh will(?) cover most of the
standard thermal-WIMP scenario

non-standard DM/cosmology with LLCP
→ next slides

- theoretical ... **SUSY?** *biased*
 - GMSB scenario: light gravitino → long-lived sleptons
 - split-SUSY: extremely heavy squarks → long-lived gluino

■ Dark Matter (especially to ameliorate DM over-abundance)

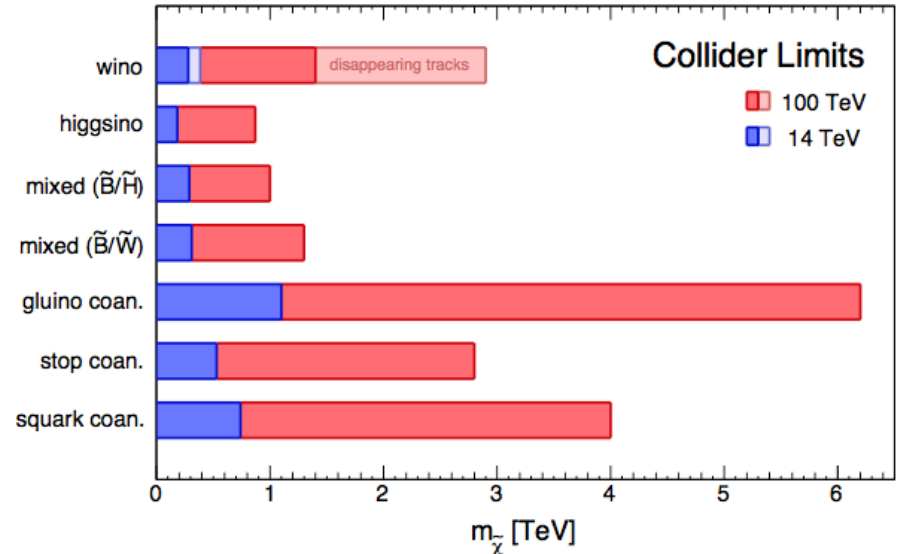
➤ co-annihilation

- $(\tilde{B} - \tilde{\tau}) \lesssim 700 \text{ GeV}$
- $(\tilde{W} - \tilde{g}) \lesssim 6-7 \text{ TeV}$
- $(\tilde{B} - \tilde{g})$ or $(\tilde{B} - \tilde{t}) \lesssim 8 \text{ TeV}$

Harigaya, Kaneta, Matsumoto [1403.0715],
Ellis, Olive, Zheng [1404.5571], etc.

➤ super-WIMP scenario

$$\rightarrow \tilde{l} > O(1) \text{ TeV}$$



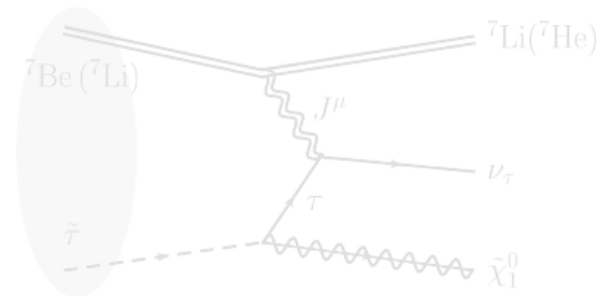
"Physics at the FCC-hh" Report [1606.00947]

■ Li problem

➤ MSSM $\tilde{\tau}$ with

$$m_{\tilde{\tau}} \sim 400 \text{ GeV}$$

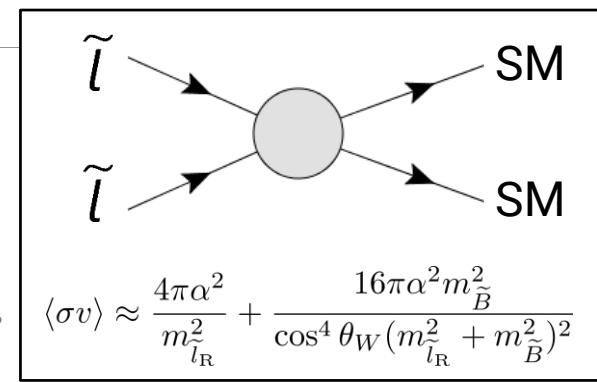
$$\Delta(m_{\tilde{\tau}} - m_{\text{LSP-DM}}) \sim 100 \text{ MeV}$$



Sato, Shimomura, Yamanaka [1604.04769]

■ Super-WIMP:

➤ NLSP slepton \tilde{l} + LSP gravitino \tilde{G}



frozen-out

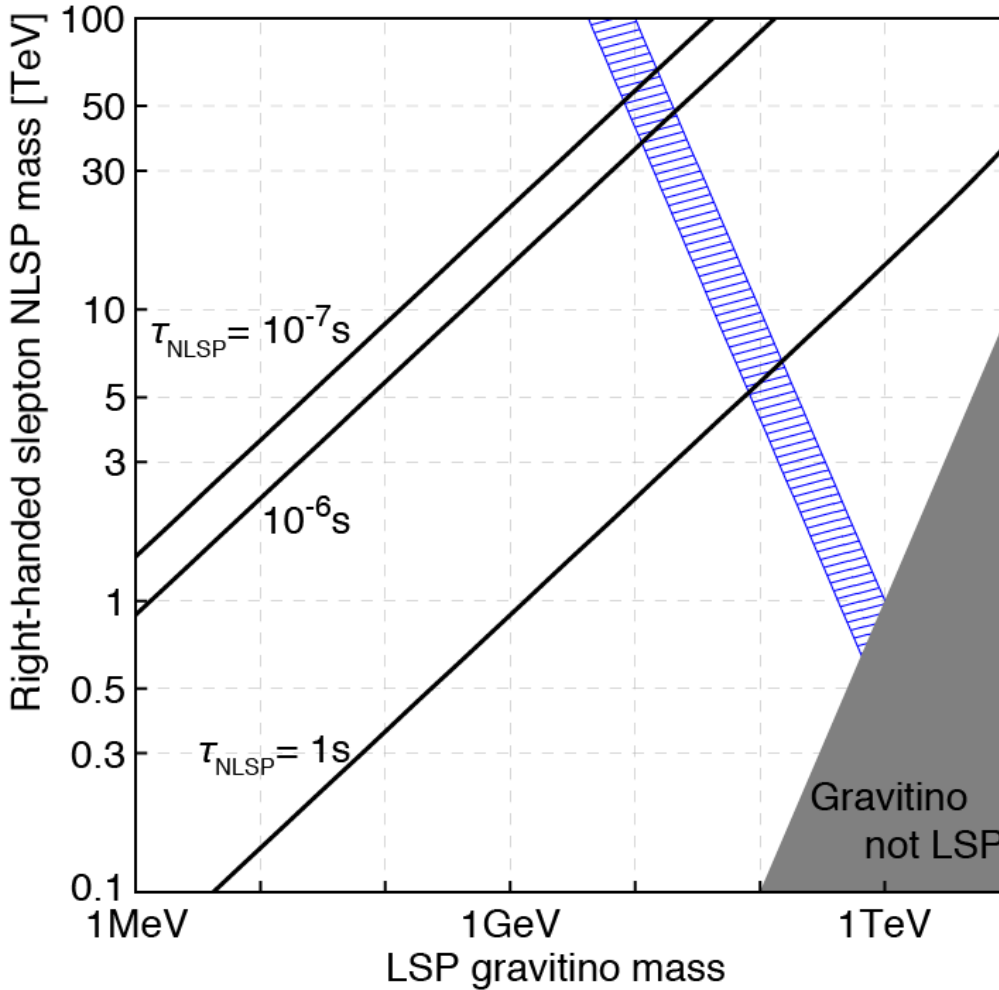
→ thermal relic with $(\Omega h^2)_{\tilde{l}} \gg 0.12$

late-time decay $\tilde{l} \rightarrow \tilde{G} + l$

$$(\Omega h^2)_{\text{DM}(\tilde{G})} = \frac{m_{\tilde{G}}}{m_{\tilde{l}}} (\Omega h^2)_{\tilde{l}}$$

$$\tau(\tilde{l} \rightarrow l\tilde{G}) = 5.7 \times 10^{-7} \text{ sec} \cdot \left(\frac{m_{\tilde{l}}}{1 \text{ TeV}}\right)^{-5} \left(\frac{m_{\tilde{G}}}{1 \text{ MeV}}\right)^2$$

■ Super-WIMP:



(short-lived)

LLCP search target

(BBN/CMB constraints are relevant.)

$$\tau(\tilde{l} \rightarrow l\tilde{G}) = 5.7 \times 10^{-7} \text{ sec} \cdot \left(\frac{m_{\tilde{l}}}{1 \text{ TeV}}\right)^{-5} \left(\frac{m_{\tilde{G}}}{1 \text{ MeV}}\right)^2$$

■ Dark Matter (especially to ameliorate DM over-abundance)

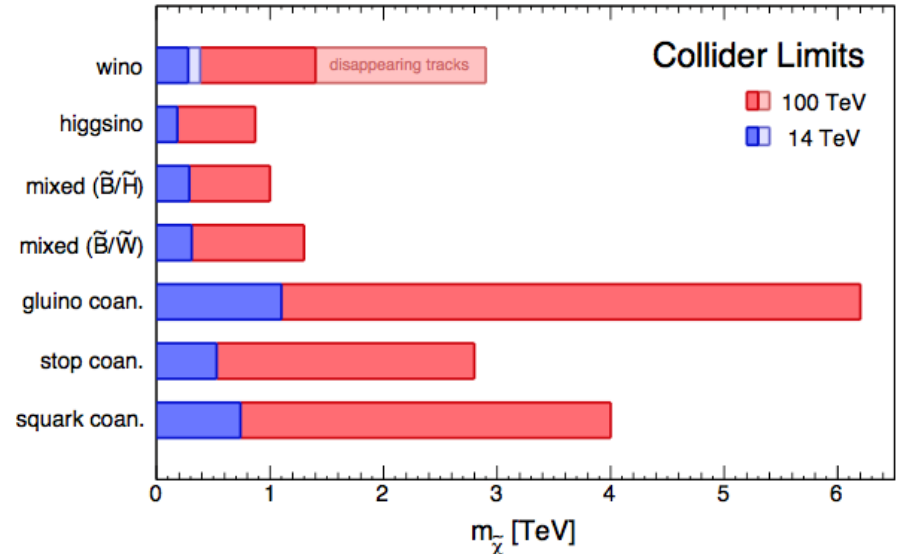
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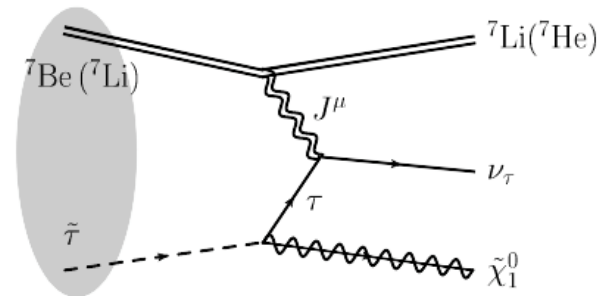
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Harigaya, Kaneta, Matsuimoto [1403.0715],
Ellis, Olive, Zheng [1404.557] etc.

non-colored LLCP

... covered in this talk

(slepton \tilde{l} / stau $\tilde{\tau}$ as benchmarks)

➤ super-WIMP scenario

→ $\tilde{l} > O(1) \text{ TeV}$
(as we see soon.)



colored LLCP

(→ hadronize: "R-hadron")

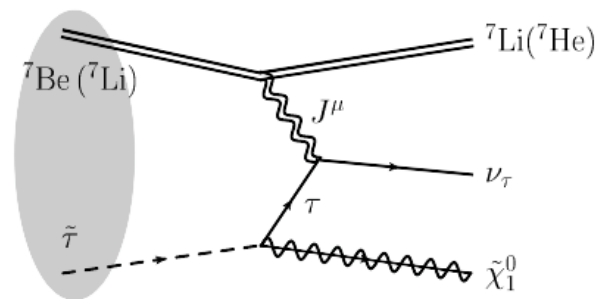


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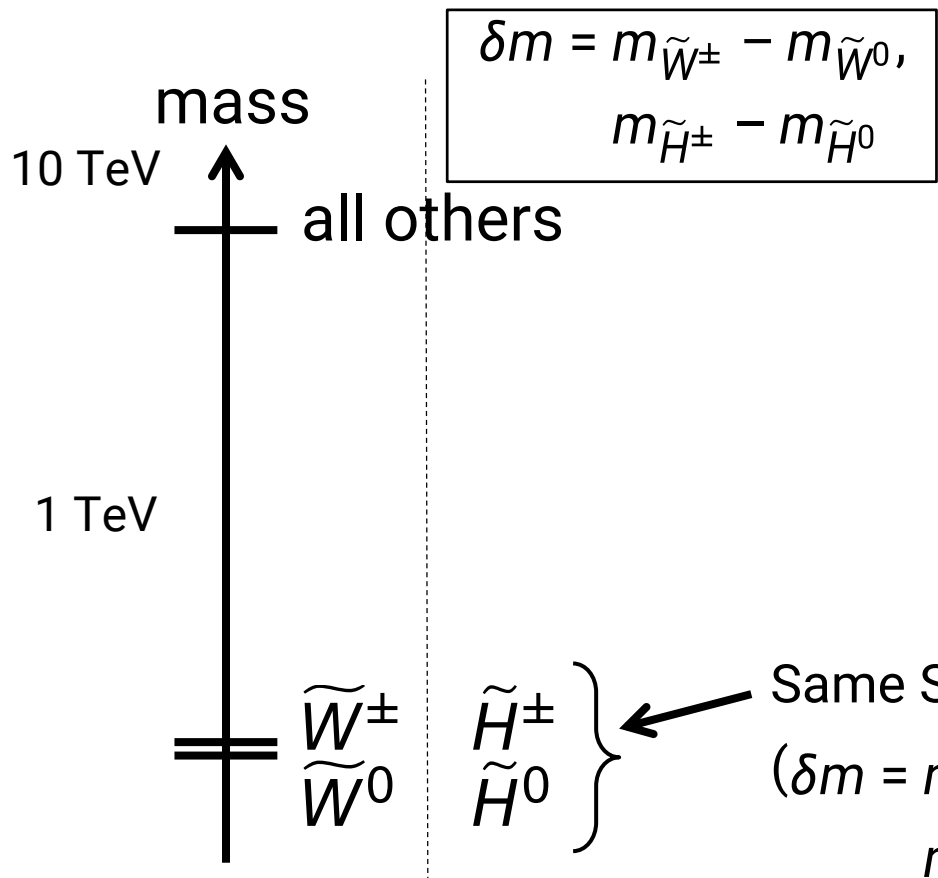
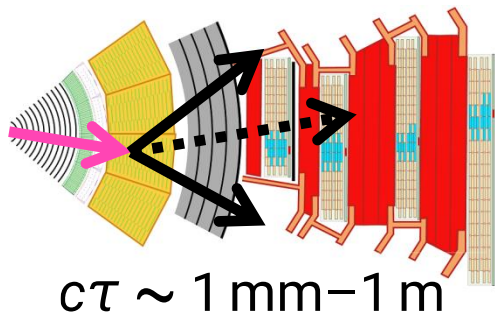
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■ Simple MSSM models

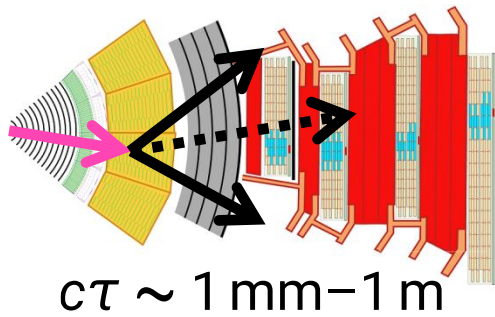
- Pure-Wino dark matter
- Pure-Higgsino dark matter



$\tilde{\chi}^\pm$ long-lived because of small Δm
($\Delta m > 0$)

■ Other MSSM models

- Gauge-Mediation
- R-parity violation



$$\delta m = m_{\tilde{W}^\pm} - m_{\tilde{W}^0},$$

$$m_{\tilde{H}^\pm} - m_{\tilde{H}^0}$$

Simple MSSM models

- Pure-Wino dark matter **60mm**
- Pure-Higgsino dark matter **7mm**

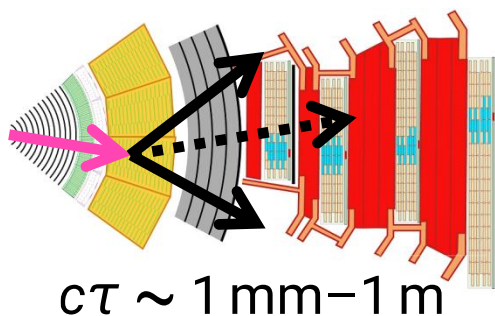


$\tilde{\chi}^\pm$ long-lived because of small δm
 ($\delta m > 0$)

$m_{\tilde{W}}$ [GeV]	200	250	300	350	400	450	500	550	600	700	800	900
δm [MeV]	159	160	161	162	162	163	163	163	163	164	164	164
$c\tau$ [mm]	71	67	64	63	62	61	60	60	59	59	59	59

➤ R-parity violation

$m_{\tilde{H}}$ [GeV]	200	250	300	350	400	450	500	550	600	700	800	900
δm [MeV]	297	306	313	319	323	326	329	331	333	336	338	340
$c\tau$ [mm]	11	10	9.4	8.9	8.5	8.2	8.0	7.8	7.7	7.4	7.2	7.1



Simple MSSM models

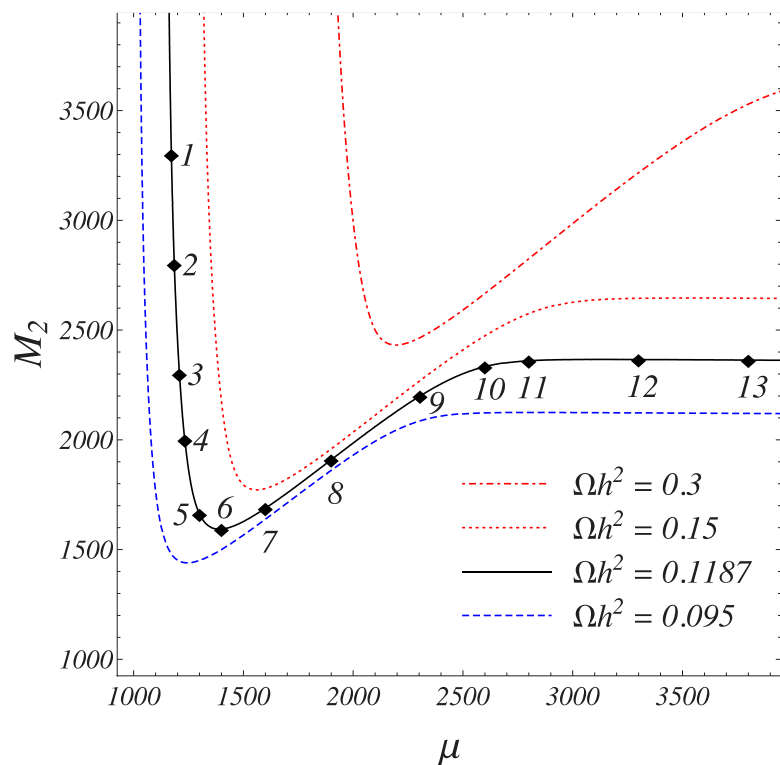
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$m_{A^0} = 500 \text{ GeV}, m_{\text{sfermion}} = 9 \text{ TeV}, \tan \beta = 15$ $\tilde{\chi}^{\pm}$ long-lived because of small δm ($\delta m > 0$)

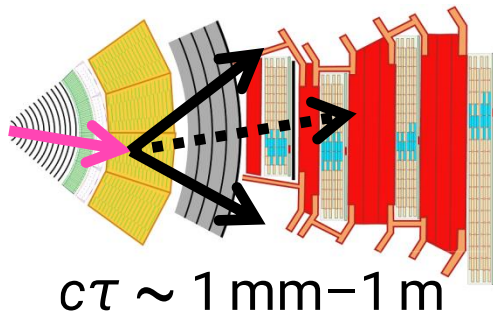
(but **DM underabundant** for "observable" region)



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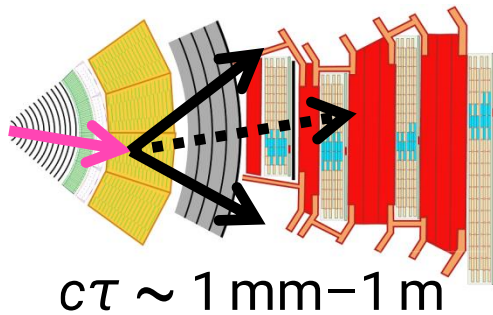
■ Other MSSM models

- Gauge-Mediation (keV \tilde{G})

- R-parity violation

with \tilde{l} -LSP

⇒ \tilde{l} long-lived
because of tiny couplings.



■ Simple MSSM models

- Pure-Wino dark matter **60mm**
- Pure-Higgsino dark matter **7mm**



$\tilde{\chi}^\pm$ long-lived because of small δm
 ($\delta m > 0$)

(but **DM underabundant**
 for “observable” region)

$$c\tau \sim 1.8 \times 10^{-5} \text{ m} \left(\frac{m_{\tilde{l}}}{100 \text{ GeV}} \right)^{-5} \left(\frac{m_{\tilde{G}}}{1 \text{ eV}} \right)^2$$

for Gauge-Mediation: “gravity is weak”,

$$0.50 \text{ m} \left(\frac{m_{\tilde{l}}}{100 \text{ GeV}} \right)^{-1} \left(\frac{\lambda_{ijk}}{10^{-8}} \right)^{-2}$$

for R-parity violation: “if RpV is tiny”.

■ Other MSSM models

- Gauge-Mediation (keV \tilde{G})
- R-parity violation

with \tilde{l} -LSP

⇒ \tilde{l} long-lived **any $c\tau$**
 because of tiny couplings.

■ Dark Matter

➤ co-annihilation

- $(\tilde{B} - \tilde{\tau}) \lesssim 700 \text{ GeV}$
- $(\tilde{W} - \tilde{g}) \lesssim 6-7 \text{ TeV}$
- $(\tilde{B} - \tilde{g})$ or $(\tilde{B} - \tilde{t}) \lesssim 8 \text{ TeV}$

$\tilde{\tau}$ -LLCP

➤ super-WIMP scenario → $\tilde{l} > O(1) \text{ TeV}$

\tilde{l} -LLCP

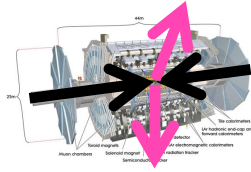
■ Li problem

➤ MSSM $\tilde{\tau}$ with

$$m_{\tilde{\tau}} \sim 400 \text{ GeV}$$

$$\Delta(m_{\tilde{\tau}} - m_{\text{LSP-DM}}) \sim 100 \text{ MeV}$$

$\tilde{\tau}$ -LLCP



■ Simple MSSM models

➤ Pure-Wino dark matter **60mm**

➤ Pure-Higgsino dark matter **7mm**



$\tilde{\chi}^{\pm}$ long-lived because of small δm
($\delta m > 0$)

(but **DM underabundant**
for “observable” region)

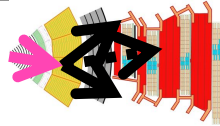
■ Other MSSM models

➤ Gauge-Mediation (keV \tilde{G})

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1. LLCPs

- Introduction & Motivations
- How to search? & Current bounds

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3. LLCP searches at FCC-hh

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Based on

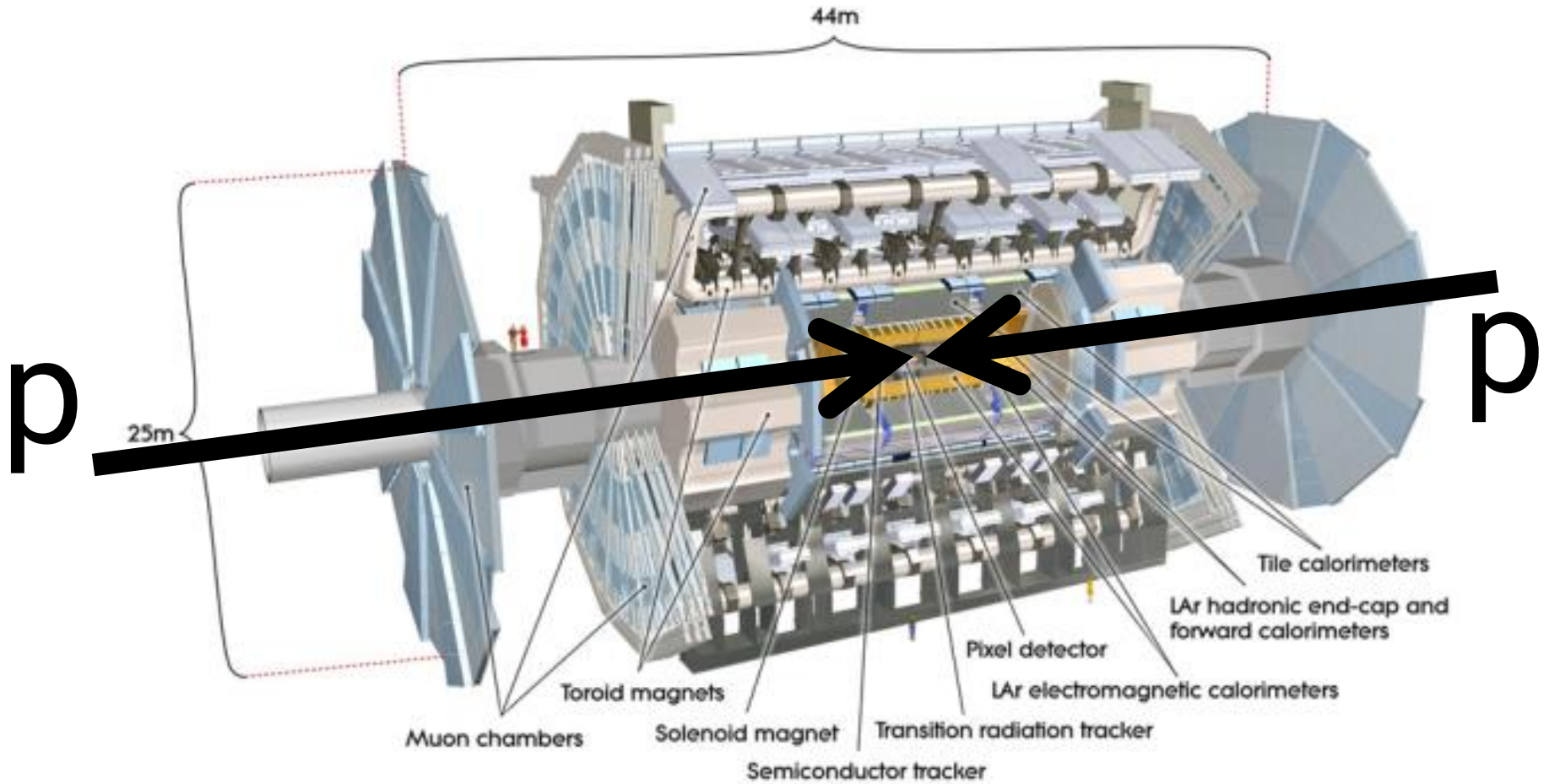
hh: Jonathan L. Feng, SI, Yael Shadmi, Shlomit Tarem [[1505.02996](#)]

(collected in FCC-hh report [[1606.00947](#)])

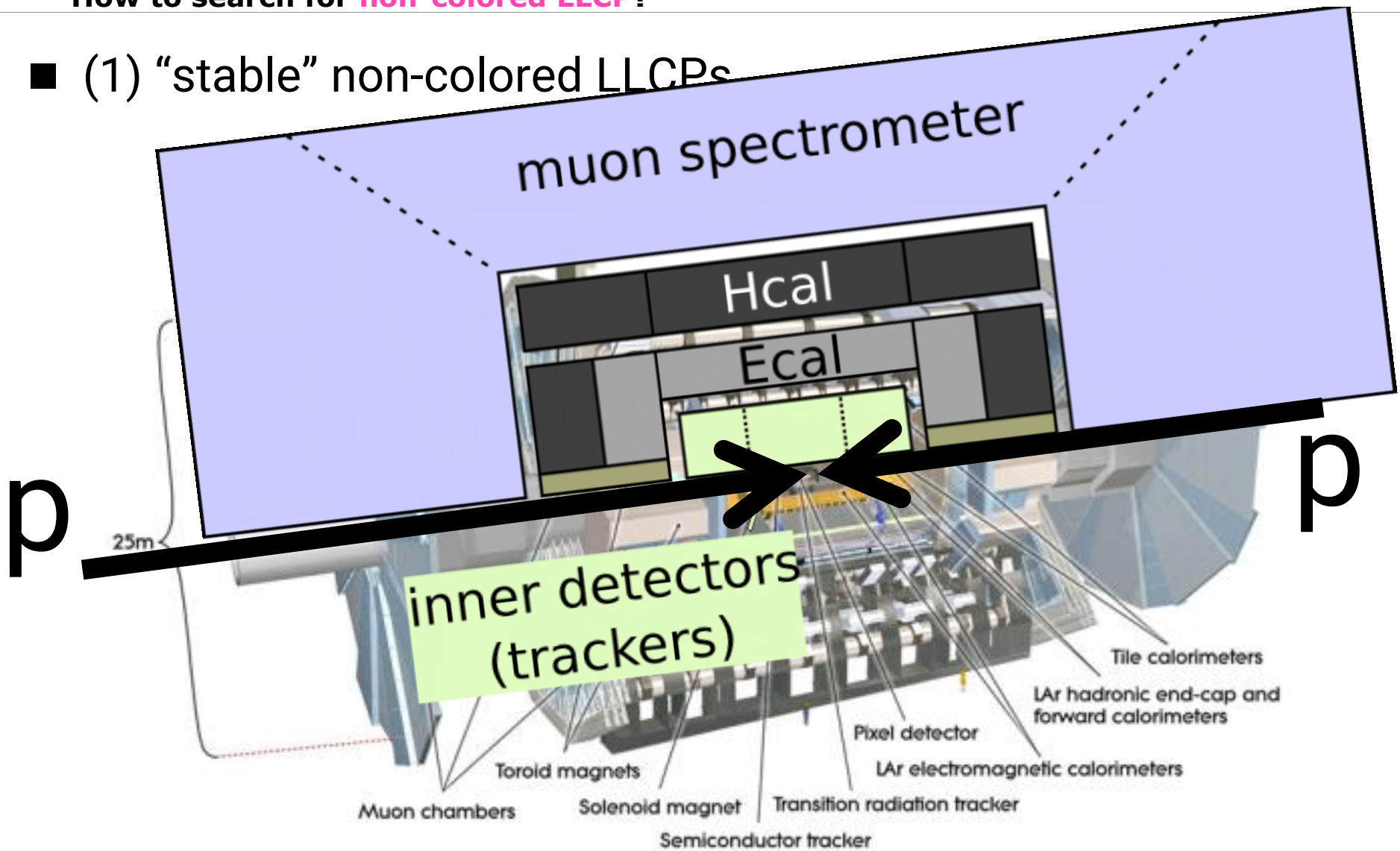
he: Kechen Wang, SI, Monica D'Onofrio, Georges Azuelos [17???.?????]

(subgroup in BSM@ep collaboration)

- (1) “stable” non-colored LLCPs

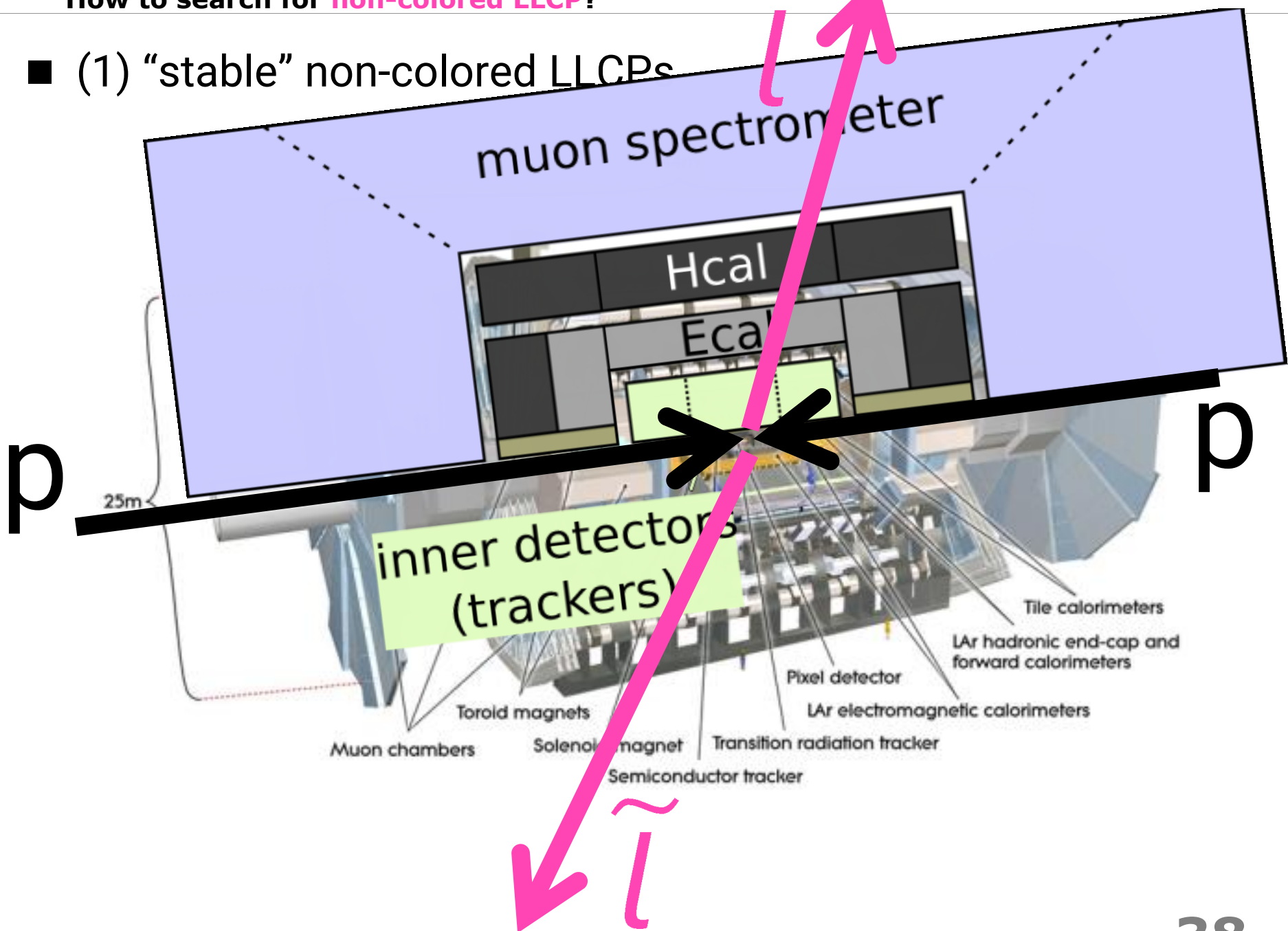


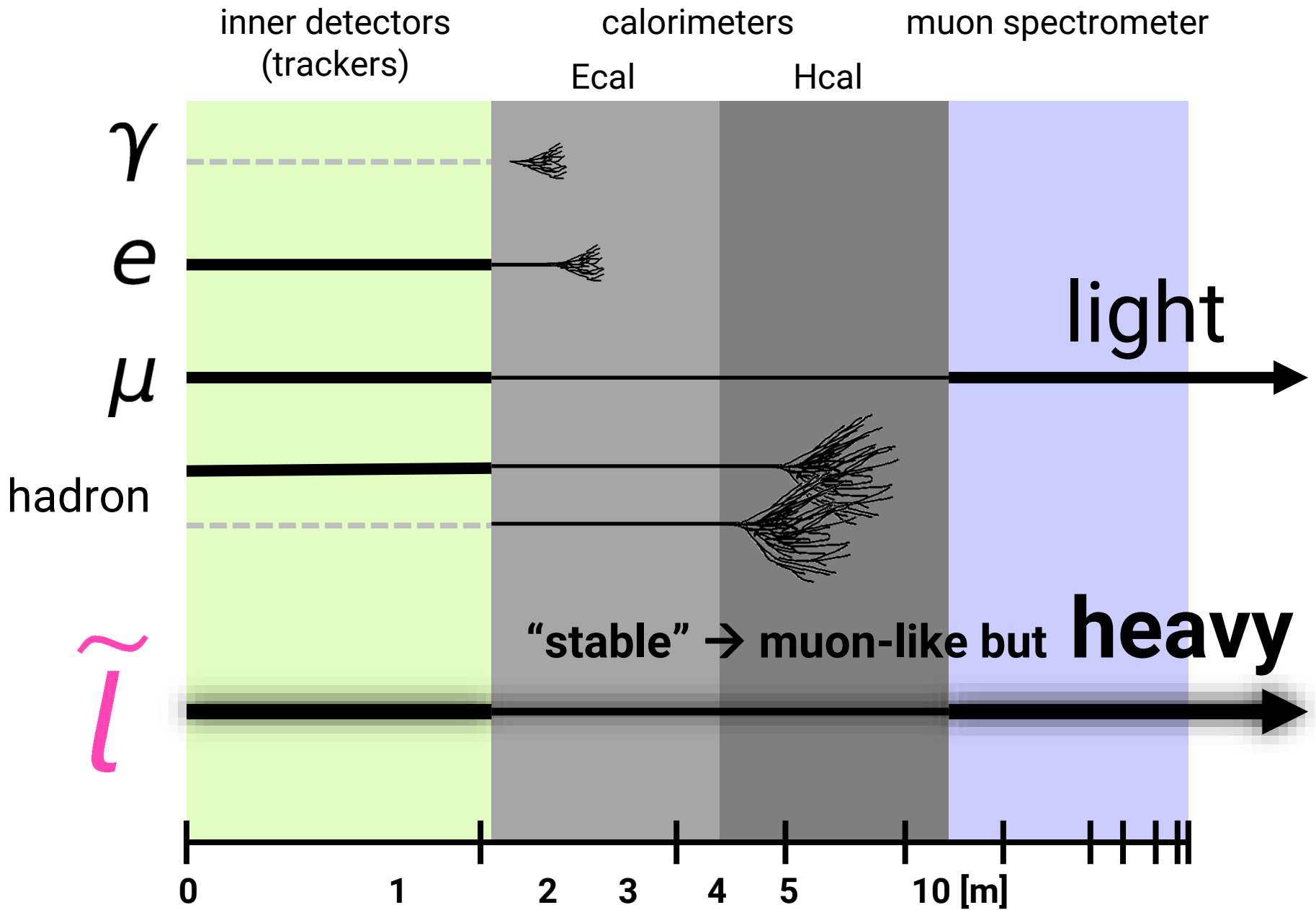
- (1) “stable” non-colored LLCPs



How to search for non-colored LLCPs?

- (1) "stable" non-colored LLCPs



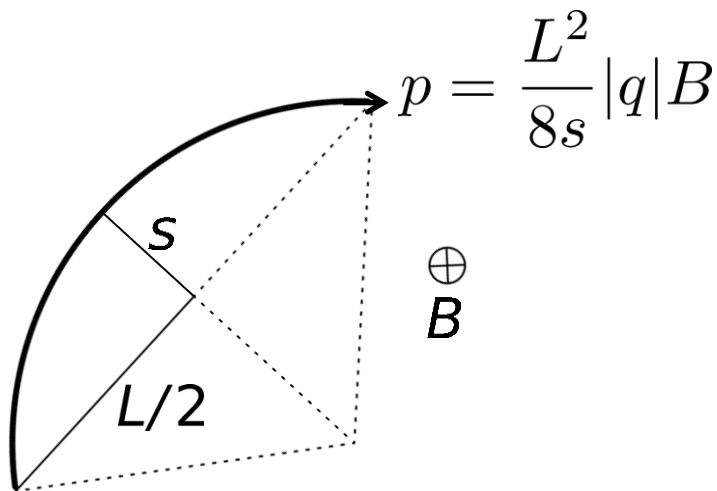


$$m = \frac{p}{\beta\gamma} = \frac{p}{\beta/\sqrt{1-\beta^2}}$$

momentum & velocity

■ **mass** measurement = **p** & **β** measurements ($\beta = v/c$)

➤ momentum



➤ velocity

- TOF [time-of-flight]
 $\beta = \Delta L / \Delta t$
- dE/dx [ionization energy loss]

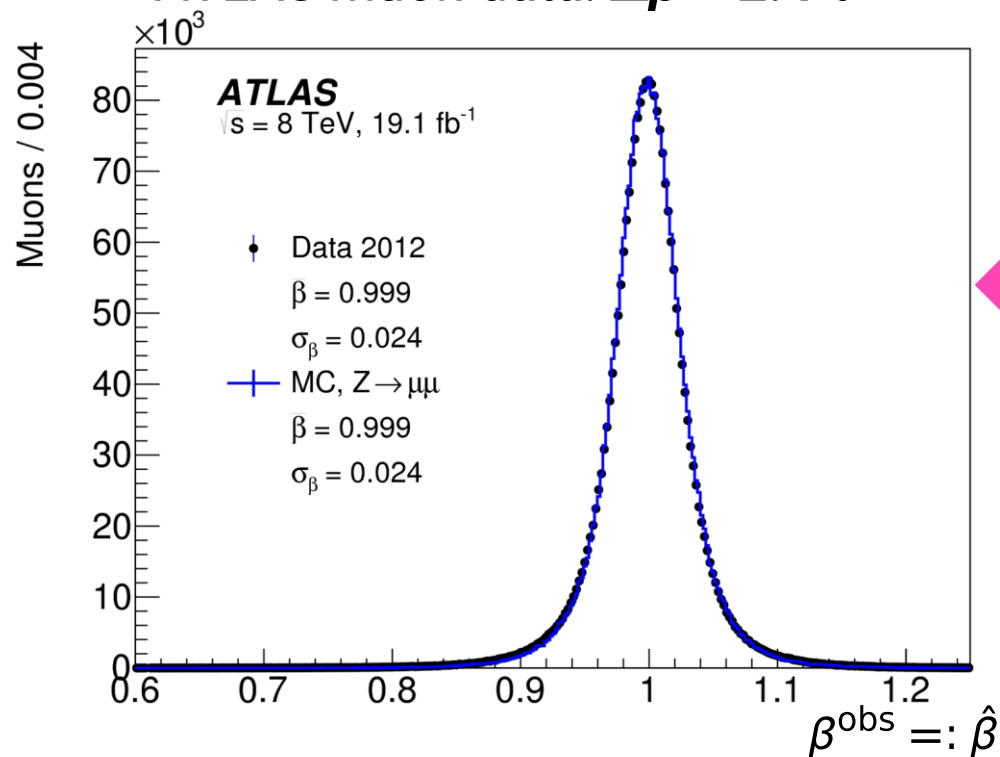
"Mass measurement" to distinguish long-lived sleptons

$$m = \frac{p}{\beta\gamma} = \frac{p}{\beta/\sqrt{1-\beta^2}}$$

momentum & velocity

■ **mass** measurement = **p** & **β** measurements ($\beta = v/c$)

ATLAS muon data: $\Delta\beta = 2.4\%$

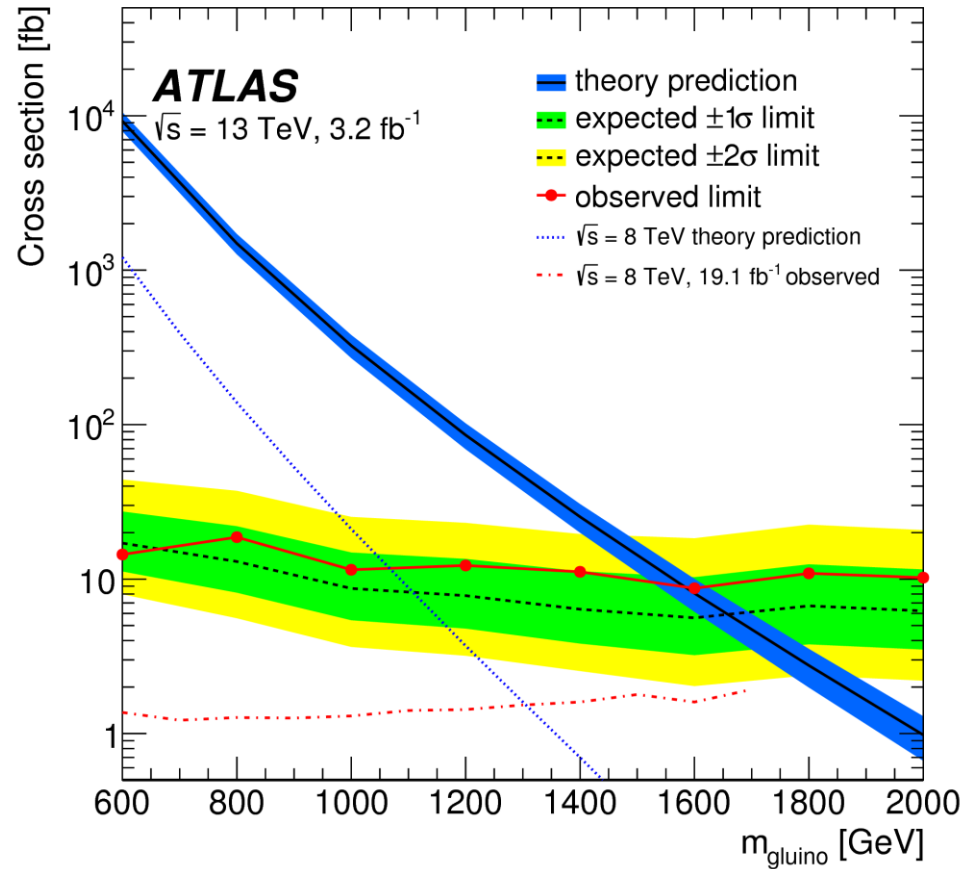
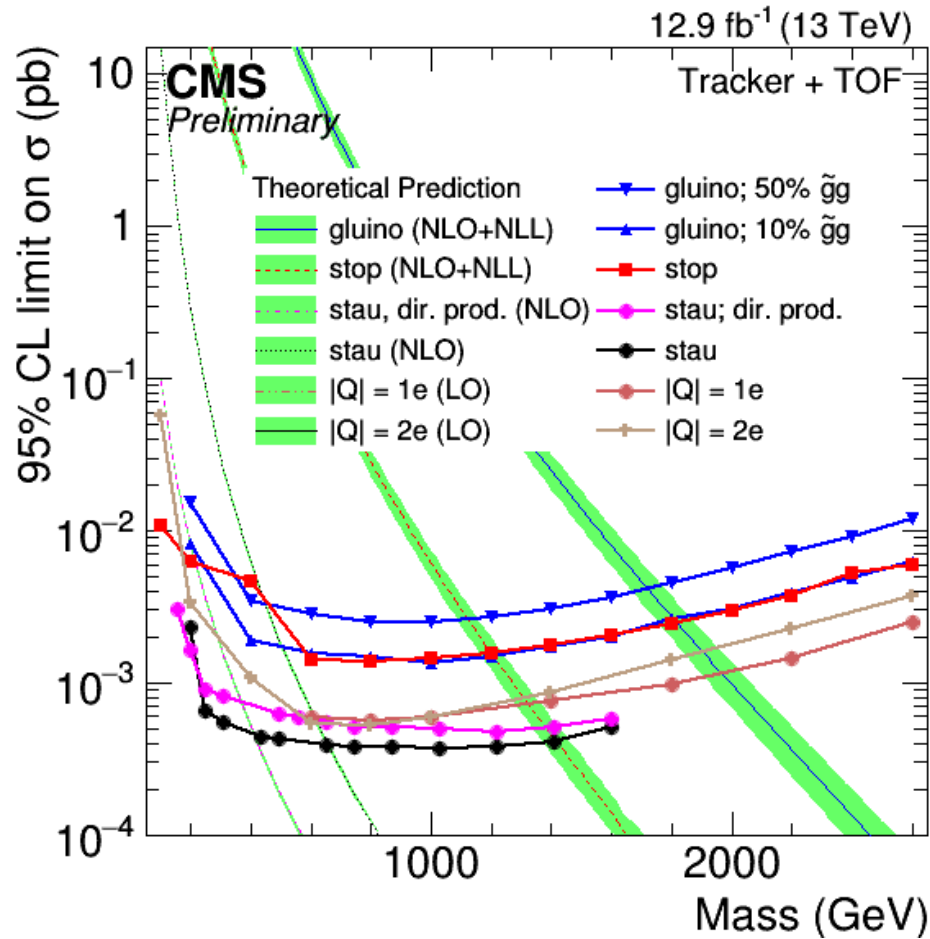


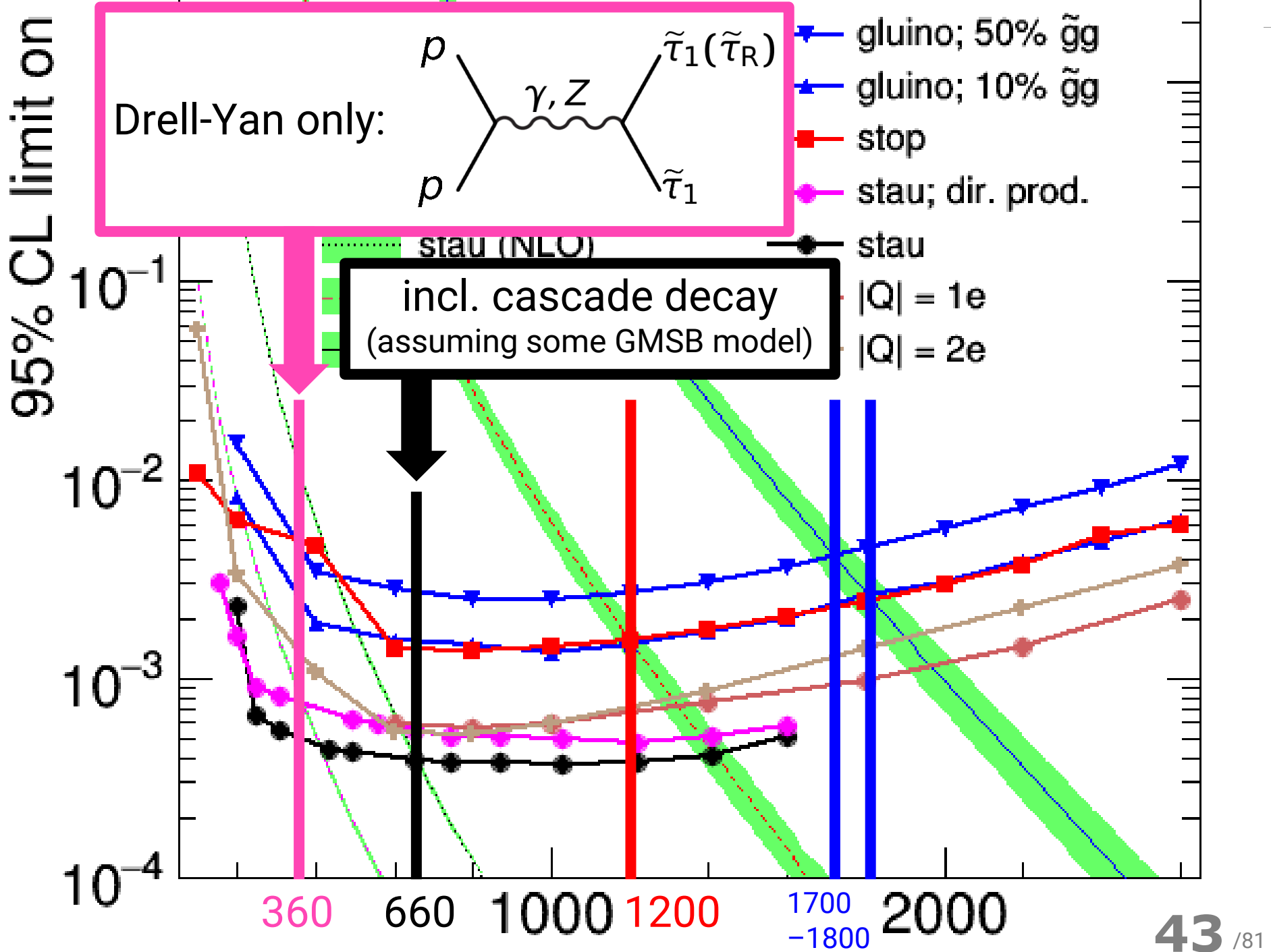
➤ velocity

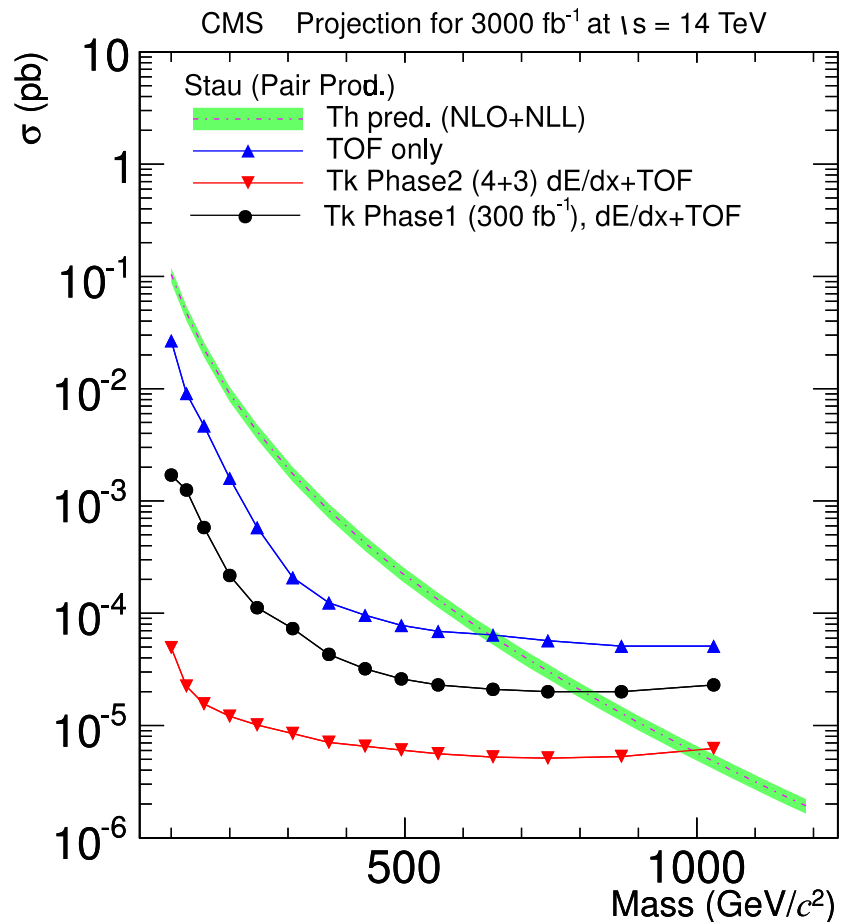
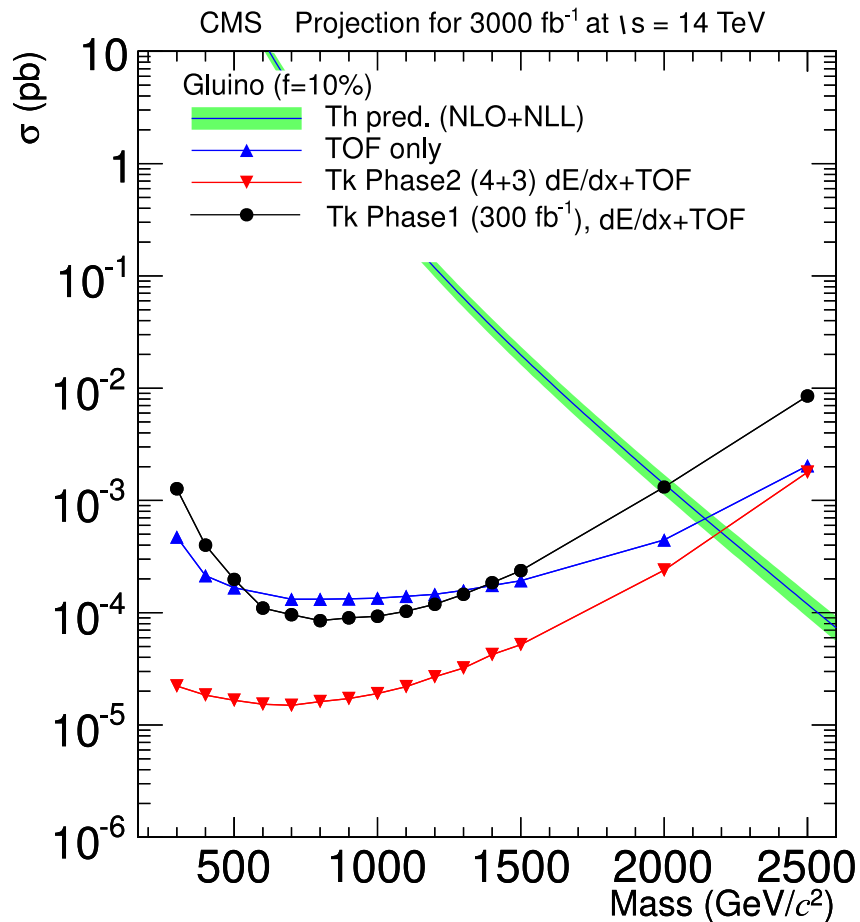
- TOF [time-of-flight]

← $\beta = \Delta L / \Delta t$

- dE/dx [ionization energy loss]



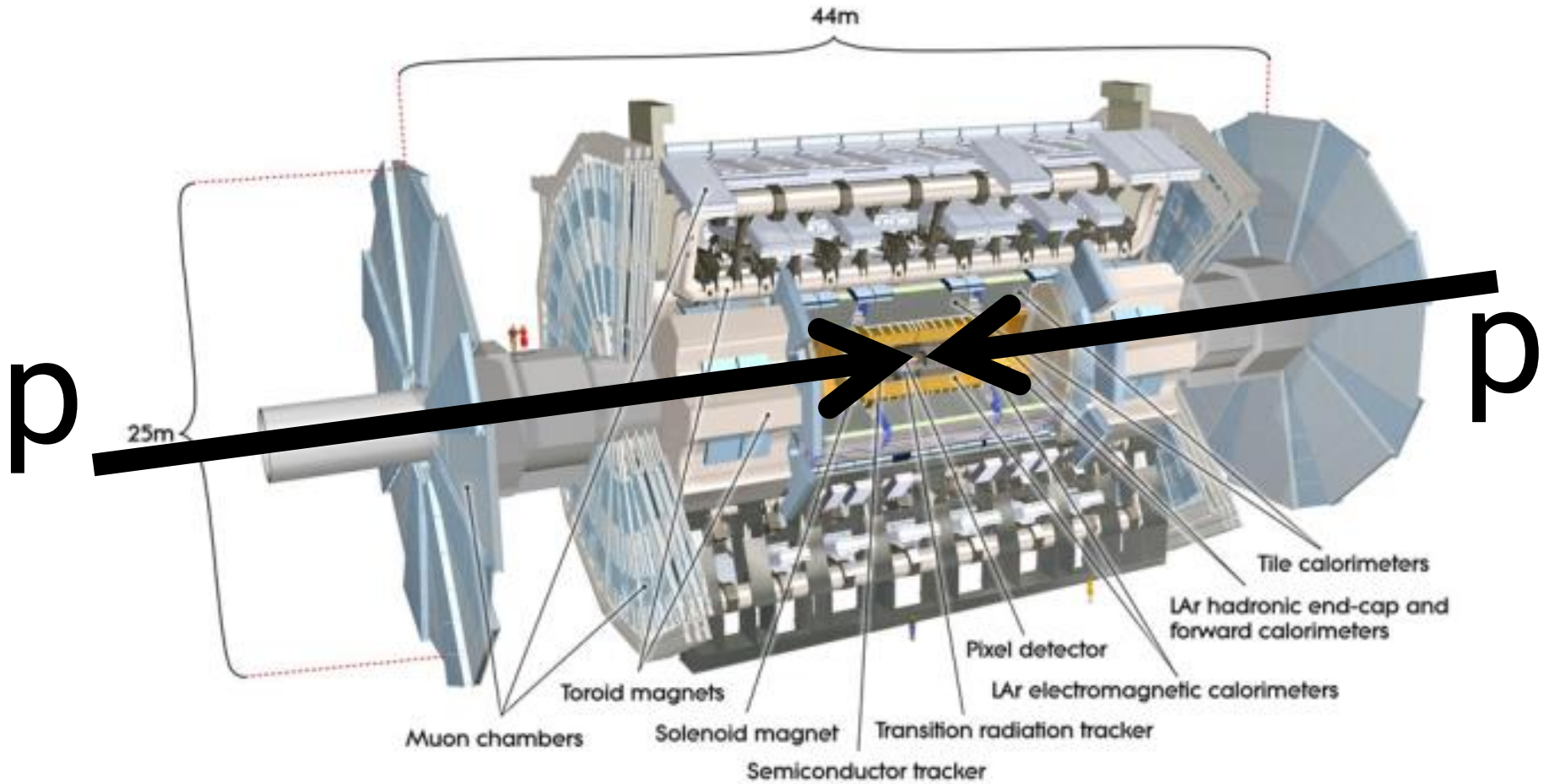




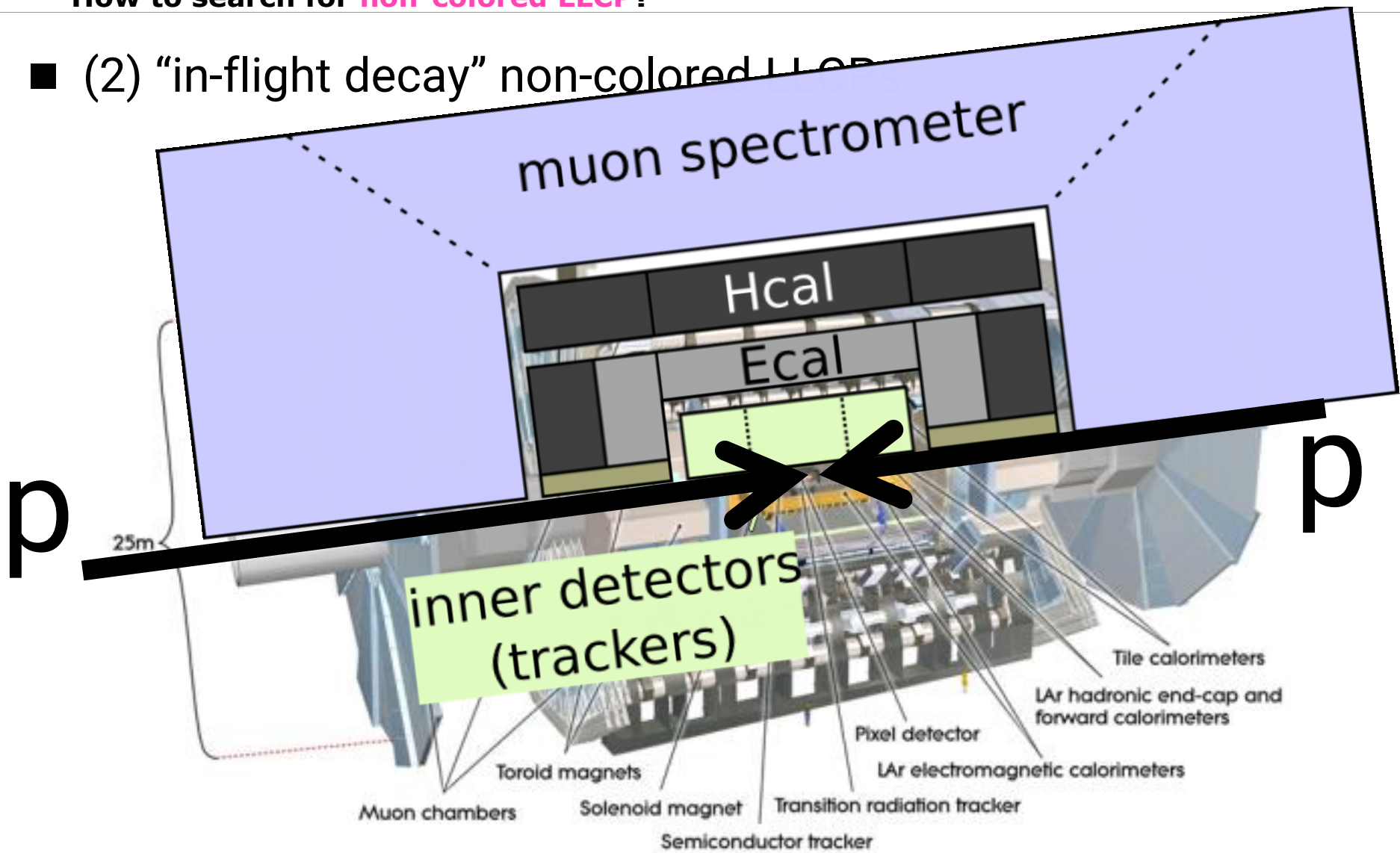
	current	HL-LHC
gluino:	1.7 TeV	→ 2.2 TeV?
stop:	1.2 TeV	→ 1.7 TeV?
stau (GMSB):	660 GeV	→ 1.2 TeV?
stau (DY):	360 GeV	→ 1.0 TeV?

(or discovery?)

- (2) “in-flight decay” non-colored LLCs

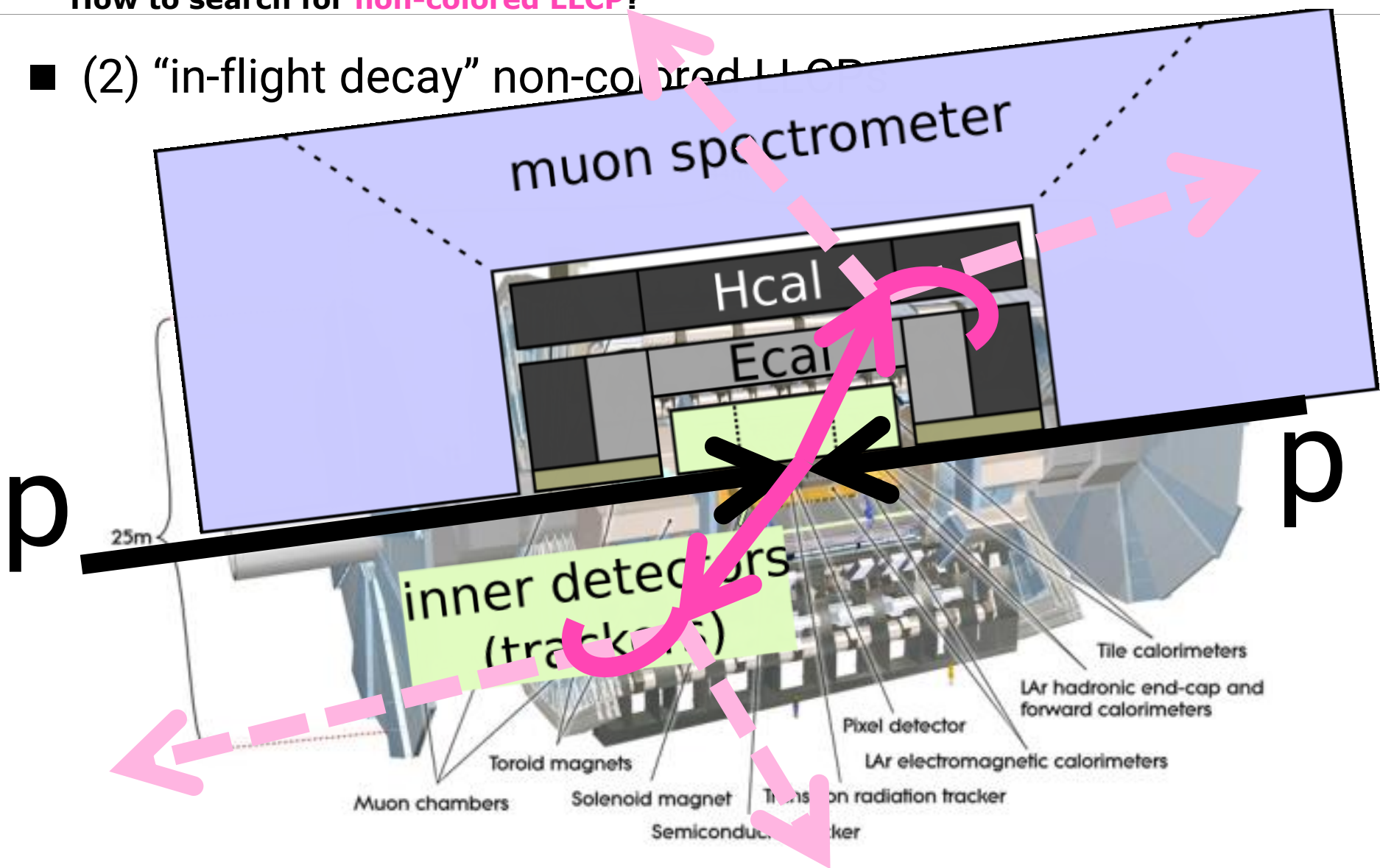


- (2) “in-flight decay” non-colored LLCP

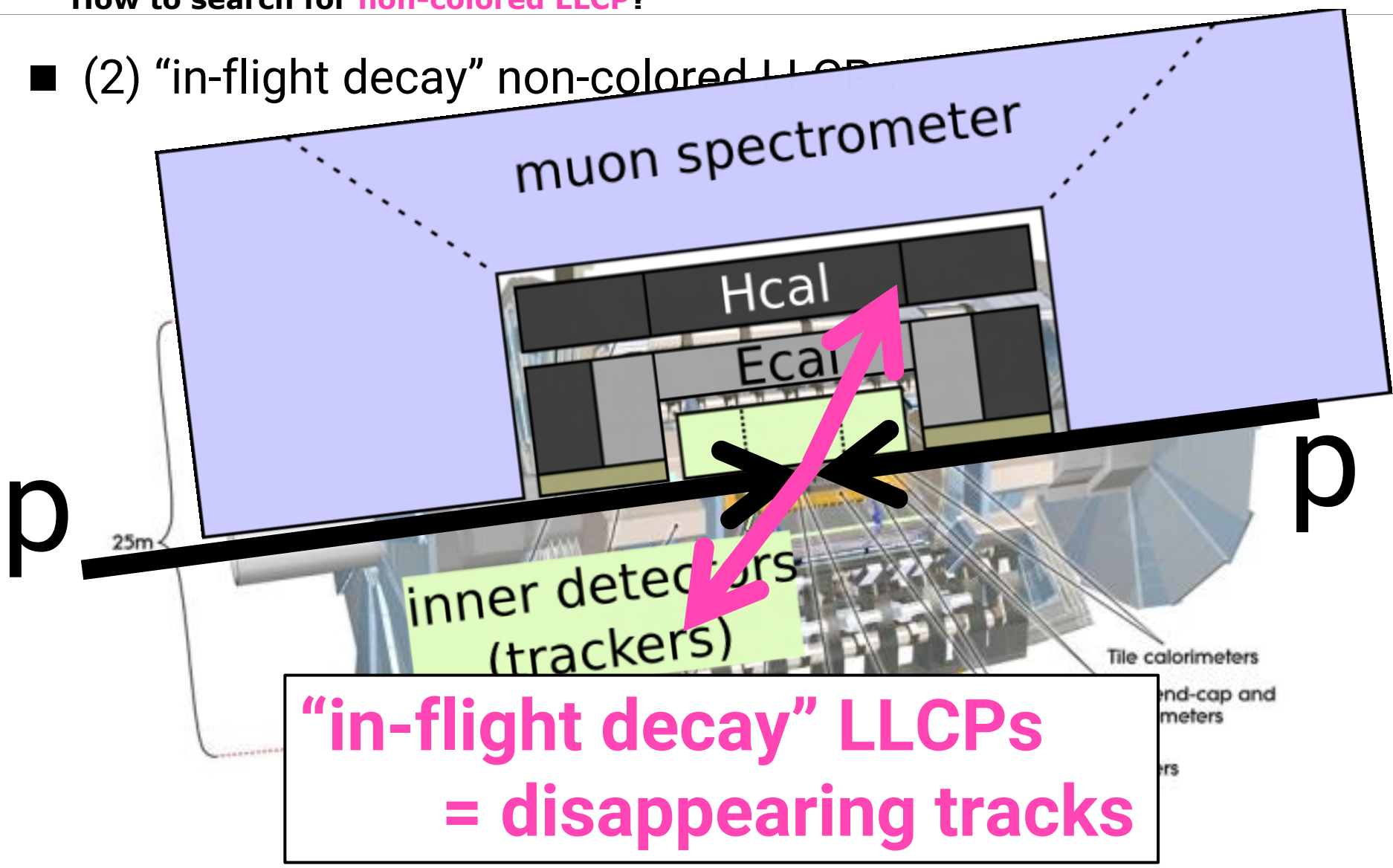


How to search for **non-colored LLCP**?

- (2) “in-flight decay” non-colored LLCP

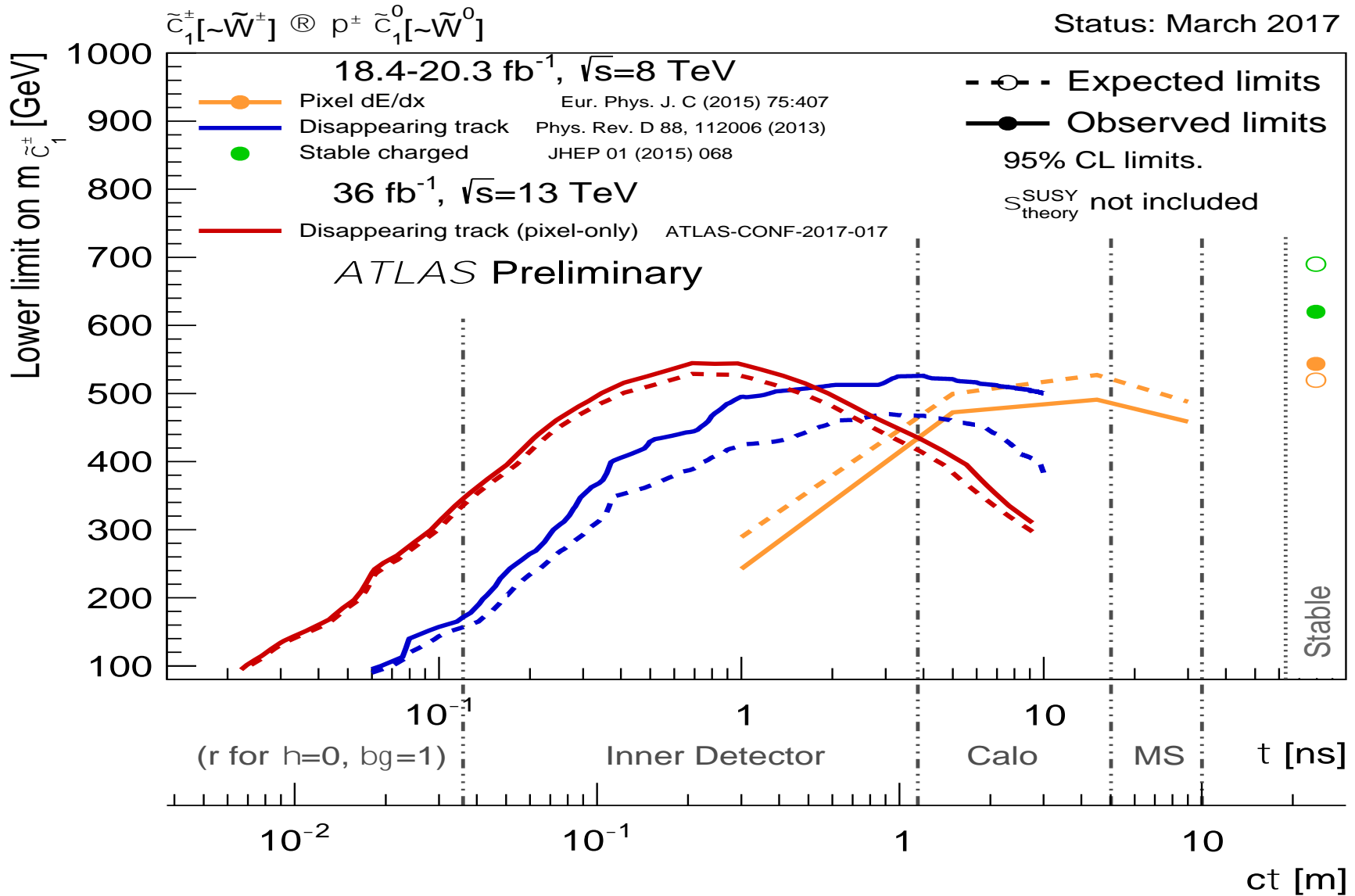


- (2) “in-flight decay” non-colored LLCP



$(\tilde{W}^\pm \text{ as the LLCP})$

Status: March 2017



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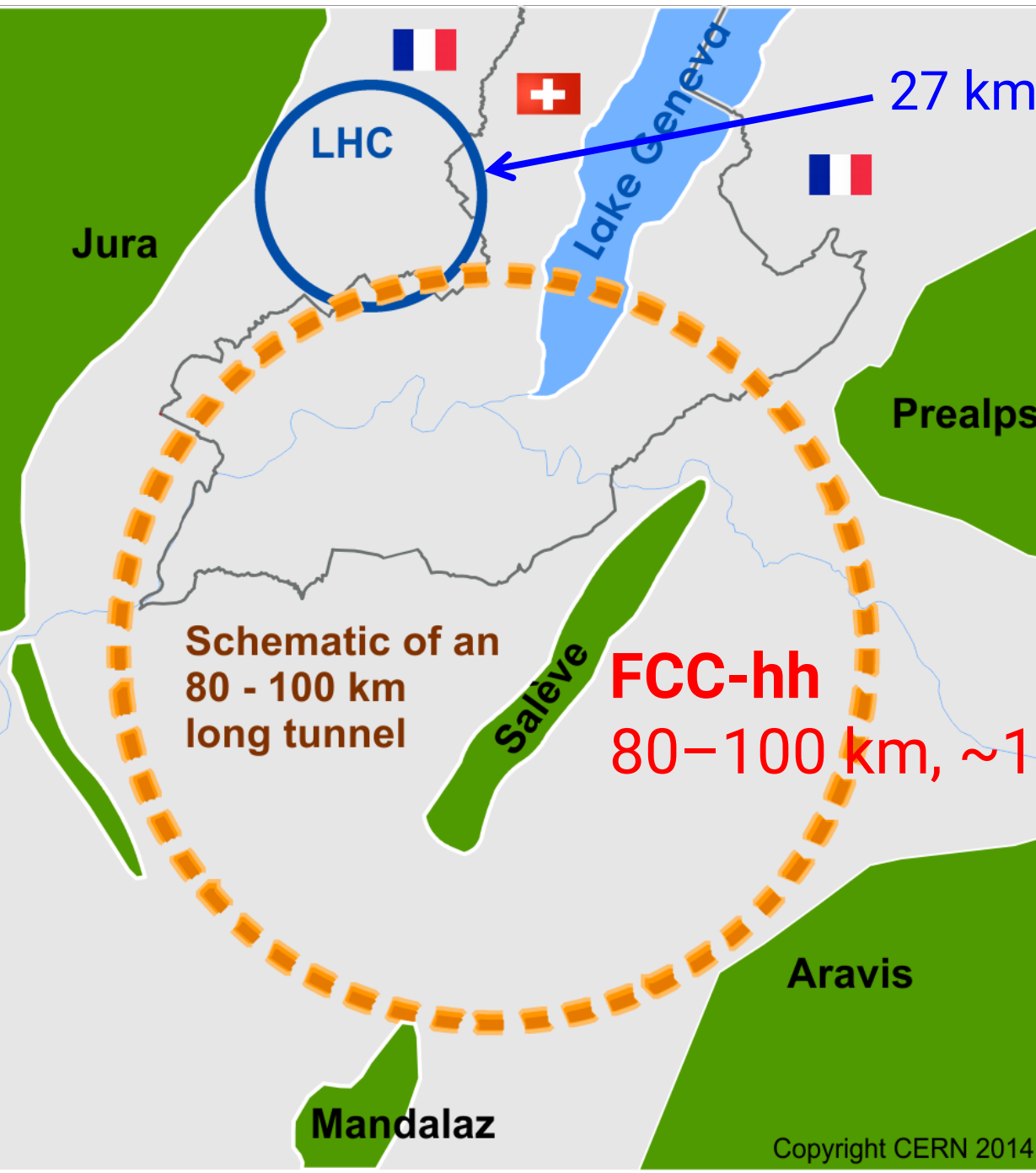
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(collected in FCC-hh report [[1606.00947](#)])

he: Kechen Wang, SI, Monica D'Onofrio, Georges Azuelos [17???.?????]

(subgroup in BSM@ep collaboration)

FCC-hh



27 km, 8.3 T, beam = 7 TeV

* **HE-LHC**
27 km, 16 T
beam = 13.5 TeV

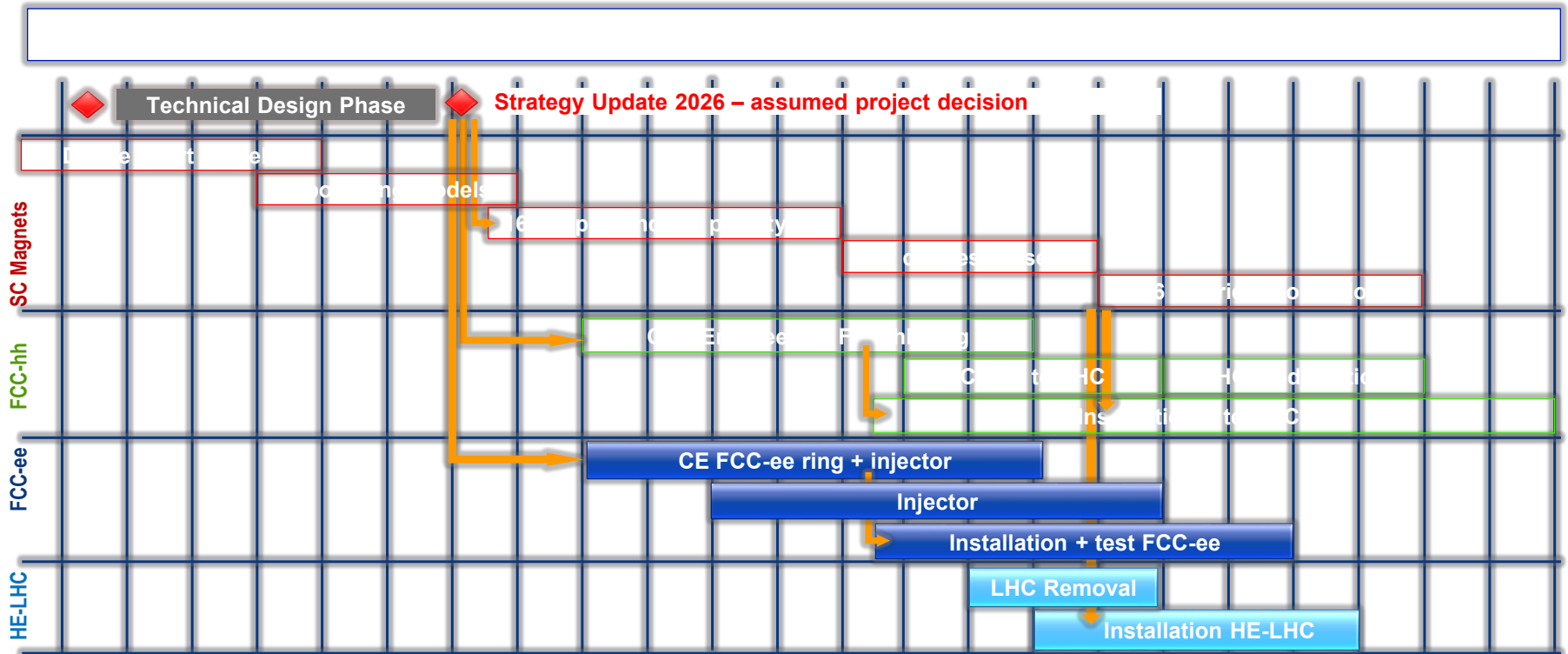
FCC-hh
80-100 km, ~16 T, beam = 50 TeV
($\int \mathcal{L} = 5-10 \text{ ab}^{-1}$)

(FCC-ee as a potential first step)

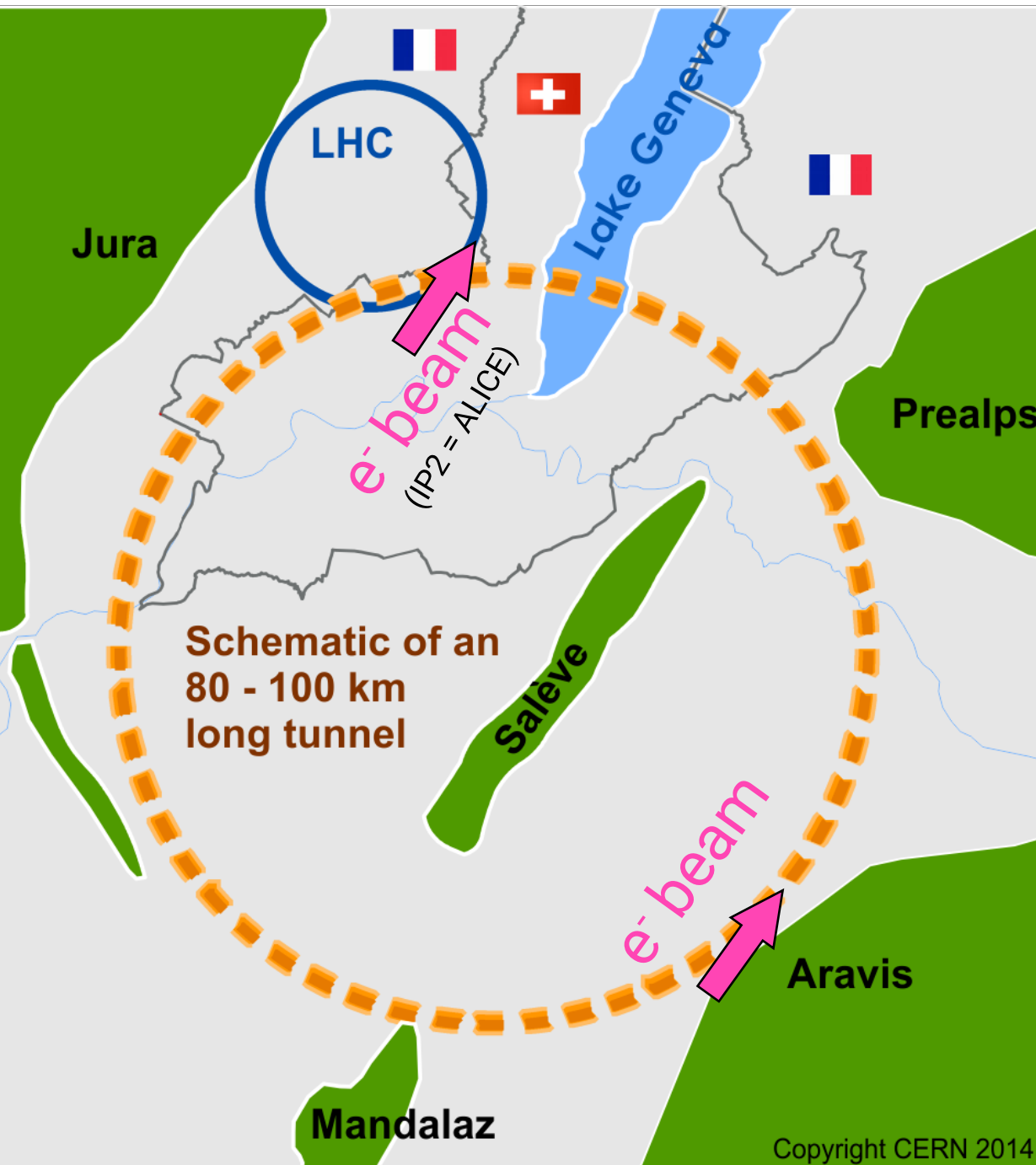
→ CDR in 2018
[conceptual design report]



Draft Schedule Considerations



FCC-he

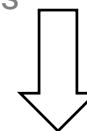


LHC + e⁻ beam = LHeC

FCC-hh + e⁻ beam
= FCC-he

$E_e = 60 \text{ GeV}$ (140 GeV?)

from "Energy Recovery
Linacs"



$\sqrt{s_{\text{LHeC}}} = 2 \times 0.65 \text{ TeV}$

$\sqrt{s_{\text{FCC-he}}} = 2 \times 1.73 \text{ TeV}$

($\int \mathcal{L}_{\text{FCC-he}} \sim 1 \text{ ab}^{-1}$)

- Mar 2015 : FCC week 2015 @ Washington D.C.
- Apr 2016 : FCC week 2016 @ Rome
- Jan 2017 : FCC physics workshop @ CERN
- May 2017 : FCC week 2017 @ Berlin
- Sep 2017 : LHeC/FCC-eh workshop @ CERN
- Jan 2018 : FCC physics workshop @ CERN
- Apr 2018 : FCC week 2018 @ Amsterdam

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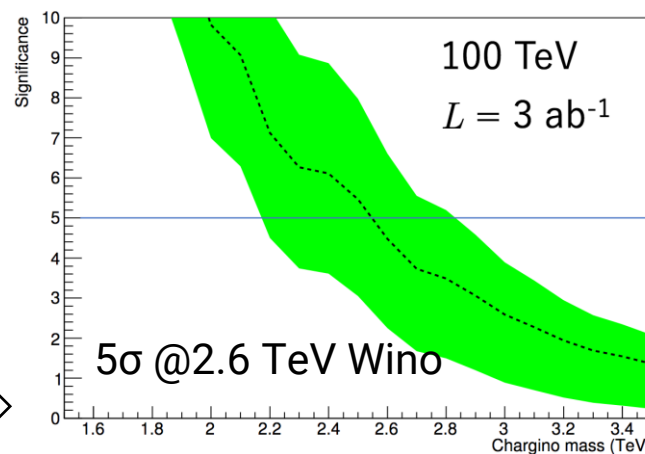
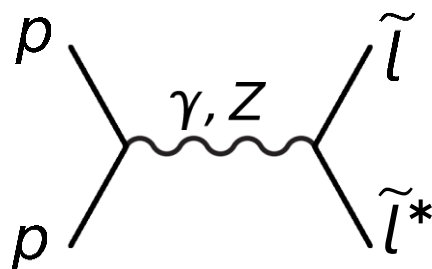
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■ LLCPs at FCC-hh \approx LLCPs at LHC

➤ same production mechanism; just with a higher energy.

- e.g., $\tilde{l} \rightarrow$ Drell–Yan process (or from cascade decay)



[R. Sawada's talk in FCC week 2017]

➤ same detection method.

- “in-flight decay” \rightarrow disappearing track.
- “stable” $\tilde{l} \rightarrow$ muon-like track + β measurement (heavy = slow)

- dE/dx (ionization energy loss)
- $\Delta L/\Delta t$ (time-of-flight)

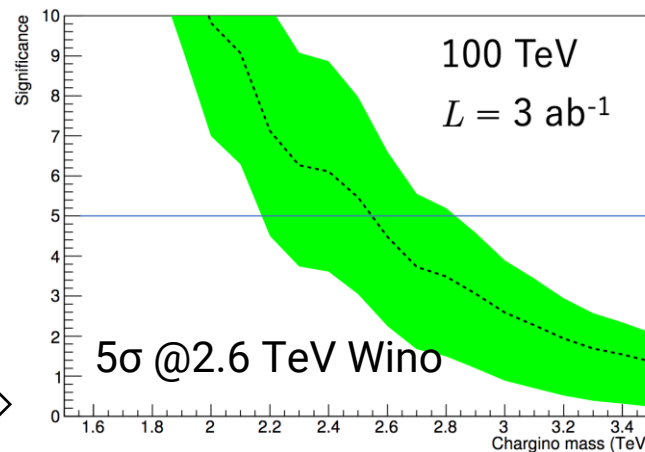
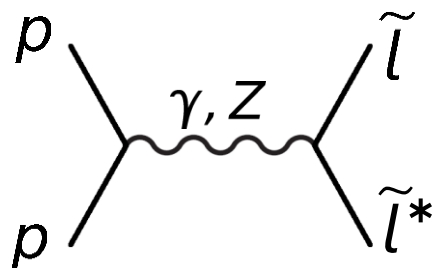
➤ Two extras:

- muon radiative energy loss
- LLCP momentum resolutions

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- **muon radiative energy loss**
- **LLCP momentum resolutions**

Muon energy loss in matter

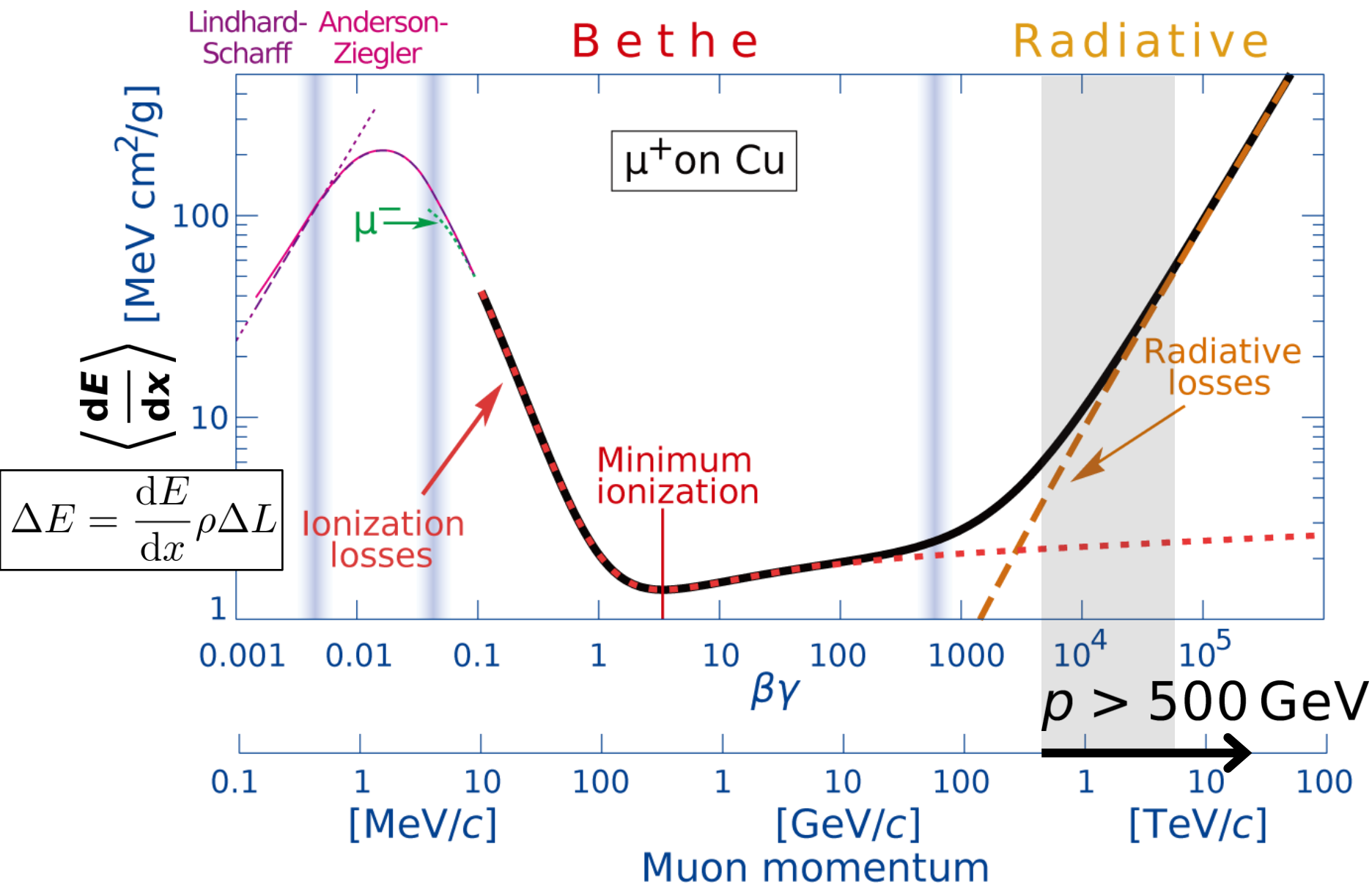
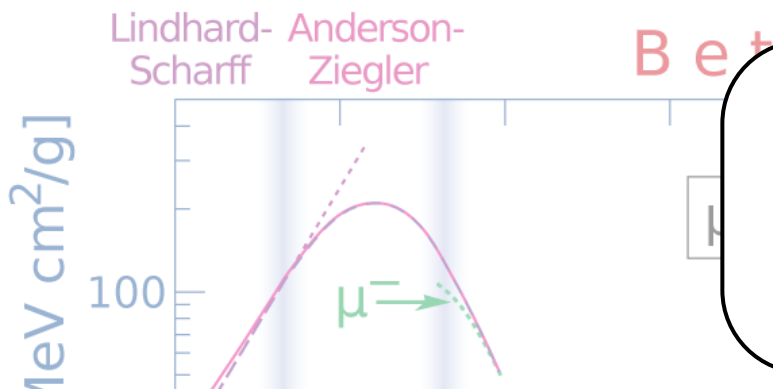


Figure from Groom, Mokhov, Striganov, Atom. Nucl. Data Tab. **78** (2001) 183-356
 [also in PDG Review "Passage of particles through matter"]

Muon energy loss in matter



Muon radiative energy loss

- Bremsstrahlung
- Photonuclear interaction
- $e^+ - e^-$ pair-production

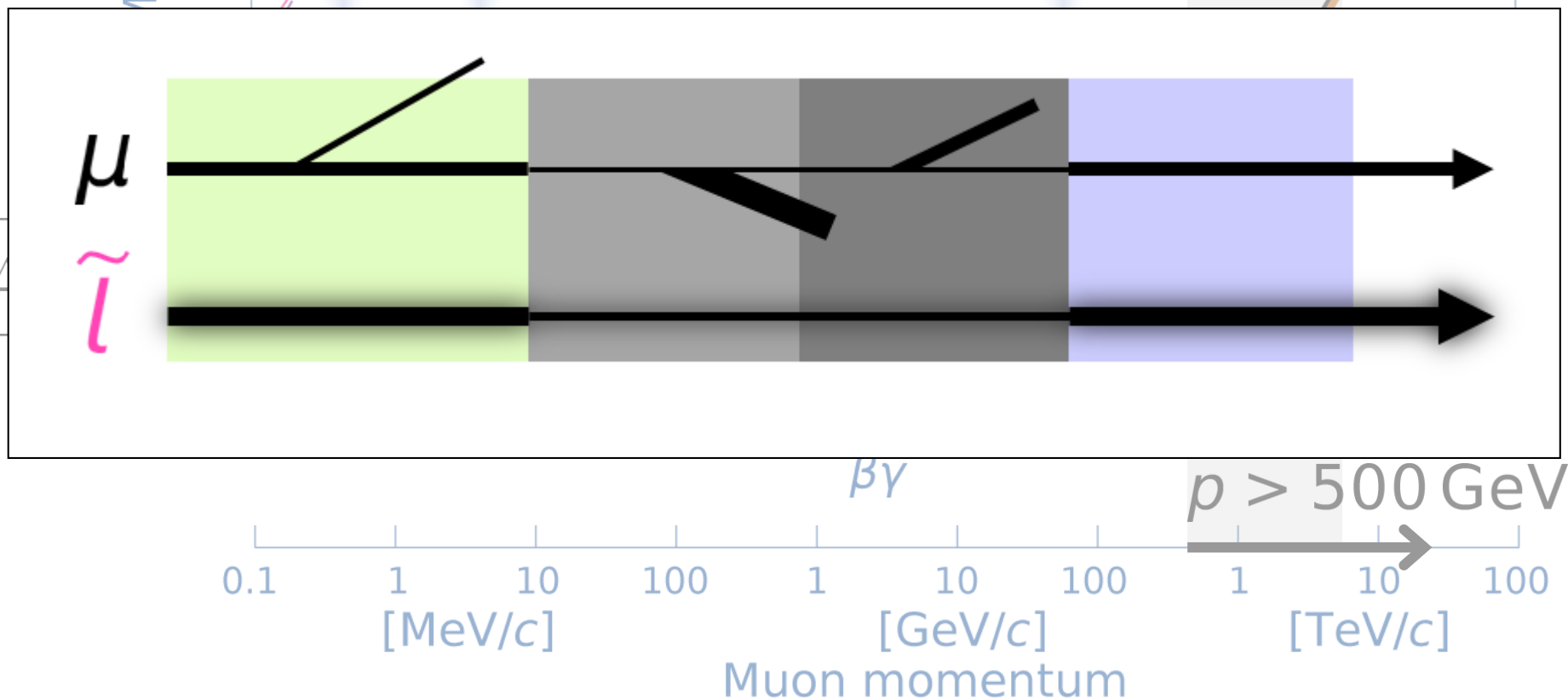
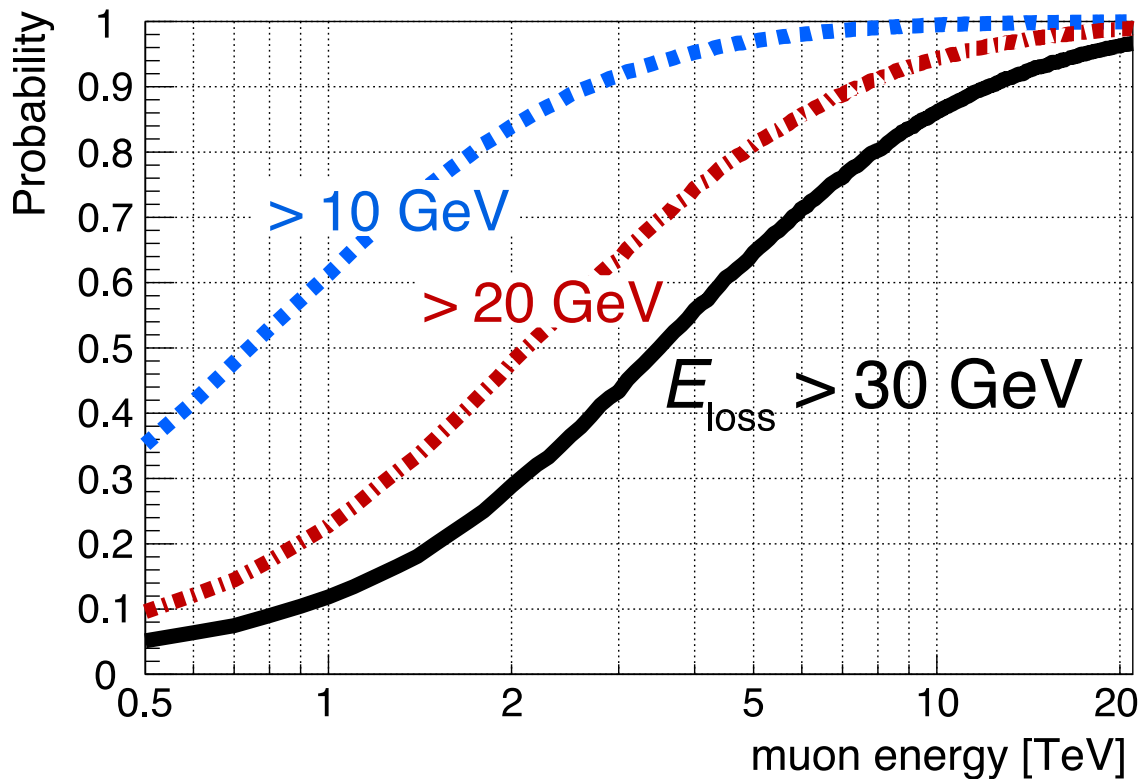
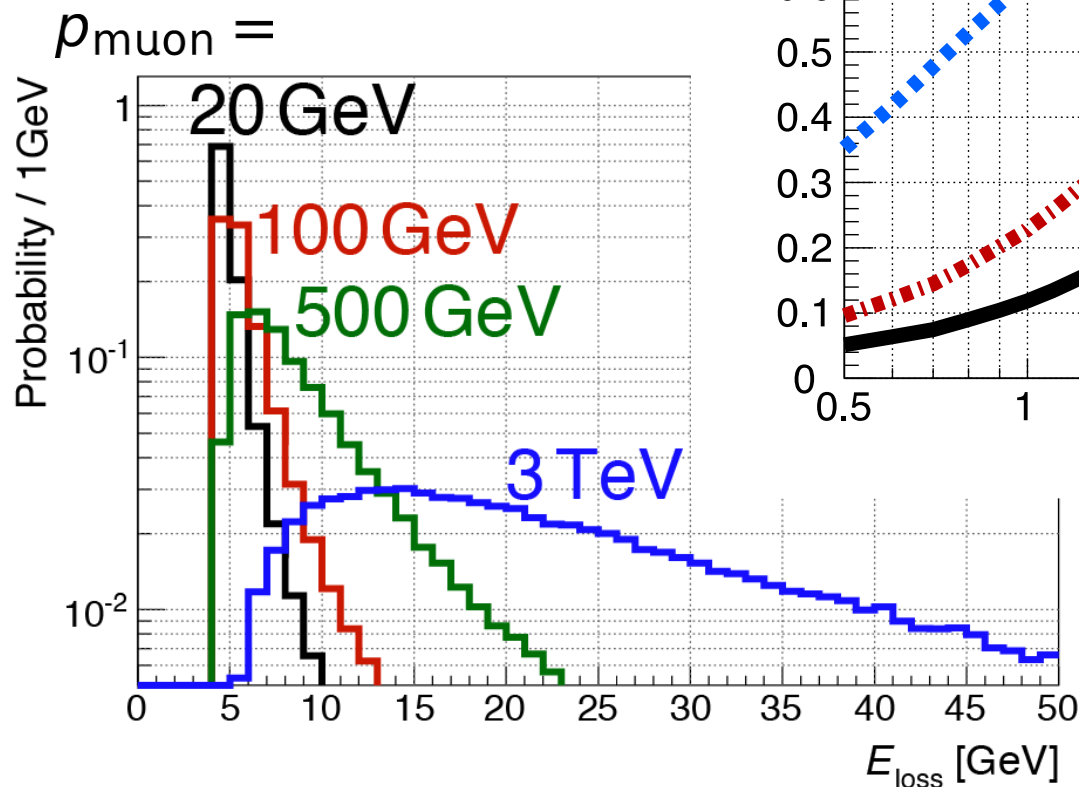


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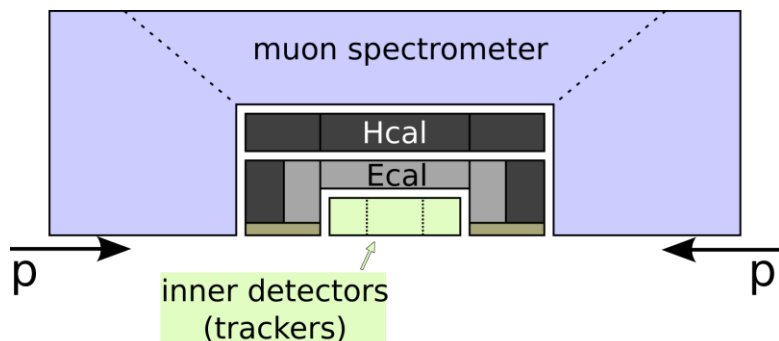
"calorimeter": approximated by iron (Fe) with 3m thickness.

→ some of μ ($P_T > 500$ GeV): > 30 GeV energy deposit.



[Simulated with GEANT 4]

- Detector
 - similar to ATLAS/CMS
 - β -resolution same as ATLAS (resolution: 2.4%)
- Signal: Madgraph5 + Pythia6 + Delphes3 (calculated at the LO)
- BKG: “Snowmass 2013”
BKG set for 100TeV
- Pile-up not considered



■ \tilde{l} -selection flow

\tilde{l} = reconstructed “muon” with

- $P_T > 500$ GeV
- $|\eta| < 2.4$
- $0.4 < \hat{\beta} < 0.95$ (from ToF)
- $E_{\text{loss}} < 30$ GeV

■ Event selection

- two \tilde{l} -candidates

LLCP selection flow ($\int L = 1 \text{ ab}^{-1}$)

	signal		SM BKG
	$\tilde{l} = 1 \text{ TeV}$	3 TeV	
total	2570	31.8	—
p_T & η	1840	28.5	9.19×10^6
β	1230	24.6	3.41×10^5
E_{loss}	1230	24.6	2.78×10^5
$\epsilon_{\text{acc}}\epsilon_{\text{eff}}$	48%	77%	—

E_{loss} reduces **34%** of BKG
 ($\because 0.82^2 = 0.66$)

$\times 0.82$

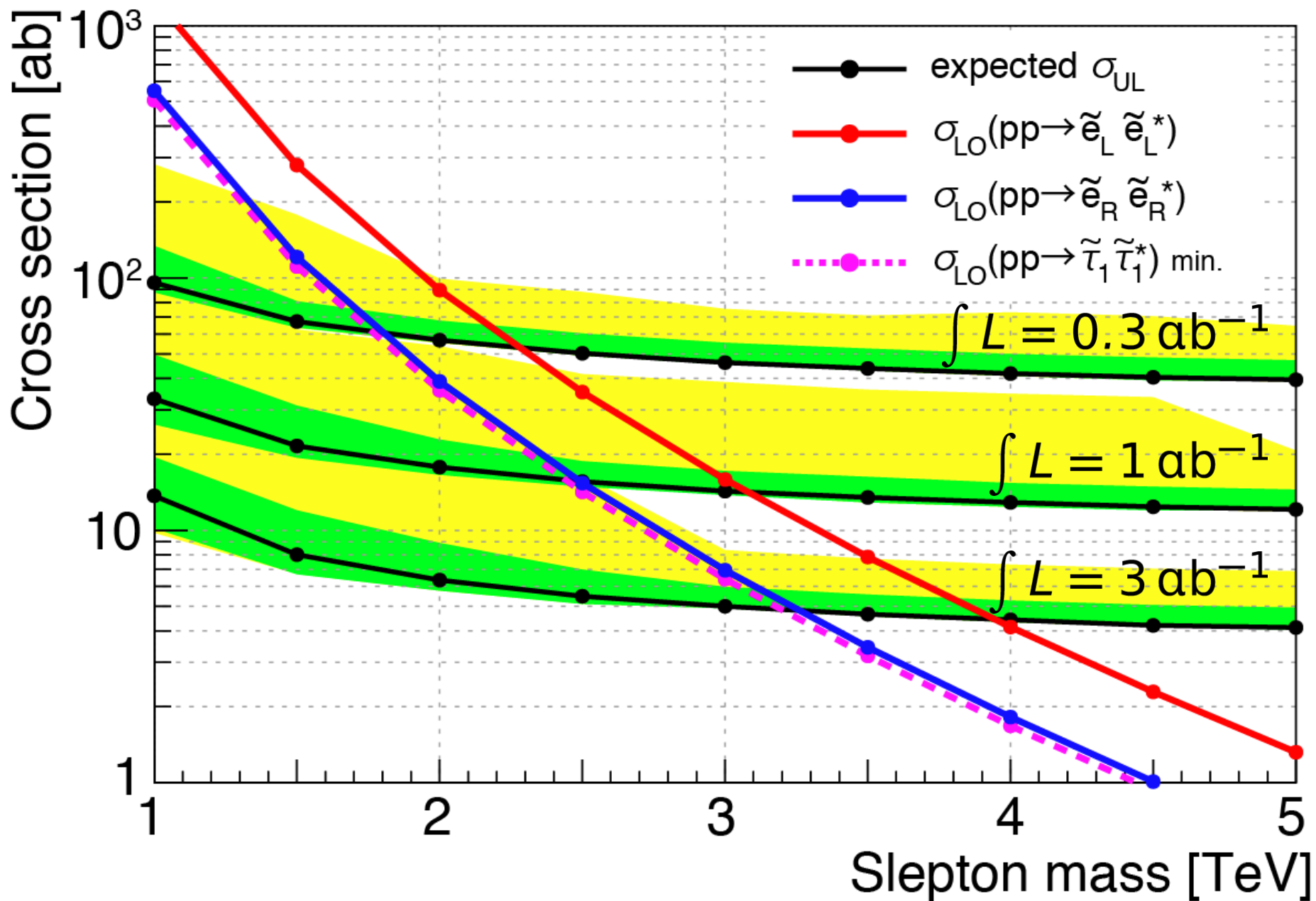
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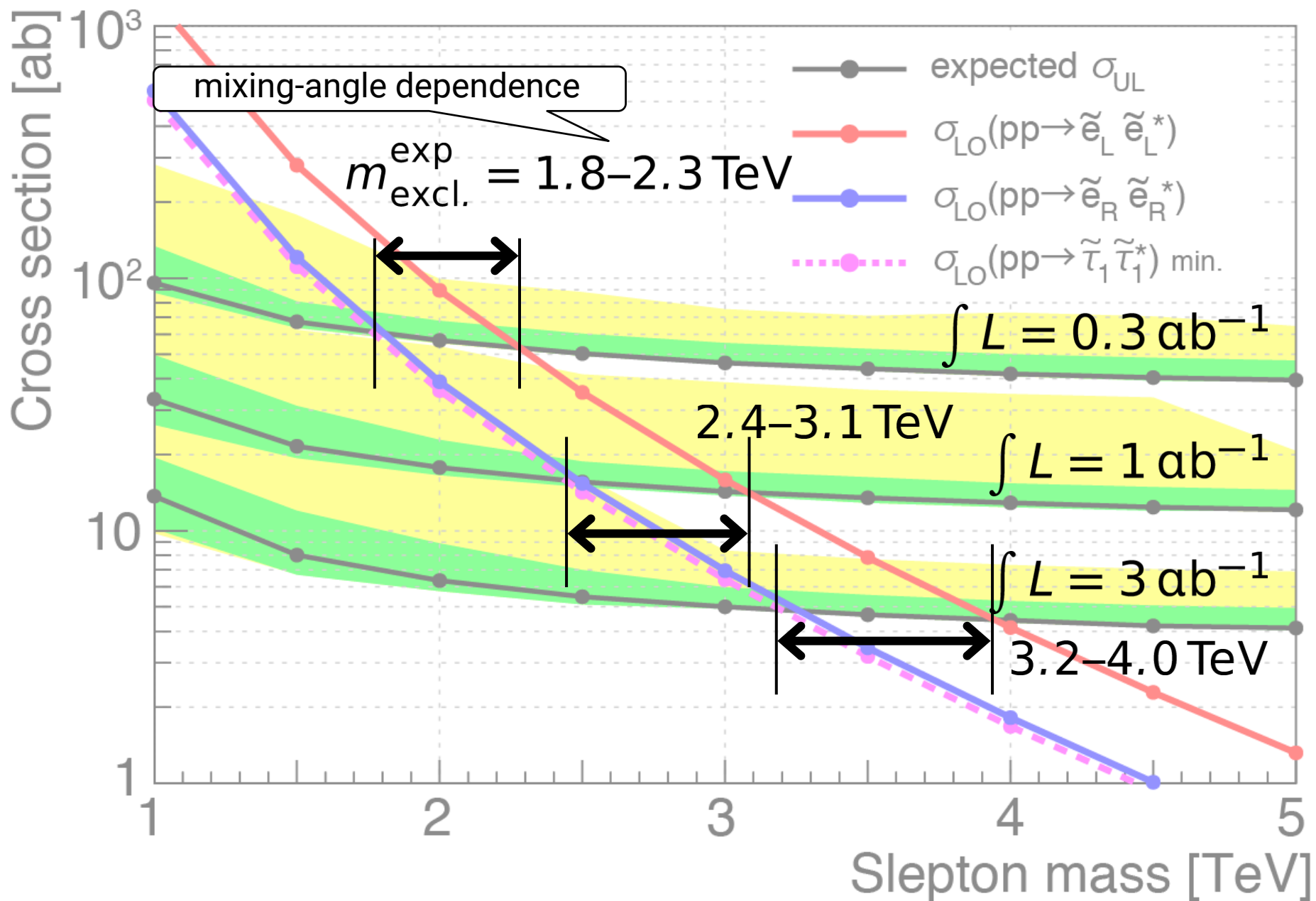
Event categorization ($\int L = 1 \text{ ab}^{-1}$)

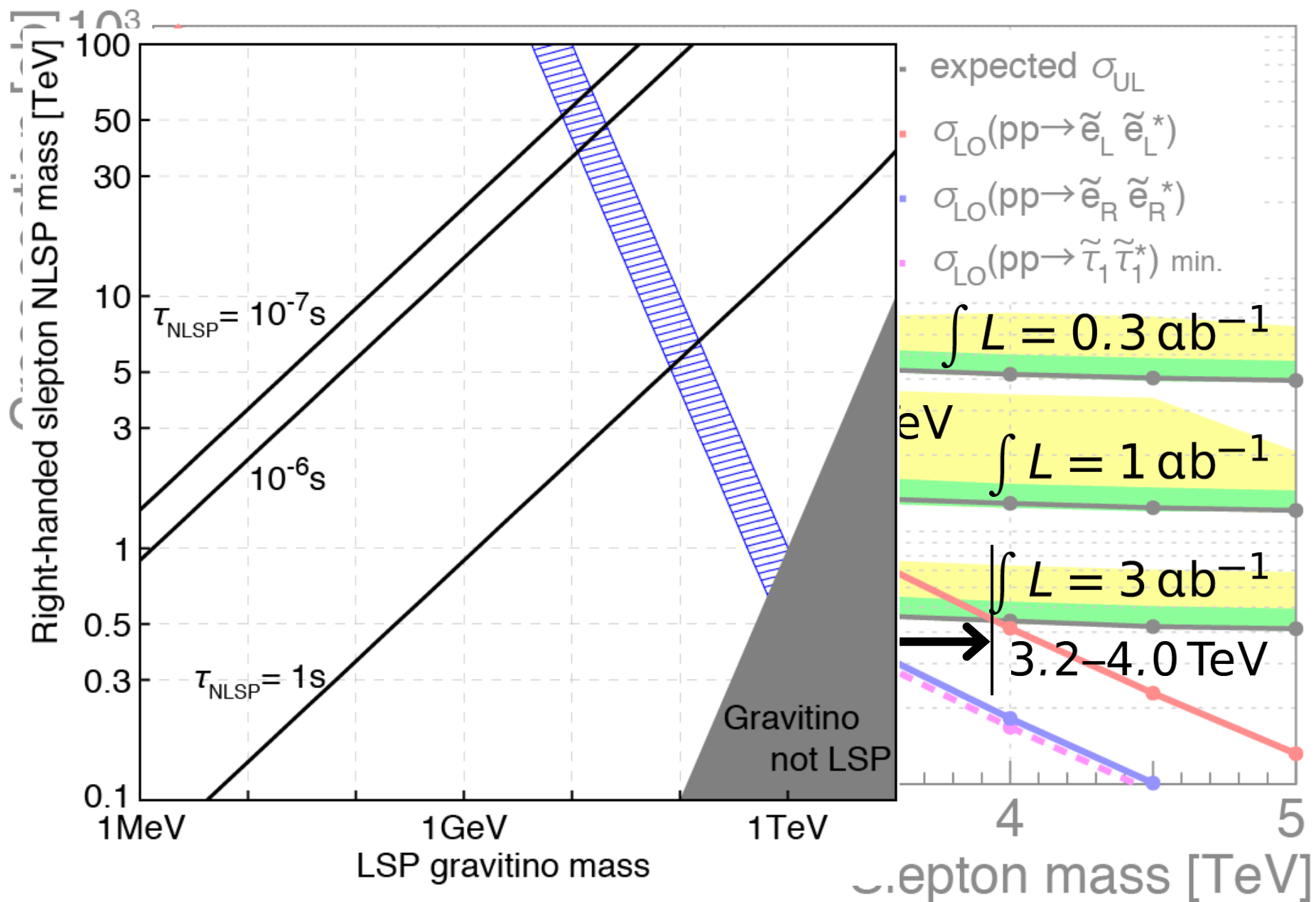
	1 TeV	3 TeV	BKG
$N_{\text{LLCP}} = 0$	483	1.34	(a lot)
$N_{\text{LLCP}} = 1$	378	4.46	2.78×10^5
$N_{\text{LLCP}} = 2$	424	10.1	34.6

SR

- Event selection
 - two \tilde{l} -candidates



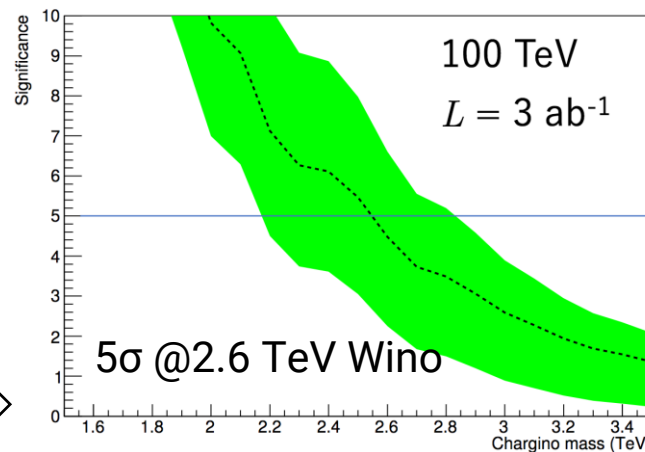
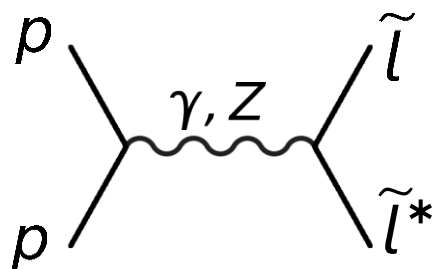




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➤ same production mechanism; just with a higher energy.

- e.g., $\tilde{l} \rightarrow$ Drell–Yan process (or from cascade decay)



[Ryu Sawada's talk in FCC week 2017]

➤ same detection method.

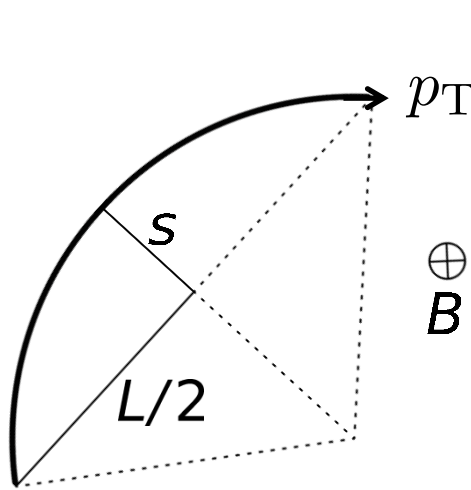
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- “stable” $\tilde{l} \rightarrow$ muon-like track + β measurement (heavy = slow)

- dE/dx (ionization energy loss)
- $\Delta L/\Delta t$ (time-of-flight)

➤ **Two extras:**

- **muon radiative energy loss**

● **LLCP momentum resolutions**



$$p_T = \frac{L^2}{8s} |q| B \implies \Delta p_T = \frac{L^2 |q| B}{8} \frac{\Delta s}{s^2}$$

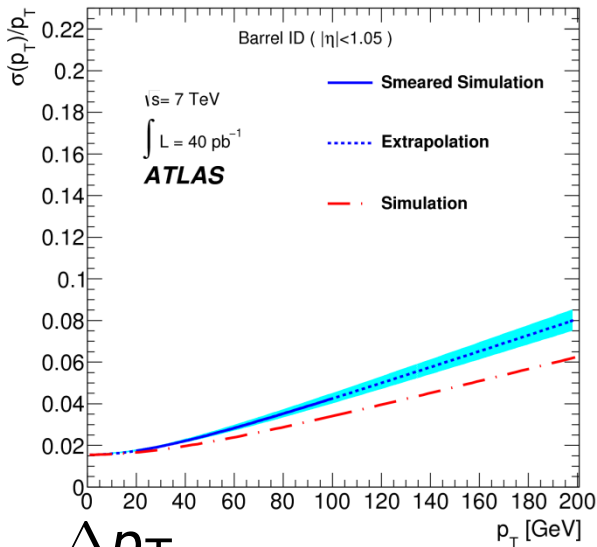
$$= \frac{8\Delta s}{L^2 |q| B} \cdot p_T^2$$

$$\therefore \Delta p_T \propto p_T^2$$

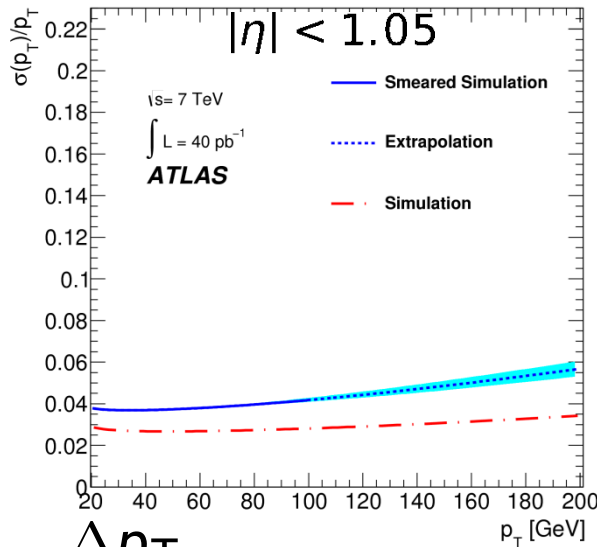
ATLAS 7 TeV results on muon momentum resolution

Inner Detector, $|\eta| < 1.05$

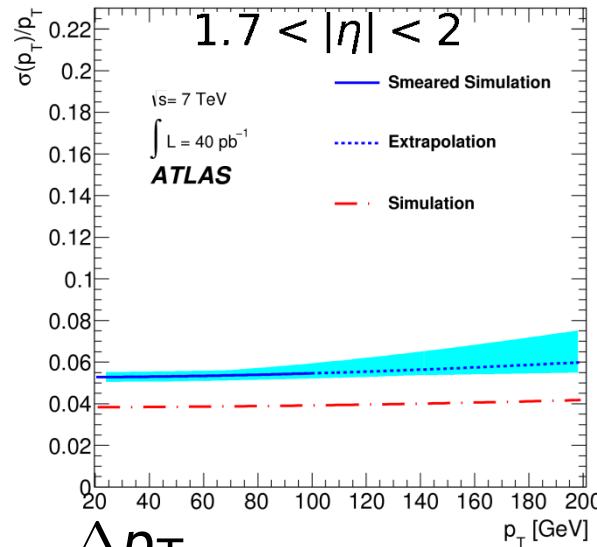
Muon Spectrometer



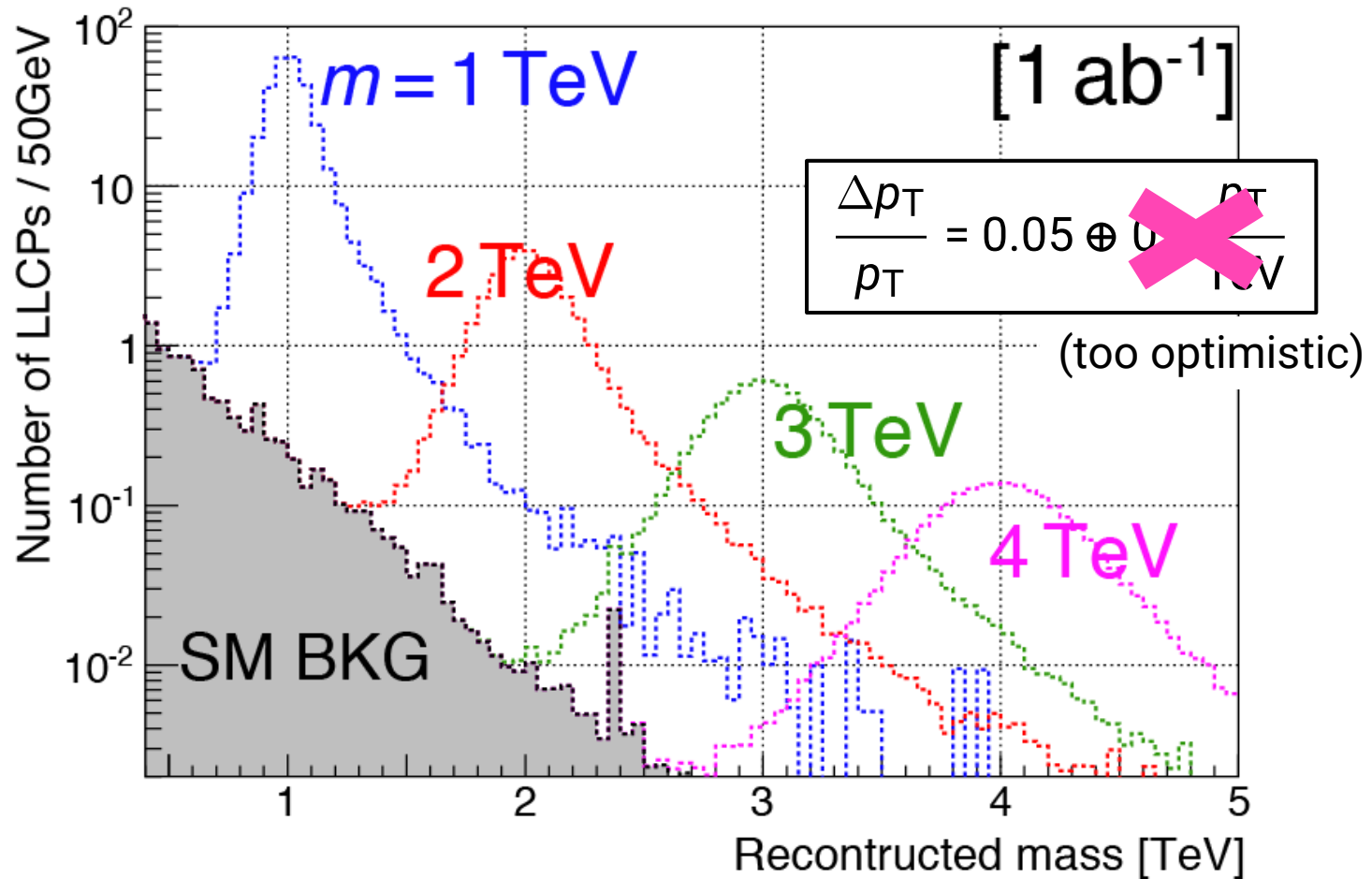
$$\frac{\Delta p_T}{p_T} \sim 0.38 p_T / \text{TeV}$$



$$\frac{\Delta p_T}{p_T} \sim 0.14 p_T / \text{TeV}$$



$$\frac{\Delta p_T}{p_T} \sim 0.06 p_T / \text{TeV}$$



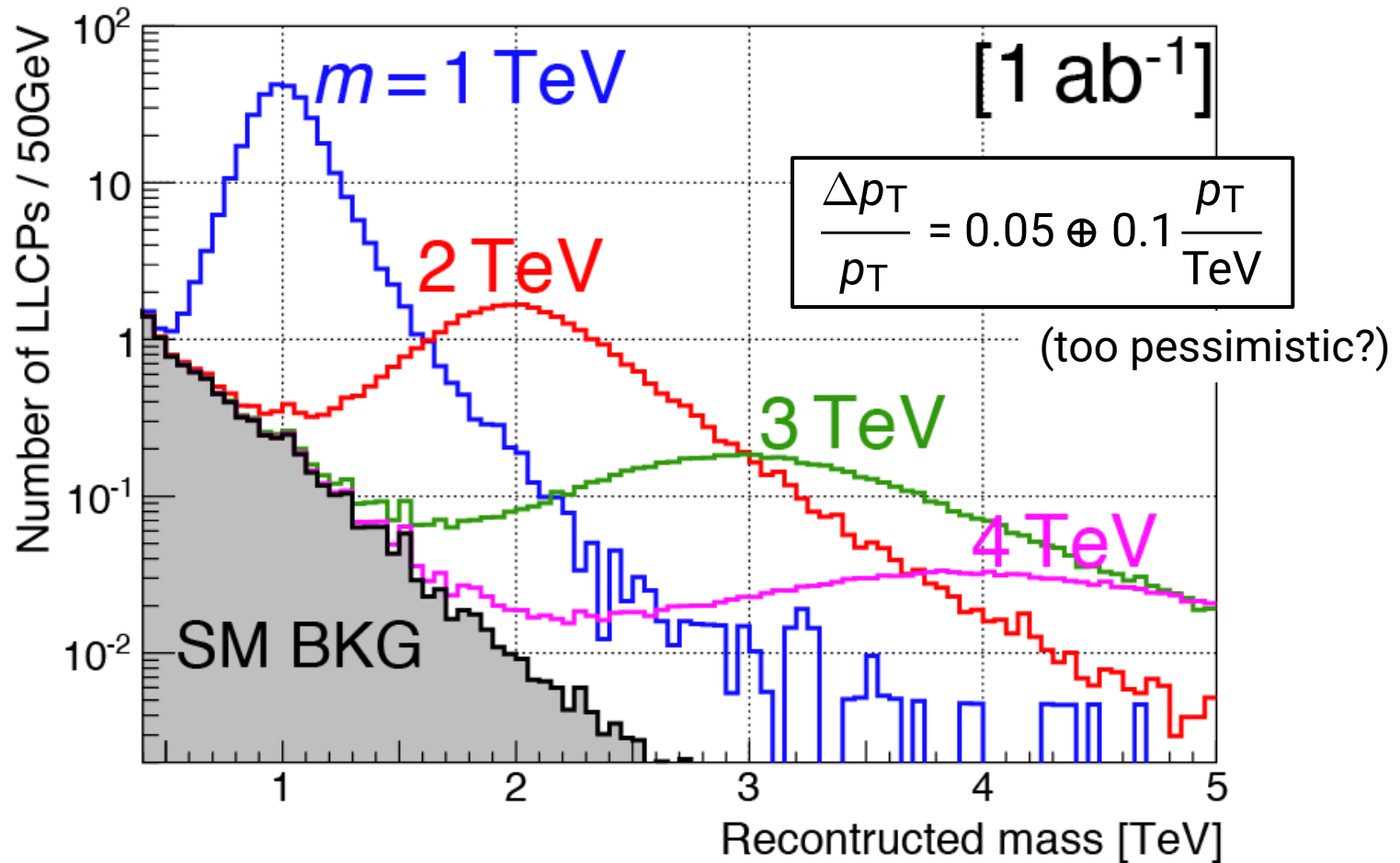
FCC-hh trk. goal: 10–20% @ 10 TeV

([Michele Selvaggi's talk](#) in FCC physics workshop)

cf. ATLAS 7 TeV commissioning:

(ID-barrel, MS-barrel, MS-extbarrel) = (38%, 14%, 6%) @ 1 TeV

$$\left(= \frac{p}{\beta\gamma} = \frac{p_T \cosh \eta}{\beta\gamma} \right)$$



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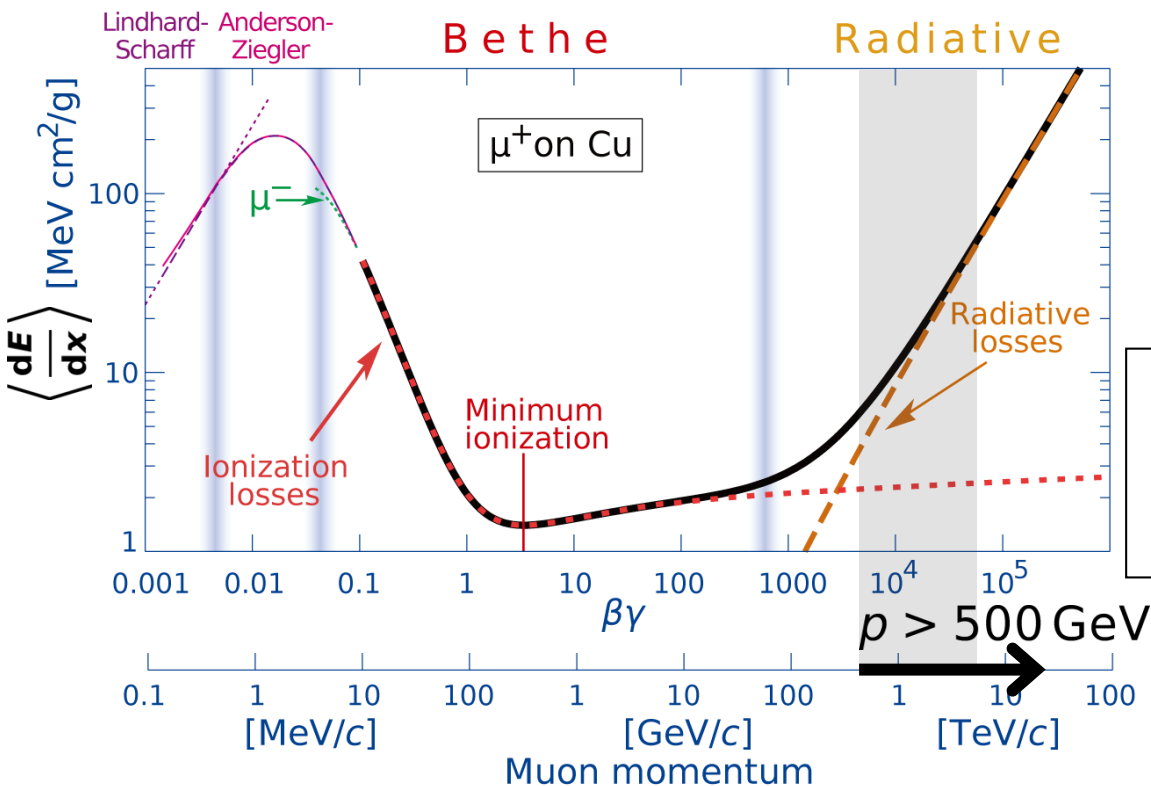
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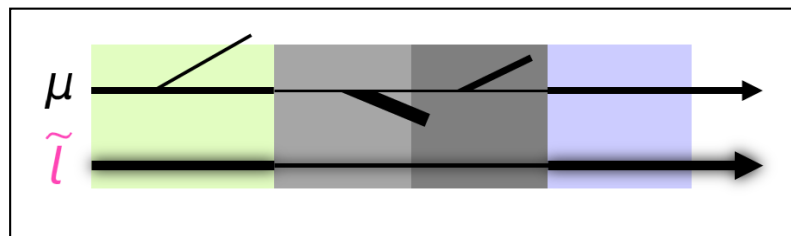
- 100 TeV FCC-hh
mass reach
(Drell-Yan \tilde{l} or $\tilde{\tau}$)

	0.3ab ⁻¹	1ab ⁻¹	3ab ⁻¹	
Exclusion	1.8–2.3	2.4–3.1	3.2–4.0	
Discovery	1.6–2.2	2.3–3.0	3.1–4.0	in TeV

“Muon radiative energy loss”



- Bremsstrahlung
- Photonuclear interaction
- e⁺-e⁻ pair-production



→ 34% of BKG reduction

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Based on

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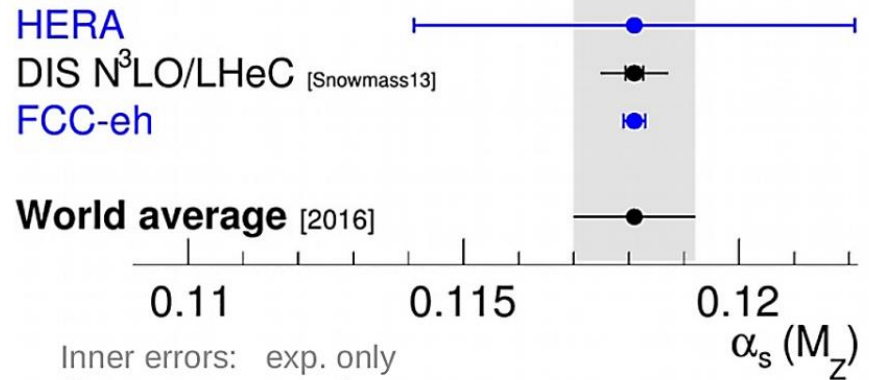
(collected in FCC-hh report [[1606.00947](#)])

he: Kechen Wang, SI, Monica D'Onofrio, Georges Azuelos [17???.?????]

(subgroup in BSM@ep collaboration)

■ FCC-he targets:

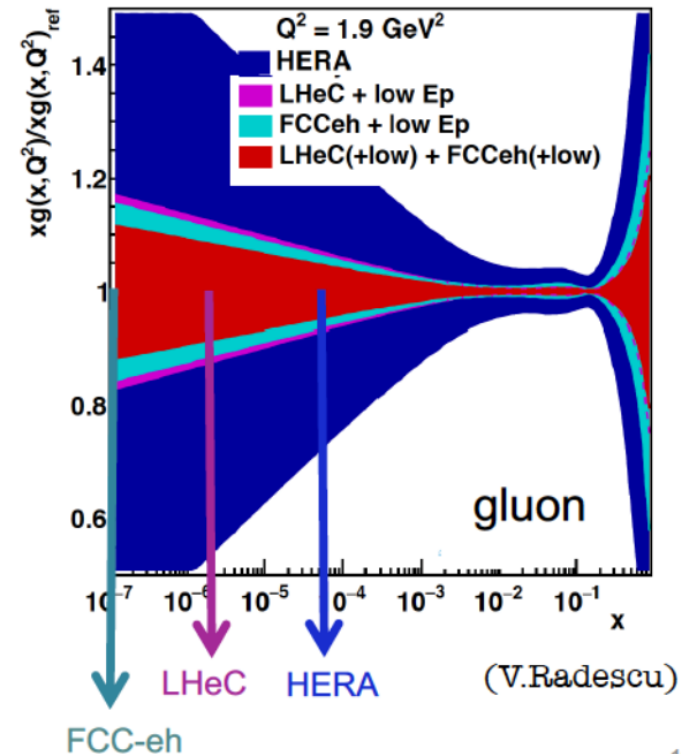
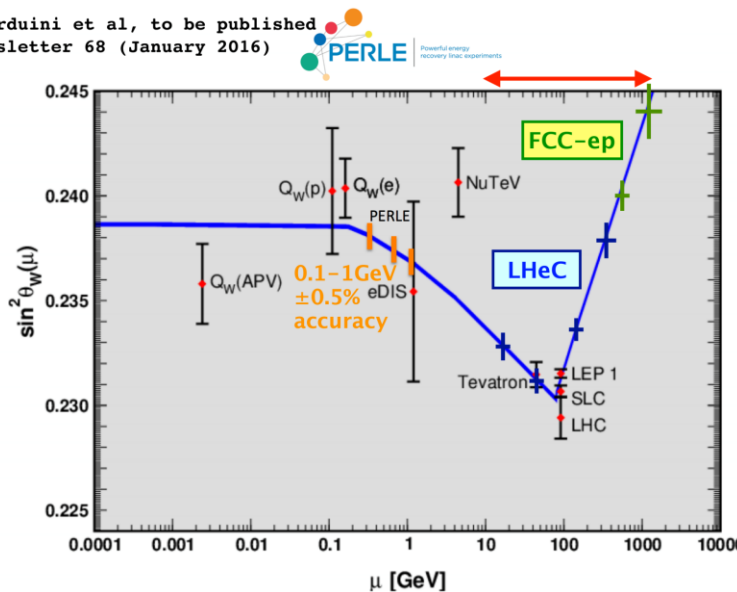
- PDFs & strong coupling
- Higgs & Electroweak physics
- QCD (heavy quark PDFs)
- low-x physics (non-linear QCD?) : $x < 10^{-6}$



Inner errors: exp. only
Outer errors: exp+theo.

Jan Kretzschmar, 11.9.2017

PERLE CDR, Arduini et al, to be published
ICFA BeamNewsletter 68 (January 2016)

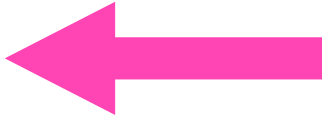


■ What's MORE?

Any power to New Physics? → BSM ep team

■ BSM ep team

★ Direct Searches

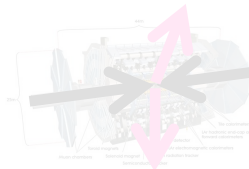
- ◆ Leptoquarks: limits, quantum # & couplings
- ◆ Contact interactions: $eeqq$
- ◆ Anomalous gauge couplings: vvv
- ◆ Vector boson scattering
- ◆ BSM in the top sector
- ◆ RPC SUSY: DM, sleptons 
- ◆ RPV SUSY: neutralinos, squarks
- ◆ BSM Higgs: exotic (invisible) decay; H^+ , H^{++}
- ◆ Sterile neutrinos

[from [a talk](#) by Ke Chen Wang @ [FCC week 2017](#)]

■ Dark Matter

➤ co-annihilation

- $(\tilde{B} - \tilde{\tau}) \lesssim 700 \text{ GeV}$
- $(\tilde{W} - \tilde{g}) \lesssim 6\text{--}7 \text{ TeV}$
- $(\tilde{B} - \tilde{g})$ or $(\tilde{B} - \tilde{t}) \lesssim 8 \text{ TeV}$



$\tilde{\tau}$ -LLCP

➤ super-WIMP scenario
→ $\tilde{l} > O(1) \text{ TeV}$

\tilde{l} -LLCP

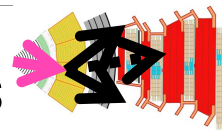
■ Li problem

➤ MSSM $\tilde{\tau}$ with

$m_{\tilde{\tau}} \sim 400 \text{ GeV}$
 $\Delta(m_{\tilde{\tau}} - m_{\text{LSP-DM}}) \sim 100 \text{ MeV}$

$\tilde{\tau}$ -LLCP

■ Simple MSSM models



➤ Pure-Wino dark matter **60mm**

➤ Pure-Higgsino dark matter **7mm**



$\tilde{\chi}^{\pm}$ long-lived because of small δm
 $(\delta m > 0)$

(but **DM underabundant**
for “observable” region)

■ Other MSSM models

➤ Gauge-Mediation (keV \tilde{G})

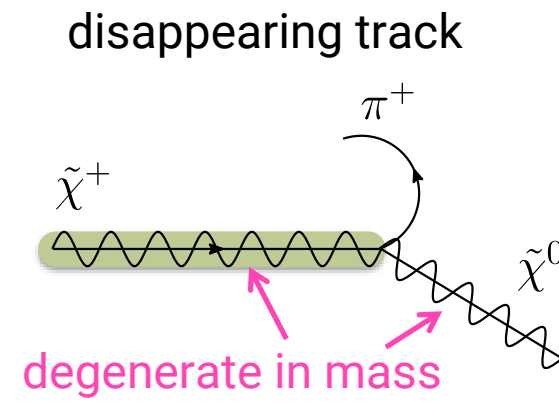
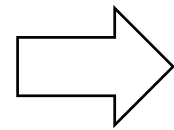
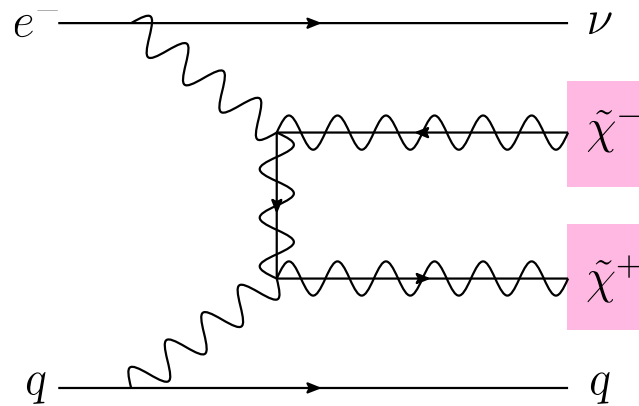
➤ R-parity violation

with \tilde{l} -LSP

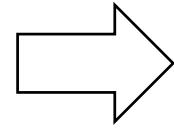
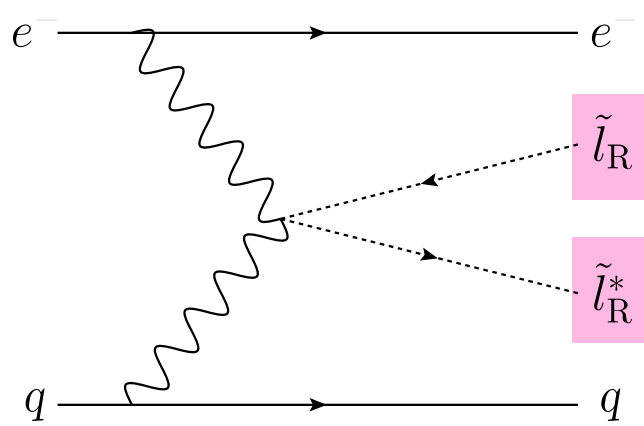
⇒ \tilde{l} long-lived **any $c\tau$**
because of tiny couplings.

■ Simplest models: 4-body production; $\sigma < 1 \text{ fb} \dots (\cdot \omega \cdot \cdot)$

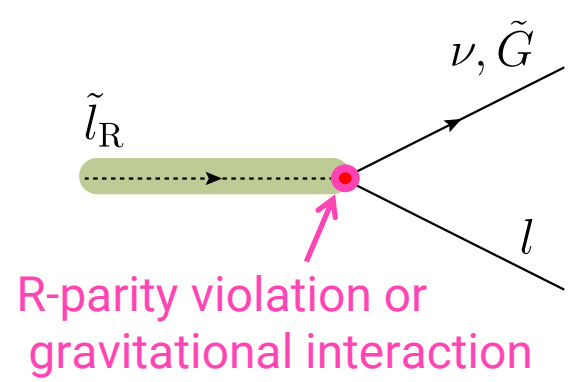
➤ Pure-Wino / Pure-Higgsino LSP



➤ Slepton LSP

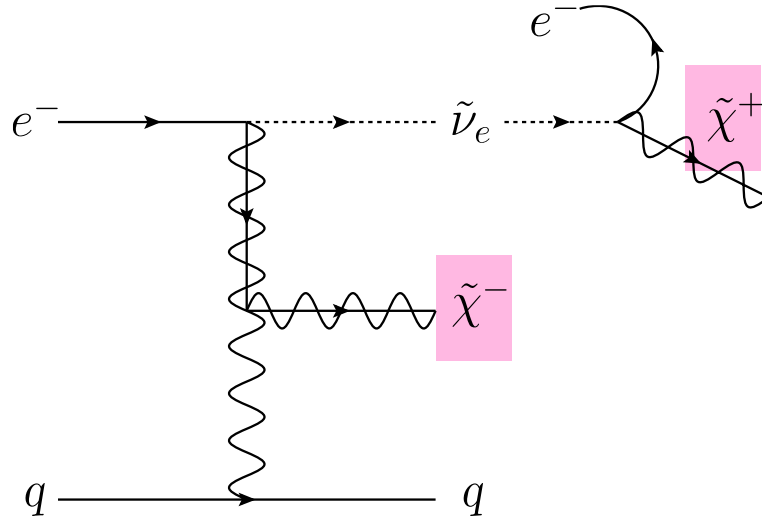


disappearing track (or "kink")

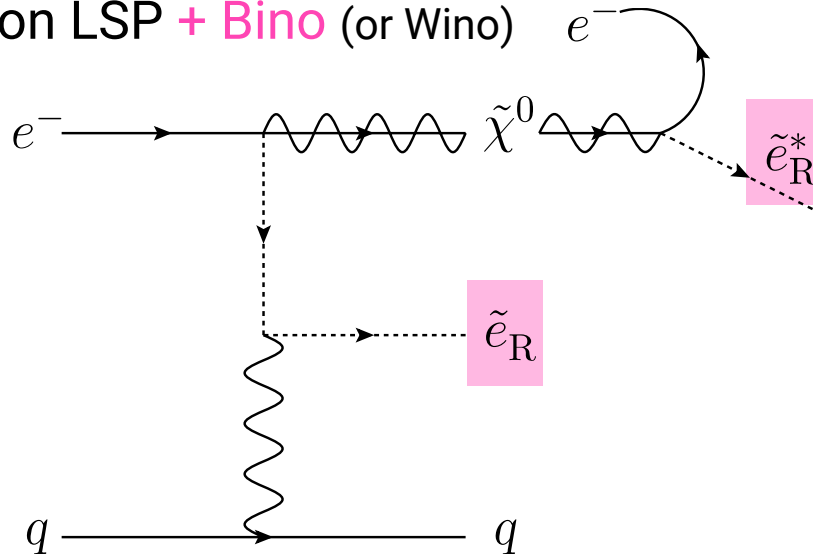


■ Introducing co-LSP allows **3-body** production

- Pure-Wino / ~~Pure-Higgsino~~ LSP + left-handed selectron

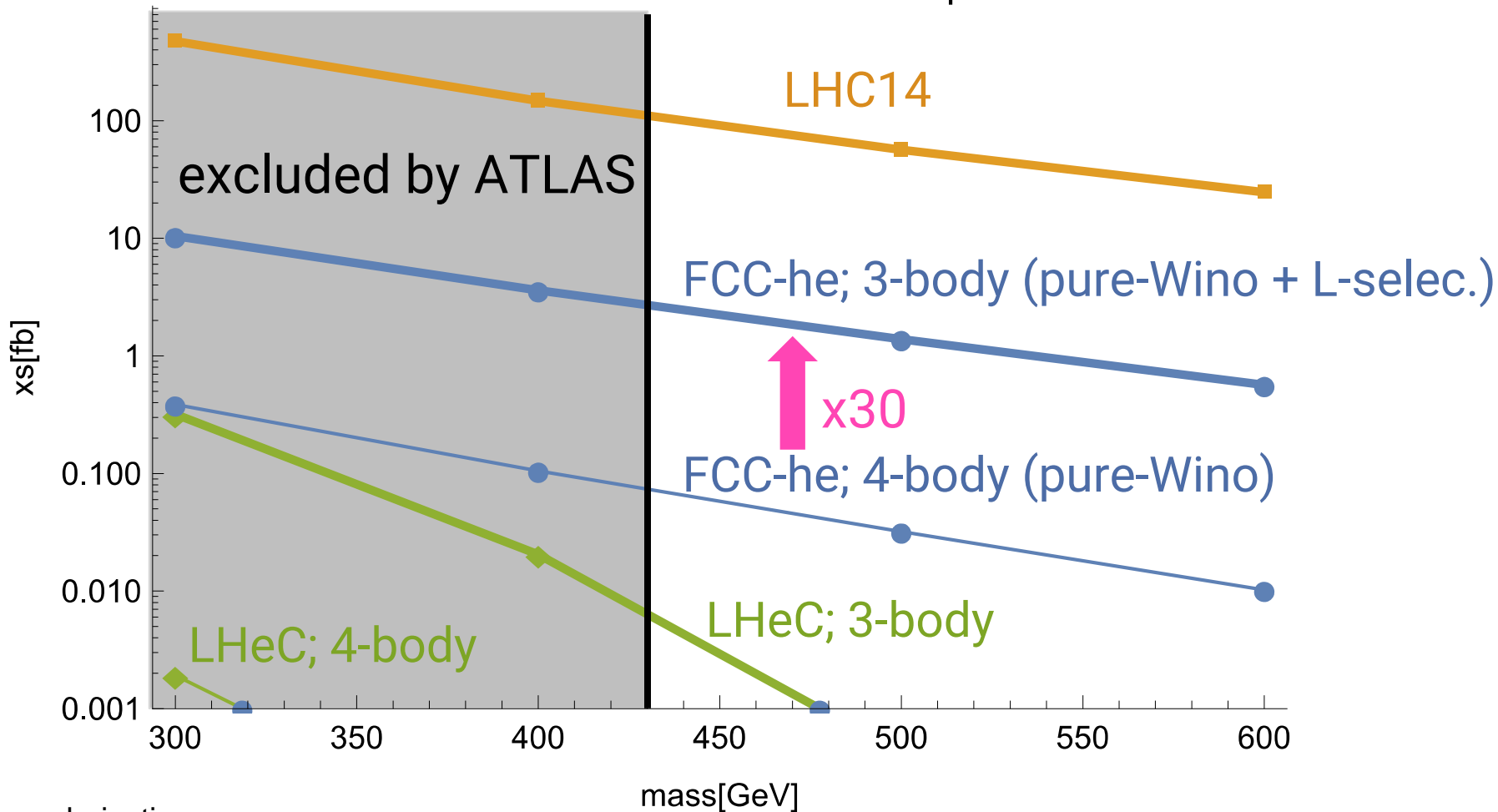


- Slepton LSP + Bino (or Wino)



■ Nominal production cross section (without acceptances / efficiencies)

Wino LSP scenarios with/without co-production



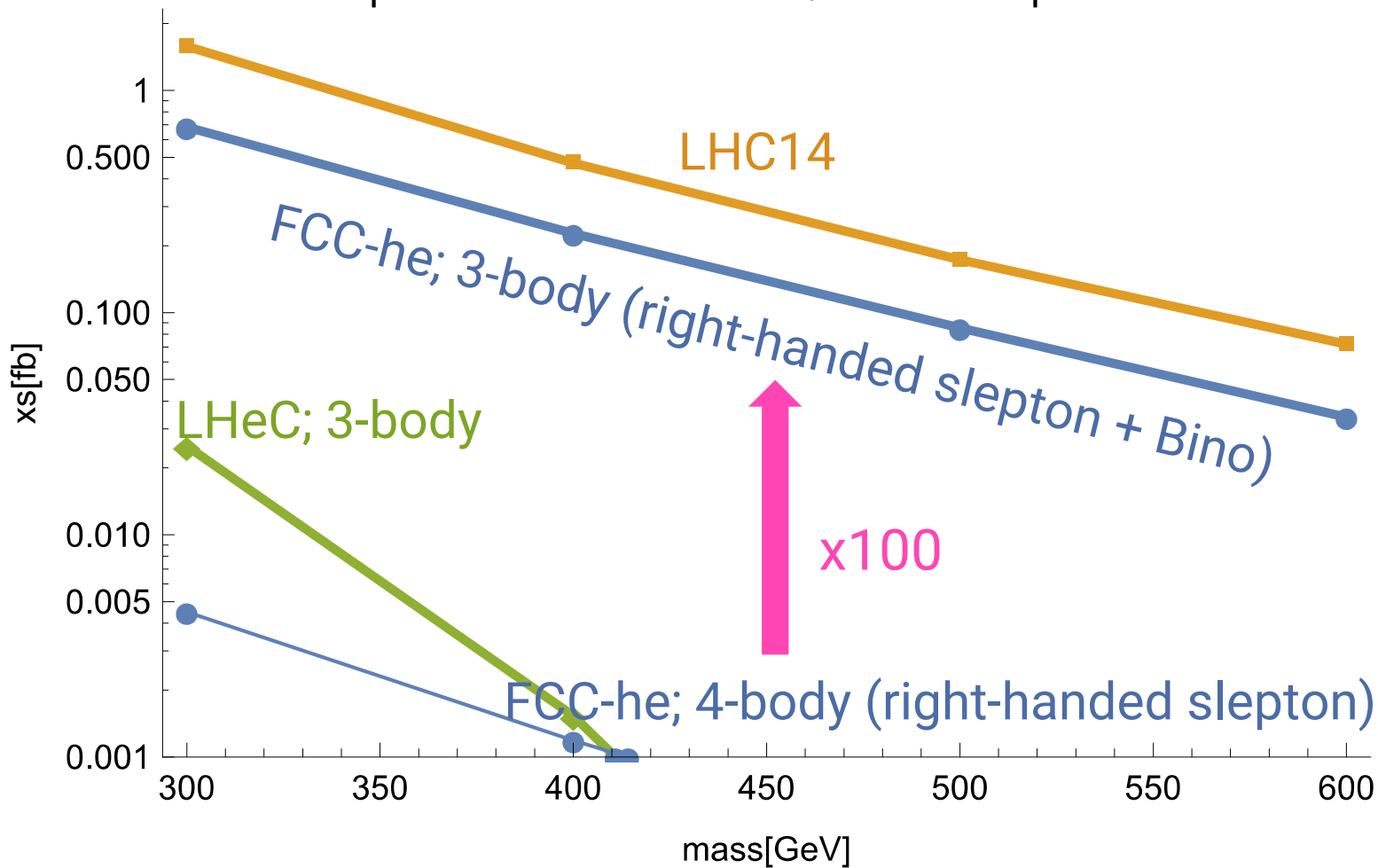
With no polarization.

Shaded region is excluded by ATLAS (13TeV, 36/fb)

"3-body" model assumes $m_{\tilde{e}_L} = m_{\tilde{\chi}_1^0} + 9 \text{ GeV}$

■ Nominal production cross section (without acceptances / efficiencies)

Slepton LSP scenarios with/without co-production

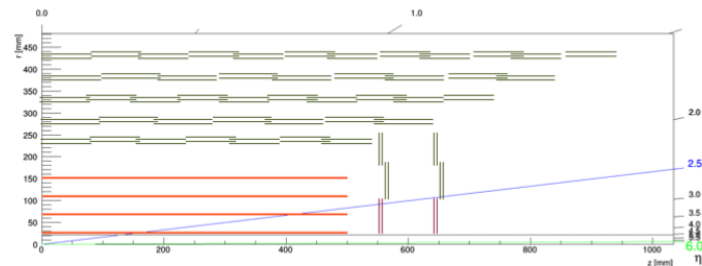


With no polarization.

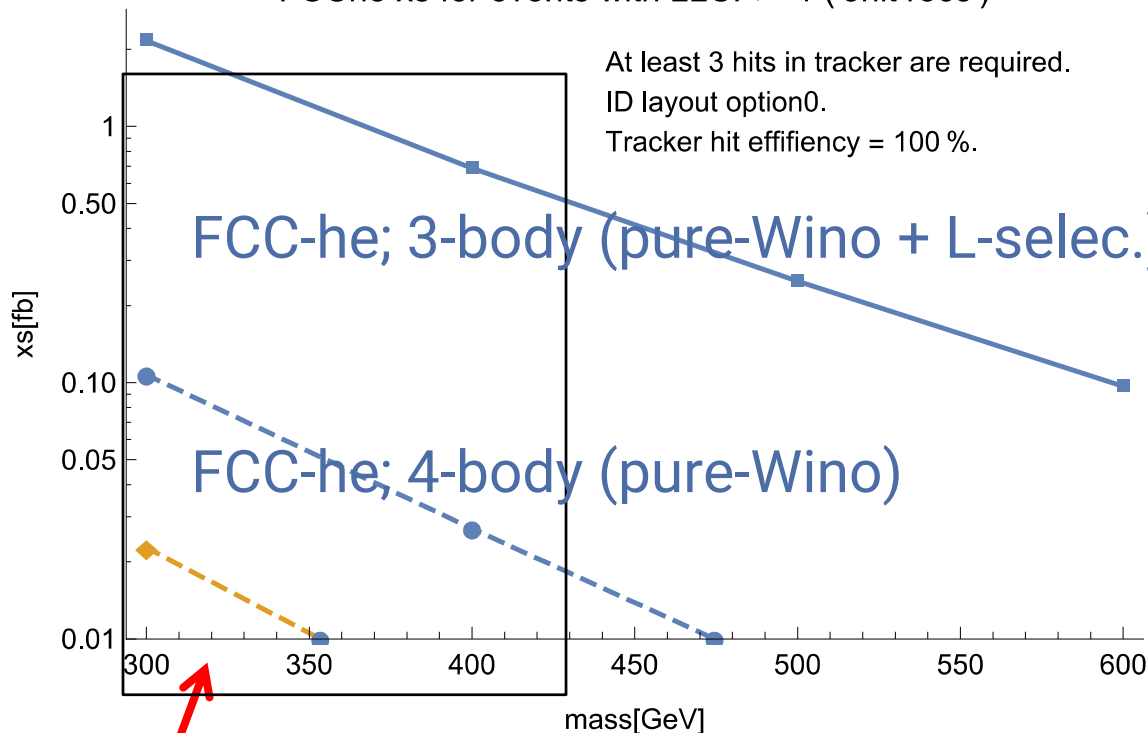
“3-body” model assumes $m_{\tilde{\chi}_1^0} = m_{\tilde{e}} + 1 \text{ GeV}$

FCC-he in-flight decay LLCPs with reconstruction efficiency

LLCPs are required to decay after 3 layers of IDs.
 (= ~ to leave 3 hits in ID)



FCChe xs for events with LLCP >= 1 ('3hit reco')



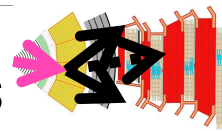
Higgsino-only scenario: less promising because of smaller $\sigma\tau$.

With no polarization.

Shaded region is excluded by ATLAS (13TeV, 36/fb)

"3-body" process assumes $m_{\tilde{e}_L} = m_{\tilde{\chi}_1^0} + 9 \text{ GeV}$.

- "Wino+slep" model is promising.
- reco eff. is governed by the innermost layers
 → closer / more-precise layers would help a lot.



“single SUSY particle”

→ tiny σ due to 4-body process;
HL-LHC will be better.

If with another sparticle for
3-body “co-production”

→ FCC-he will be competitive
with HL-LHC.

- Pure-Wino LSP + \tilde{e}_L
- Slepton LSP + Bino (or Wino)

■ Simple MSSM models

- Pure-Wino dark matter **60mm**
- ~~Pure Higgsino dark matter~~
~~**7mm**~~

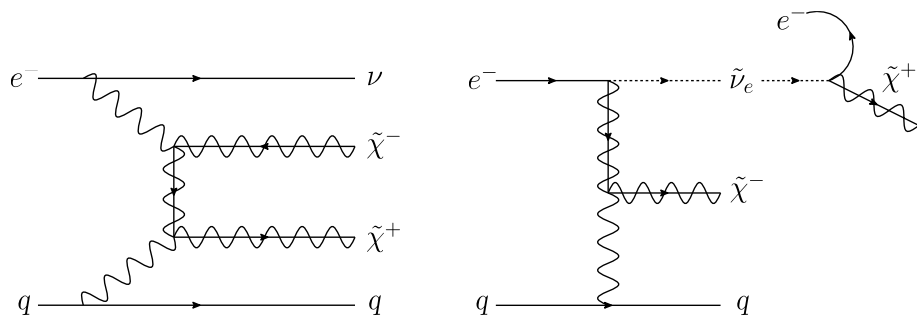
lifetime too short

→ trying to improve by
identifying daughter particles...

■ Other MSSM models

- Gauge-Mediation (keV \tilde{G})
- R-parity violation
with \tilde{l} -LSP

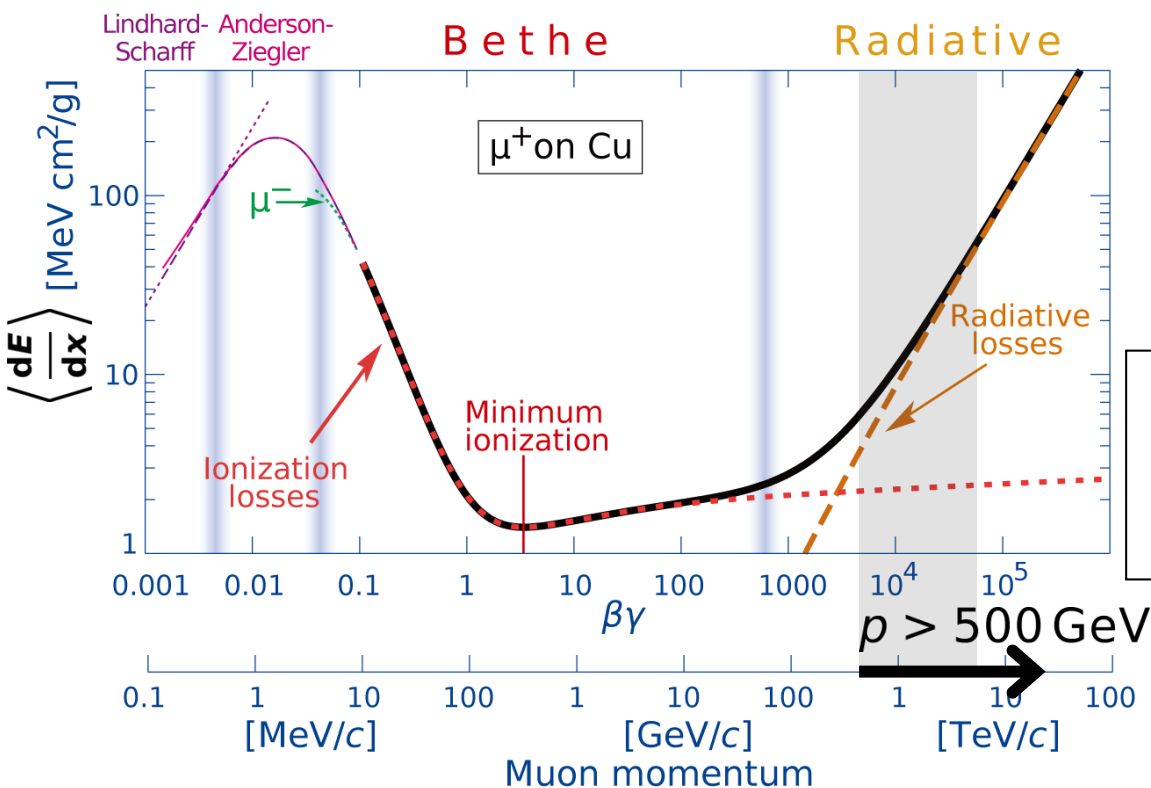
any $c\tau$



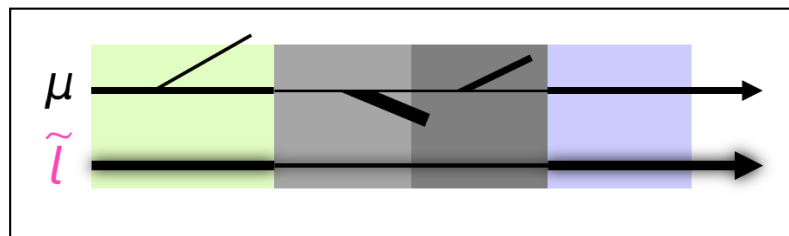
- 100 TeV FCC-hh
mass reach
(Drell-Yan \tilde{l} or $\tilde{\tau}$)

	0.3ab ⁻¹	1ab ⁻¹	3ab ⁻¹	
Exclusion	1.8–2.3	2.4–3.1	3.2–4.0	
Discovery	1.6–2.2	2.3–3.0	3.1–4.0	in TeV

“Muon radiative energy loss”



- Bremsstrahlung
- Photonuclear interaction
- e⁺-e⁻ pair-production



→ 34% of BKG reduction