



# Long-lived charged particles at future $hh$ and $he$ colliders

[Sho IWAMOTO](#) (岩本 祥)

18 Sep. 2017

Seminar @ National Taiwan University

Based on

$hh$ : Jonathan L. Feng, SI, Yael Shadmi, Shlomit Tarem [[1505.02996](#)]

(collected in FCC-hh report [[1606.00947](#)])

$he$ : Kechen Wang, SI, Monica D'Onofrio, Georges Azuelos [17???.?????]

(subgroup in BSM@ep collaboration)

## 1. LLCPs

- Introduction & Motivations
- How to search? & Current bounds

## 2. Future colliders (FCC-hh and FCC-he)

## 3. LLCP searches at FCC-hh

## 4. LLCP searches at FCC-he

Based on

**hh**: Jonathan L. Feng, SI, Yael Shadmi, Shlomit Tarem [[1505.02996](#)]

(collected in FCC-hh report [[1606.00947](#)])

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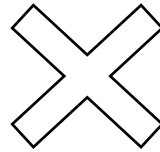
**hh**: Jonathan L. Feng, SI, Yael Shadmi, Shlomit Tarem [[1505.02996](#)]

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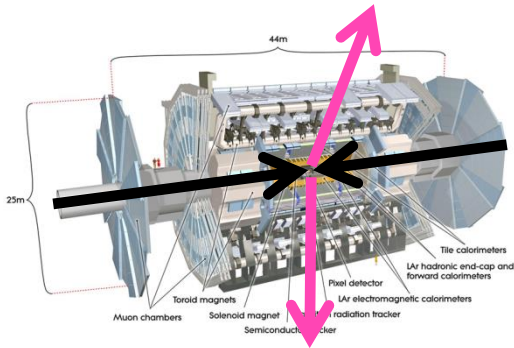
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- Particle property  $m > O(100)\text{GeV}$ 
  - Heavy colored  $\rightarrow$  hadronize  $\rightarrow$   $\left\{ \begin{array}{l} \text{R-hadron /} \\ \text{stopping particles} \end{array} \right.$
  - **Heavy non-colored**
  - Light non-colored, milli-charged  $\rightarrow$  dedicated searches



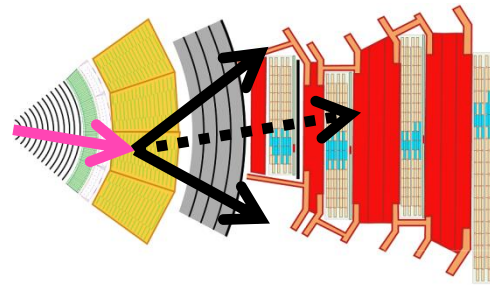
## ■ Lifetime

➤ “stable”



$$c\tau \gtrsim 1 \text{ m}$$

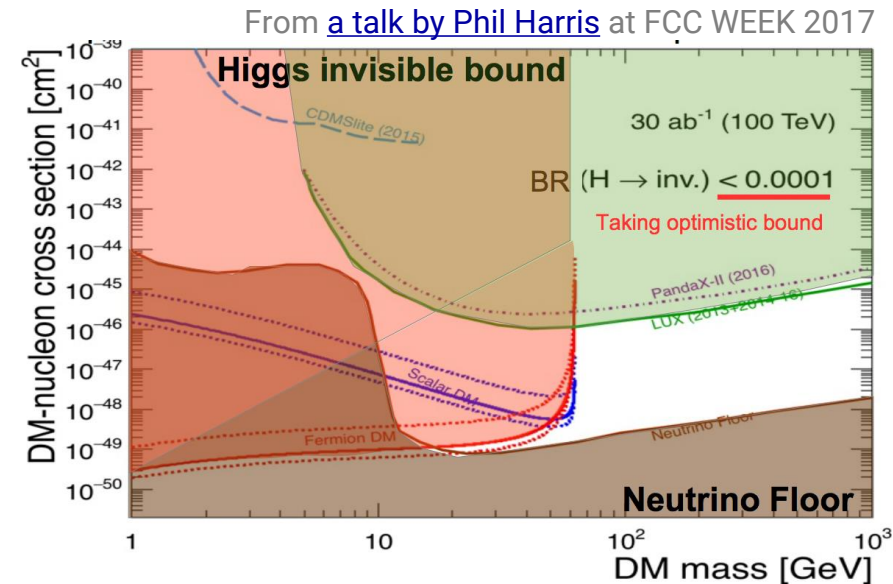
➤ “in-flight decay”



$$c\tau \sim 1 \text{ mm} - 1 \text{ m}$$

# Why should we search for LLCP? – three viewpoints

- experimental
  - ...why not?
  - relevant for detector design
- phenomenology



long lifetime  
→ an actor in early Universe

FCC-hh will(?) cover most of the  
standard thermal-WIMP scenario

**non-standard DM/cosmology with LLCP**  
→ next slides

- theoretical ... **SUSY?** *biased*
  - GMSB scenario: light gravitino → long-lived sleptons
  - split-SUSY: extremely heavy squarks → long-lived gluino

## ■ Dark Matter (especially to ameliorate DM over-abundance)

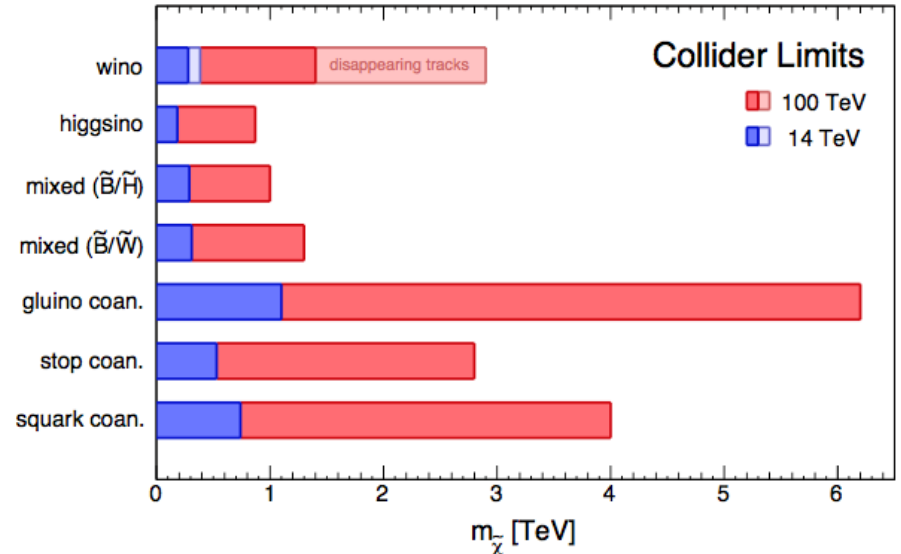
### ➤ co-annihilation

- $(\tilde{B} - \tilde{\tau}) \lesssim 700 \text{ GeV}$
- $(\tilde{W} - \tilde{g}) \lesssim 6\text{--}7 \text{ TeV}$
- $(\tilde{B} - \tilde{g})$  or  $(\tilde{B} - \tilde{t}) \lesssim 8 \text{ TeV}$

Harigaya, Kaneta, Matsumoto [1403.0715],  
Ellis, Olive, Zheng [1404.5571], etc.

### ➤ super-WIMP scenario

→  $\tilde{l} > O(1) \text{ TeV}$



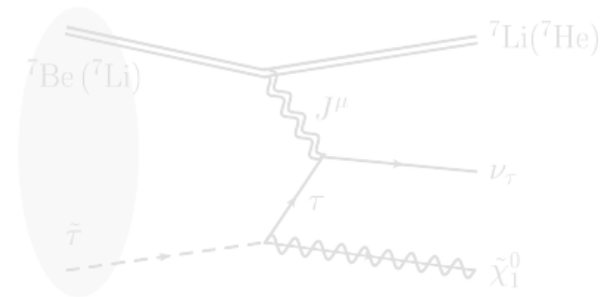
"Physics at the FCC-hh" Report [1606.00947]

## ■ Li problem

### ➤ MSSM $\tilde{\tau}$ with

$$m_{\tilde{\tau}} \sim 400 \text{ GeV}$$

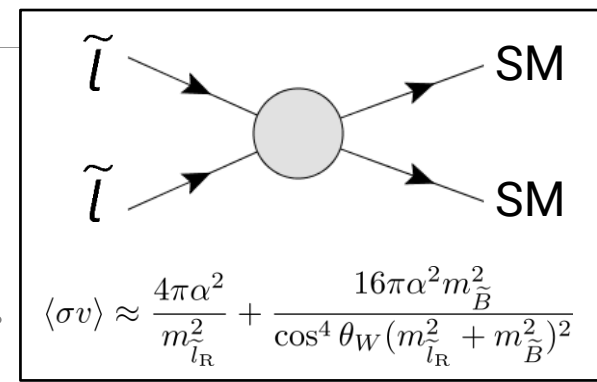
$$\Delta(m_{\tilde{\tau}} - m_{\text{LSP-DM}}) \sim 100 \text{ MeV}$$



Sato, Shimomura, Yamanaka [1604.04769]

■ Super-WIMP:

➤ NLSP slepton  $\tilde{l}$  + LSP gravitino  $\tilde{G}$



frozen-out

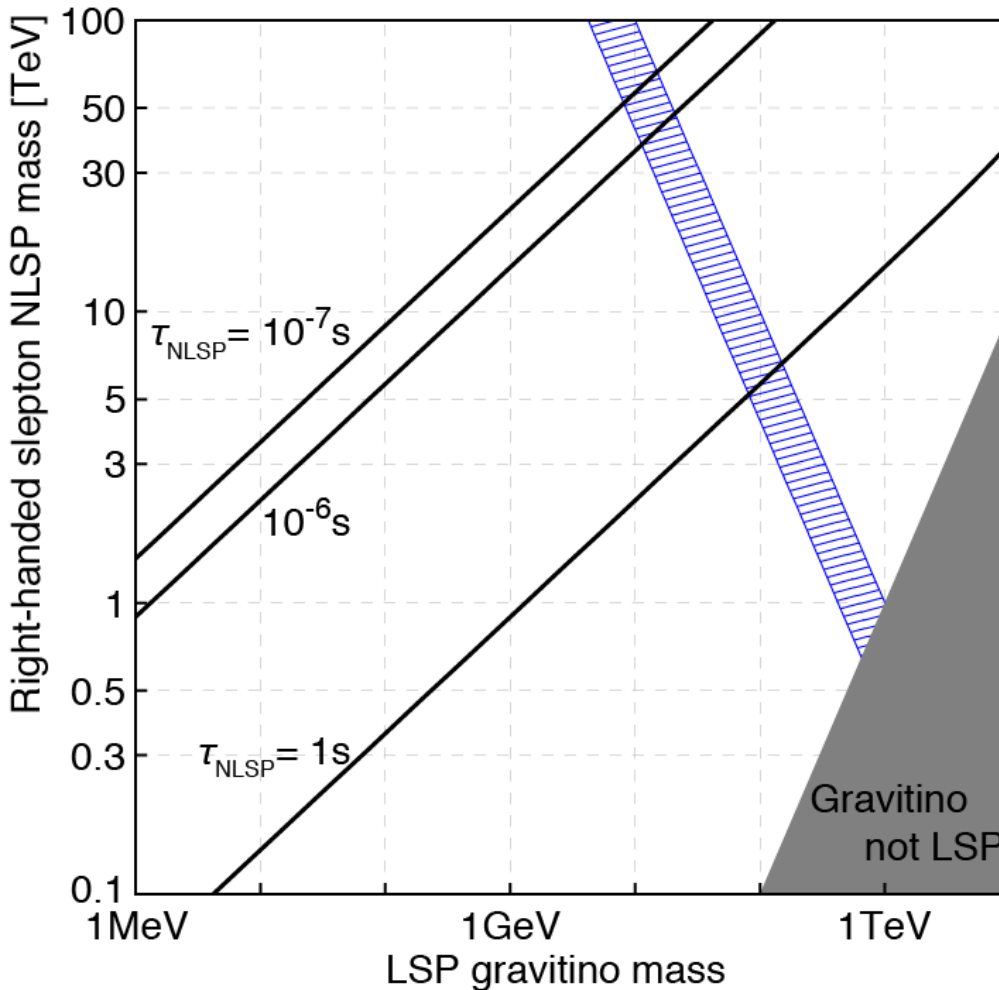
→ thermal relic with  $(\Omega h^2)_{\tilde{l}} \gg 0.12$

late-time decay  $\tilde{l} \rightarrow \tilde{G} + l$

$$(\Omega h^2)_{\text{DM}(\tilde{G})} = \frac{m_{\tilde{G}}}{m_{\tilde{l}}} (\Omega h^2)_{\tilde{l}}$$

$$\tau(\tilde{l} \rightarrow l\tilde{G}) = 5.7 \times 10^{-7} \text{ sec} \cdot \left(\frac{m_{\tilde{l}}}{1 \text{ TeV}}\right)^{-5} \left(\frac{m_{\tilde{G}}}{1 \text{ MeV}}\right)^2$$

■ Super-WIMP:



(short-lived)

**LLCP search target**

(BBN/CMB constraints are relevant.)

$$\tau(\tilde{l} \rightarrow l\tilde{G}) = 5.7 \times 10^{-7} \text{ sec} \cdot \left(\frac{m_{\tilde{l}}}{1 \text{ TeV}}\right)^{-5} \left(\frac{m_{\tilde{G}}}{1 \text{ MeV}}\right)^2$$



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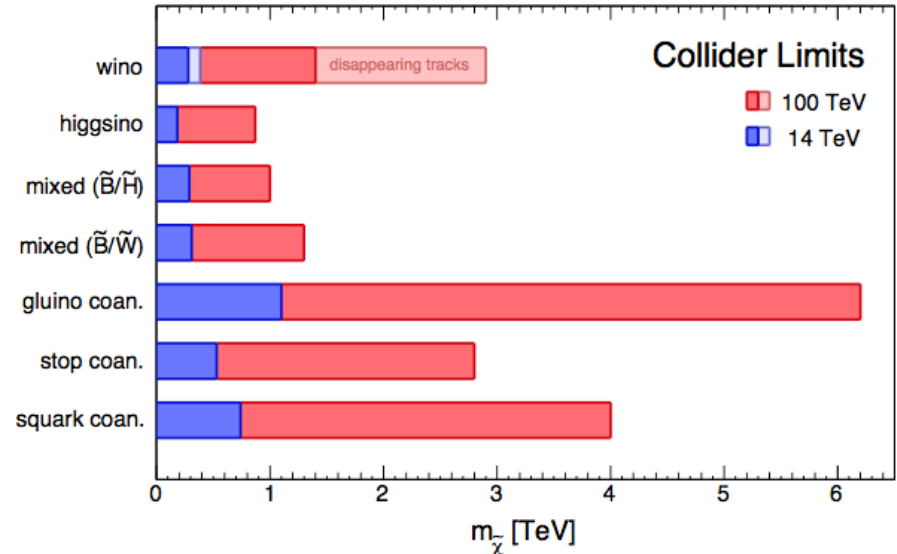
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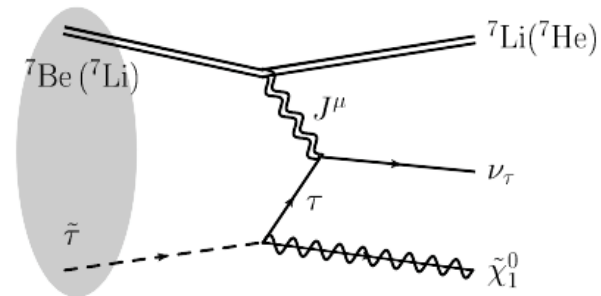
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Harigaya, Kaneta, Masuhira [1403.0715],  
Ellis, Olive, Zheng [1404.557] etc.

**non-colored LLCP**

**... covered in this talk**

(slepton  $\tilde{l}$  / stau  $\tilde{\tau}$  as benchmarks)

➤ super-WIMP scenario

→  $\tilde{l} > O(1) \text{ TeV}$

(as we see soon.)

**colored LLCP**

(→ hadronize: "R-hadron")



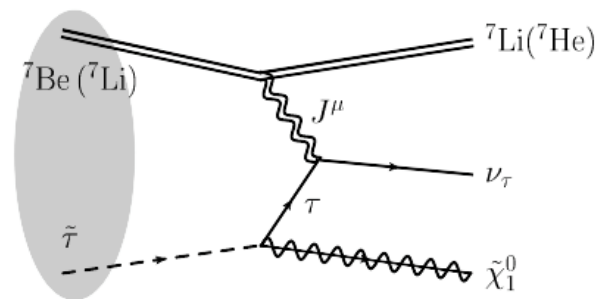
[1606.00947]

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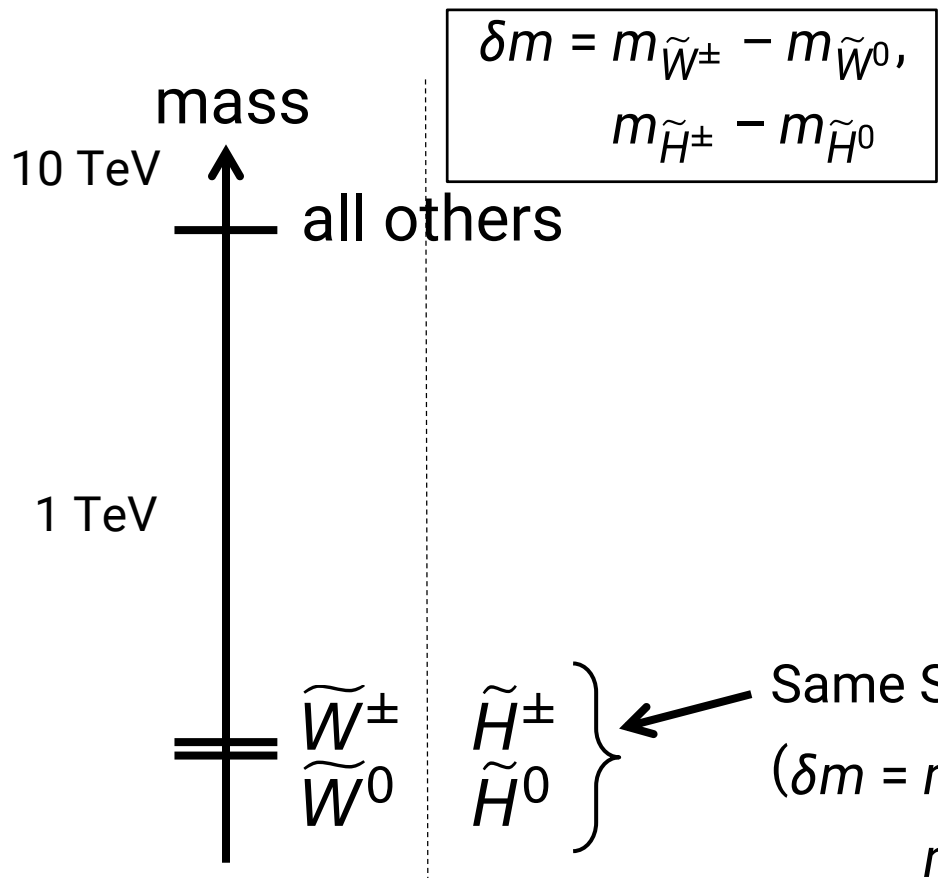
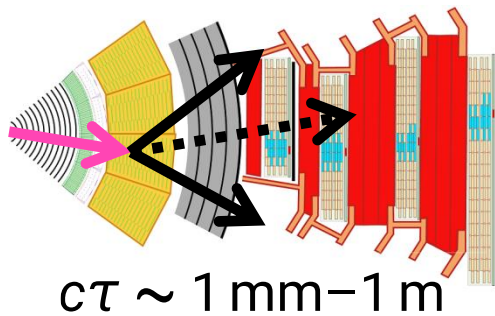
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## ■ Simple MSSM models

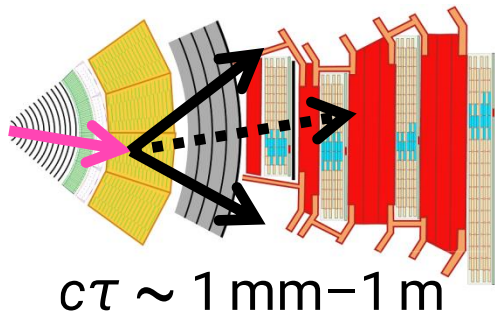
- Pure-Wino dark matter
- Pure-Higgsino dark matter



$\tilde{\chi}^\pm$  long-lived because of small  $\delta m$   
( $\delta m > 0$ )

## ■ Other MSSM models

- Gauge-Mediation
- R-parity violation



$$\delta m = m_{\tilde{W}^\pm} - m_{\tilde{W}^0},$$

$$m_{\tilde{H}^\pm} - m_{\tilde{H}^0}$$

## ■ Simple MSSM models

- Pure-Wino dark matter **60mm**
- Pure-Higgsino dark matter **7mm**

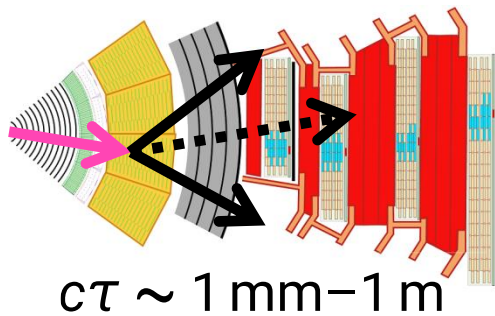


$\tilde{\chi}^\pm$  long-lived because of small  $\delta m$   
 ( $\delta m > 0$ )

$m_{\tilde{W}}$ [GeV]	200	250	300	350	400	450	500	550	600	700	800	900
$\delta m$ [MeV]	159	160	161	162	162	163	163	163	163	164	164	164
$c\tau$ [mm]	71	67	64	63	62	61	60	60	59	59	59	59

### ➤ R-parity violation

$m_{\tilde{H}}$ [GeV]	200	250	300	350	400	450	500	550	600	700	800	900
$\delta m$ [MeV]	297	306	313	319	323	326	329	331	333	336	338	340
$c\tau$ [mm]	11	10	9.4	8.9	8.5	8.2	8.0	7.8	7.7	7.4	7.2	7.1



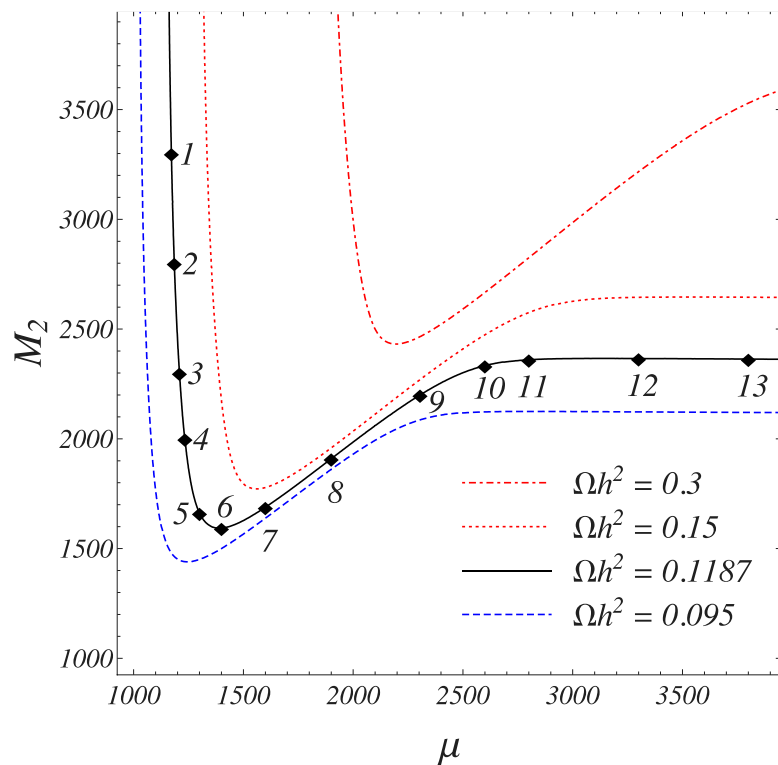
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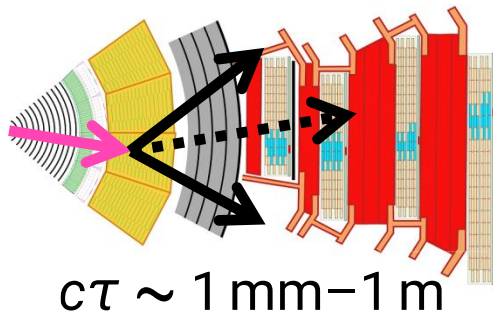
$m_{A^0} = 500 \text{ GeV}$ ,  $m_{\text{sfermion}} = 9 \text{ TeV}$ ,  $\tan \beta = 15$   $\tilde{\chi}^{\pm}$  long-lived because of small  $\delta m$  ( $\delta m > 0$ )

(but **DM underabundant** for "observable" region)



## Other MSSM models

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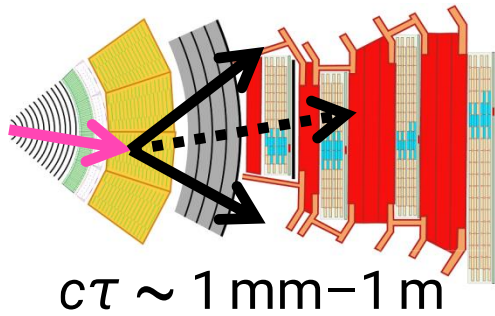
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## ■ Other MSSM models

- Gauge-Mediation (keV  $\tilde{G}$ )
- R-parity violation

with  $\tilde{l}$ -LSP

⇒  $\tilde{l}$  long-lived  
 because of tiny couplings.



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$$c\tau \sim 1.8 \times 10^{-5} \text{ m} \left( \frac{m_{\tilde{l}}}{100 \text{ GeV}} \right)^{-5} \left( \frac{m_{\tilde{G}}}{1 \text{ eV}} \right)^2$$

for Gauge-Mediation: “gravity is weak”,

$$0.50 \text{ m} \left( \frac{m_{\tilde{l}}}{100 \text{ GeV}} \right)^{-1} \left( \frac{\lambda_{ijk}}{10^{-8}} \right)^{-2}$$

for R-parity violation: “if RpV is tiny”.

## ■ Other MSSM models

- Gauge-Mediation (keV  $\tilde{G}$ )
- R-parity violation

with  $\tilde{l}$ -LSP

⇒  $\tilde{l}$  long-lived **any  $c\tau$**   
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$\tilde{\tau}$ -LLCP

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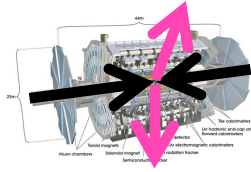
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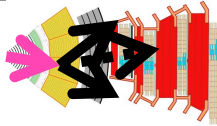
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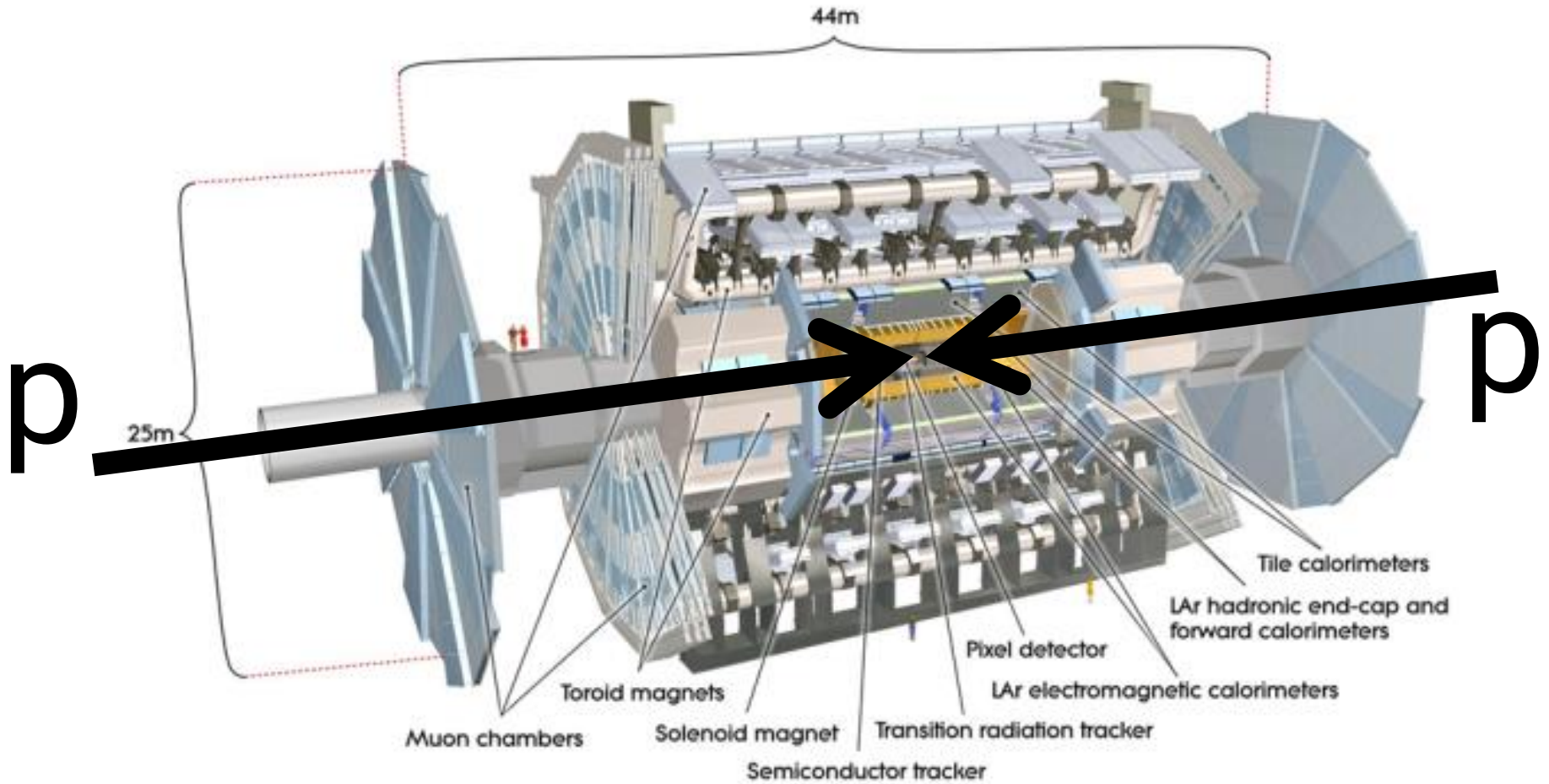
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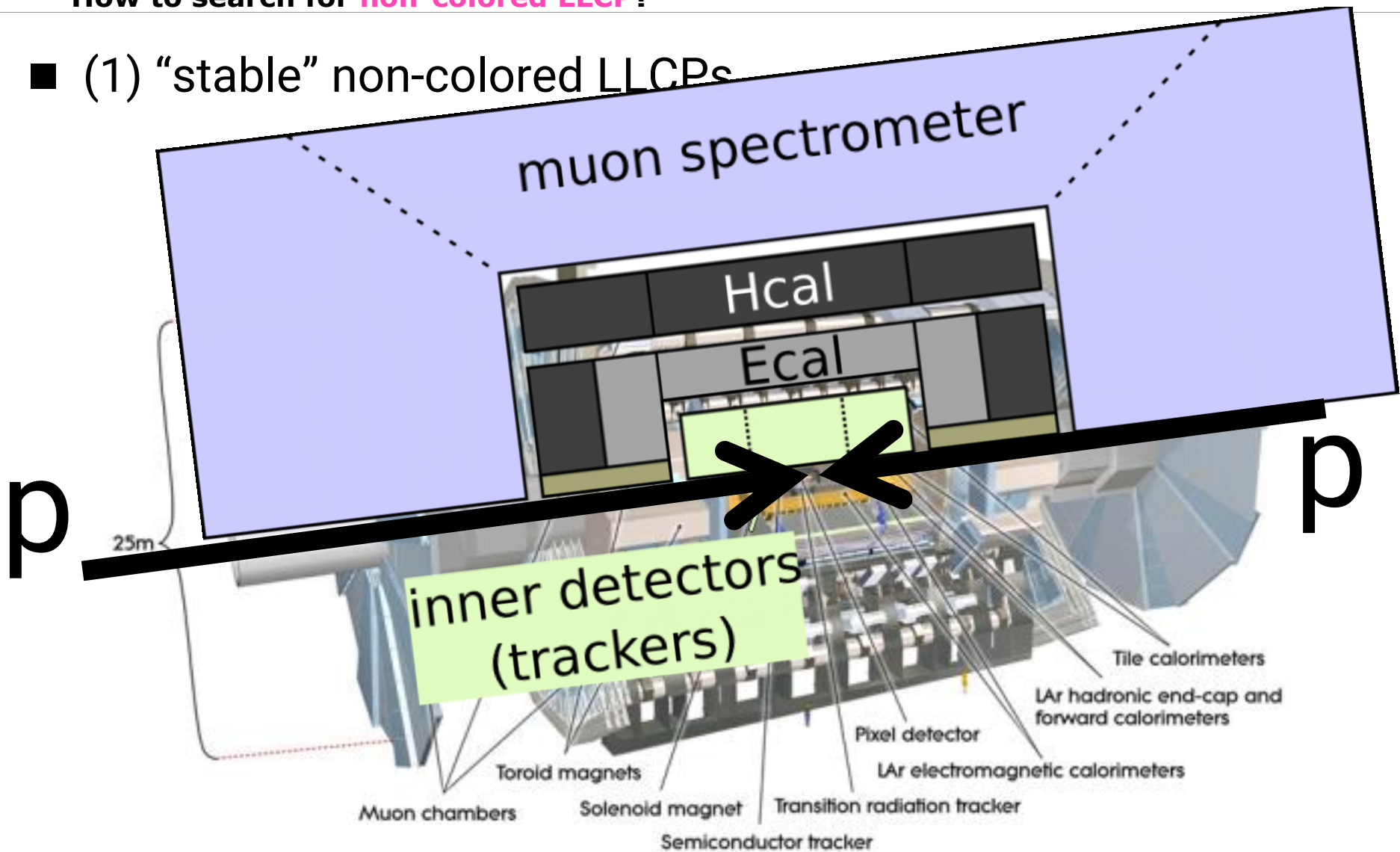
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- (1) “stable” non-colored LLCPs

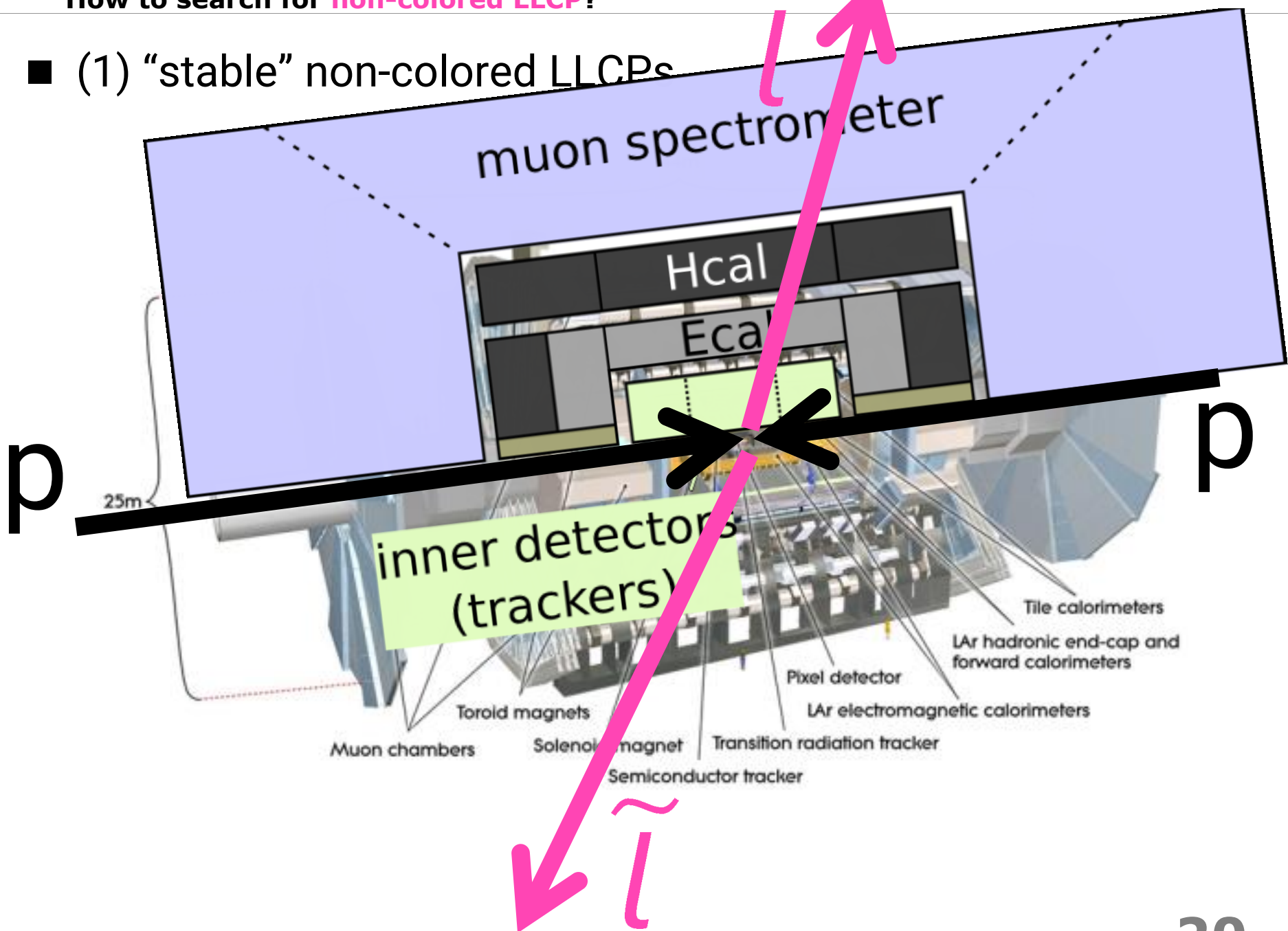


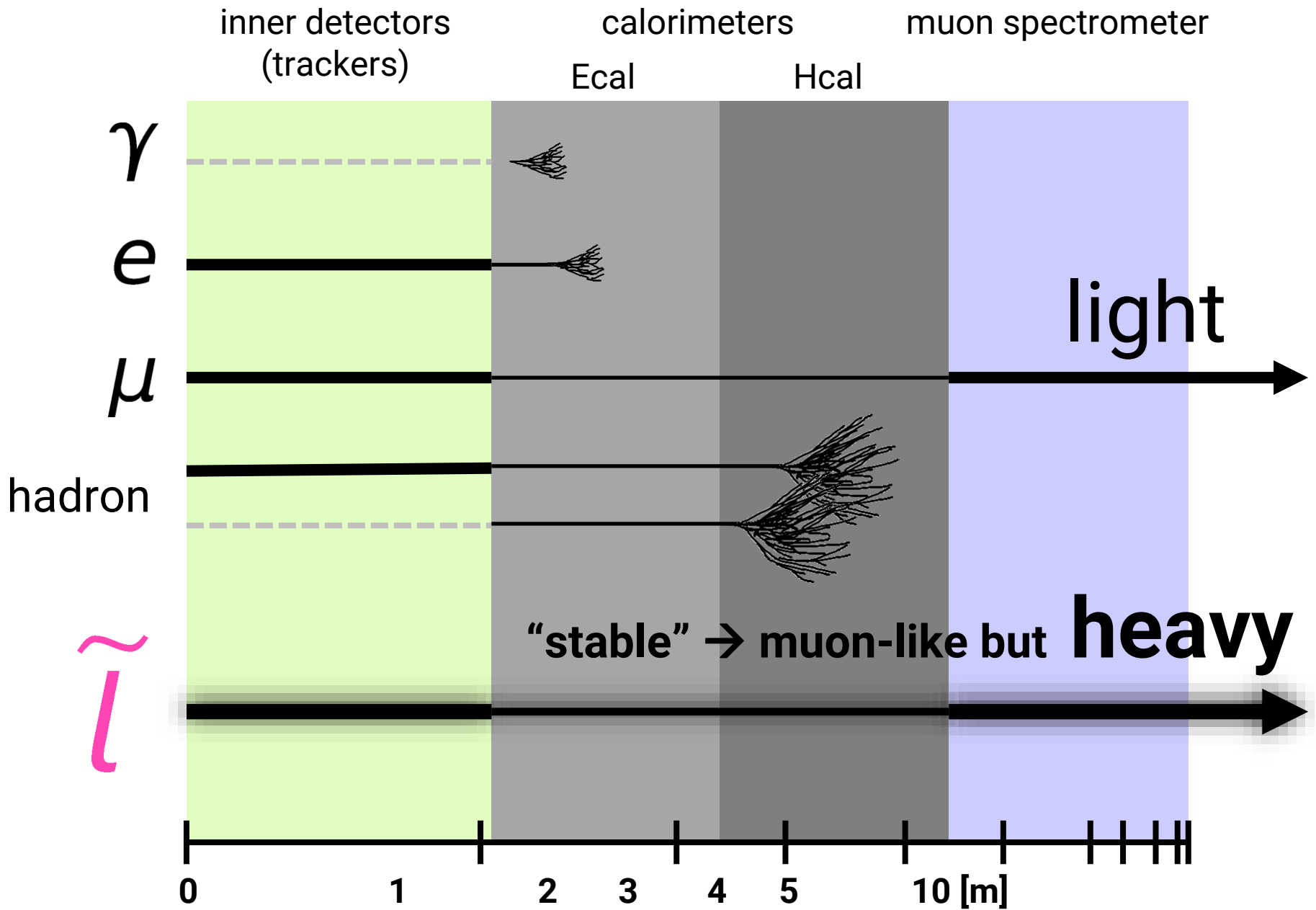
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# How to search for **non-colored LLCPs**?

- (1) “stable” non-colored LLCPs



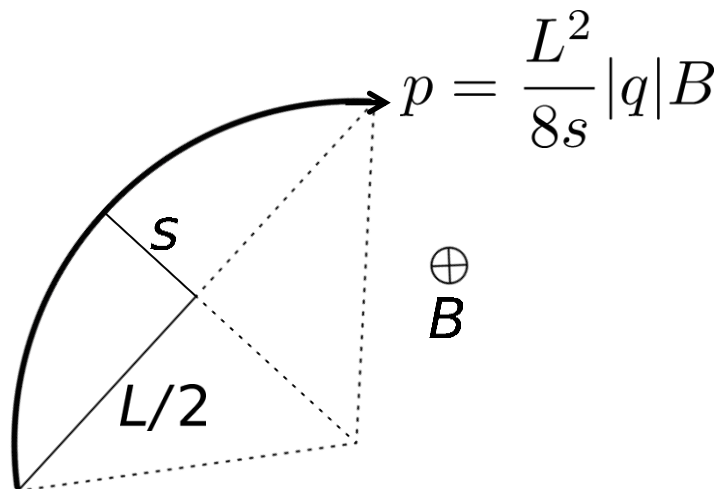


$$m = \frac{p}{\beta\gamma} = \frac{p}{\beta/\sqrt{1-\beta^2}}$$

momentum & velocity

■ **mass** measurement =  **$p$**  &  **$\beta$**  measurements ( $\beta = v/c$ )

➤ momentum



➤ velocity

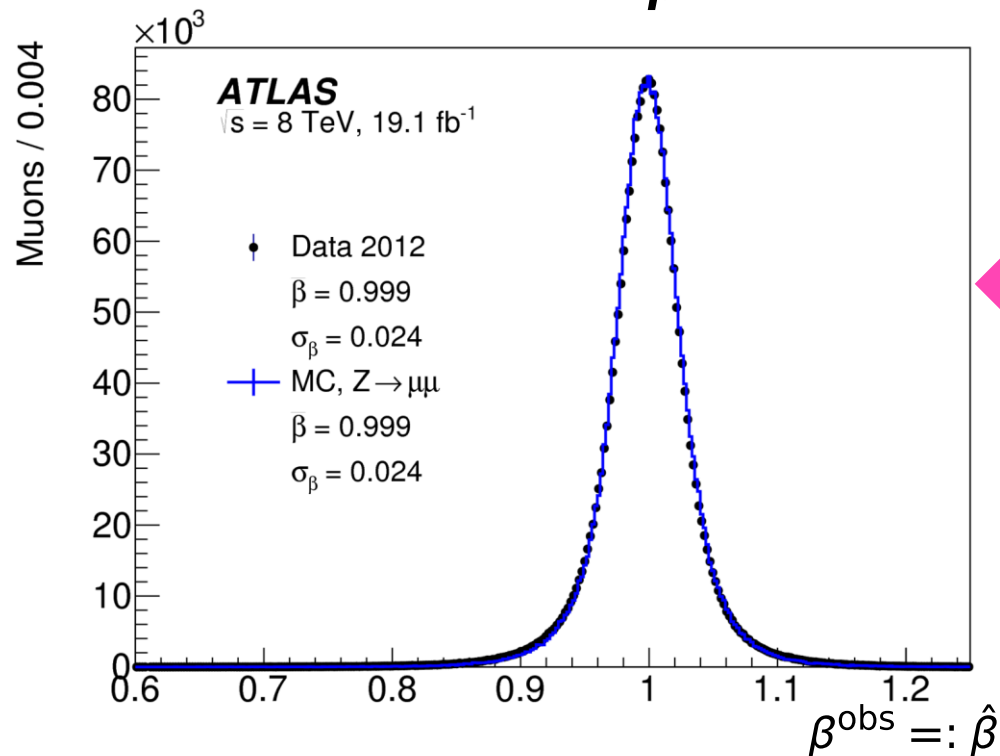
- TOF [time-of-flight]  
 $\beta = \Delta L / \Delta t$
- $dE/dx$  [ionization energy loss]

$$m = \frac{p}{\beta\gamma} = \frac{p}{\beta/\sqrt{1-\beta^2}}$$

momentum & velocity

■ **mass** measurement =  **$p$**  &  **$\beta$**  measurements ( $\beta = v/c$ )

ATLAS muon data:  $\Delta\beta = 2.4\%$

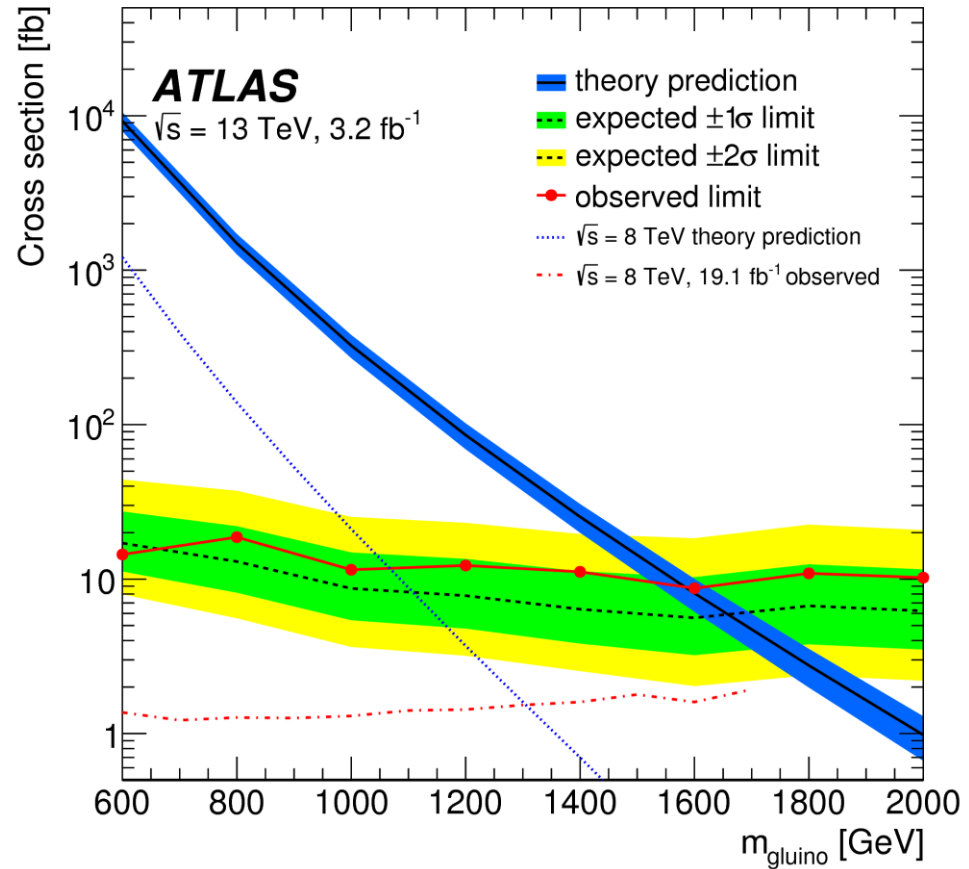
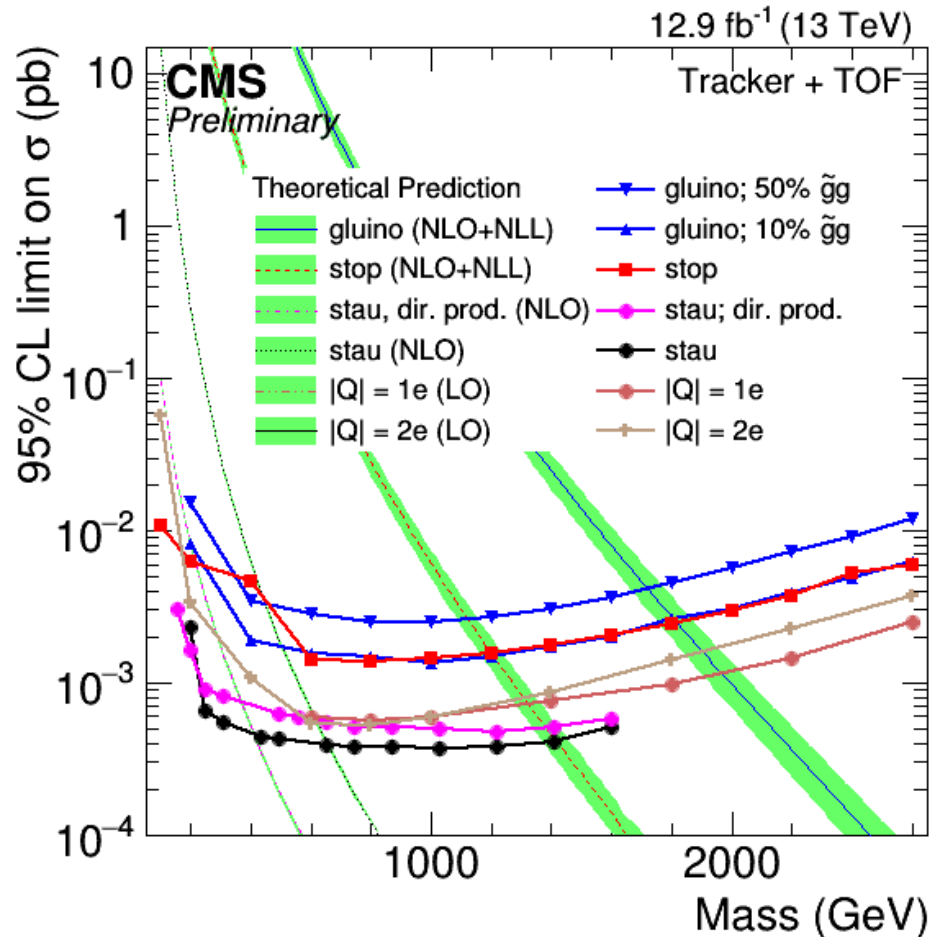


➤ velocity

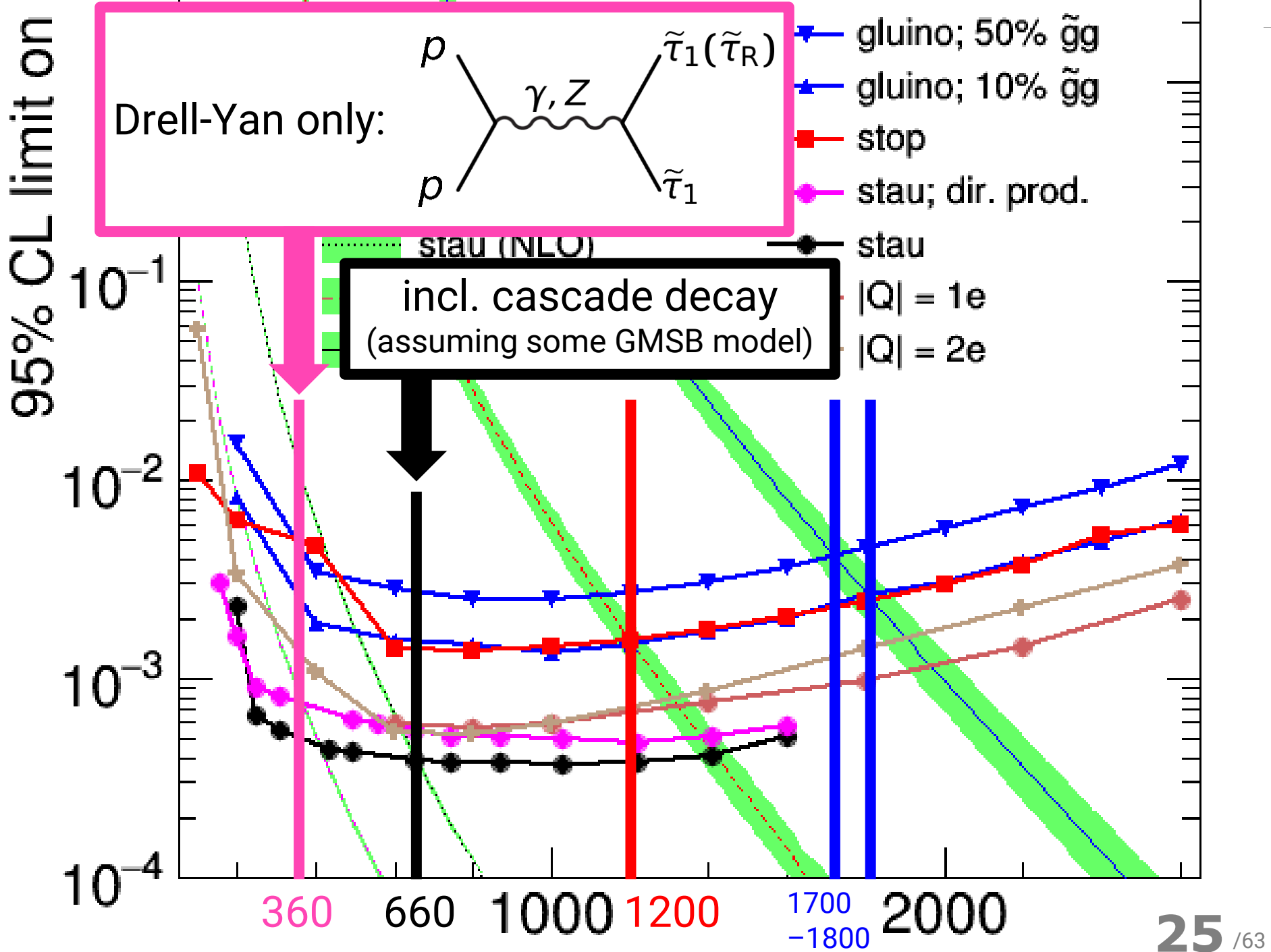
- TOF [time-of-flight]

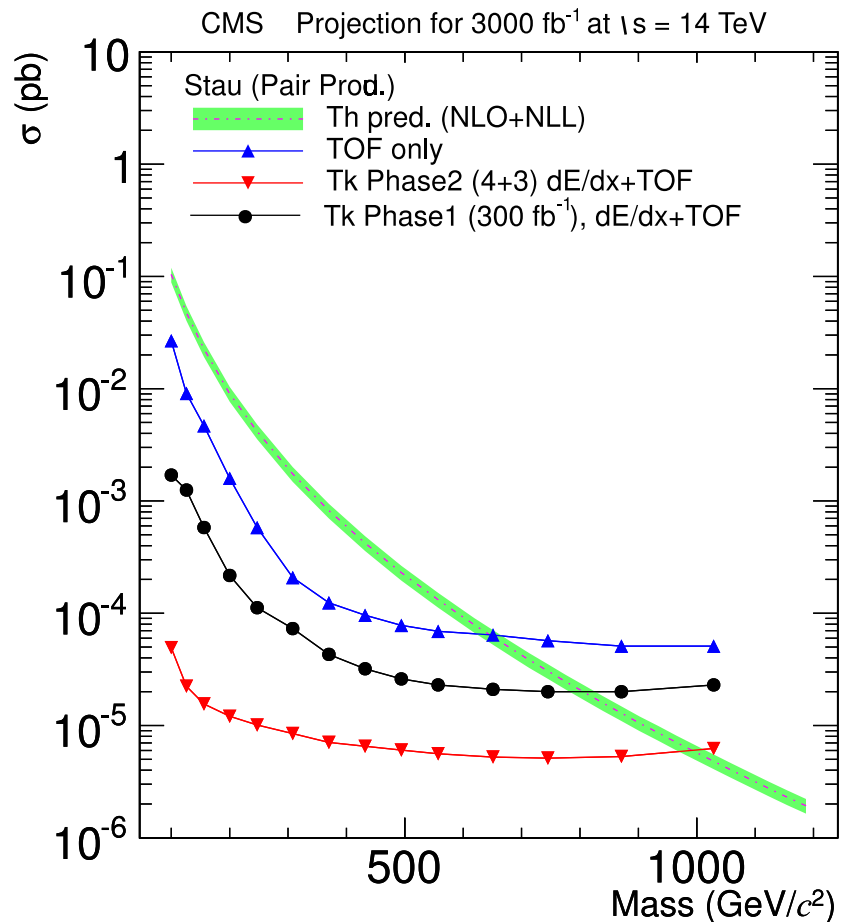
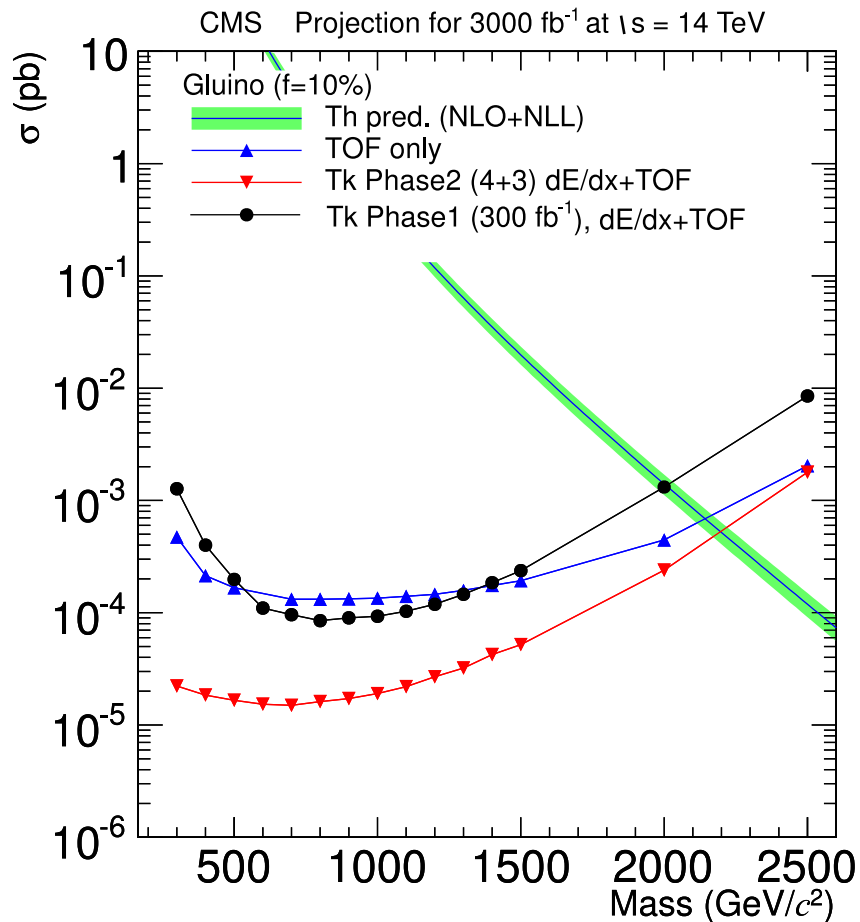
←  $\beta = \Delta L / \Delta t$

- $dE/dx$  [ionization energy loss]





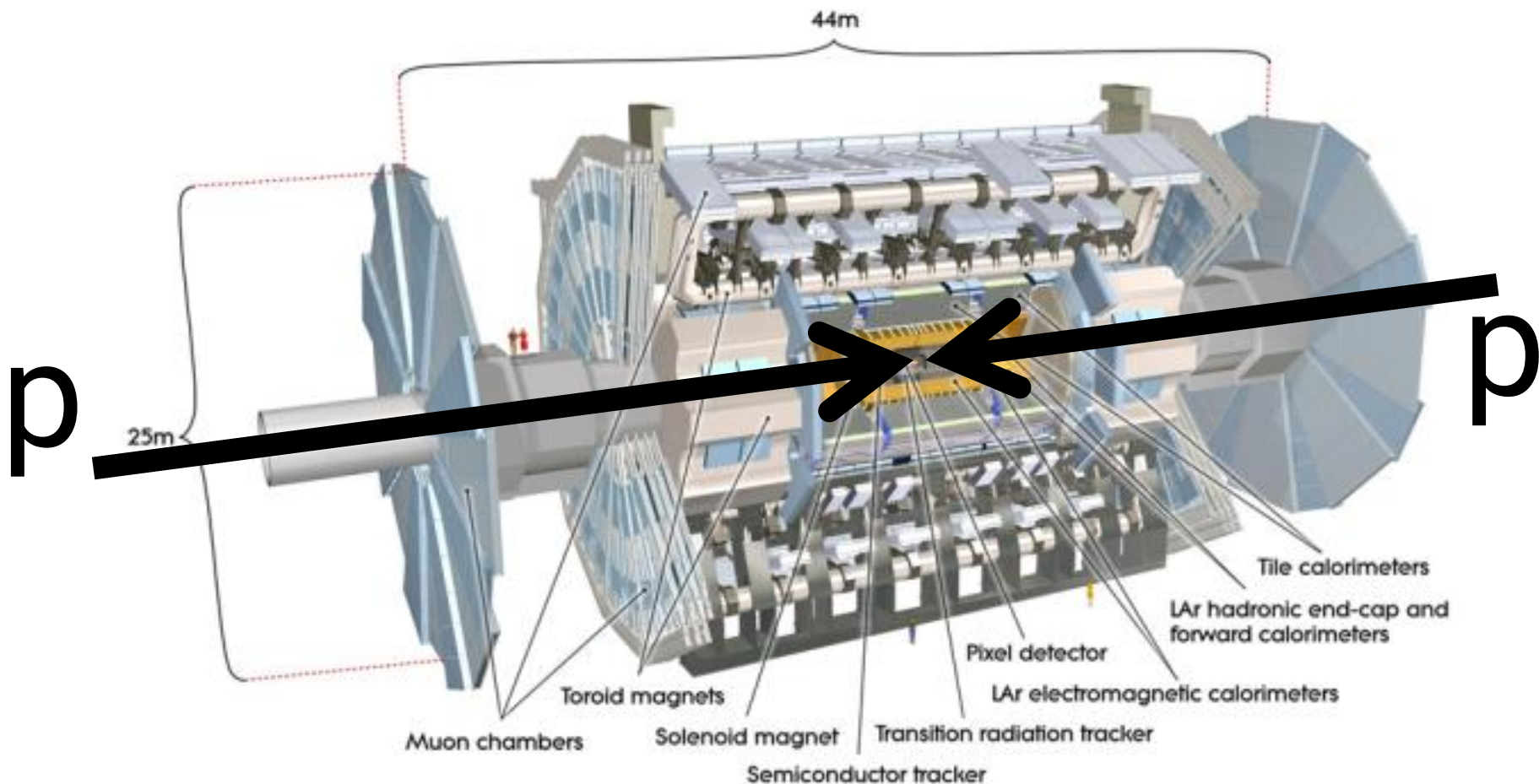




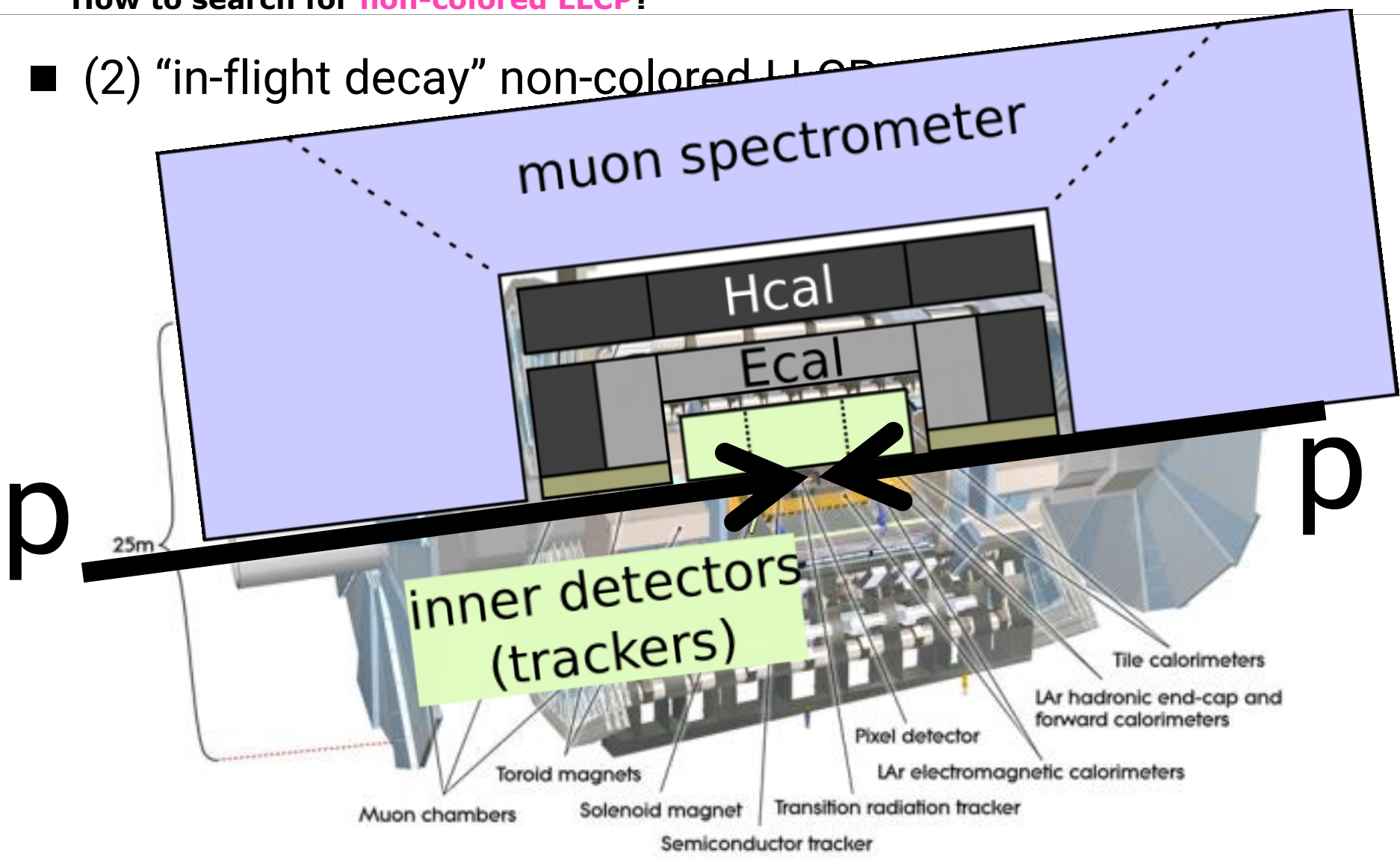
	current	HL-LHC
gluino:	1.7 TeV	→ 2.2 TeV?
stop:	1.2 TeV	→ 1.7 TeV?
stau (GMSB):	660 GeV	→ 1.2 TeV?
stau (DY):	360 GeV	→ 1.0 TeV?

(or discovery?)

- (2) “in-flight decay” non-colored LLCPs

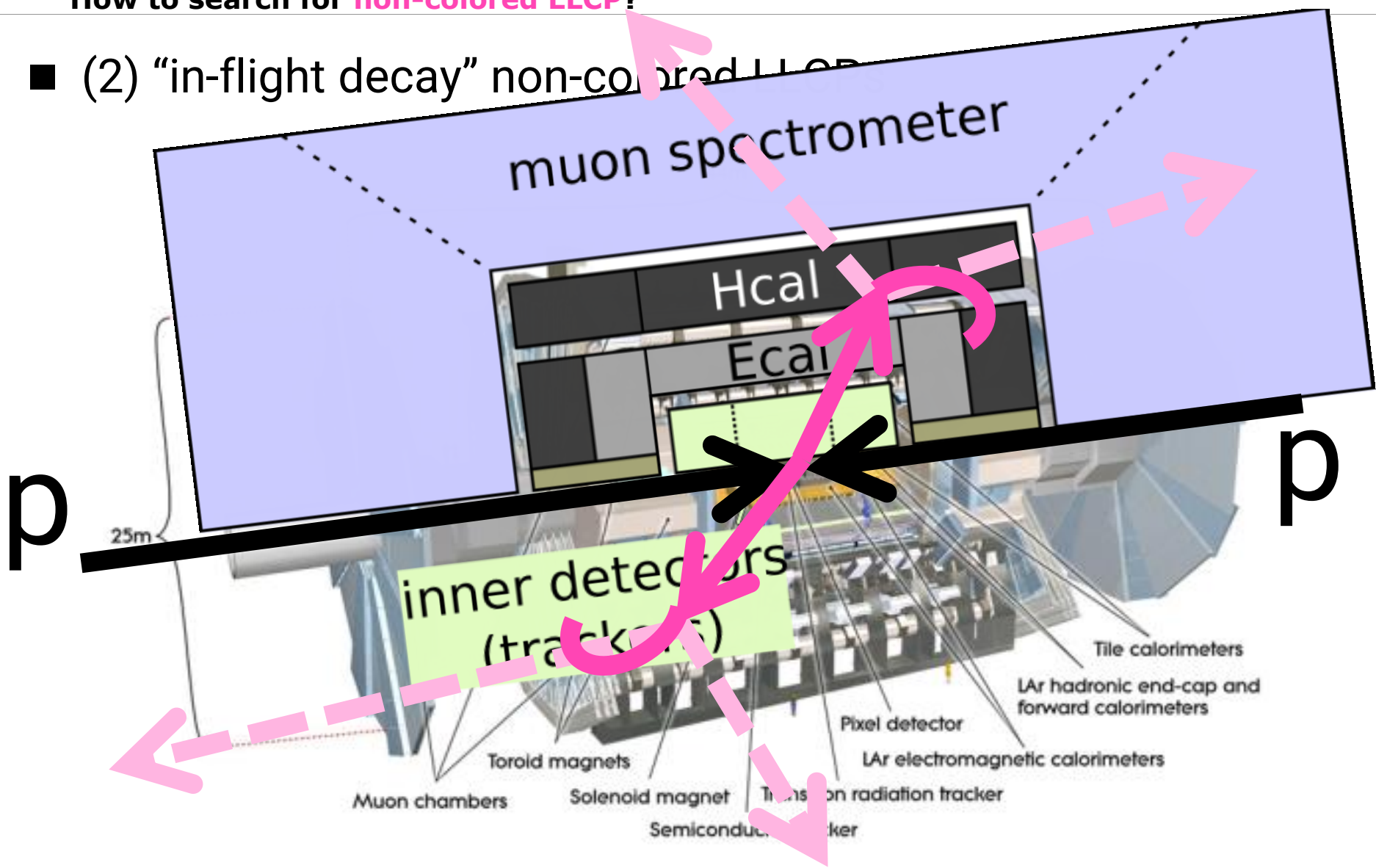


- (2) “in-flight decay” non-colored LLCP

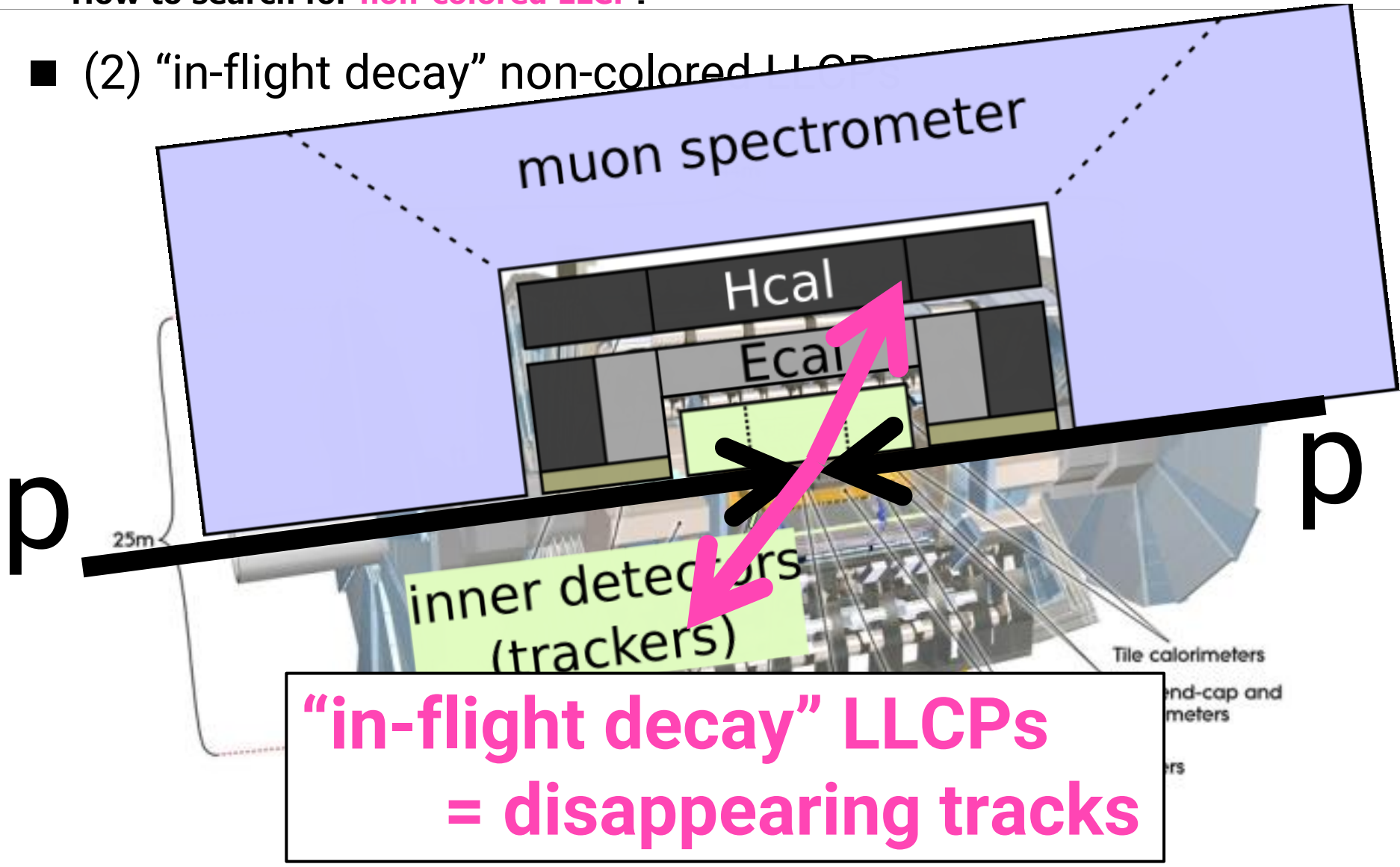


# How to search for non-colored LLCP?

- (2) "in-flight decay" non-colored LLCP

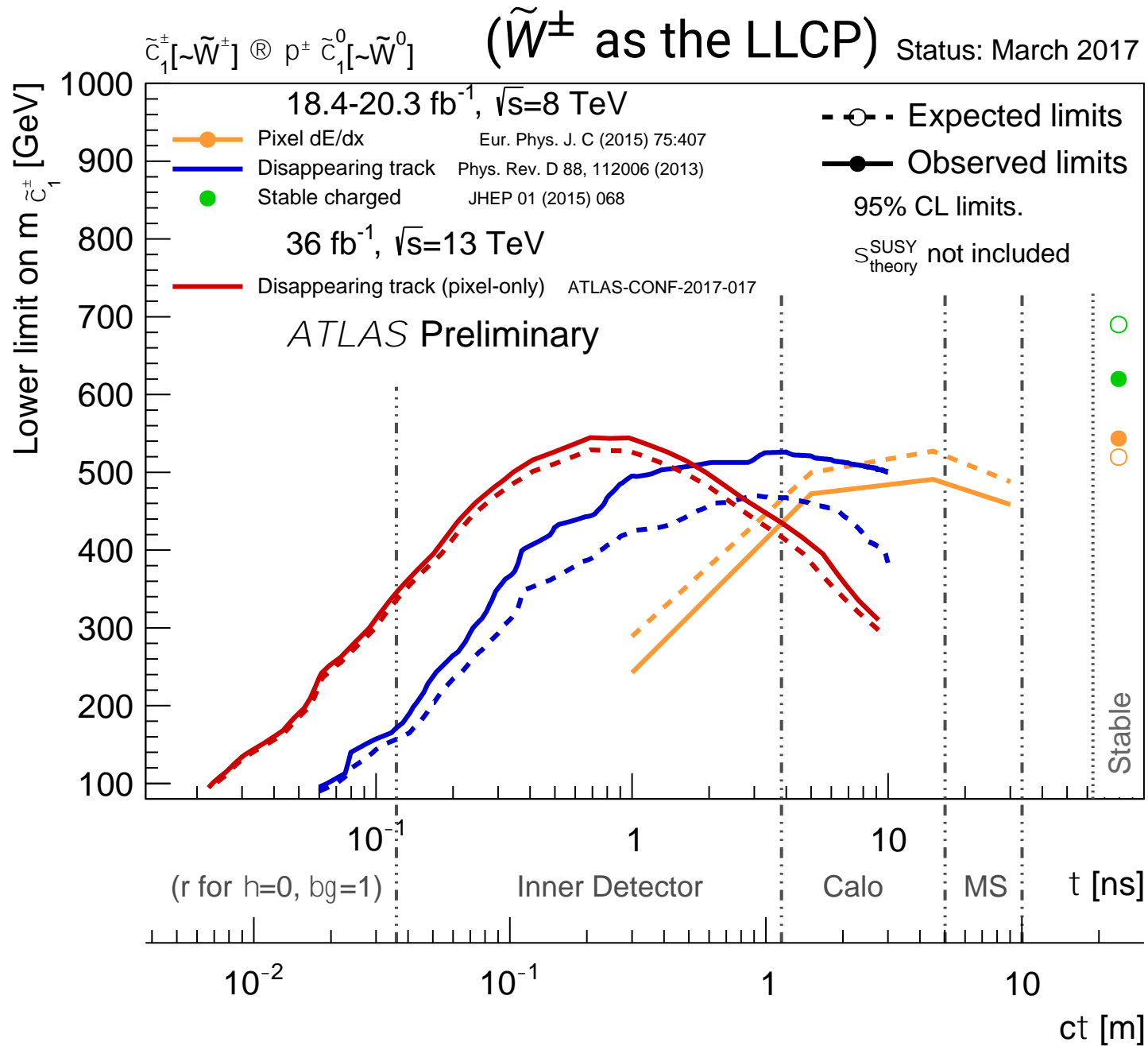


- (2) “in-flight decay” non-colored LLCP



**“in-flight decay” LLCPs  
= disappearing tracks**





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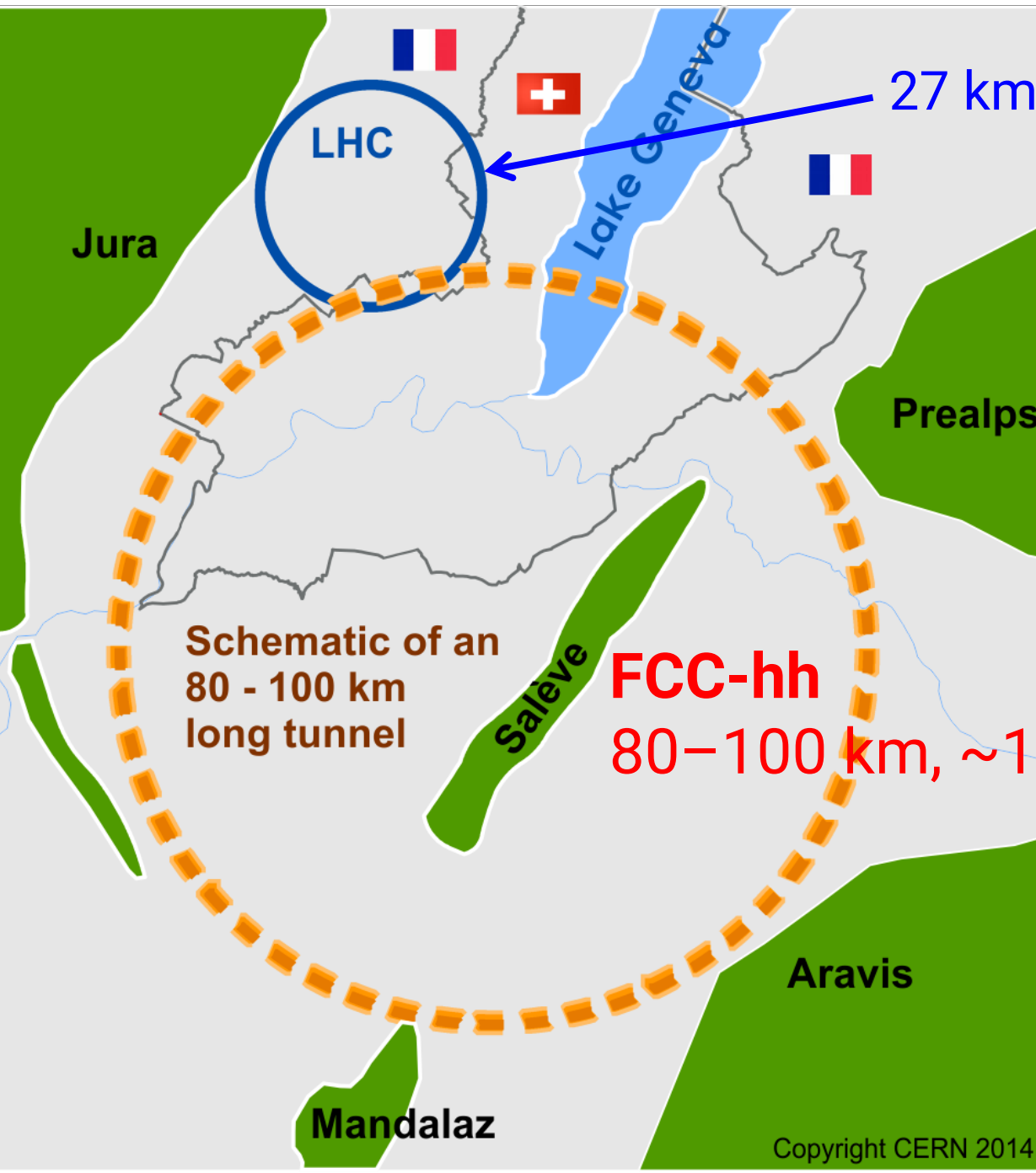
(collected in FCC-hh report [[1606.00947](#)])

**he**: Kechen Wang, SI, Monica D'Onofrio, Georges Azuelos [17???.?????]

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# FCC-hh



27 km, 8.3 T, beam = 7 TeV

\* **HE-LHC**  
27 km, 16 T  
beam = 13.5 TeV

Schematic of an  
80 - 100 km  
long tunnel

**FCC-hh**  
80-100 km, ~16 T, beam = 50 TeV  
( $\int \mathcal{L} = 5-10 \text{ ab}^{-1}$ )

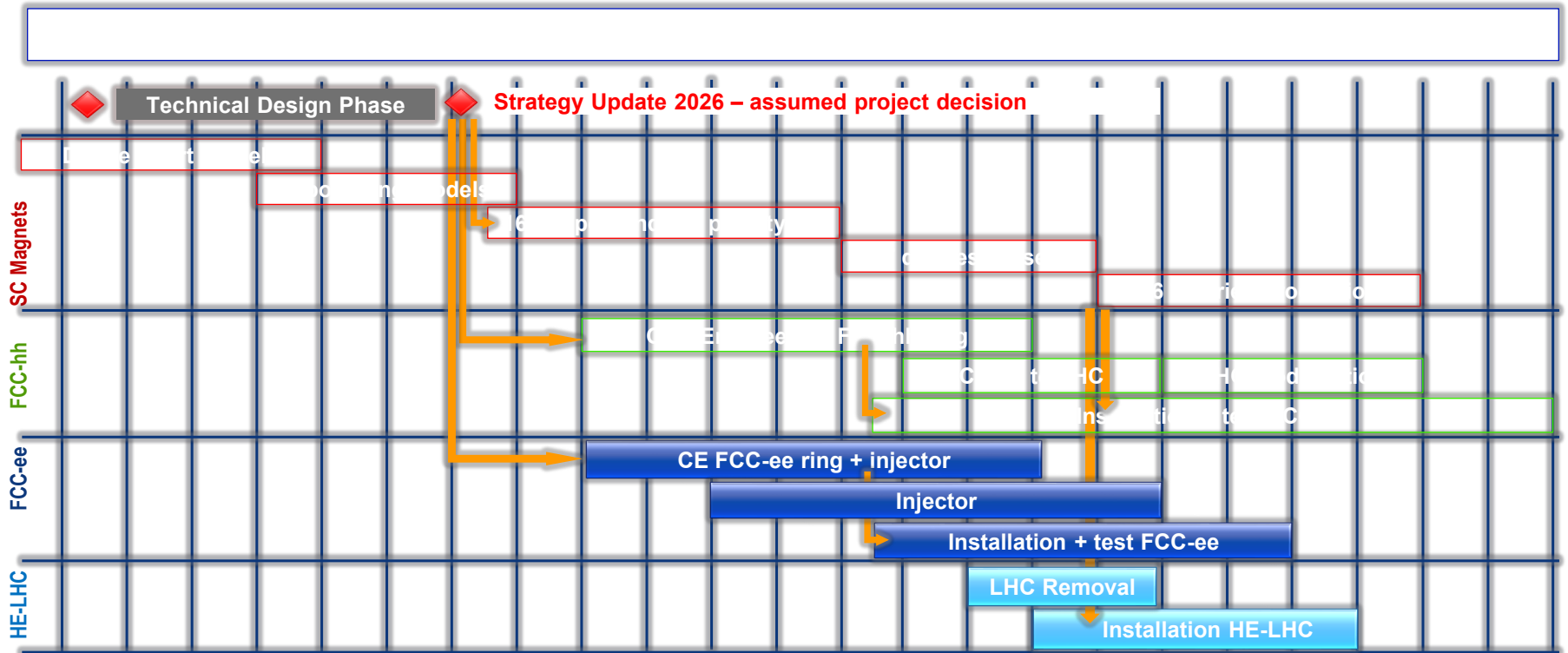
(FCC-ee as a potential first step)

→ CDR in 2018

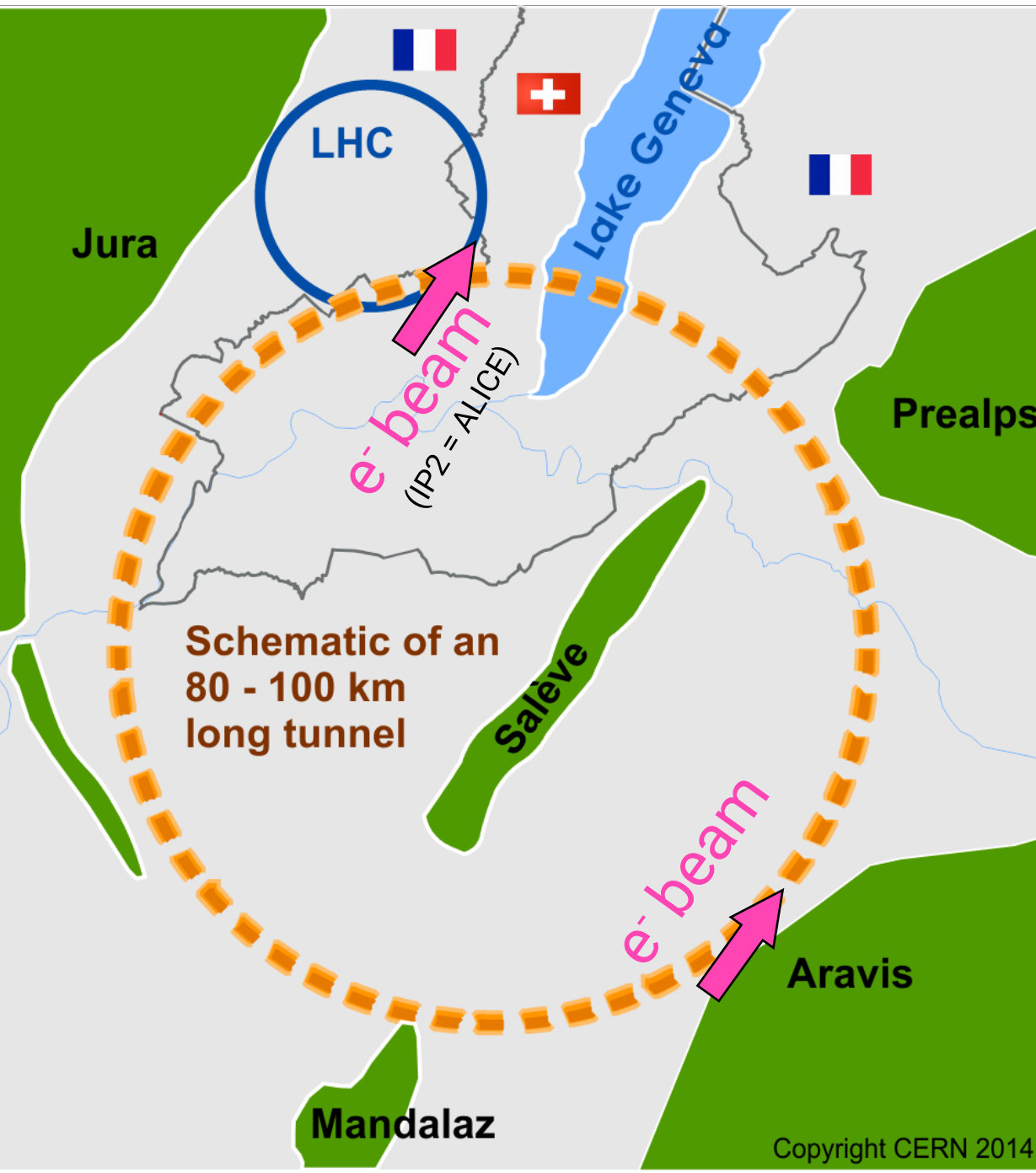
[conceptual design report]



# Draft Schedule Considerations



# FCC-he

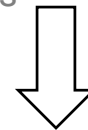


LHC +  $e^-$  beam = LHeC

FCC-hh +  $e^-$  beam  
= FCC-he

$E_e = 60 \text{ GeV}$  (140 GeV?)

from "Energy Recovery  
Linacs"



$\sqrt{s_{\text{LHeC}}} = 2 \times 0.65 \text{ TeV}$

$\sqrt{s_{\text{FCC-he}}} = 2 \times 1.73 \text{ TeV}$

( $\int \mathcal{L}_{\text{FCC-he}} \sim 1 \text{ ab}^{-1}$ )

- Mar 2015 : FCC week 2015 @ Washington D.C.
- Apr 2016 : FCC week 2016 @ Rome
- Jan 2017 : FCC physics workshop @ CERN
- May 2017 : FCC week 2017 @ Berlin
- Sep 2017 : LHeC/FCC-eh workshop @ CERN
- Jan 2018 : FCC physics workshop @ CERN
- Apr 2018 : FCC week 2018 @ Amsterdam

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Based on

**hh**: Jonathan L. Feng, SI, Yael Shadmi, Shlomit Tarem [[1505.02996](#)]

(collected in FCC-hh report [[1606.00947](#)])

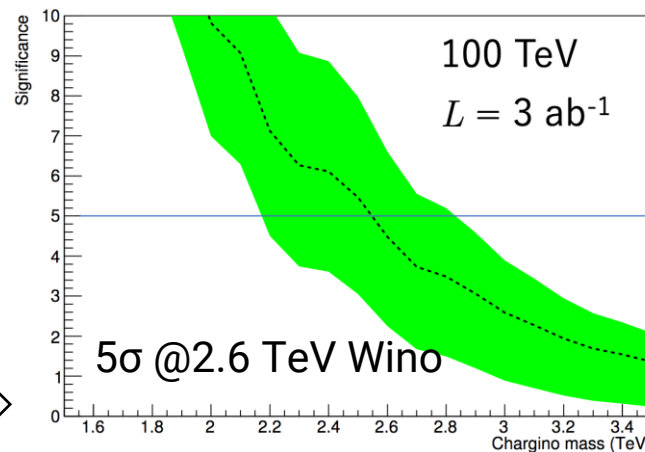
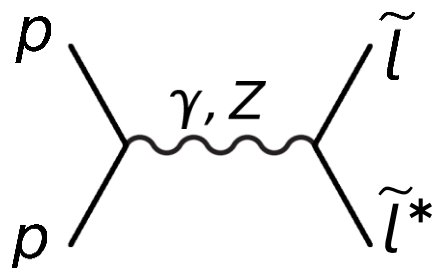
**he**: Kechen Wang, SI, Monica D'Onofrio, Georges Azuelos [17???.?????]

(subgroup in BSM@ep collaboration)

■ LLCPs at FCC-hh  $\approx$  LLCPs at LHC

➤ same production mechanism; just with a higher energy.

- e.g.,  $\tilde{l} \rightarrow$  Drell–Yan process (or from cascade decay)



[R. Sawada's talk in FCC week 2017]

➤ same detection method.

- “in-flight decay”  $\rightarrow$  disappearing track.
- “stable”  $\tilde{l} \rightarrow$  muon-like track +  $\beta$  measurement (heavy = slow)

- $dE/dx$  (ionization energy loss)
- $\Delta L/\Delta t$  (time-of-flight)

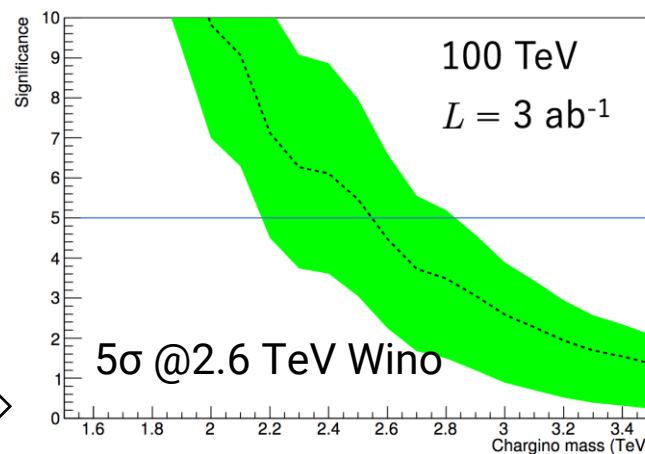
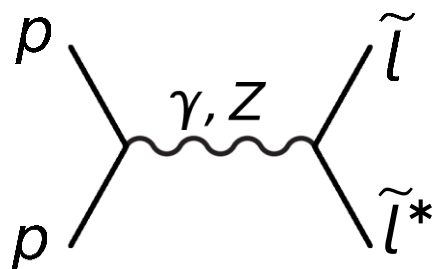
➤ Two extras:

- muon radiative energy loss
- LLCP momentum resolutions

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➤ **Two extras:**

- **muon radiative energy loss**
- **LLCP momentum resolutions**

# Muon energy loss in matter

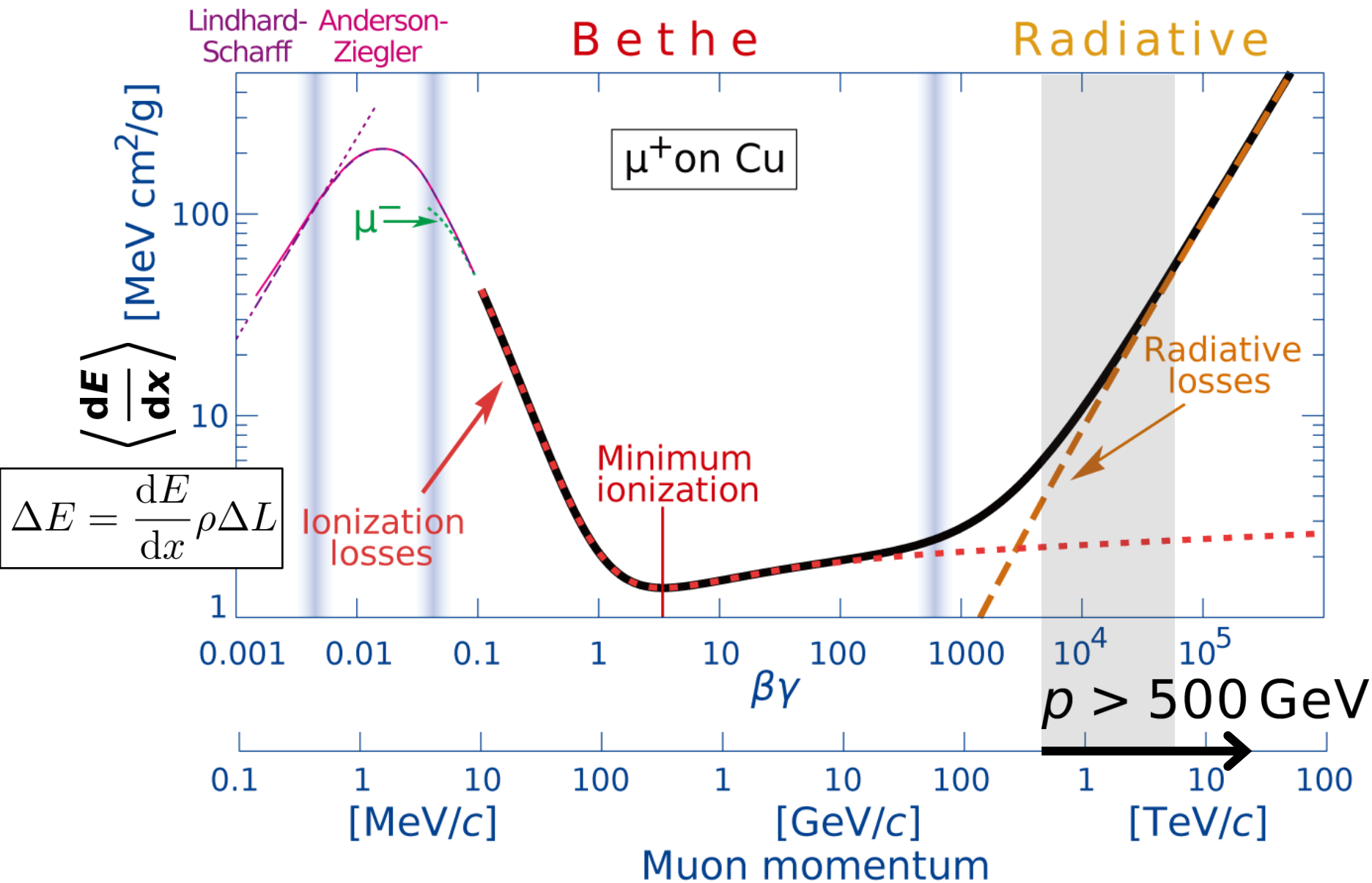
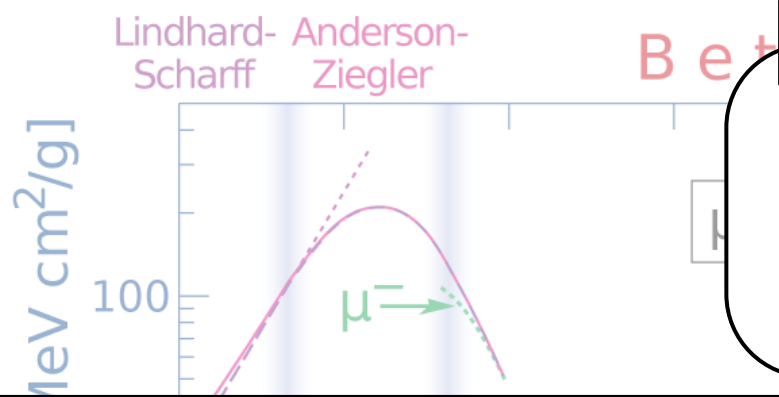


Figure from Groom, Mokhov, Striganov, Atom. Nucl. Data Tab. **78** (2001) 183-356  
 [also in PDG Review "Passage of particles through matter"]



# Muon energy loss in matter



## Muon radiative energy loss

- Bremsstrahlung
- Photonuclear interaction
- $e^+ - e^-$  pair-production

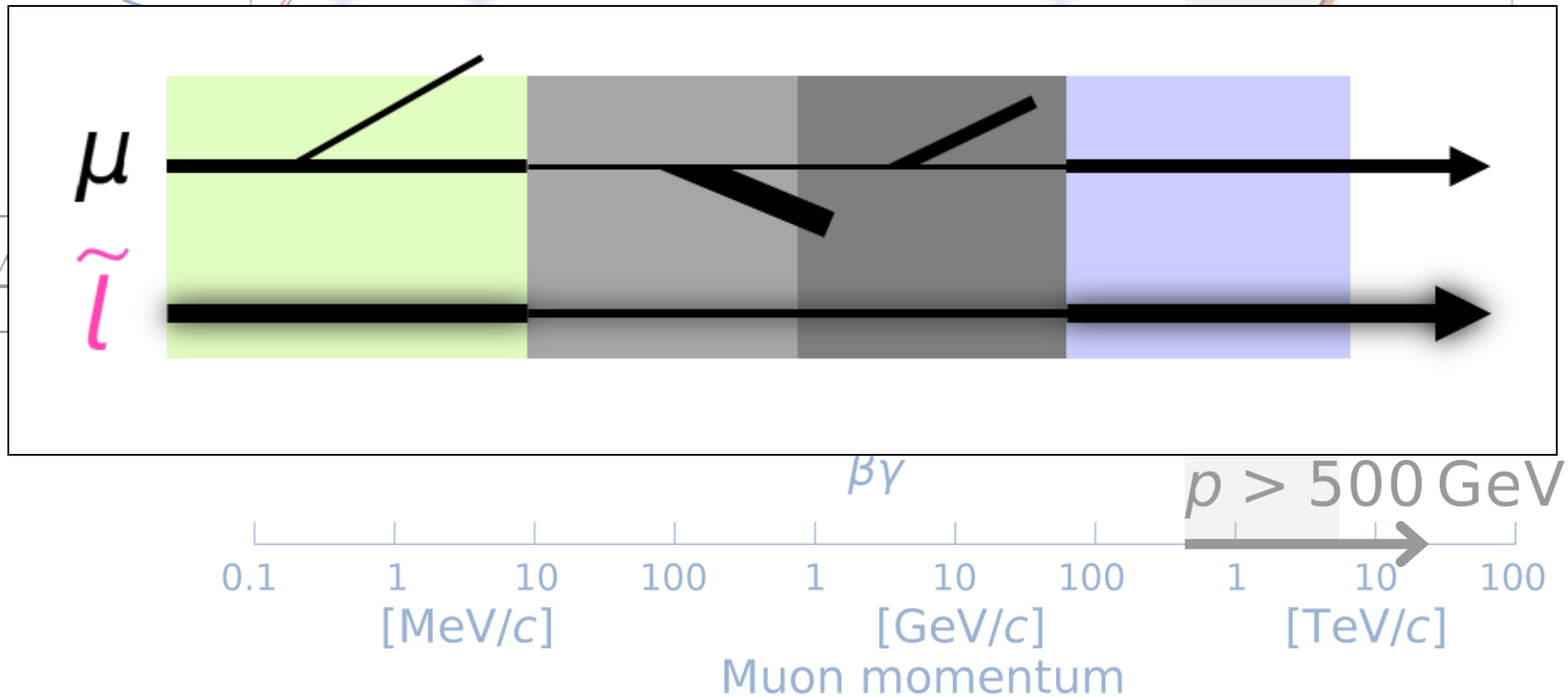
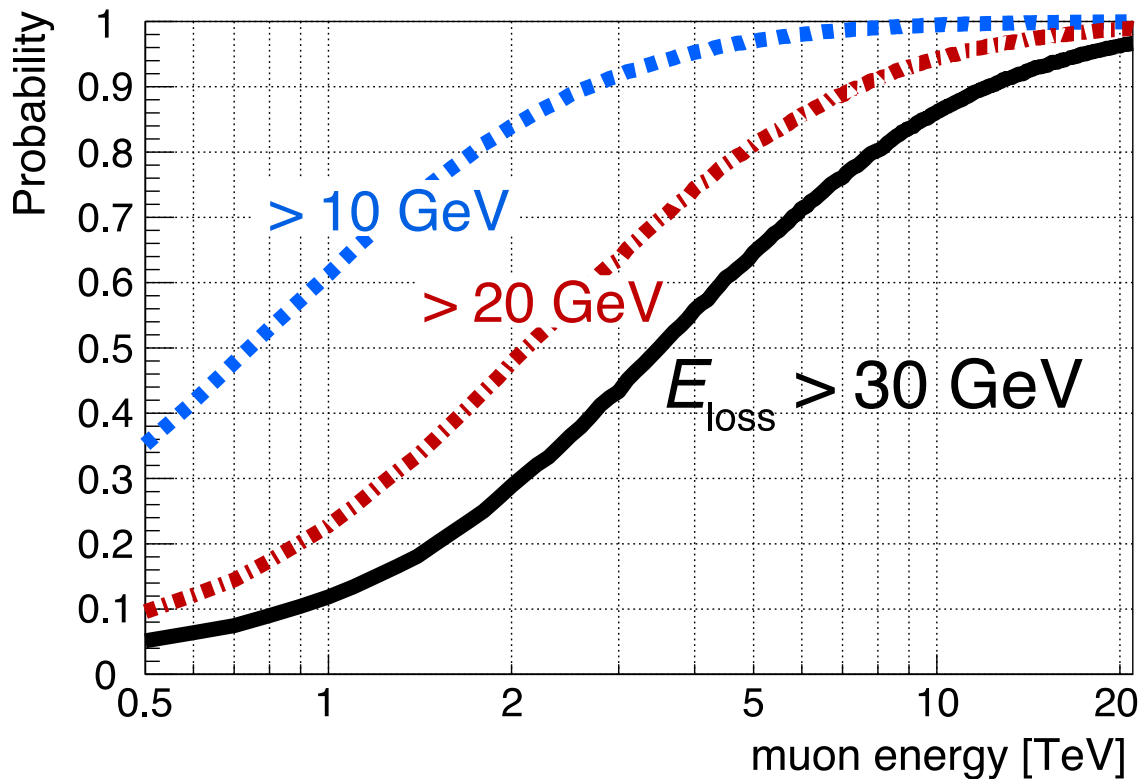
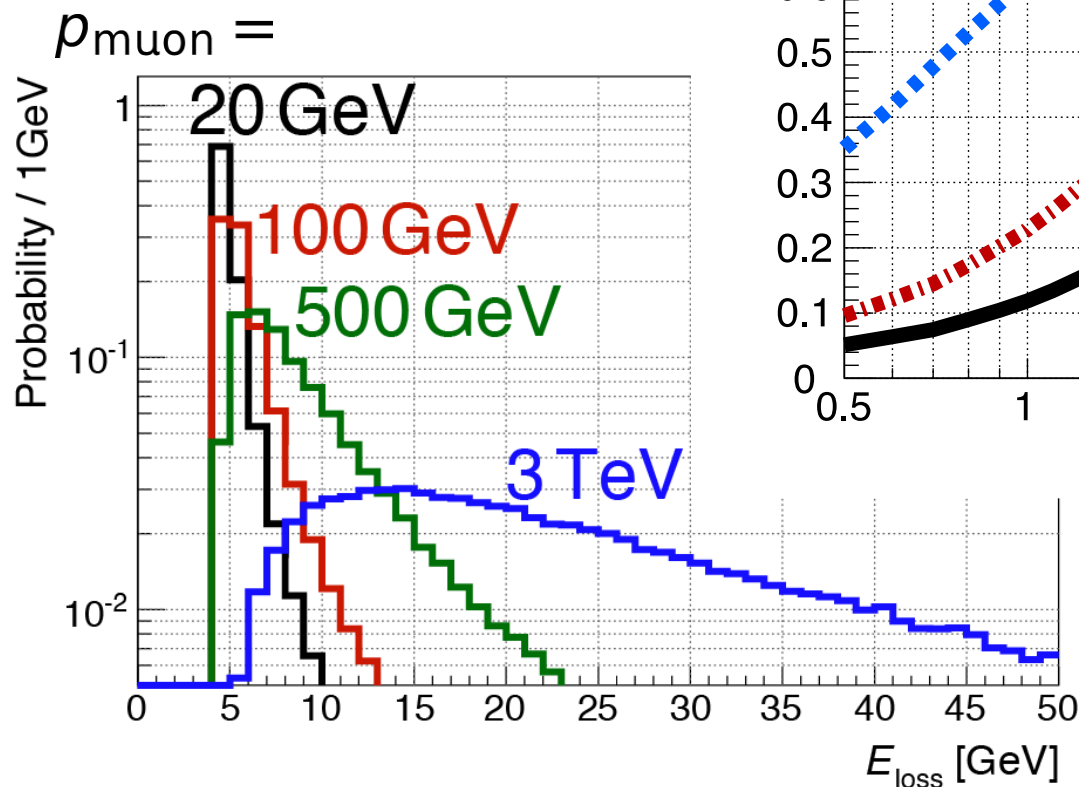


Figure from Groom, Mokhov, Striganov, *Atom. Nucl. Data Tab.* **78** (2001) 183-356  
 [also in PDG Review "Passage of particles through matter"]

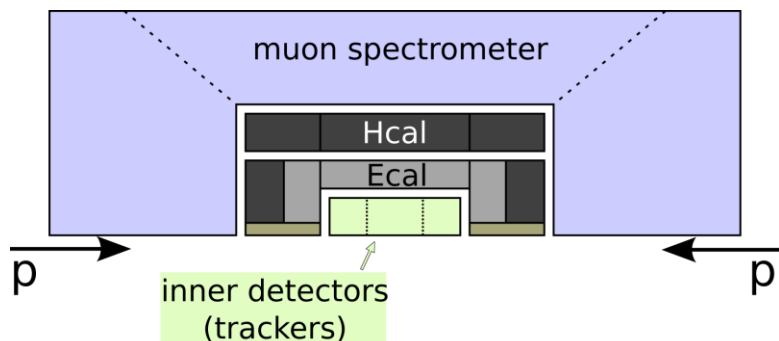
"calorimeter": approximated by iron (Fe) with 3m thickness.

→ some of  $\mu$  ( $P_T > 500$  GeV):  $> 30$  GeV energy deposit.



[Simulated with GEANT 4]

- Detector
  - similar to ATLAS/CMS
  - $\beta$ -resolution same as ATLAS (resolution: 2.4%)
- Signal: Madgraph5 + Pythia6 + Delphes3 (calculated at the LO)
- BKG: “Snowmass 2013” BKG set for 100TeV
- Pile-up not considered



## ■ $\tilde{l}$ -selection flow

$\tilde{l}$  = reconstructed “muon” with

- $P_T > 500$  GeV
- $|\eta| < 2.4$
- $0.4 < \hat{\beta} < 0.95$  (from ToF)
- $E_{\text{loss}} < 30$  GeV

## ■ Event selection

- two  $\tilde{l}$ -candidates

LLCP selection flow ( $\int L = 1 \text{ ab}^{-1}$ )

	signal		SM BKG
	$\tilde{l} = 1 \text{ TeV}$	3 TeV	
total	2570	31.8	—
$p_T$ & $\eta$	1840	28.5	$9.19 \times 10^6$
$\beta$	1230	24.6	$3.41 \times 10^5$
$E_{\text{loss}}$	1230	24.6	$2.78 \times 10^5$
$\epsilon_{\text{acc}}\epsilon_{\text{eff}}$	48%	77%	—

$E_{\text{loss}}$  reduces **34%** of BKG  
 ( $\because 0.82^2 = 0.66$ )

$\times 0.82$

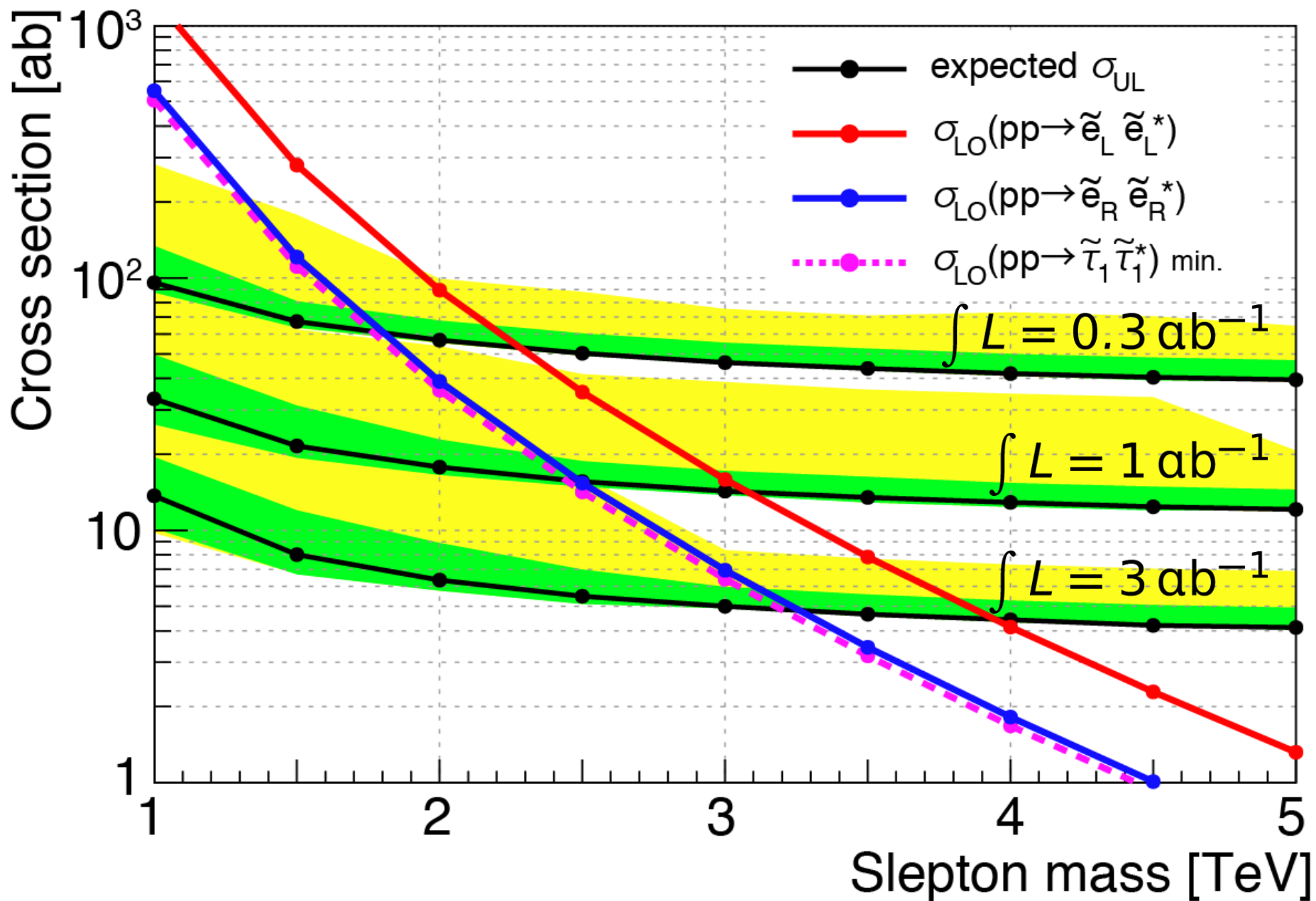
- $|\eta| < 2.4$
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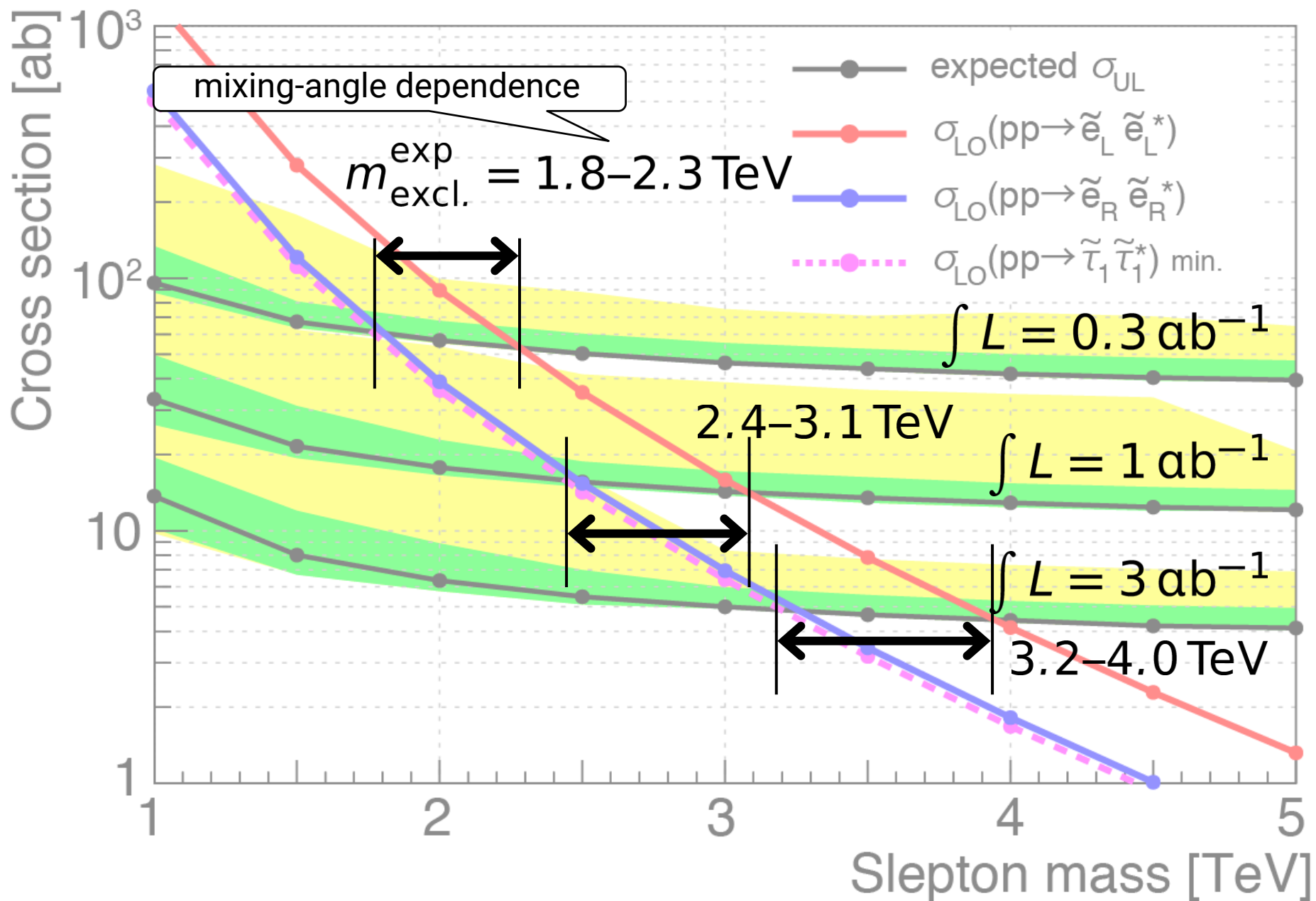
Event categorization ( $\int L = 1 \text{ ab}^{-1}$ )

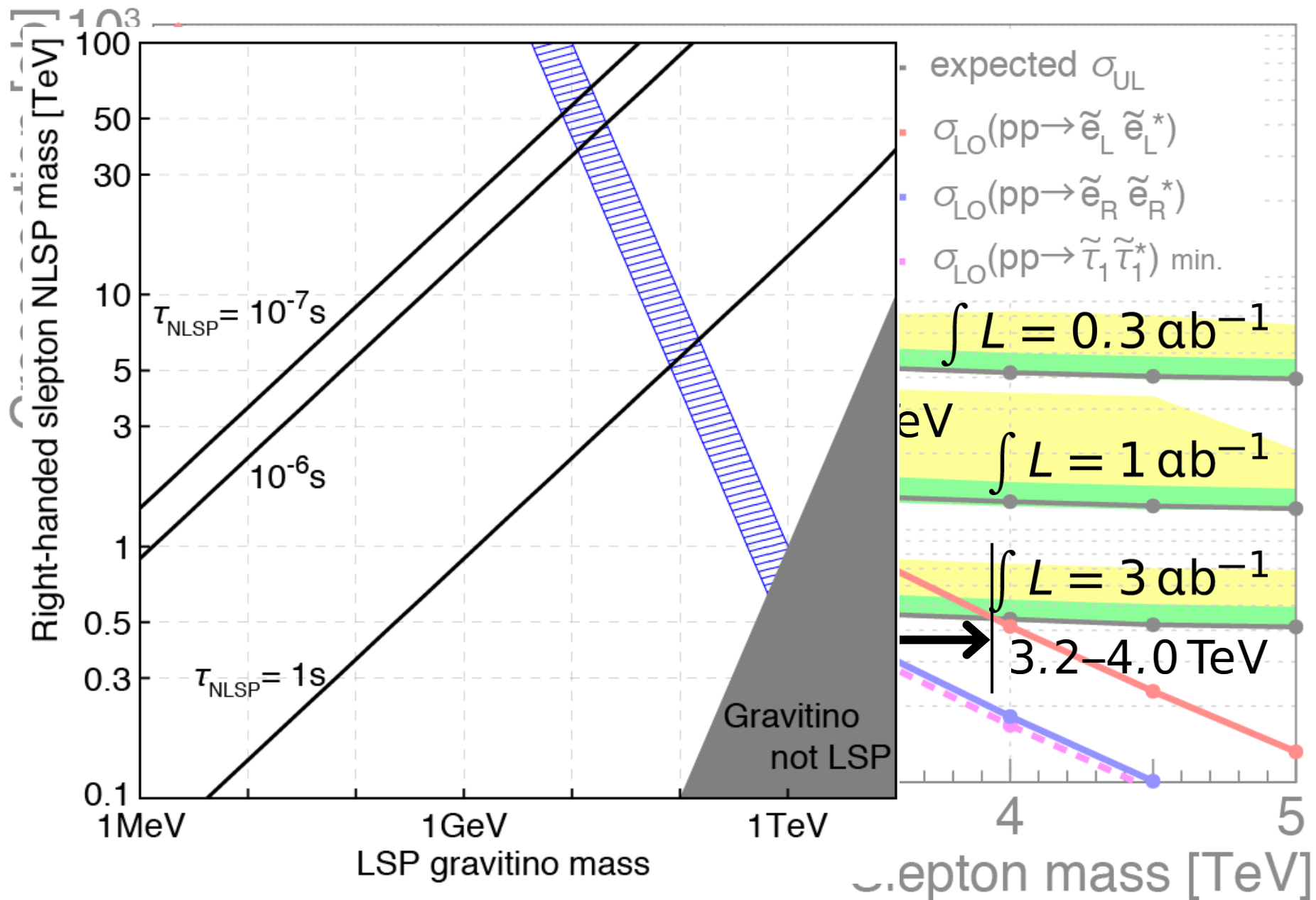
	1 TeV	3 TeV	BKG
$N_{\text{LLCP}} = 0$	483	1.34	(a lot)
$N_{\text{LLCP}} = 1$	378	4.46	$2.78 \times 10^5$
$N_{\text{LLCP}} = 2$	424	10.1	34.6

SR

- Event selection
  - two  $\tilde{l}$ -candidates



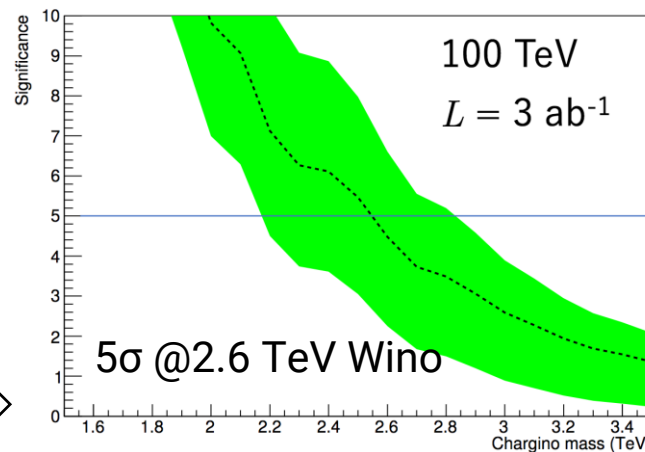
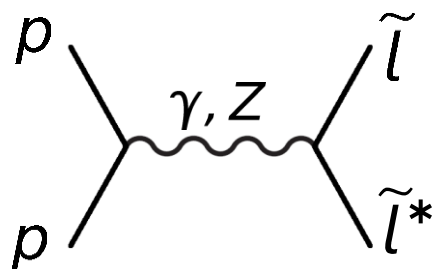




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[Ryu Sawada's talk in FCC week 2017]

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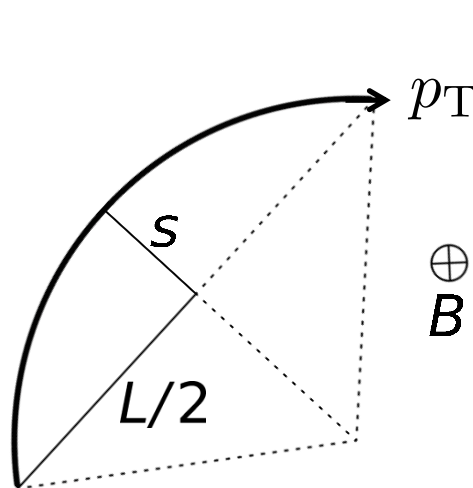
➤ **Two extras:**

- muon radiative energy loss

● **LLCP momentum resolutions**

- $dE/dx$  (ionization energy loss)
- $\Delta L/\Delta t$  (time-of-flight)





$$p_T = \frac{L^2}{8s} |q| B \implies \Delta p_T = \frac{L^2 |q| B}{8} \frac{\Delta s}{s^2}$$

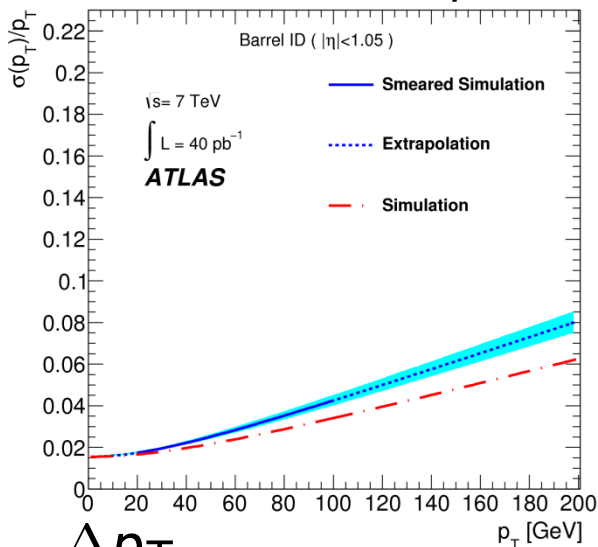
$$= \frac{8\Delta s}{L^2 |q| B} \cdot p_T^2$$

$$\therefore \Delta p_T \propto p_T^2$$

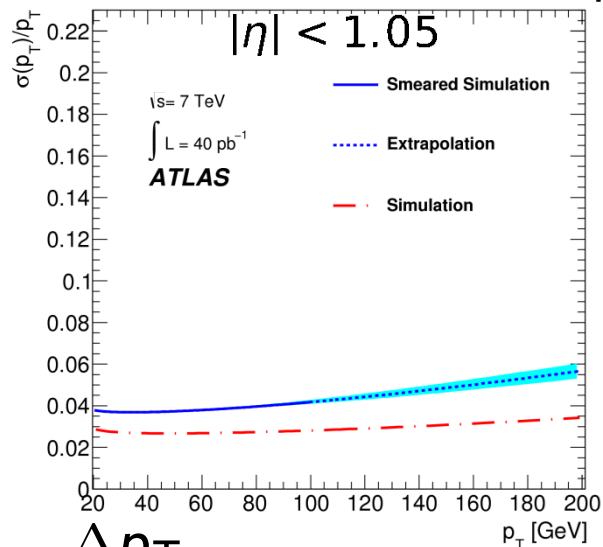
## ATLAS 7 TeV results on muon momentum resolution

Inner Detector,  $|\eta| < 1.05$

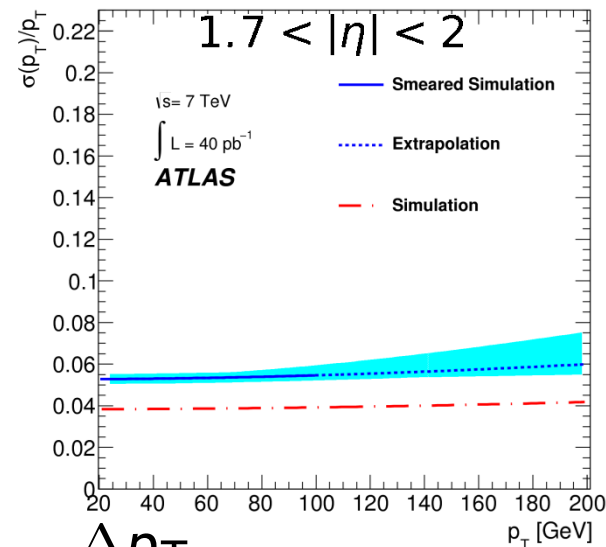
Muon Spectrometer



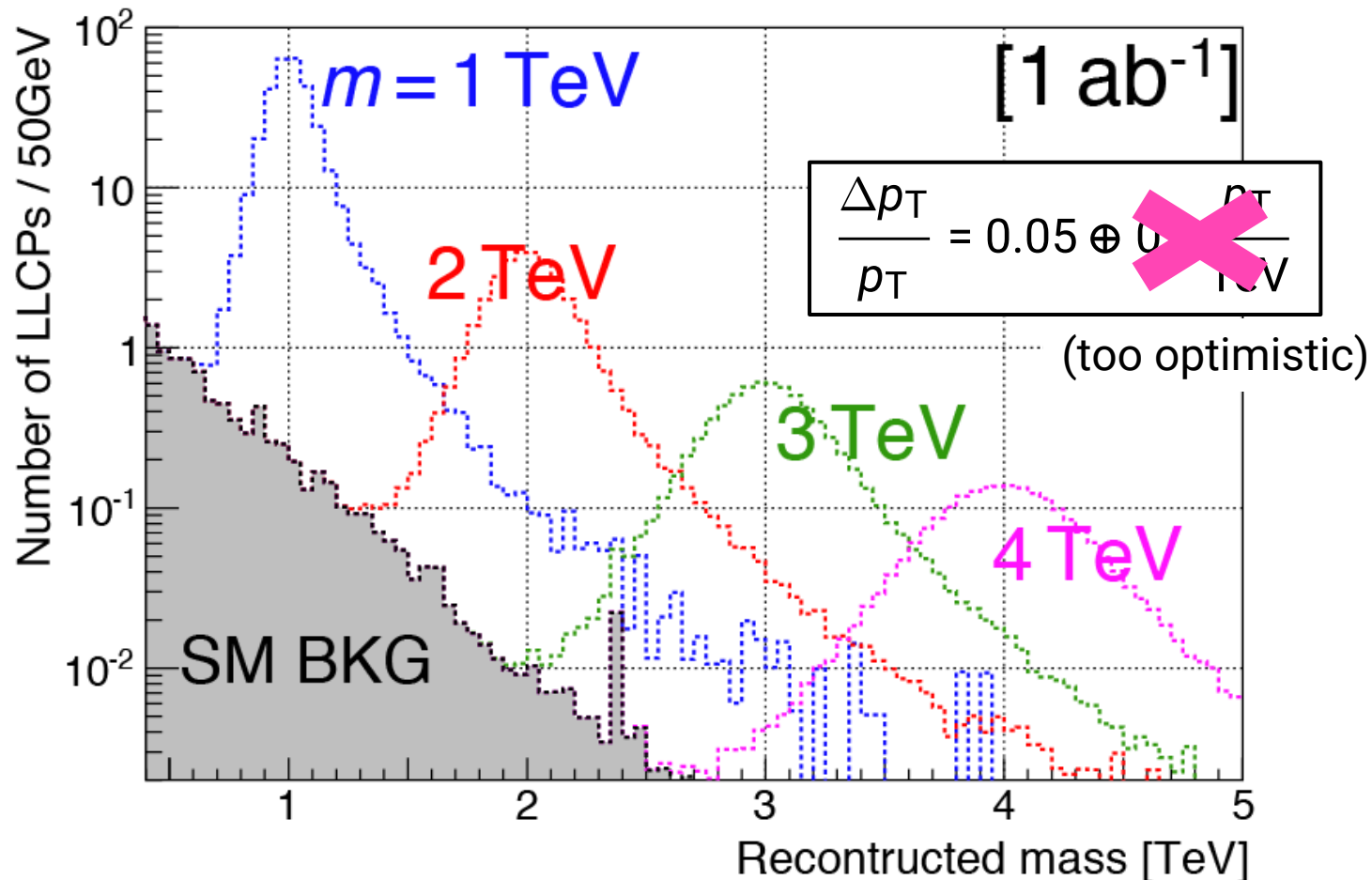
$$\frac{\Delta p_T}{p_T} \sim 0.38 p_T / \text{TeV}$$



$$\frac{\Delta p_T}{p_T} \sim 0.14 p_T / \text{TeV}$$



$$\frac{\Delta p_T}{p_T} \sim 0.06 p_T / \text{TeV}$$



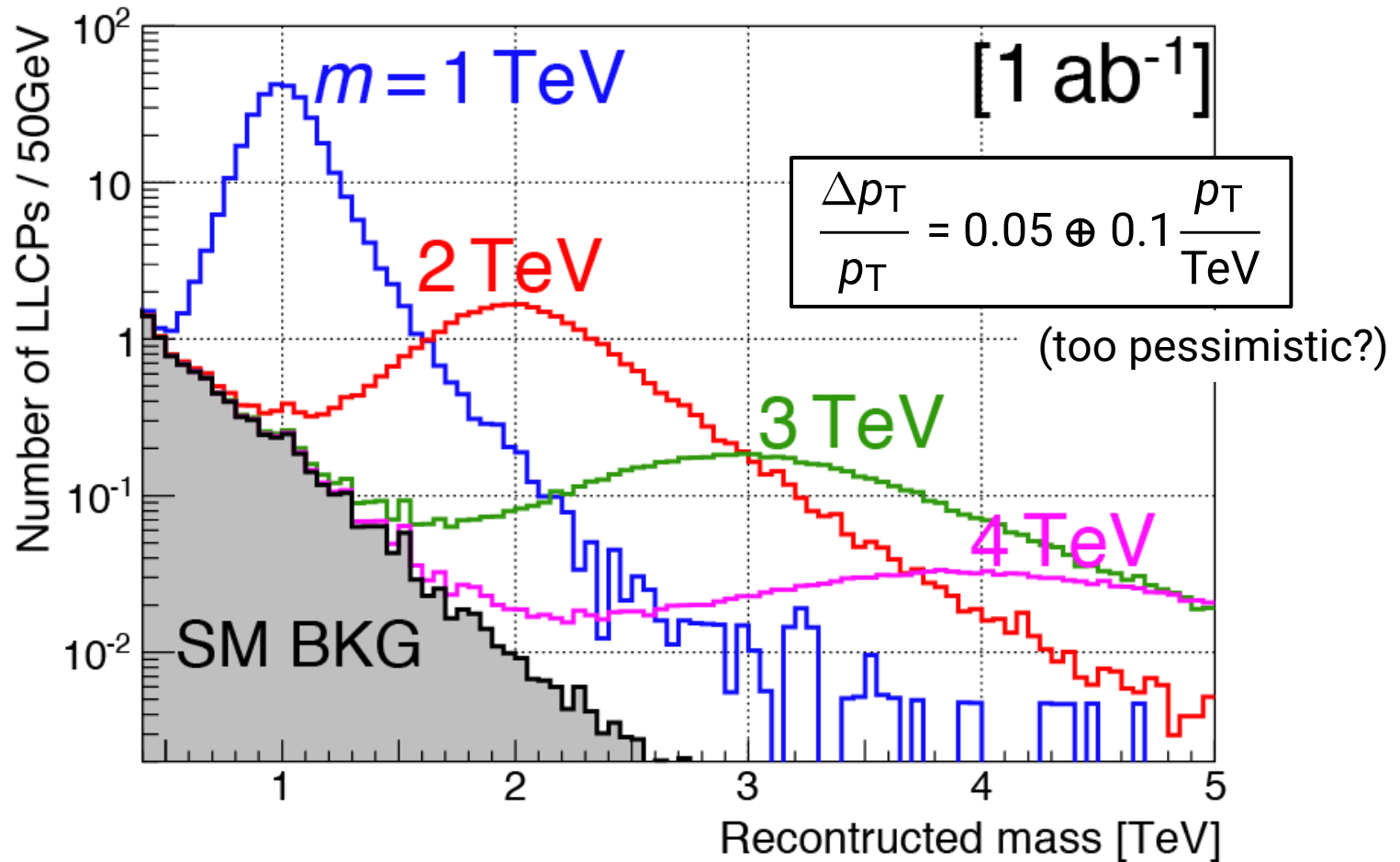
FCC-hh trk. goal: 10–20% @ 10 TeV

([Michele Selvaggi's talk](#) in FCC physics workshop)

cf. ATLAS 7 TeV commissioning:

(ID-barrel, MS-barrel, MS-extbarrel) = (38%, 14%, 6%) @ 1 TeV

$$\left( = \frac{p}{\beta\gamma} = \frac{p_T \cosh \eta}{\beta\gamma} \right)$$



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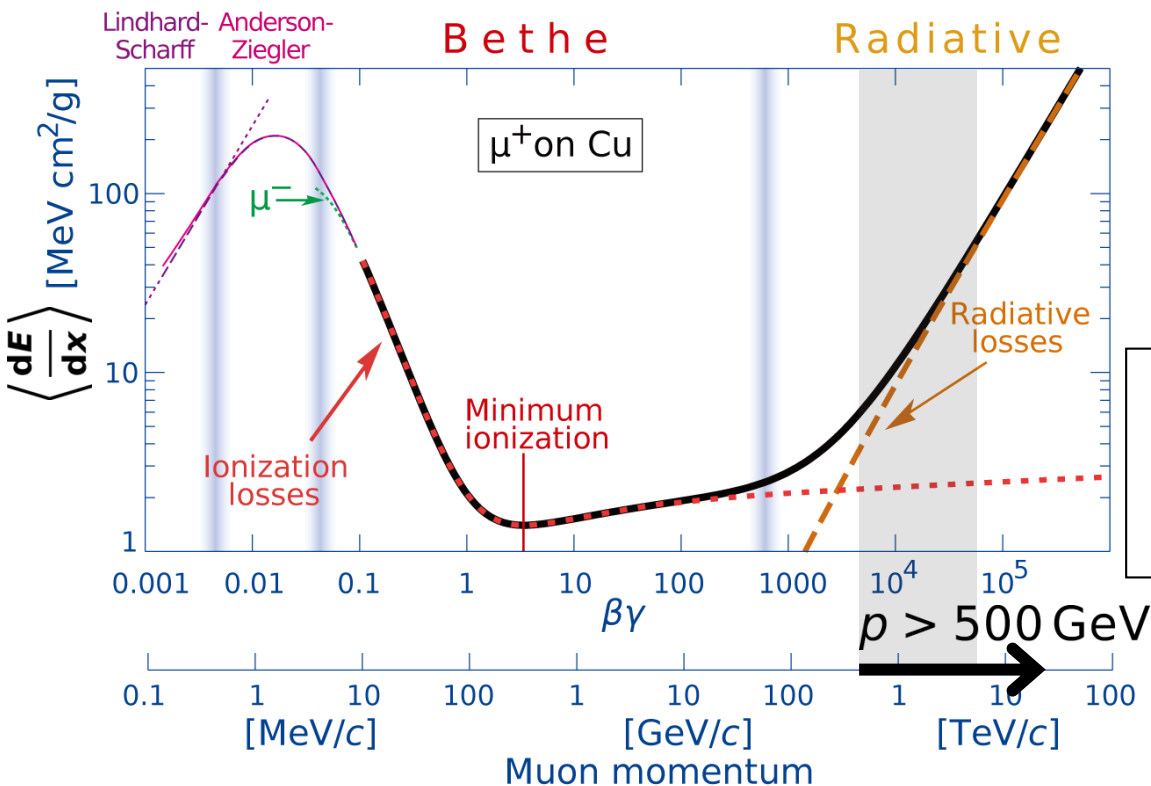
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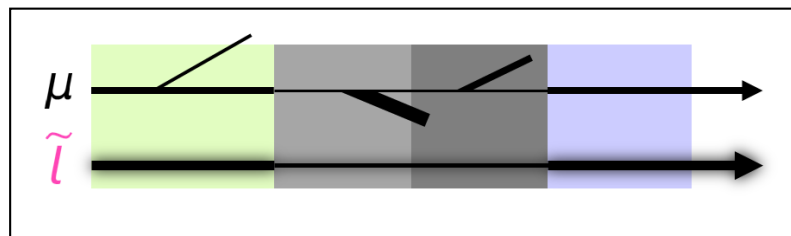
- 100 TeV FCC-hh  
mass reach  
(Drell-Yan  $\tilde{l}$  or  $\tilde{\tau}$ )

	0.3ab <sup>-1</sup>	1ab <sup>-1</sup>	3ab <sup>-1</sup>	
Exclusion	1.8–2.3	2.4–3.1	3.2–4.0	
Discovery	1.6–2.2	2.3–3.0	3.1–4.0	in TeV

## “Muon radiative energy loss”



- Bremsstrahlung
- Photonuclear interaction
- $e^+e^-$  pair-production



→ 34% of BKG reduction

# 1. LLCPs

- Introduction & Motivations
- How to search? & Current bounds

# 2. Future colliders (FCC-hh and FCC-he)

# 3. LLCP searches at FCC-hh

# 4. LLCP searches at FCC-he

Based on

**hh**: Jonathan L. Feng, SI, Yael Shadmi, Shlomit Tarem [[1505.02996](#)]

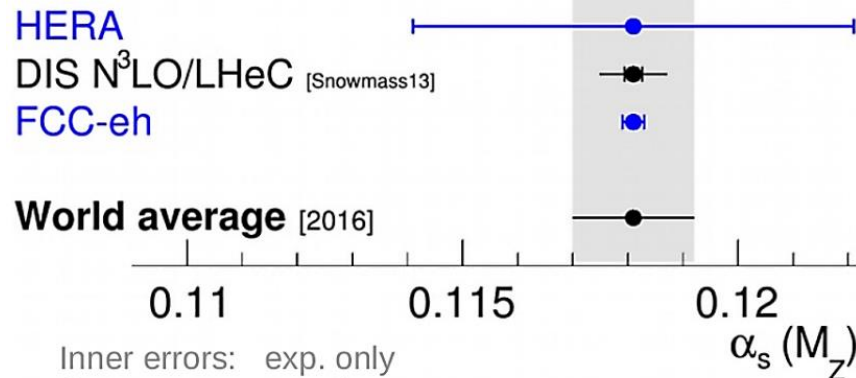
(collected in FCC-hh report [[1606.00947](#)])

**he**: Kechen Wang, SI, Monica D'Onofrio, Georges Azuelos [17???.?????]

(subgroup in BSM@ep collaboration)

■ FCC-he targets:

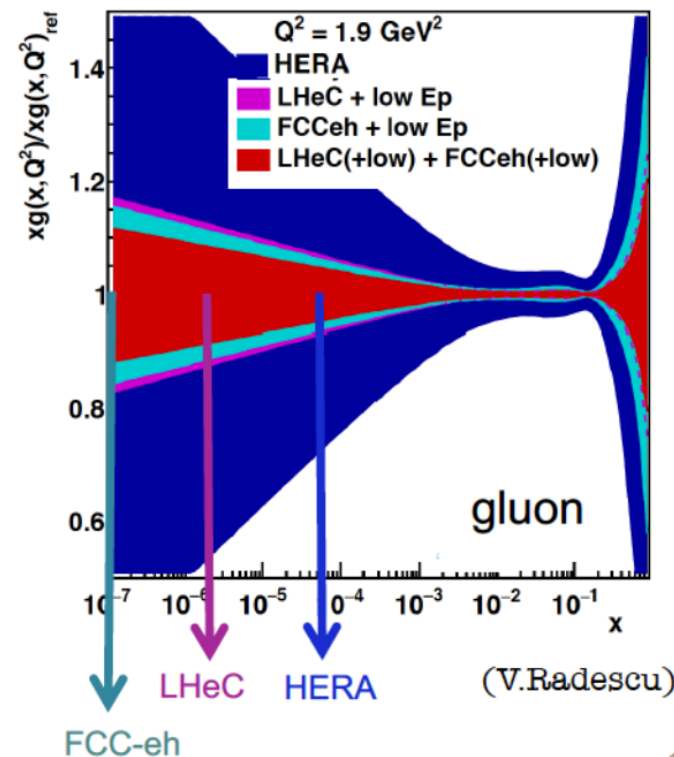
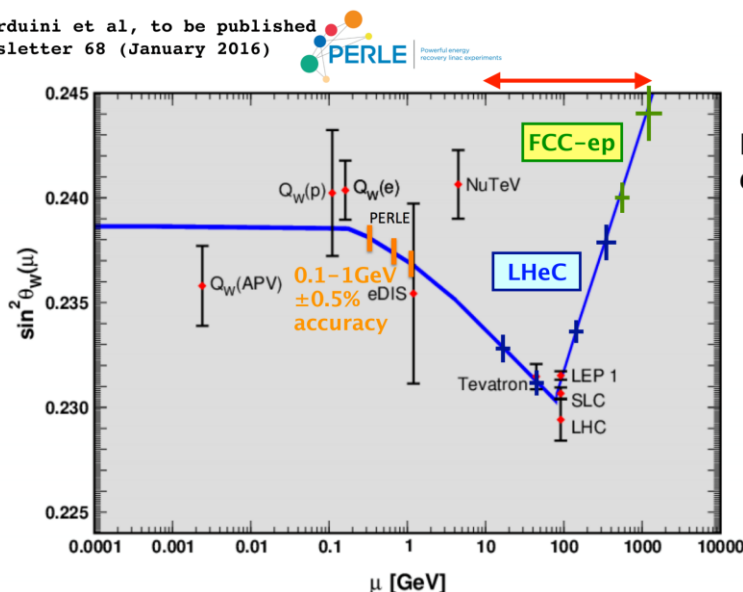
- PDFs & strong coupling
- Higgs & Electroweak physics
- QCD (heavy quark PDFs)
- low-x physics (non-linear QCD?) :  $x < 10^{-6}$



Inner errors: exp. only  
Outer errors: exp+theo.

Jan Kretzschmar, 11.9.2017

PERLE CDR, Arduini et al, to be published  
ICFA BeamNewsletter 68 (January 2016)




■ What's MORE?

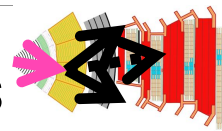
Any power to New Physics? → BSM ep team

## ■ BSM ep team

## ★ Direct Searches

- ◆ Leptoquarks: limits, quantum # & couplings
- ◆ Contact interactions:  $eeqq$
- ◆ Anomalous gauge couplings:  $vvv$
- ◆ Vector boson scattering
- ◆ BSM in the top sector
- ◆ RPC SUSY: DM, sleptons 
- ◆ RPV SUSY: neutralinos, squarks
- ◆ BSM Higgs: exotic (invisible) decay;  $H^+$ ,  $H^{++}$
- ◆ Sterile neutrinos

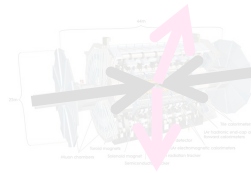
[from [a talk](#) by Ke Chen Wang @ [FCC week 2017](#)]



■ Dark Matter

➤ co-annihilation

- $(\tilde{B} - \tilde{\tau}) \lesssim 700 \text{ GeV}$
- $(\tilde{W} - \tilde{g}) \lesssim 6-7 \text{ TeV}$
- $(\tilde{B} - \tilde{g})$  or  $(\tilde{B} - \tilde{t}) \lesssim 8 \text{ TeV}$



$\tilde{\tau}$ -LLCP

➤ super-WIMP scenario  
→  $\tilde{l} > O(1) \text{ TeV}$

$\tilde{l}$ -LLCP

■ Li problem

➤ MSSM  $\tilde{\tau}$  with

$m_{\tilde{\tau}} \sim 400 \text{ GeV}$   
 $\Delta(m_{\tilde{\tau}} - m_{\text{LSP-DM}}) \sim 100 \text{ MeV}$

$\tilde{\tau}$ -LLCP

■ Simple MSSM models

- Pure-Wino dark matter **60mm**
- Pure-Higgsino dark matter **7mm**



$\tilde{\chi}^{\pm}$  long-lived because of small  $\delta m$   
 ( $\delta m > 0$ )

(but **DM underabundant**  
 for “observable” region)

■ Other MSSM models

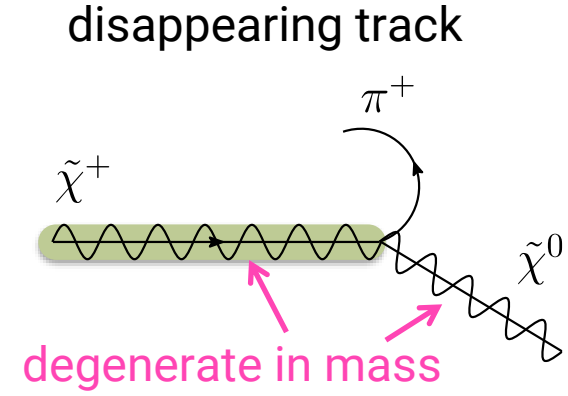
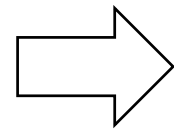
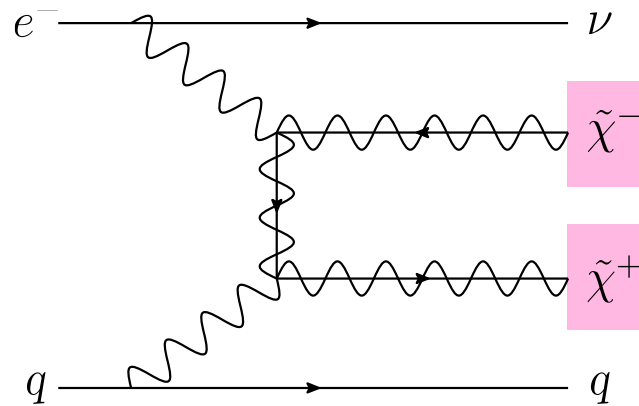
- Gauge-Mediation (keV  $\tilde{G}$ )
- R-parity violation  
 with  $\tilde{l}$  -LSP

⇒  $\tilde{l}$  long-lived **any  $c\tau$**   
 because of tiny couplings.

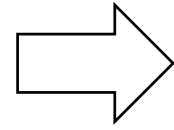
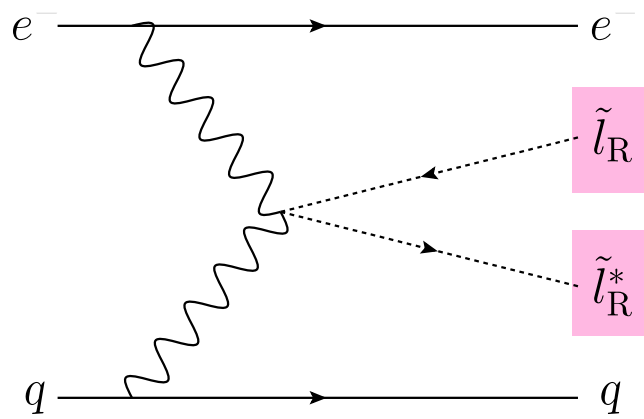


■ Simplest models: 4-body production;  $\sigma < 1 \text{ fb} \dots (\cdot \omega \cdot \cdot)$

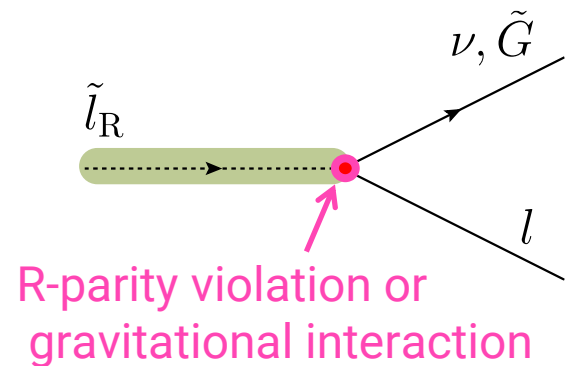
➤ Pure-Wino / Pure-Higgsino LSP



➤ Slepton LSP

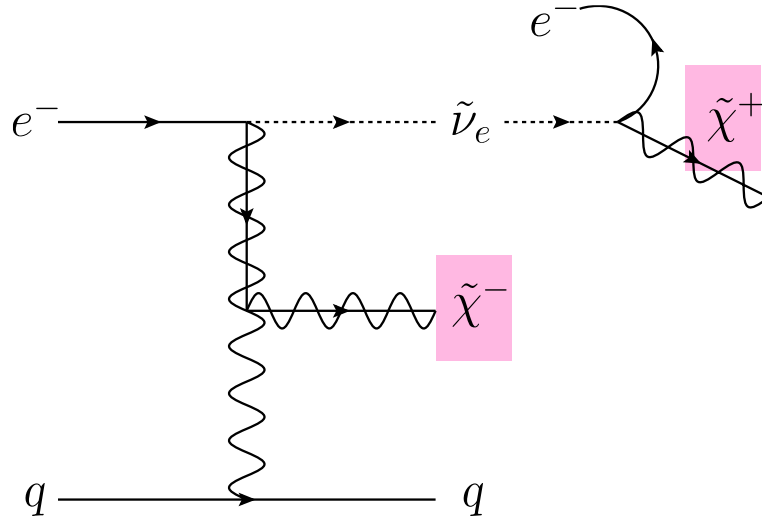


disappearing track (or "kink")

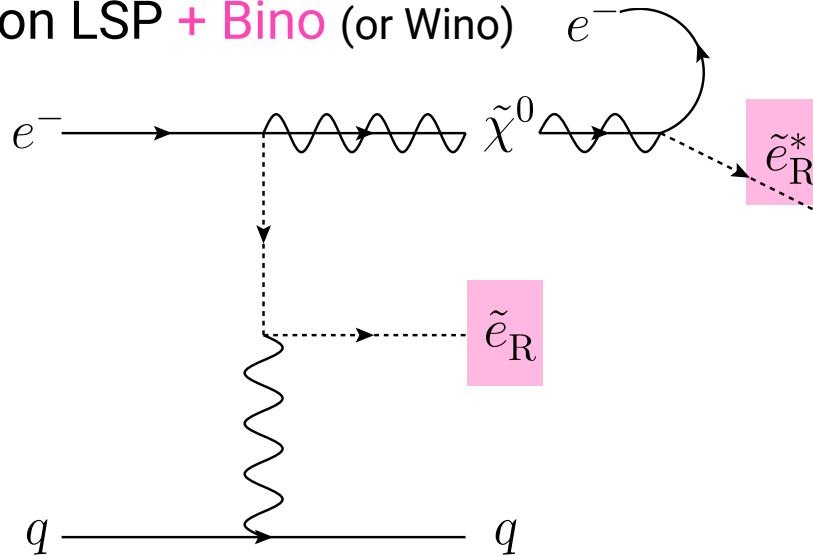


■ Introducing co-LSP allows **3-body** production

- ~~Pure-Wino / Pure-Higgsino LSP~~ + left-handed selectron

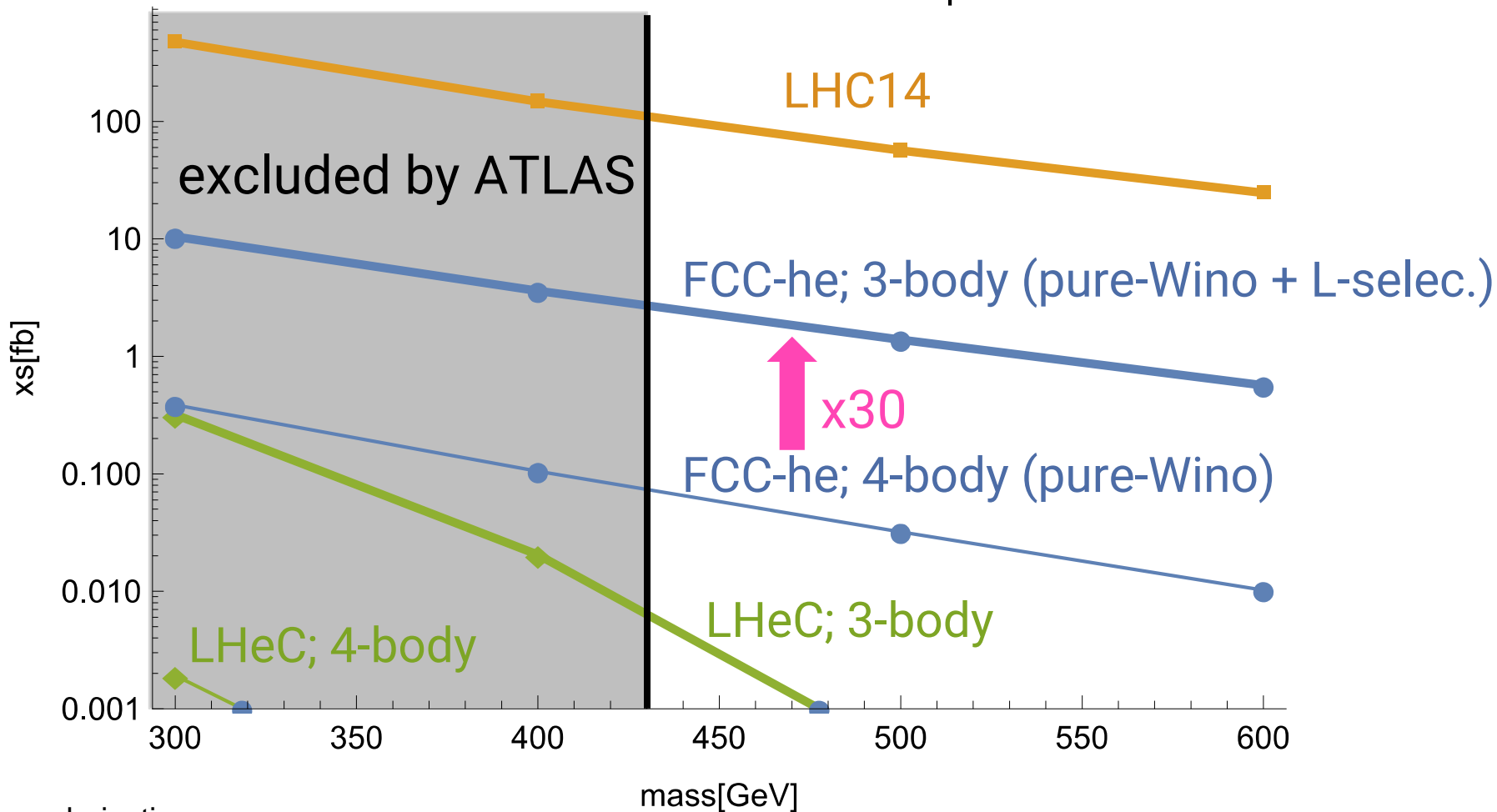


- Slepton LSP + Bino (or Wino)



■ Nominal production cross section (without acceptances / efficiencies)

Wino LSP scenarios with/without co-production



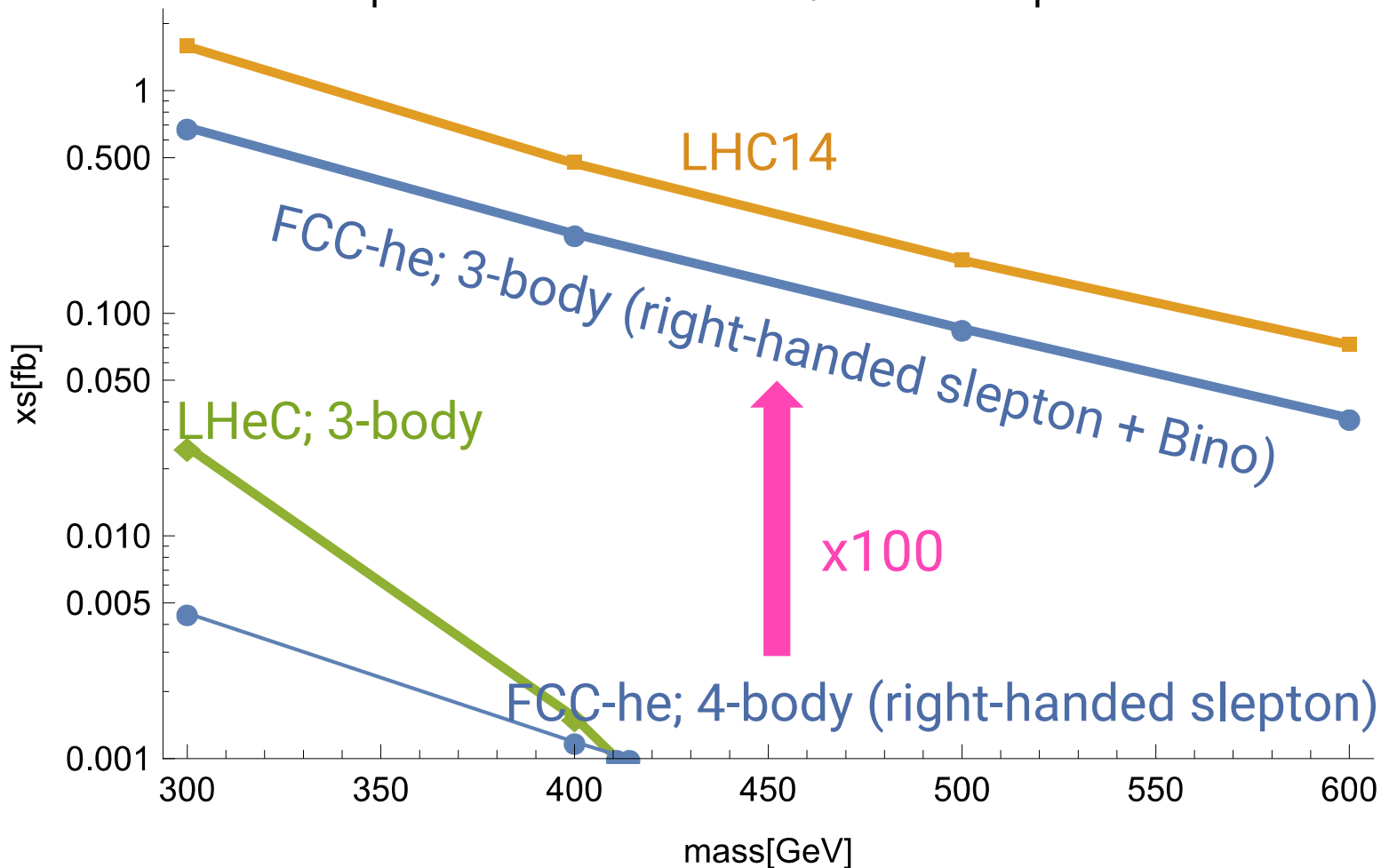
With no polarization.

Shaded region is excluded by ATLAS (13TeV, 36/fb)

“3-body” model assumes  $m_{\tilde{e}_L} = m_{\tilde{\chi}_1^0} + 9 \text{ GeV}$

■ Nominal production cross section (without acceptances / efficiencies)

Slepton LSP scenarios with/without co-production

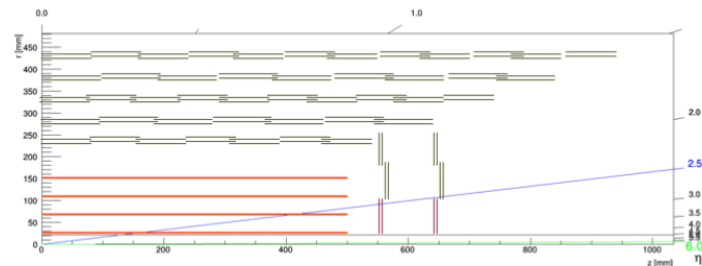


With no polarization.

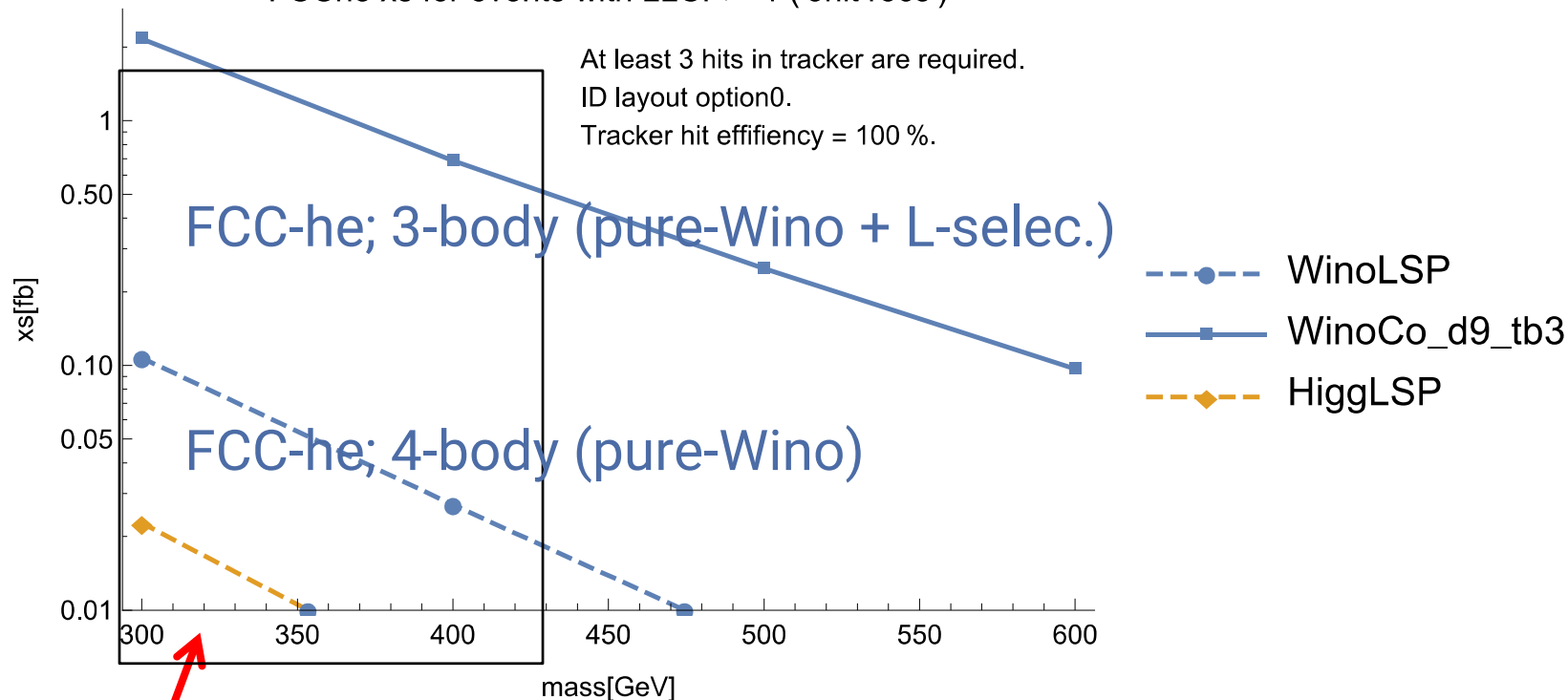
“3-body” model assumes  $m_{\tilde{\chi}_1^0} = m_{\tilde{e}} + 1 \text{ GeV}$

# FCC-he in-flight decay LLCPs with reconstruction efficiency

LLCPs are required to decay after 3 layers of IDs.  
 (= ~ to leave 3 hits in ID)



FCChe xs for events with LLCP  $\geq 1$  ('3hit reco')



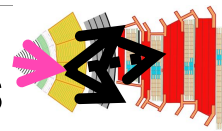
Higgsino-only scenario: less promising because of smaller  $\sigma\tau$ .

With no polarization.

Shaded region is excluded by ATLAS (13TeV, 36/fb)

"3-body" process assumes  $m_{\tilde{e}_L} = m_{\tilde{\chi}_1^0} + 9 \text{ GeV}$ .

- "Wino+slep" model is promising.
- reco eff. is governed by the innermost layers  
 → closer / more-precise layers would help a lot.



**“single SUSY particle”**

→ tiny  $\sigma$  due to 4-body process;  
HL-LHC will be better.

If with another sparticle for  
**3-body “co-production”**

→ FCC-he will be competitive  
with HL-LHC.

- Pure-Wino LSP +  $\tilde{e}_L$
- Slepton LSP + Bino (or Wino)

■ Simple MSSM models

- Pure-Wino dark matter **60mm**
- ~~Pure Higgsino dark matter~~  
~~**7mm**~~

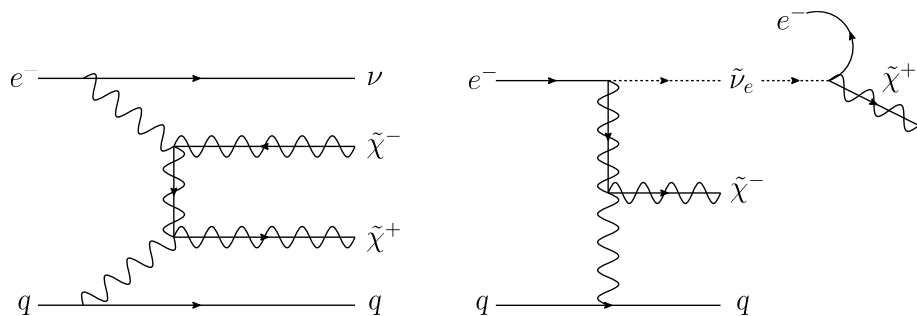
**lifetime too short**

→ trying to improve by  
identifying daughter particles...

■ Other MSSM models

- Gauge-Mediation (keV  $\tilde{G}$ )
- R-parity violation  
with  $\tilde{l}$ -LSP

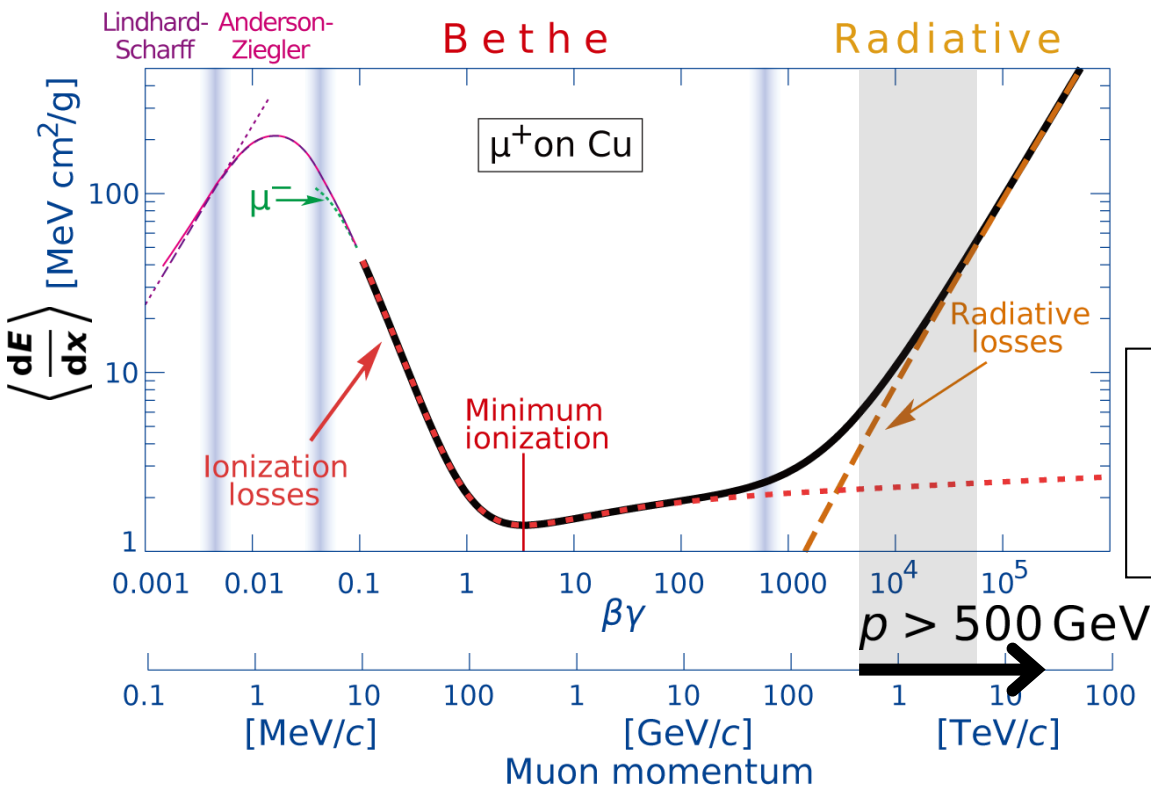
**any  $c\tau$**



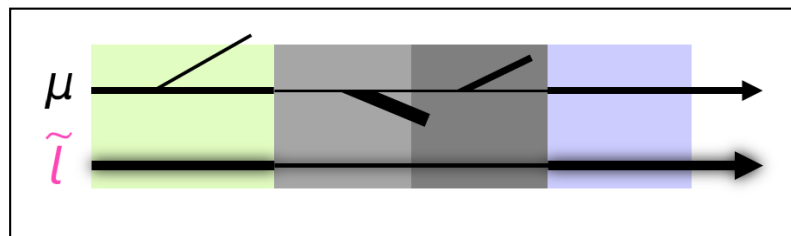
- 100 TeV FCC-hh  
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(Drell-Yan  $\tilde{l}$  or  $\tilde{\tau}$ )

	0.3ab <sup>-1</sup>	1ab <sup>-1</sup>	3ab <sup>-1</sup>	
Exclusion	1.8–2.3	2.4–3.1	3.2–4.0	
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## “Muon radiative energy loss”



- Bremsstrahlung
- Photonuclear interaction
- $e^+e^-$  pair-production



→ 34% of BKG reduction

# Lithium problem

---



# Lithium7-underabundant problem

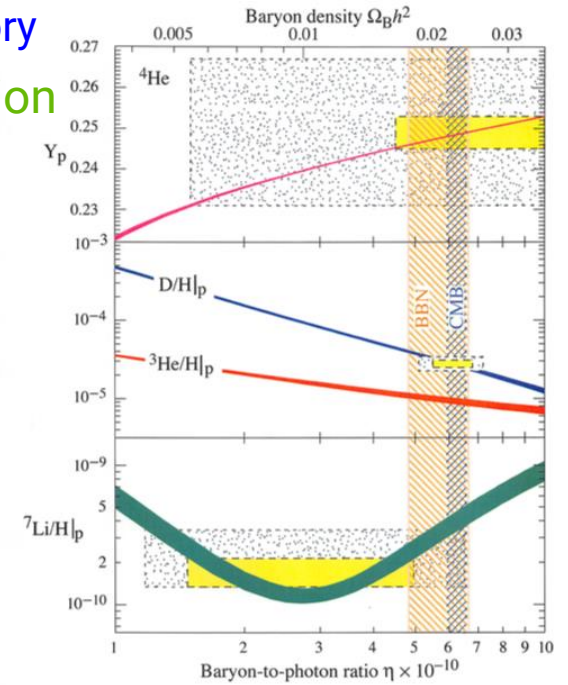
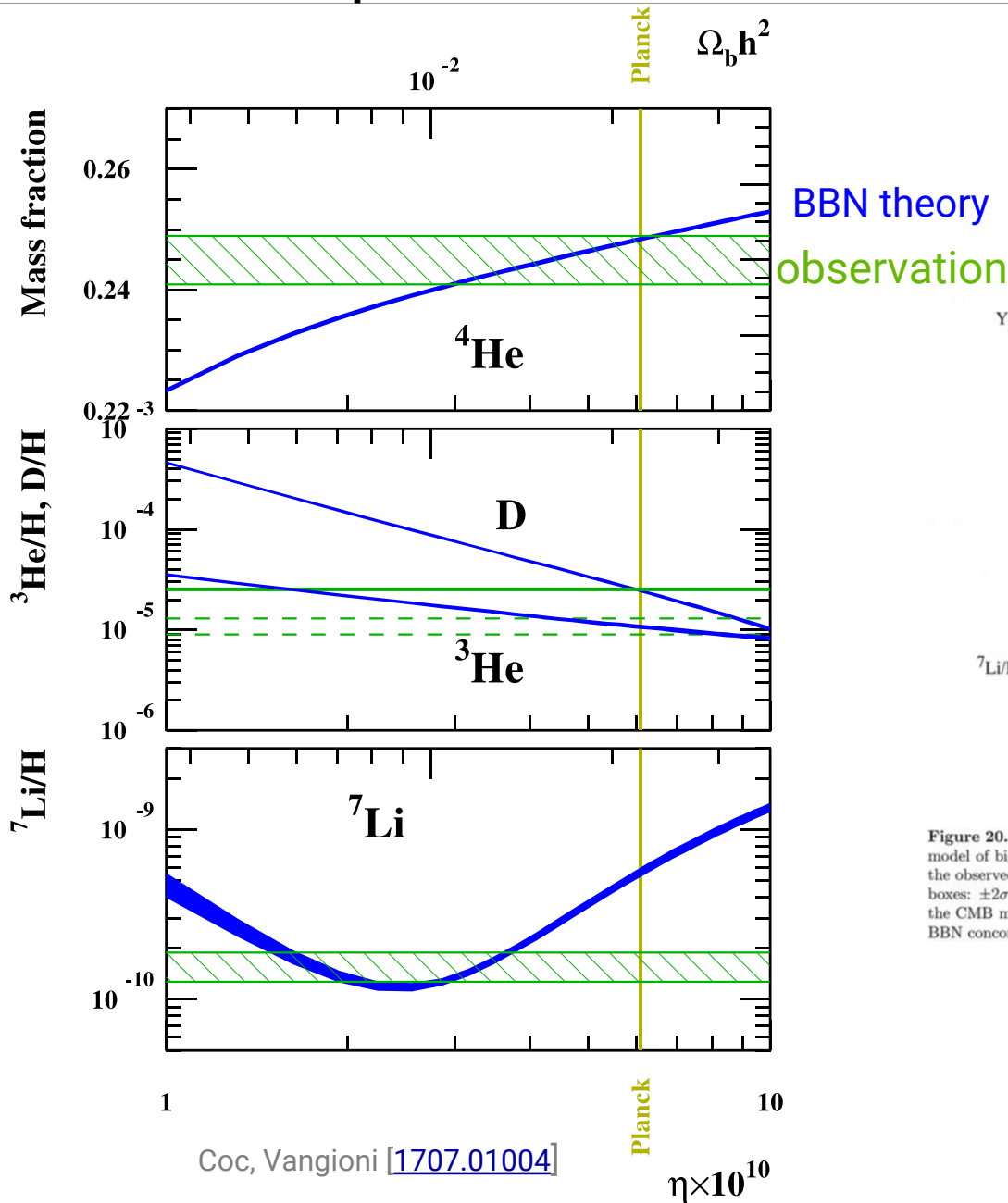


Figure 20.1: The abundances of  $^4\text{He}$ , D,  $^3\text{He}$  and  $^7\text{Li}$  as predicted by the standard model of big-bang nucleosynthesis — the bands the 95% CL range. Boxes indicate the observed light element abundances (smaller boxes:  $\pm 2\sigma$  statistical errors; larger boxes:  $\pm 2\sigma$  statistical and systematic errors). The narrow vertical band indicates the CMB measure of the cosmic baryon density, while the wider band indicates the BBN concordance range (both at 95% CL). Color version at end of book.

## ■ Dark Matter (especially to am...)

### ➤ co-annihilation

- $(\tilde{B} - \tilde{\tau}) \lesssim 700 \text{ GeV}$
- $(\tilde{W} - \tilde{g}) \lesssim 6-7 \text{ TeV}$
- $(\tilde{B} - \tilde{g})$  or  $(\tilde{B} - \tilde{t}) \lesssim 8 \text{ TeV}$

Harigaya, Kaneta, M... [1403.0715],  
Ellis, Olive, Zheng [1404.55...]

## non-colored LLCP

... covered in this talk

(slepton  $\tilde{l}$  / stau  $\tilde{\tau}$  as benchmarks)

### ➤ super-WIMP scenario

→  $\tilde{l} > O(1) \text{ TeV}$

(as we see soon.)

## colored LLCP

(→ hadronize: "R-hadron")



[1606.00947]

## ■ Li problem

### ➤ MSSM $\tilde{\tau}$ with

$m_{\tilde{\tau}} \sim 400 \text{ GeV}$

$\Delta(m_{\tilde{\tau}} - m_{\text{LSP-DM}}) \sim$

### "Extrapolations"

$\tilde{g}$ (-like) R-hadron: 11–20 TeV?

$\tilde{t}$  (-like) R-hadron: 7 TeV

[cf. Barnard, Cox, Gherghetta, Spray, [1510.06405](#)]

→ Dedicated studies after detector specs(?)

Why  $\beta > 0.4$ ? (slepton  $dE/dx$ )

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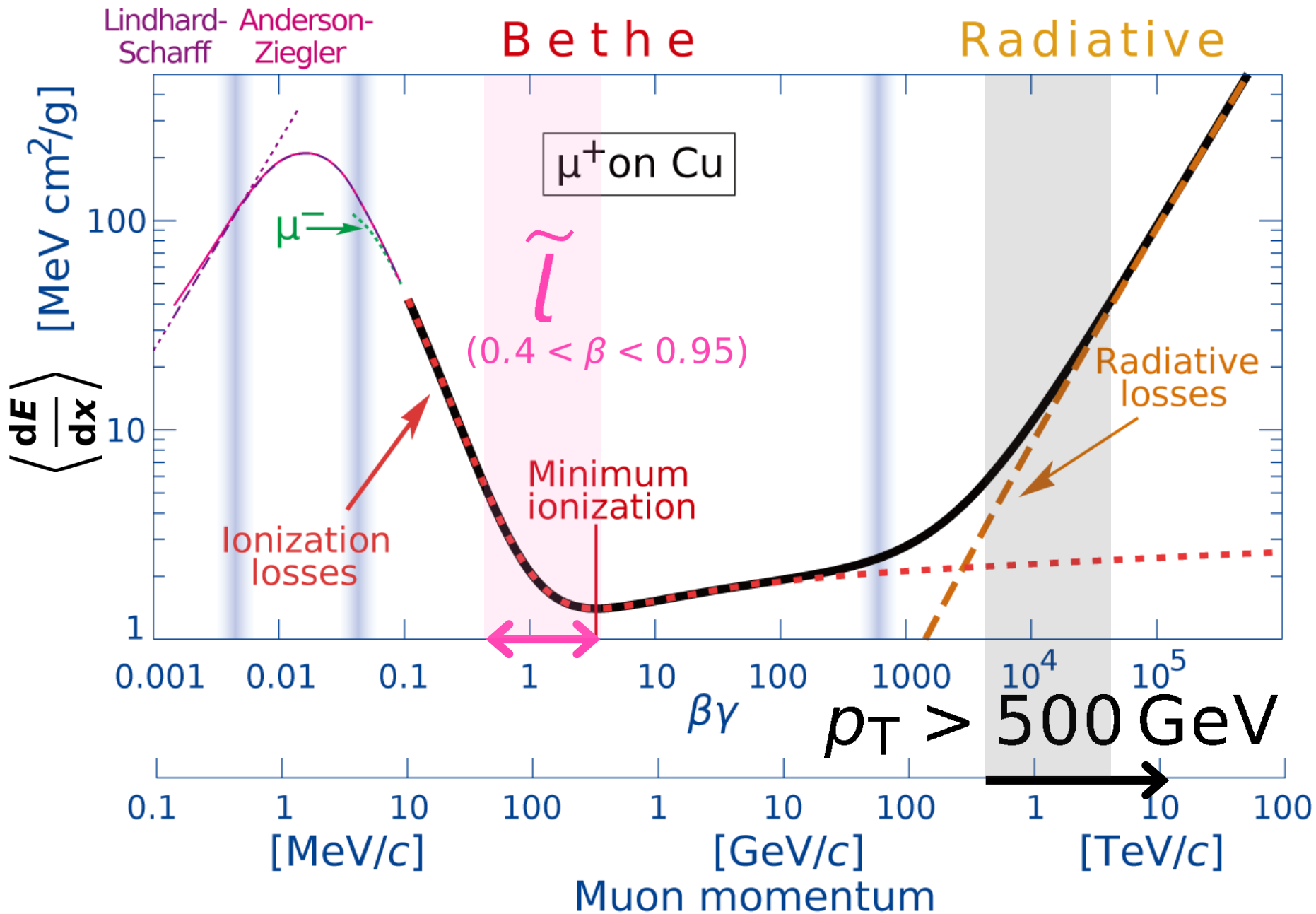


Figure from Groom, Mokhov, Striganov, *Atom. Nucl. Data Tab.* **78** (2001) 183-356  
 [also in PDG Review "Passage of particles through matter"]

# HL-LHC

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## ■ Detector

- similar to ATLAS/CMS
- $\beta$ -resolution same as ATLAS  
(resolution: 2.4%)

## ■ Signal: Madgraph5 + Pythia6 + Delphes3 (calculated at the LO)

## ■ BKG: “Snowmass 2013” BKG set for 14 TeV (publicly available)

## ■ Pile-up not considered

## ■ $\tilde{l}$ -selection flow

reconstructed “muon” w.

- $p_T > 100$  GeV

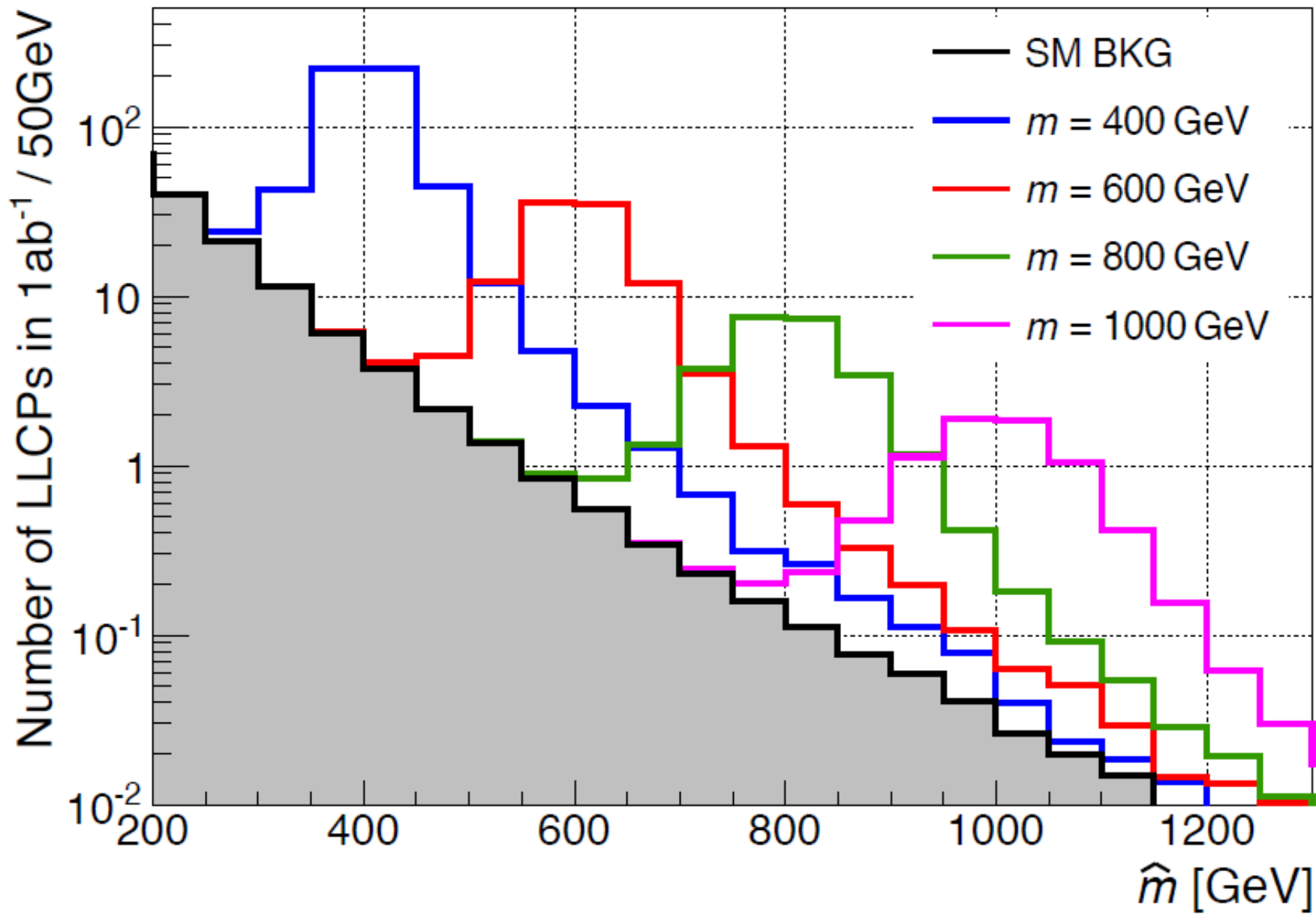
- $|\eta| < 2.4$

- $0.3 < \hat{\beta} < 0.95$

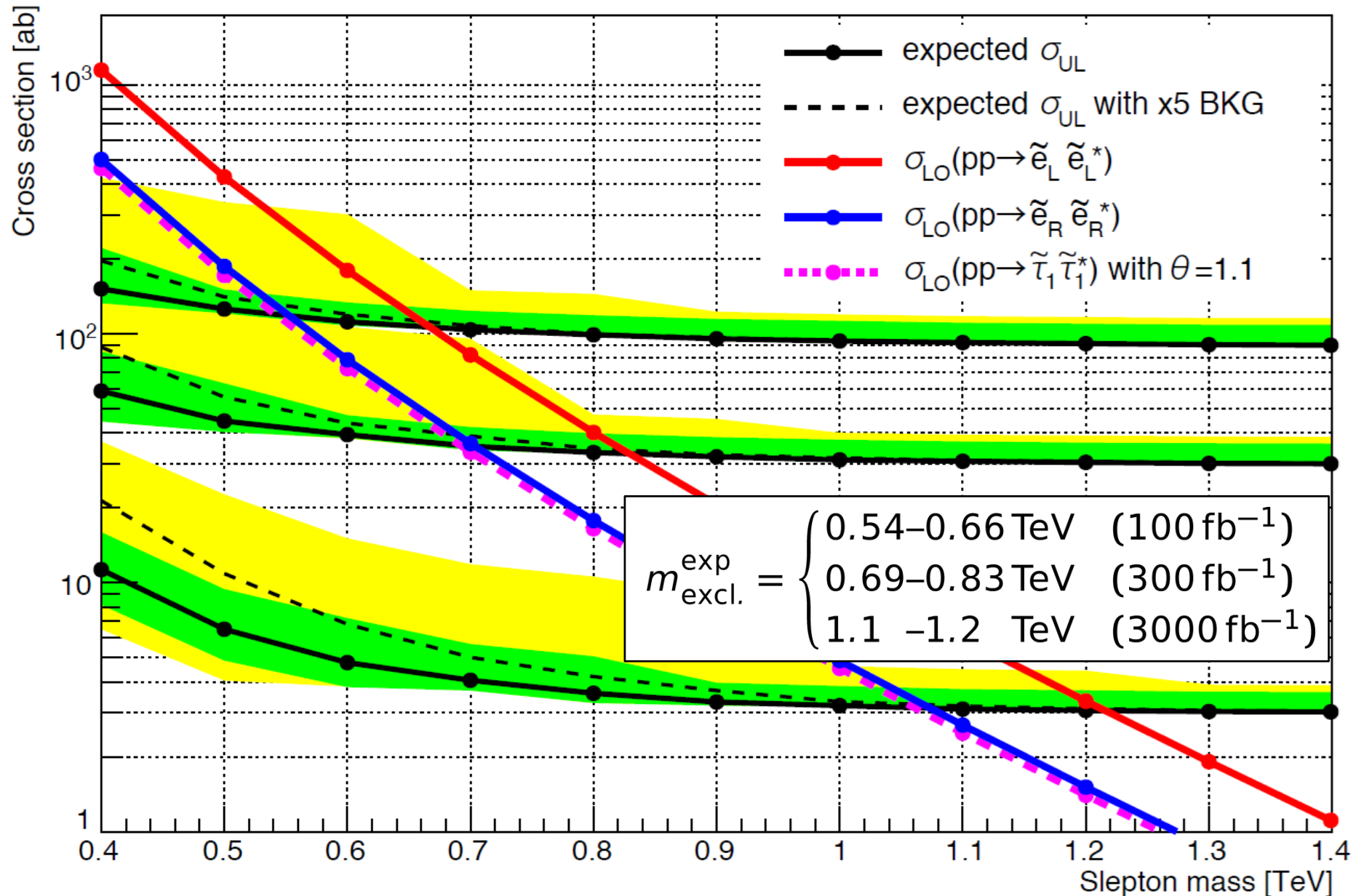
- ~~•  $E_{\text{loss}} < 30$  GeV~~

## ■ Event selection

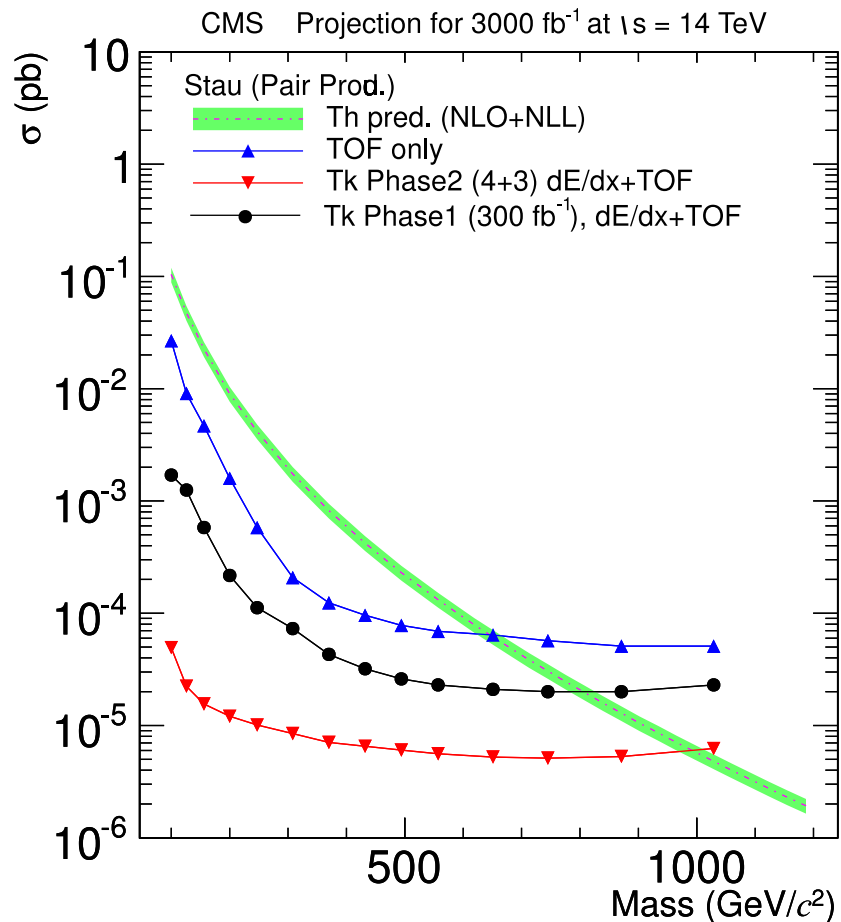
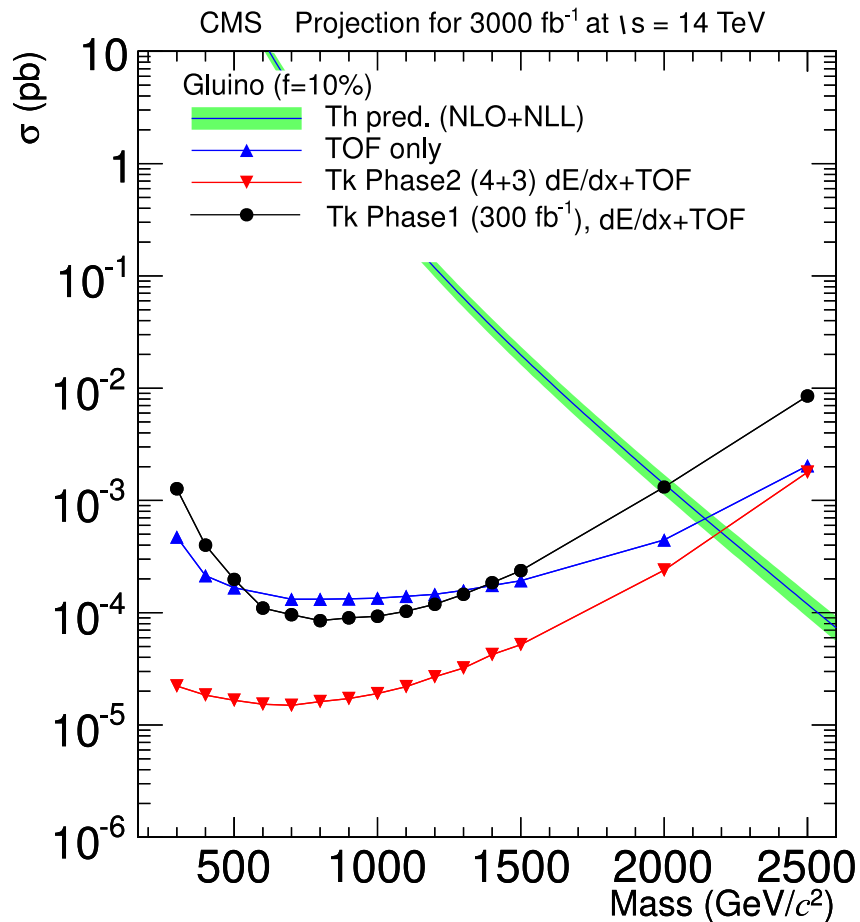
- two  $\tilde{l}$ -candidates



14 TeV LHC prospects are also studied in [1106.0764] & [1203.1581] by J. Heisig and J. Kersten.

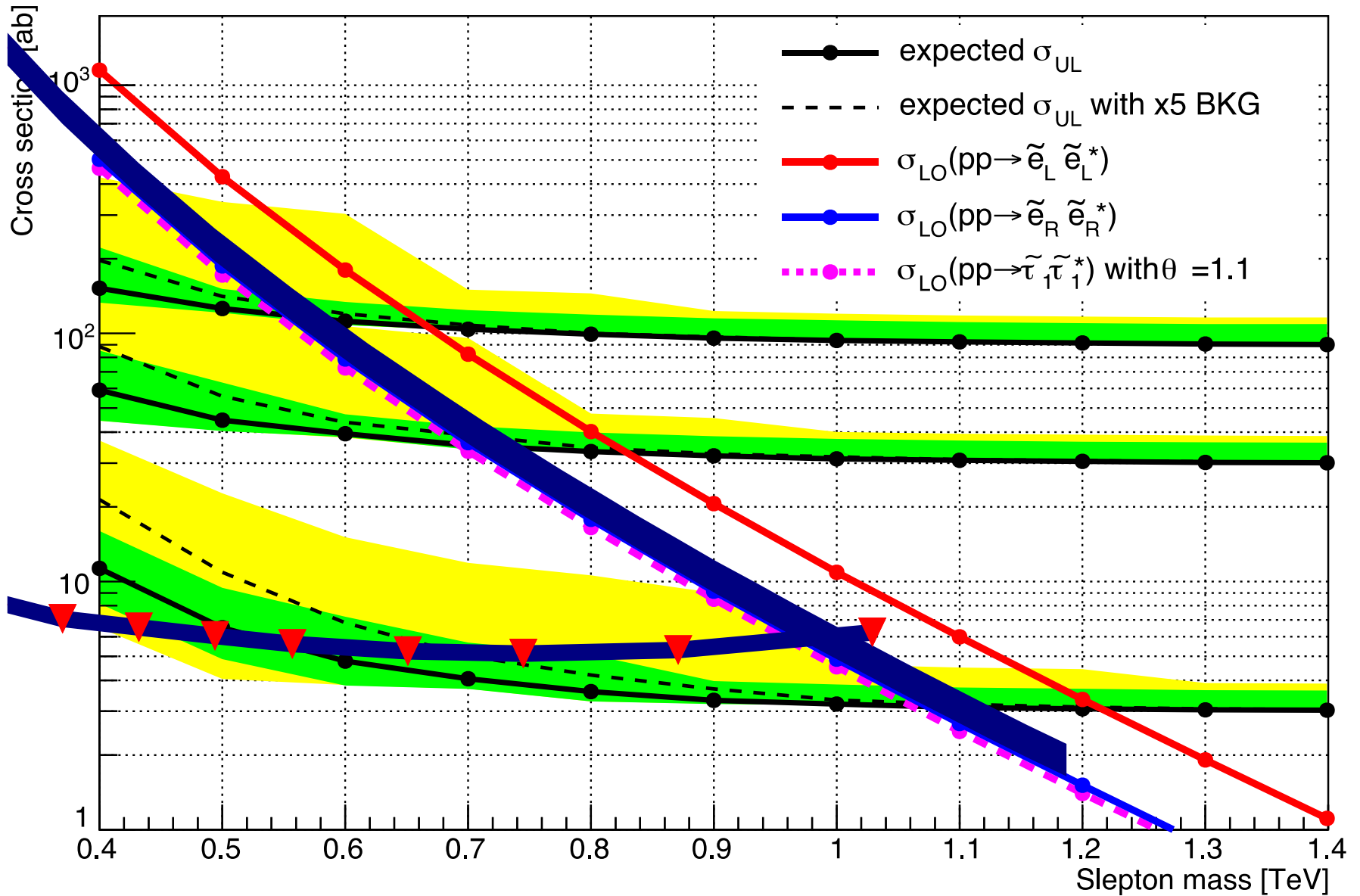


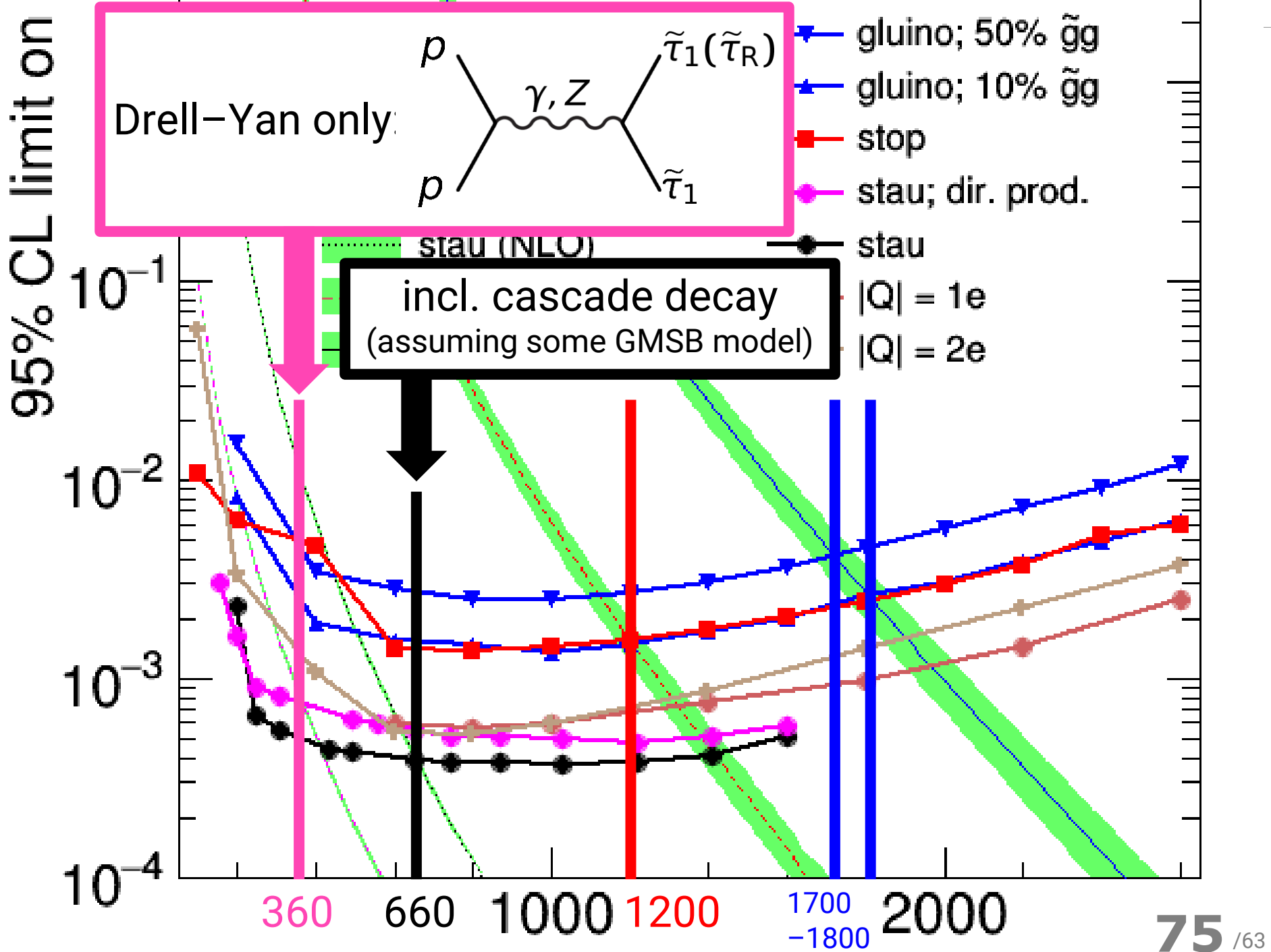




	current	HL-LHC
gluino:	1.7 TeV	→ 2.2 TeV?
stop:	1.2 TeV	→ 1.7 TeV?
stau (GMSB):	660 GeV	→ 1.2 TeV?
stau (DY):	360 GeV	→ 1.0 TeV?

(or discovery?)





Why  $\beta > 0.4$ ? (slepton  $dE/dx$ )

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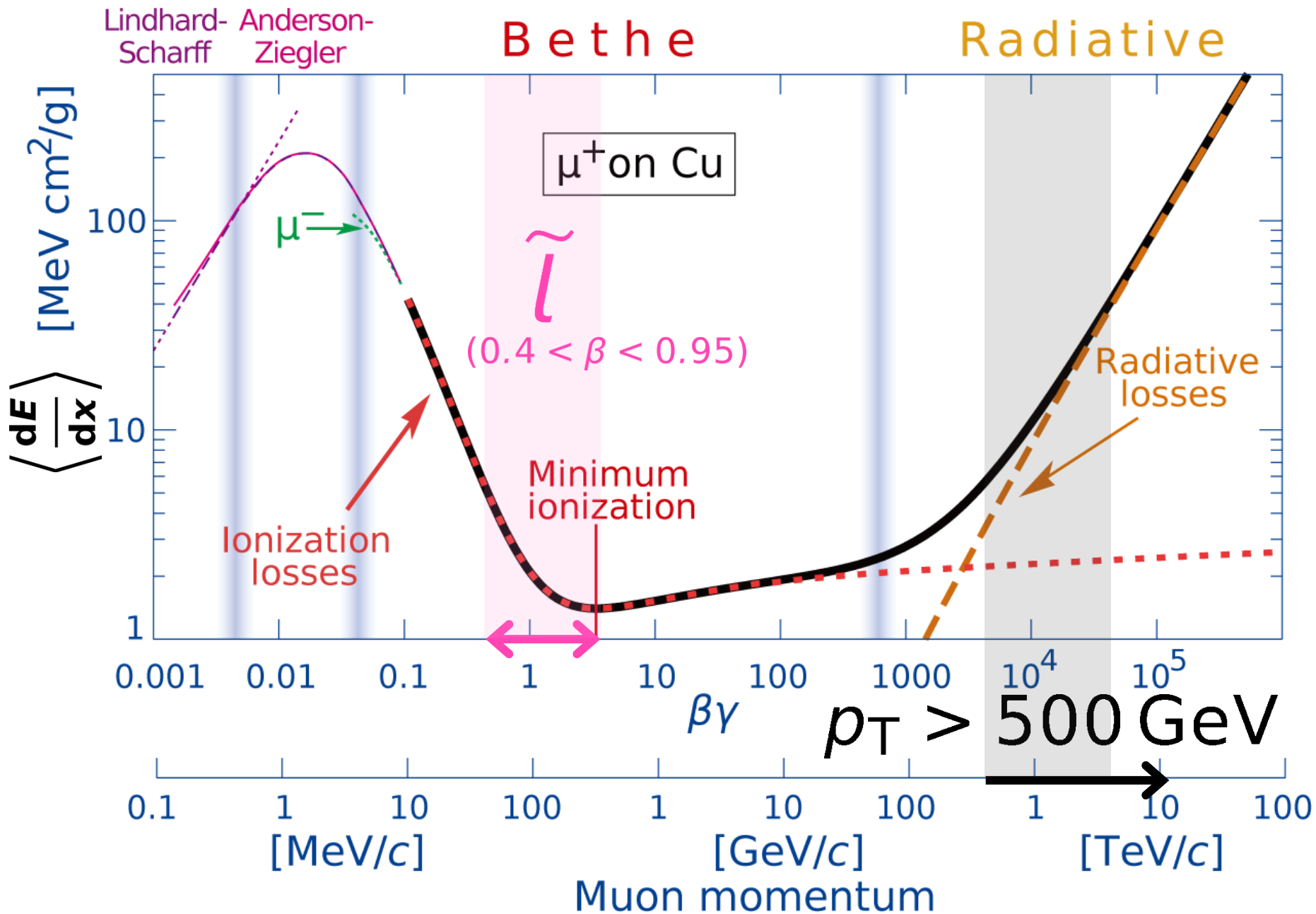


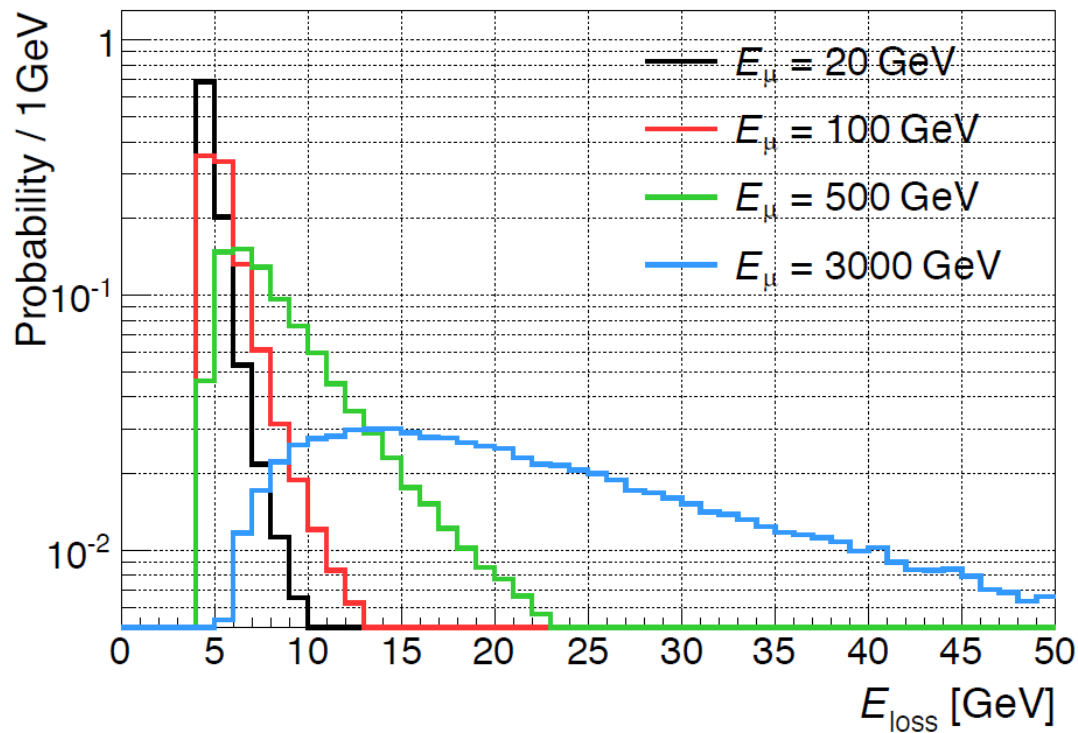
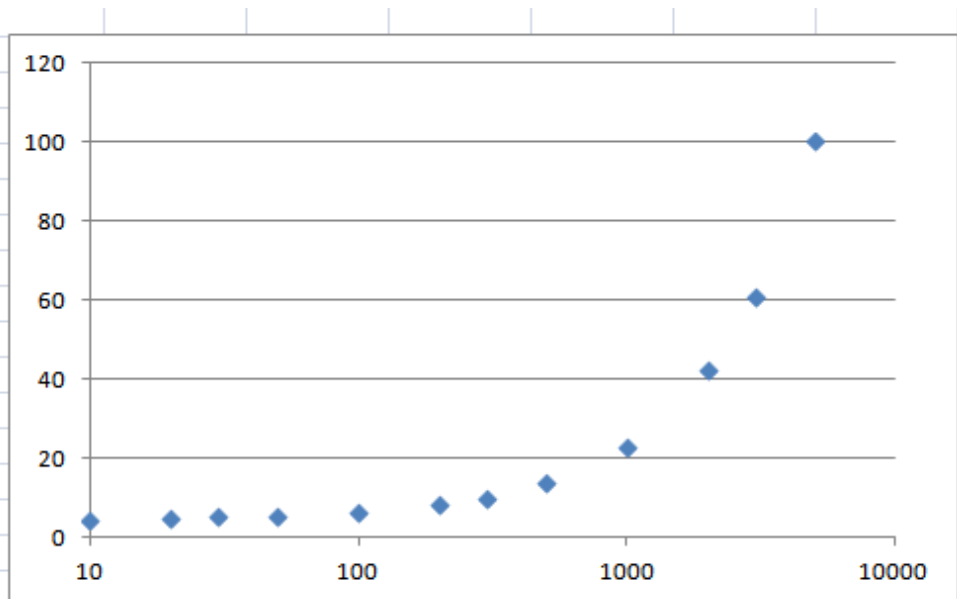
Figure from Groom, Mokhov, Striganov, Atom. Nucl. Data Tab. **78** (2001) 183-356  
 [also in PDG Review "Passage of particles through matter"]

Mean value of  $E_{\text{loss}}$ ?

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# Averaged muon energy loss in 3m iron (internal)

10	4.64883
20	5.0253
30	5.27343
50	5.68943
100	6.60542
200	8.43546
300	10.2127
500	13.9577
1000	23.231
2000	42.3777
3000	61.1561
5000	100.336



Note that the mean is much larger than the median because of its long long long tail.

$dE/dx$  to measure  $\beta$

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# Mass measurement = Measurement of velocity $\beta$

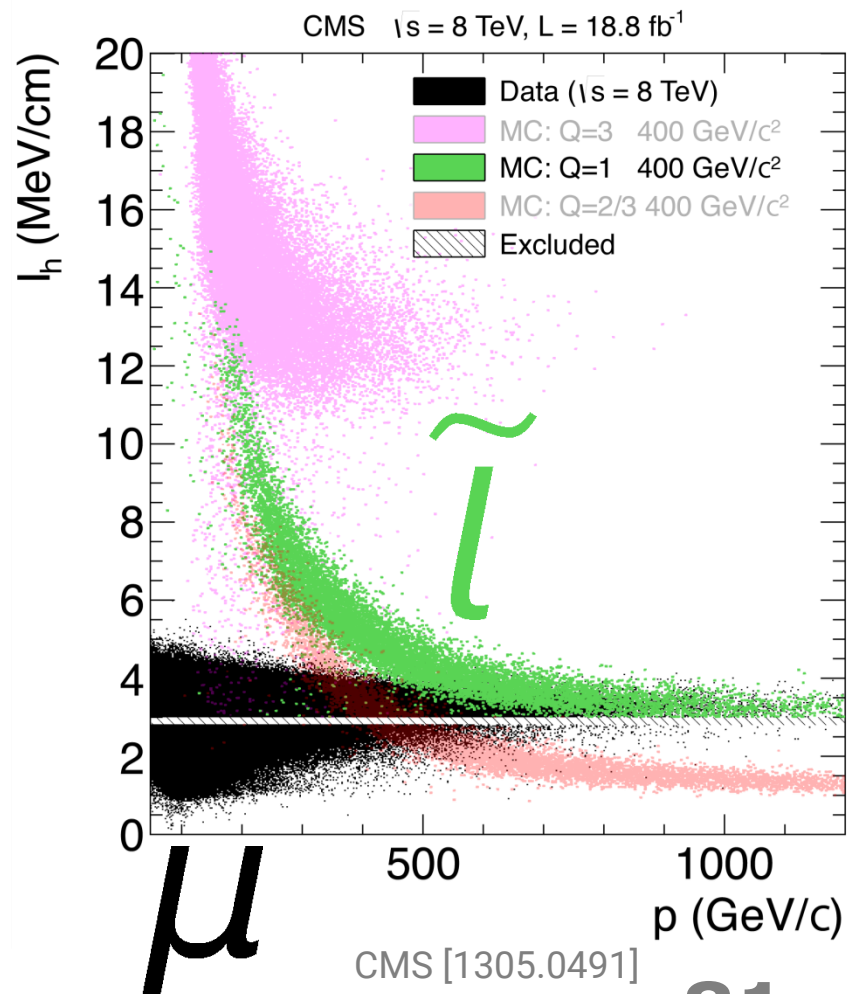
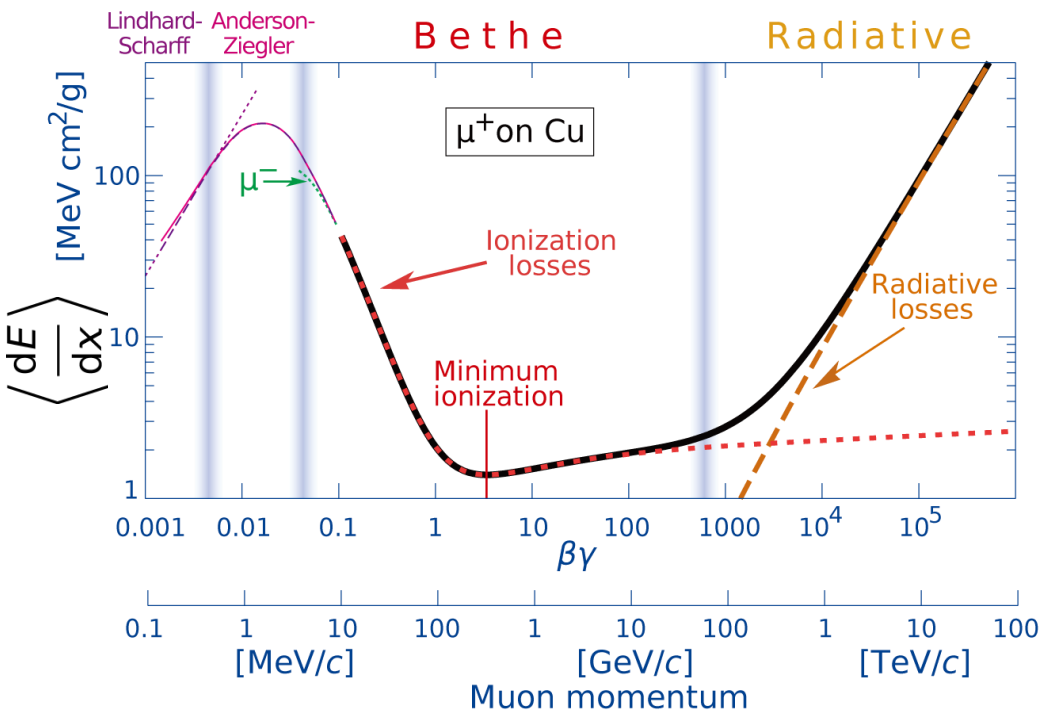
- TOF : time-of-flight

$$\beta = \Delta L / \Delta t$$

- $dE/dx$  : ionization energy loss

$$m = \frac{p}{\beta\gamma} = \frac{p}{\beta/\sqrt{1-\beta^2}}$$

$$I_h = \rho \cdot \frac{dE}{dx} \Big|_{\text{estimated}}$$

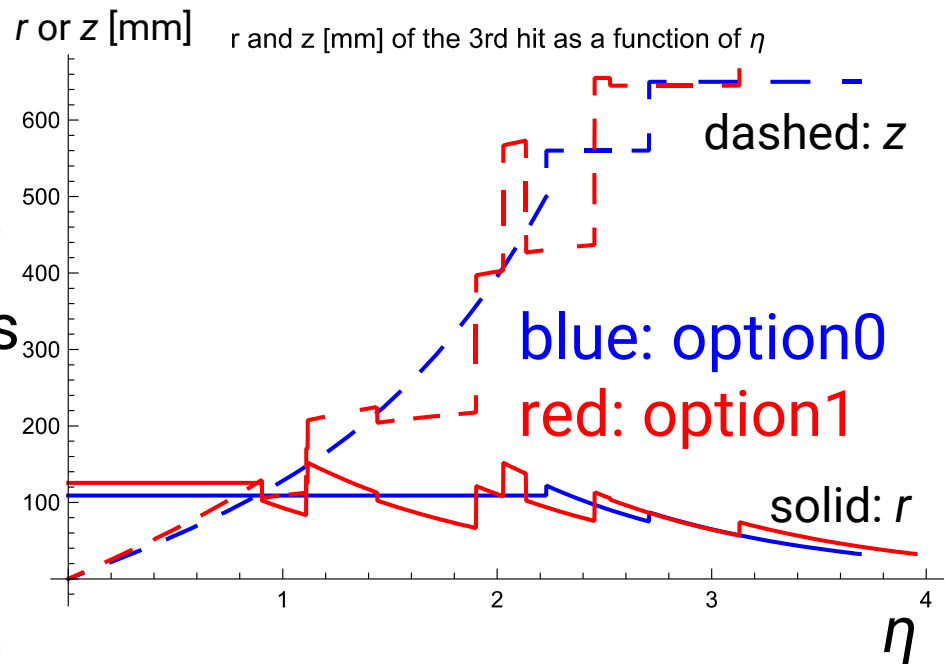
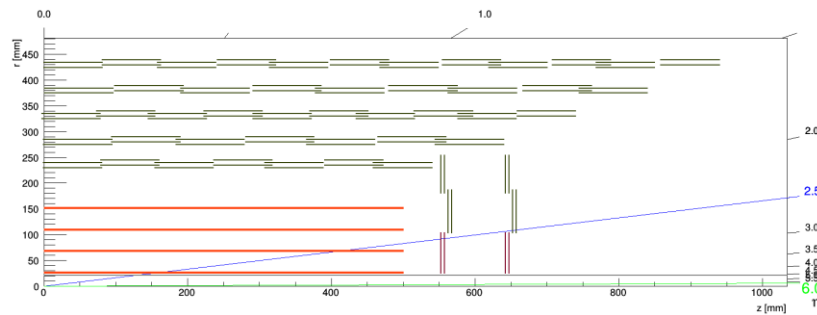


CMS [1305.0491]

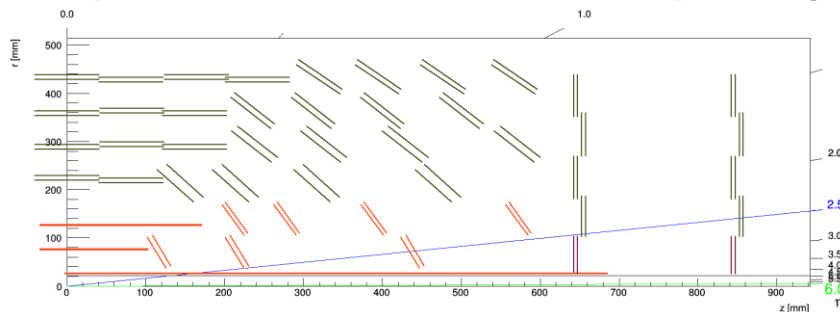
# FCC-he ID options

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## ■ Option 0 – FCC-Berlin



## ■ Option 1 - tilted endcap-rings



The position of “3<sup>rd</sup> layer” is not very different.  
Acceptance will be similar for both options.