

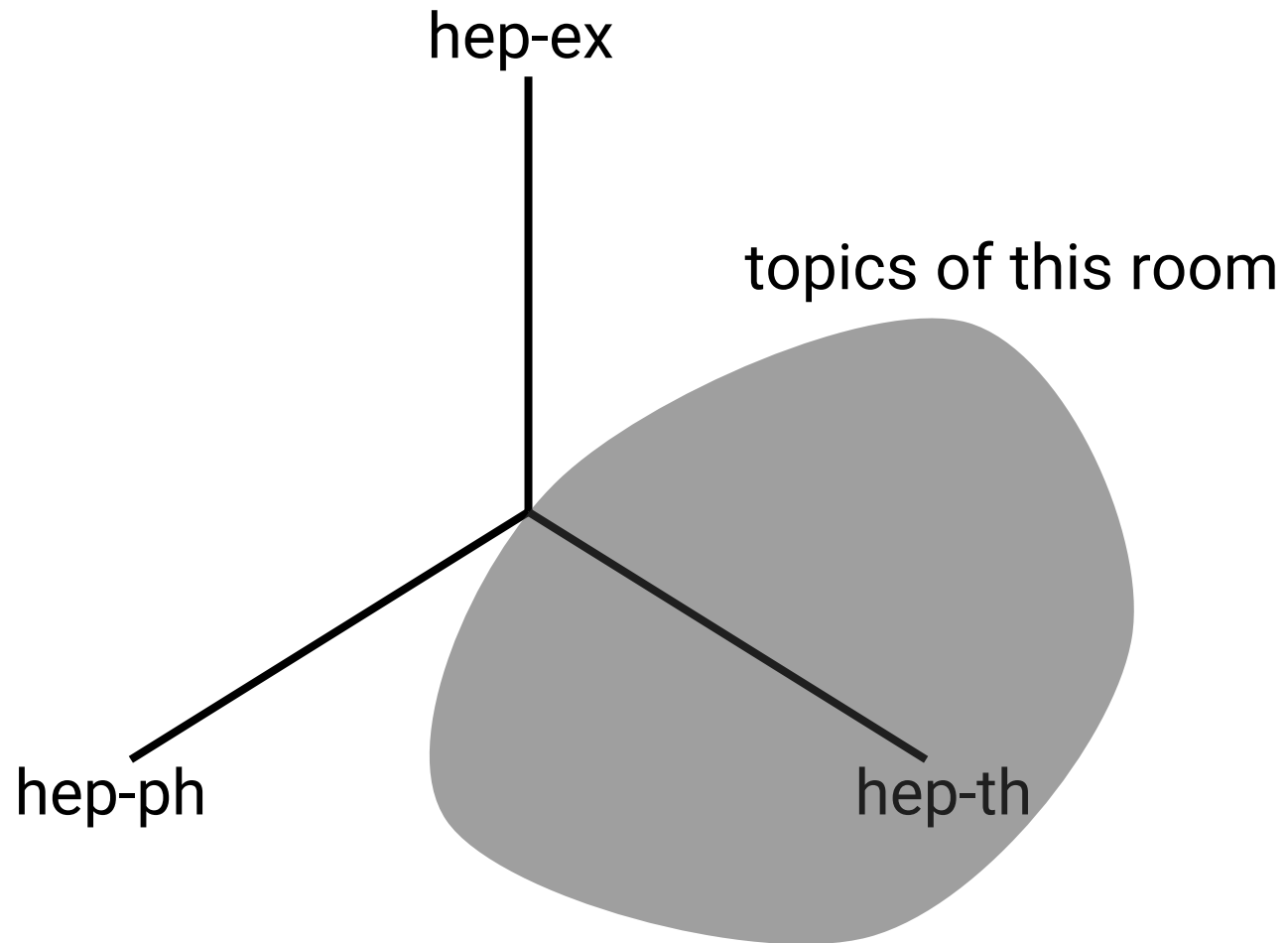
# MSSM 4G<sup>📶</sup> scenario

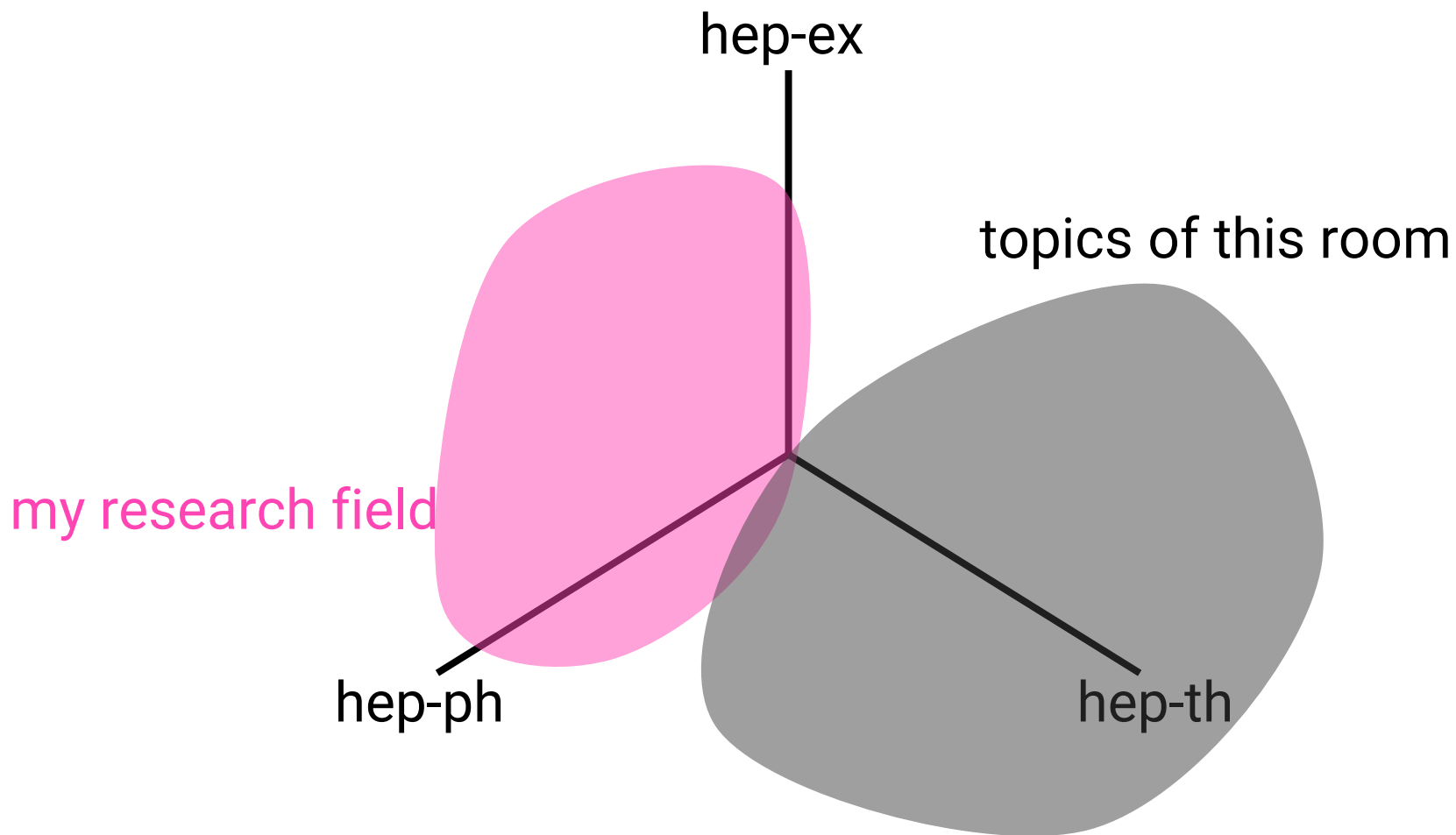
[Sho IWAMOTO](#) (岩本 祥)

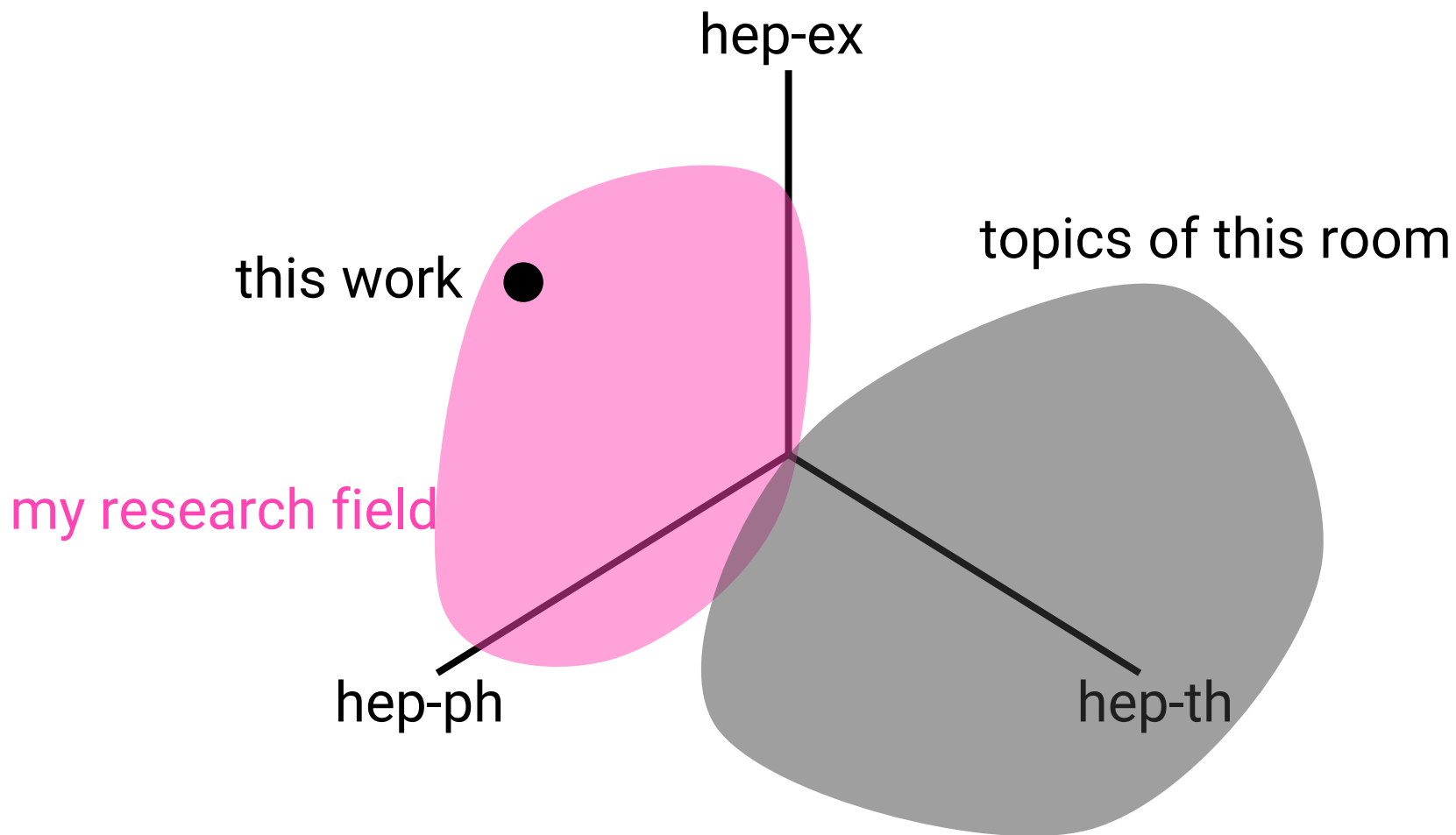
25 Dec. 2016

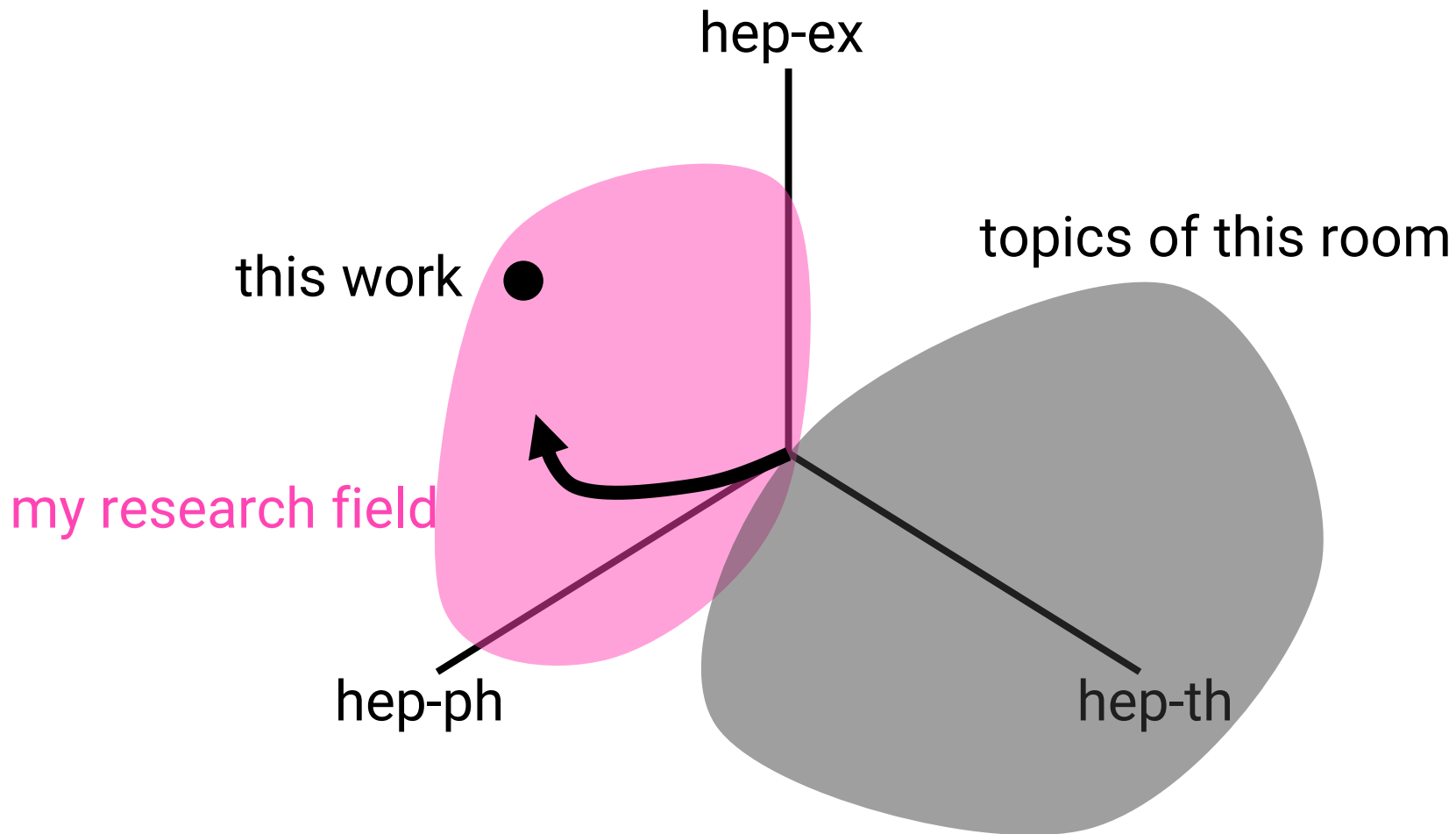
[IPS 62nd annual meeting](#) @ Tel-Aviv University

Based on [[1608.00283](#)] in collaboration with  
**M. Abdullah, J. L. Feng, and B. Lillard** (UC Irvine)









## Problems of the Standard Model = Hints of “New Physics”

- Anomalies in  $B$ -physics
- Proton charge radius
- Beryllium anomaly
- Muon “ $g-2$ ”
- Higgs mass (“naturalness”)
- Neutrino mass
- Dark matter
- Unification of 3 forces
- Dark energy
- Gravity

## New Physics Candidates

- SUSY [supersymmetry]

- 
- 
- 

Please fill this list  
with your models  
/ models you like

etc...

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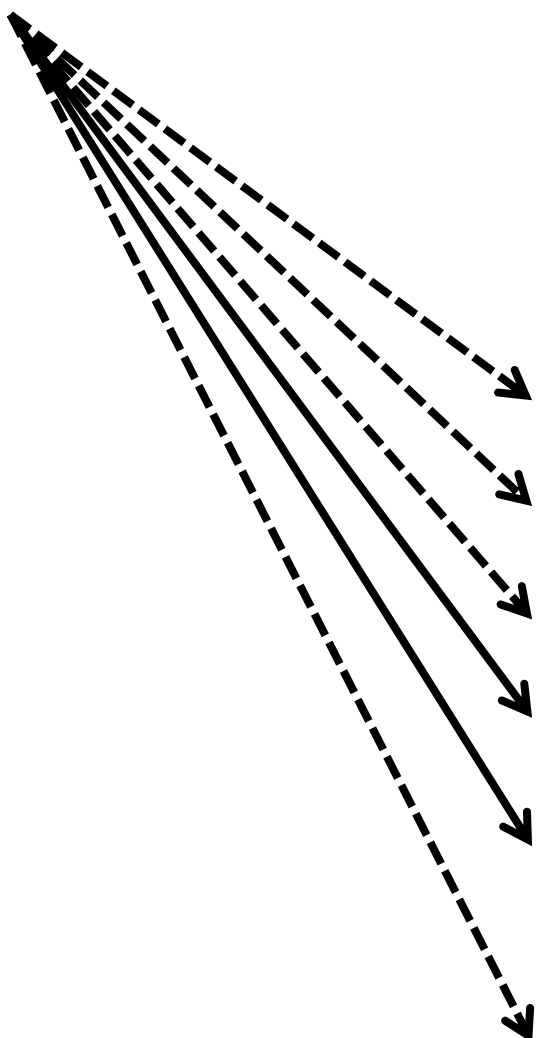
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■

■

■

etc...

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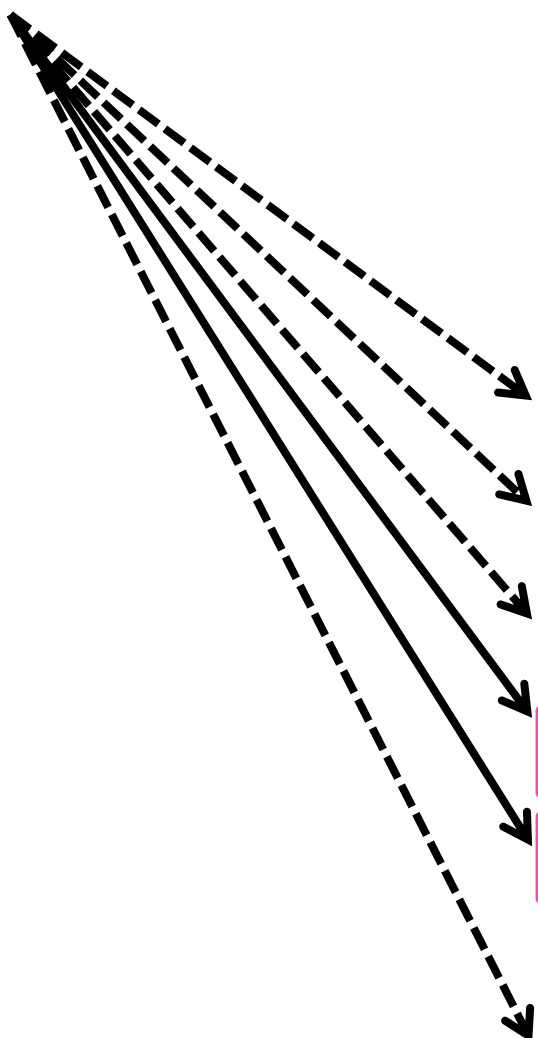
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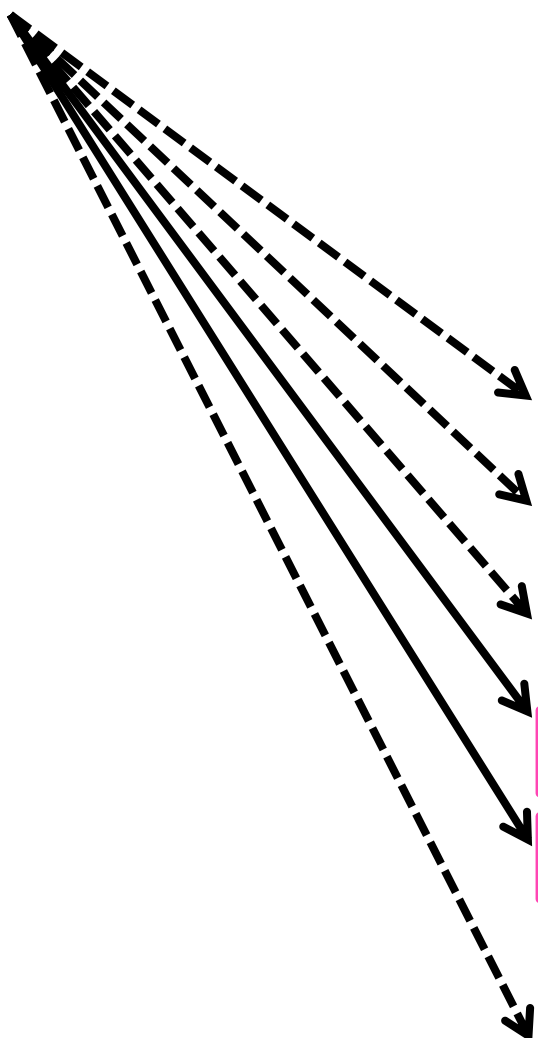
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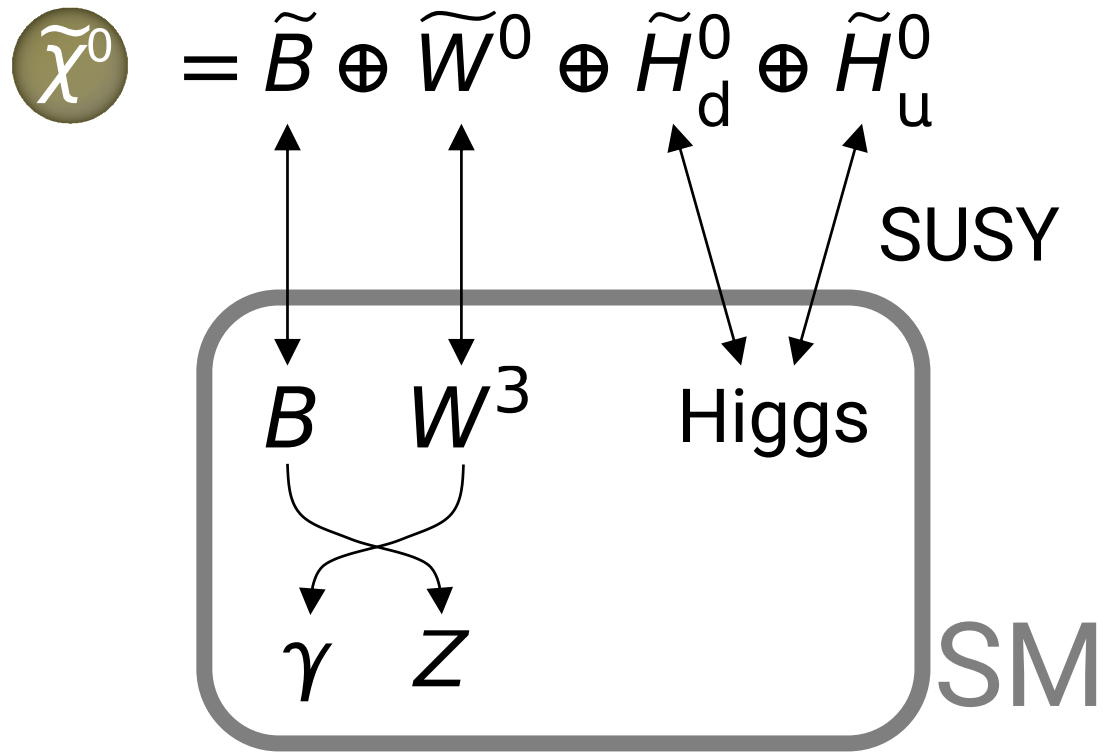
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■ MSSM  $\ni$  Dark matter candidate



- MSSM  $\ni$  Dark matter candidate

$$\tilde{\chi}^0 = \tilde{B} \oplus \tilde{W}^0 \oplus \tilde{H}_d^0 \oplus \tilde{H}_u^0$$

- Pure- $\tilde{B}$  dark matter (i.e., DM is  $\tilde{\chi}^0$  and it is purely  $\tilde{B}$ -like)

✓ theoretically motivated & simple

✗ “thermal overabundance” problem

→ MSSM4G model as one solution

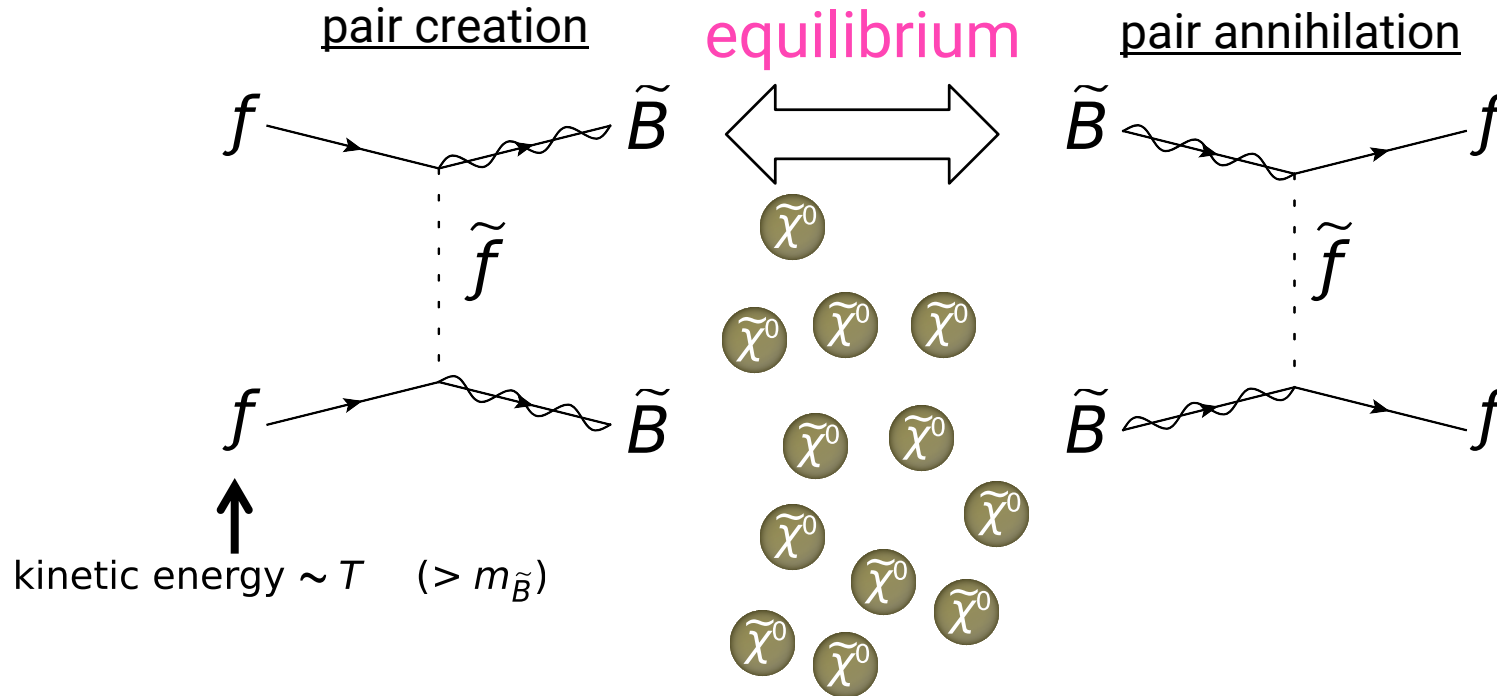
Introduction: Overabundant problem

Model: **MSSM4G** 

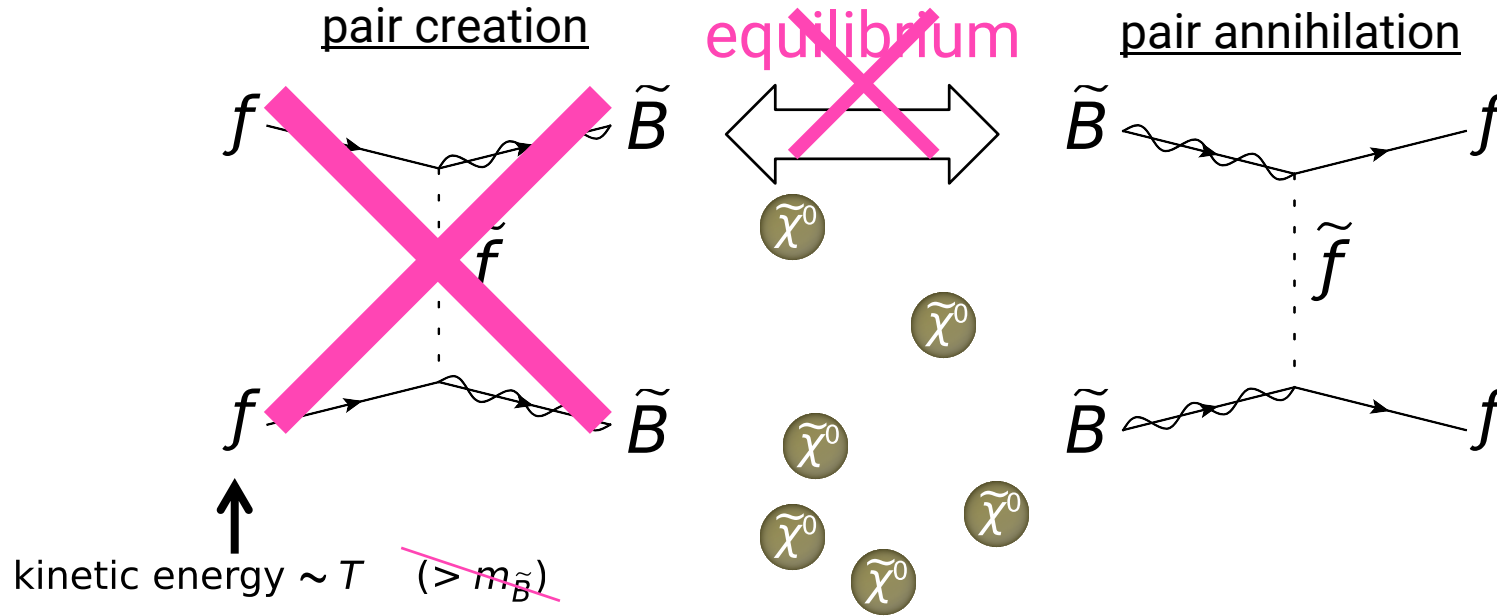
Phenomenology: Gamma-ray obs. & LHC

Summary

■ Early Universe with  $T > m_{\tilde{B}}$



■ Early Universe with  $T \lesssim m_{\tilde{B}}$



## ■ Early Universe with $T \lesssim m_{\tilde{B}}/20$

pair creation

“relic density”

$\tilde{\chi}^0$

$\tilde{\chi}^0$

$\tilde{\chi}^0$

$\tilde{\chi}^0$

pair annihilation

$\tilde{B}$   $f$

$\tilde{f}$

- far apart due to
- pair annihilation
  - Universe's expansion

$\tilde{B}$   $f$



■ Early Universe with  $T \lesssim m_{\tilde{B}}/20$

pair creation

pair annihilation

“relic density”

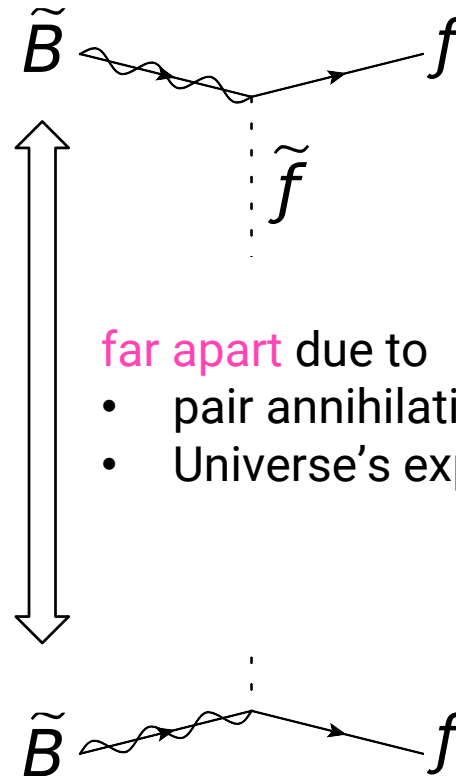
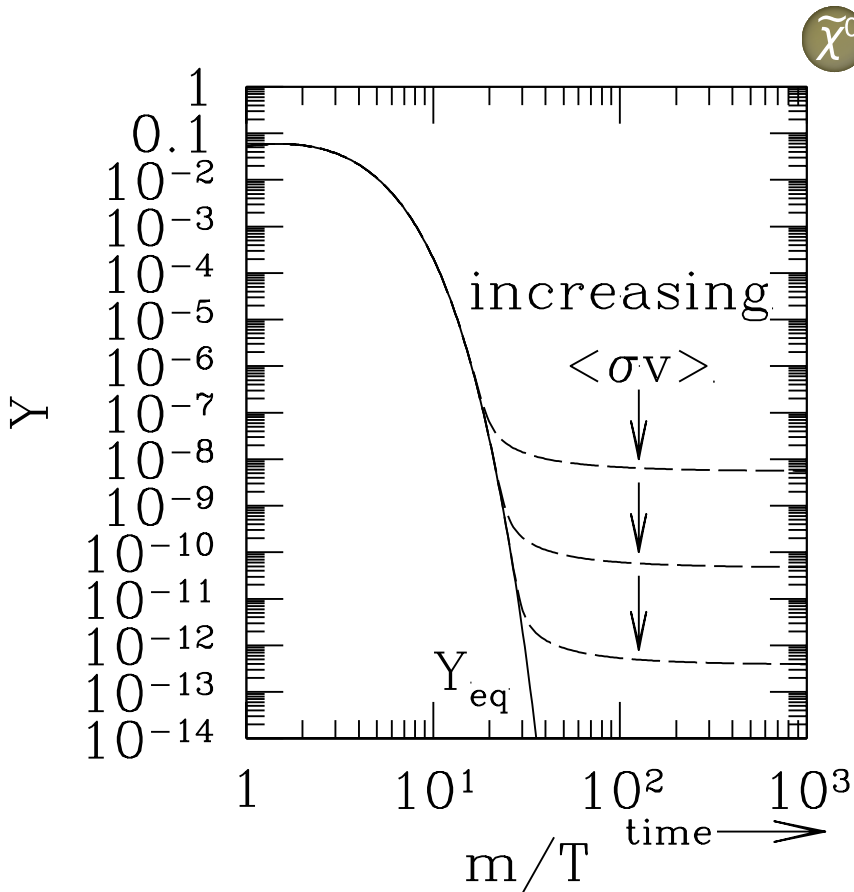
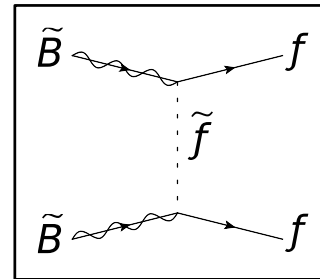
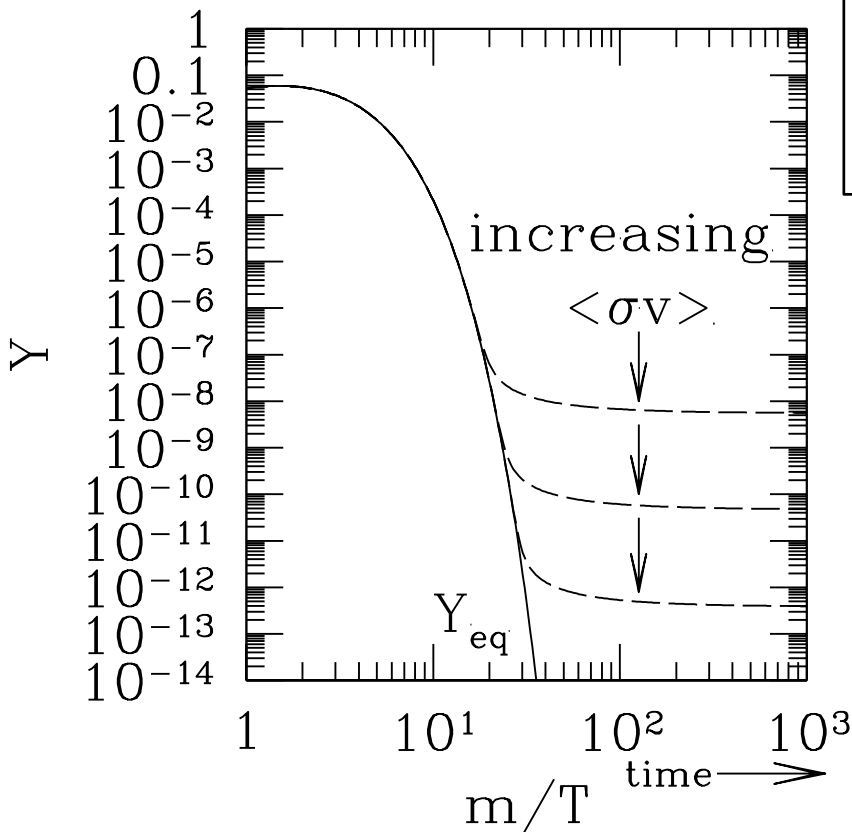


Figure from Gelmini and Gondolo, [1009.3690](https://arxiv.org/abs/1009.3690)

■ pure- $\tilde{B}$  DM

$m_{\tilde{B}} \lesssim 100$  GeV: DM density (“relic density”)  $\sim$  observation

$m_{\tilde{B}} \gtrsim 100$  GeV: smaller crosssection  $\rightarrow$  larger DM density



“overabundant problem”  
of Bino thermal relic DM

Figure from Gelmini and Gondolo, [1009.3690](#)

### ■ pure- $\tilde{B}$ DM

$m_{\tilde{B}} \lesssim 100$  GeV: DM density (“relic density”)  $\sim$  observation

$m_{\tilde{B}} \gtrsim 100$  GeV: smaller crosssection  $\rightarrow$  larger DM density

### ● DM is not pure $\tilde{B}$ ?

➤ pure- $\tilde{W}^0$  DM

➤  $\tilde{B}$ - $\tilde{H}$  mixing

“overabundant problem”

of Bino thermal relic DM

### ● Other annihilation channels?

➤ co-annihilation [Griest, Seckel, 1991]

➤ **MSSM4G** [Abdullah, Feng, 2015]

Introduction: Overabundant problem

Model: **MSSM4G** 

Phenomenology: Gamma-ray obs. & LHC

Summary

■ SM = 3Generations

■ MSSM = 3Generations

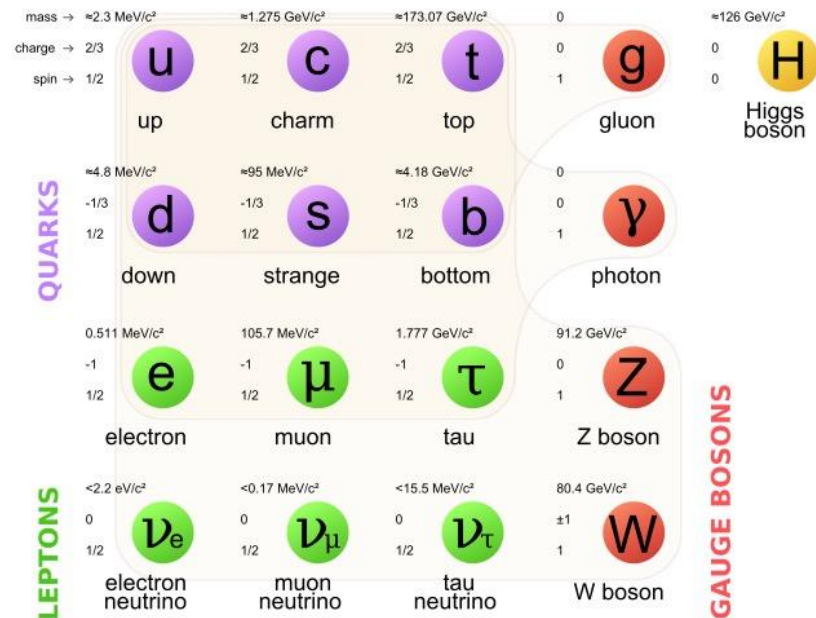


extra vector-like

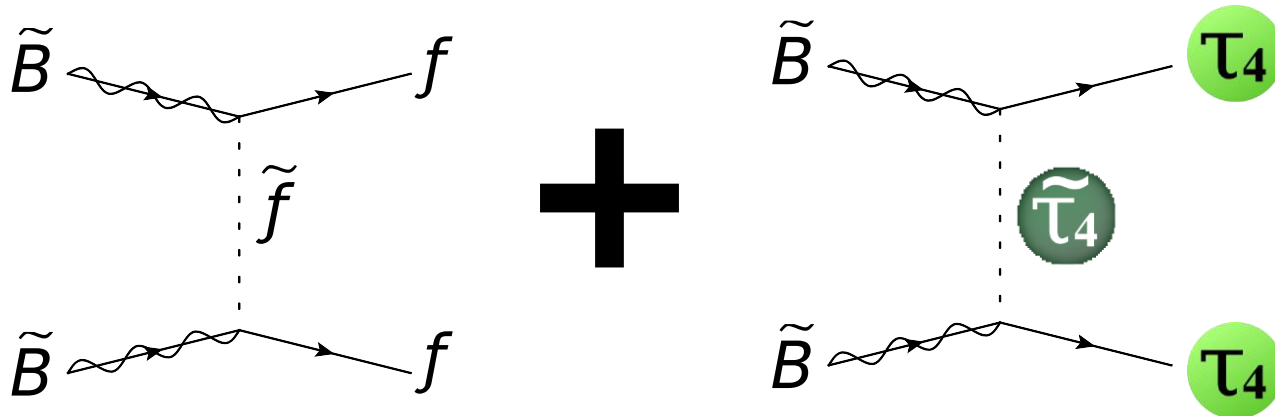
4<sup>th</sup>-Generation lepton



**MSSM4G**



- A new solution to  $\tilde{B}$ -overabundant problem: MSSM4G



extra annihilation channel

if  $\tau_4 \gtrsim \tilde{B} > \tau_4$

$$\langle \sigma v \rangle = \frac{g_Y^4 Y_L^2 Y_R^2 m_f^2}{2\pi m_{\tilde{B}}} \frac{\sqrt{m_{\tilde{B}}^2 - m_f^2}}{(m_{\tilde{B}}^2 + m_{\tilde{f}}^2 - m_f^2)^2}$$

■ MSSM +  $E\bar{E}$  → breaks coupling unification

■ QUE model : MSSM +  $Q\bar{Q}U\bar{U}E\bar{E}$

✓ gauge coupling unification

✓ SU(5) GUT

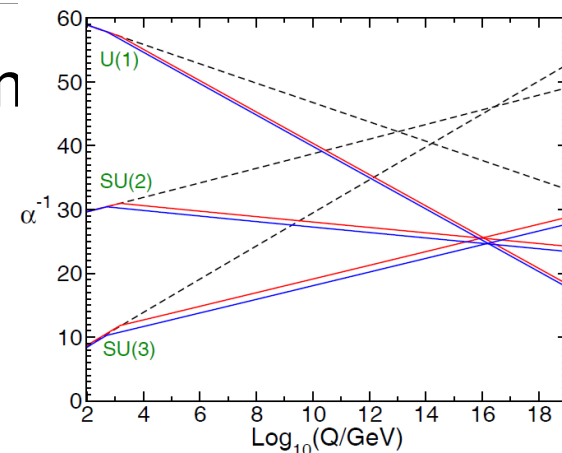
➤ extra  $H_u Q_4 \bar{U}_4$  interaction →  $m_h$  **UP**

■ QDEE model : MSSM +  $Q\bar{Q}D\bar{D}E\bar{E}E\bar{E}$

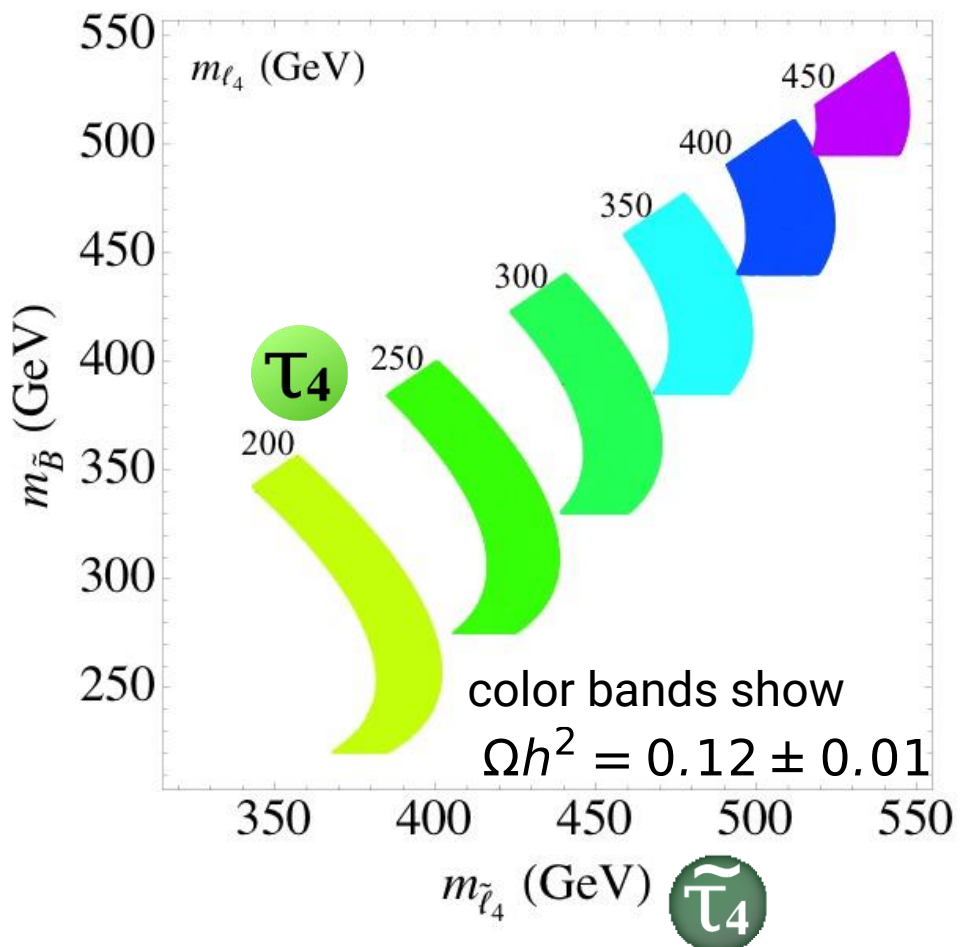
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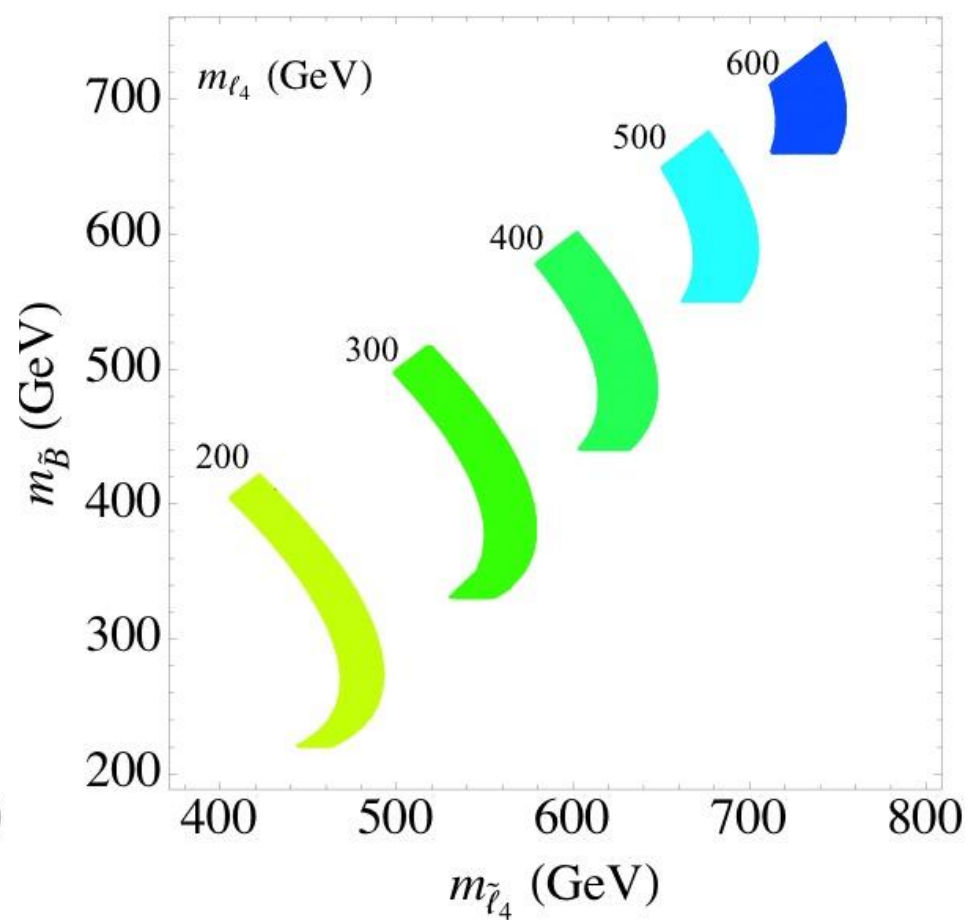
➤ extra  $H_d Q_4 \bar{D}_4$  coupling →  $m_h$  slightly **UP**



### QUE model



### QDEE model



$$\tilde{\tau}_4 \gtrsim \tilde{B} > \tau_4$$



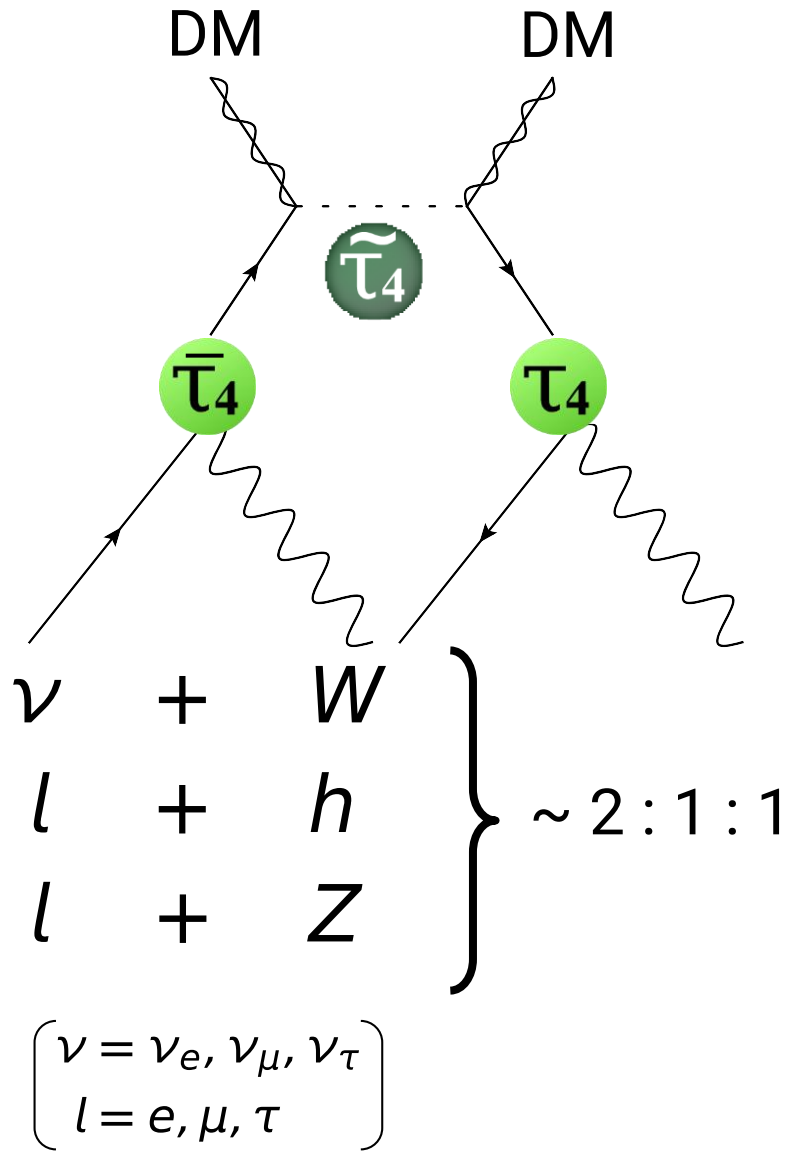
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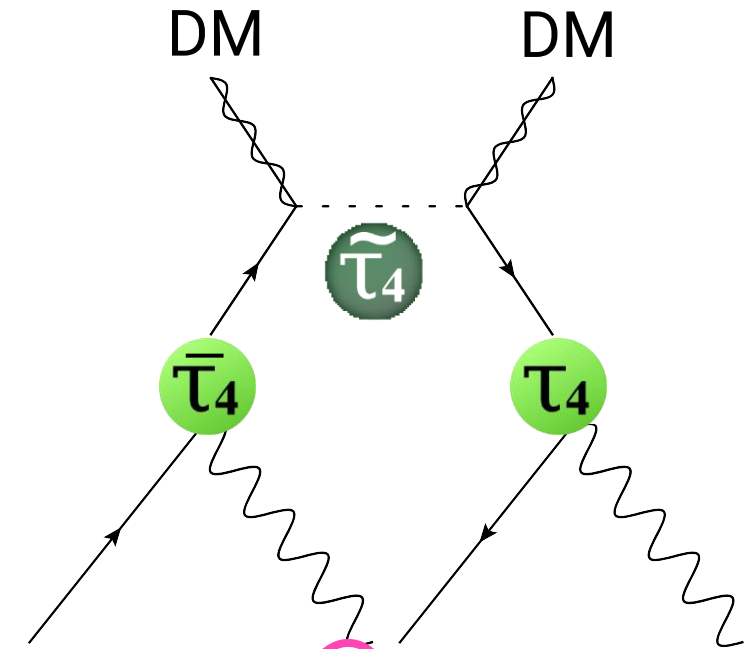
Phenomenology: Gamma-ray obs. & LHC

Summary

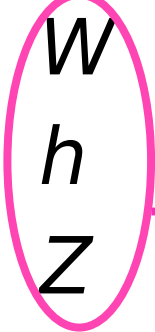
■ DM indirect detection by Gamma-ray observation



## DM indirect detection by Gamma-ray observation



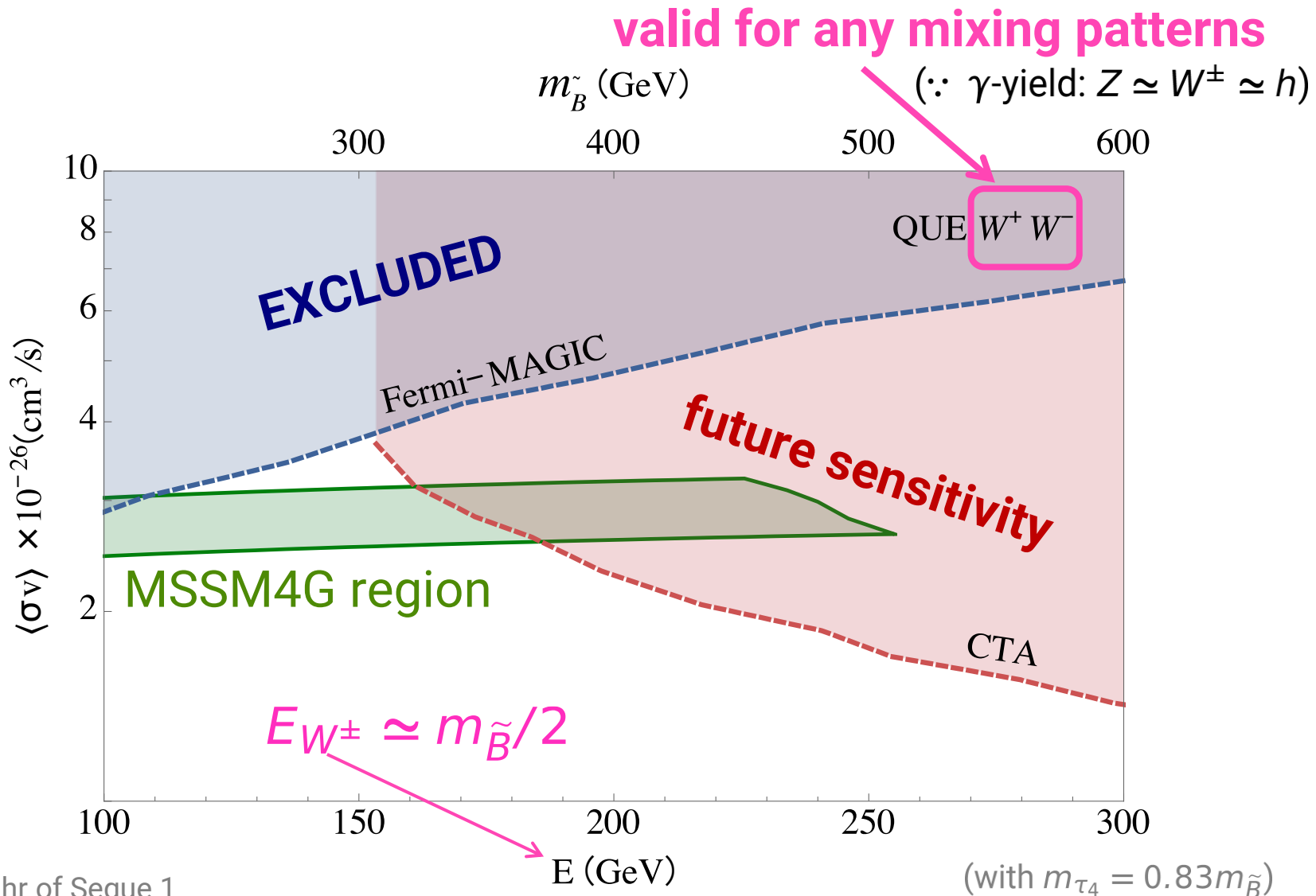
$\nu$  +  $W$   
 $l$  +  $h$   
 $l$  +  $Z$



$\left( \begin{array}{l} \nu = \nu_e, \nu_\mu, \nu_\tau \\ l = e, \mu, \tau \end{array} \right)$



- Fermi-LAT (satellite)
- MAGIC (Air Cherenkov telescope)
- CTA (future A. C. Telescope)



MAGIC: 158 hr of Segue 1

Fermi-LAT: 6 yr of 15 dSph (incl. Segue 1)

DM profile: NFW

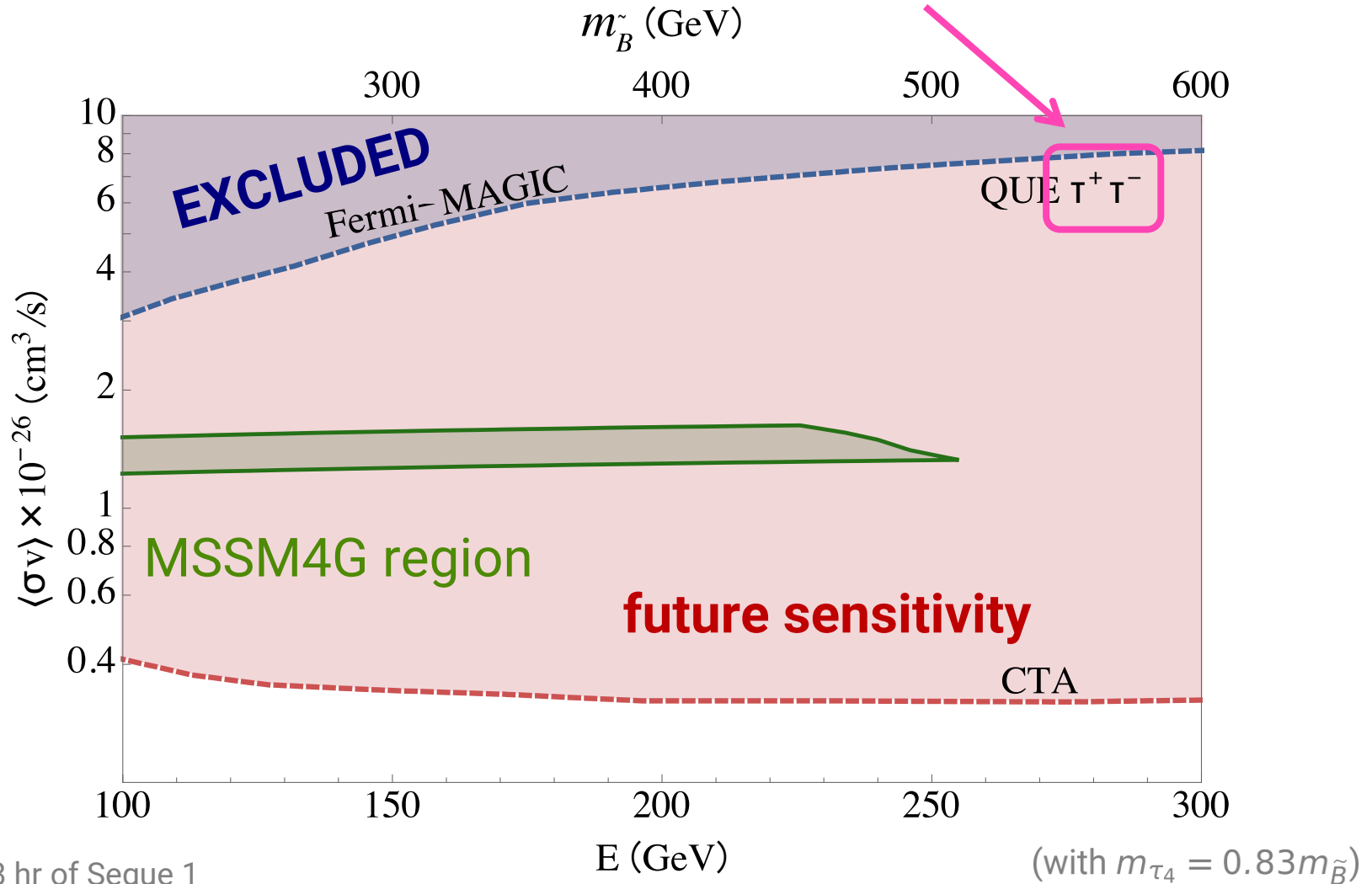
Fermi-LAT dominates MAGIC in almost all  $E$ -range.

CTA prospect : 500hr of Milky Way

DM profile: Einasto

No syst. unc. (stat only)

if 4G lepton decays to tau-lepton



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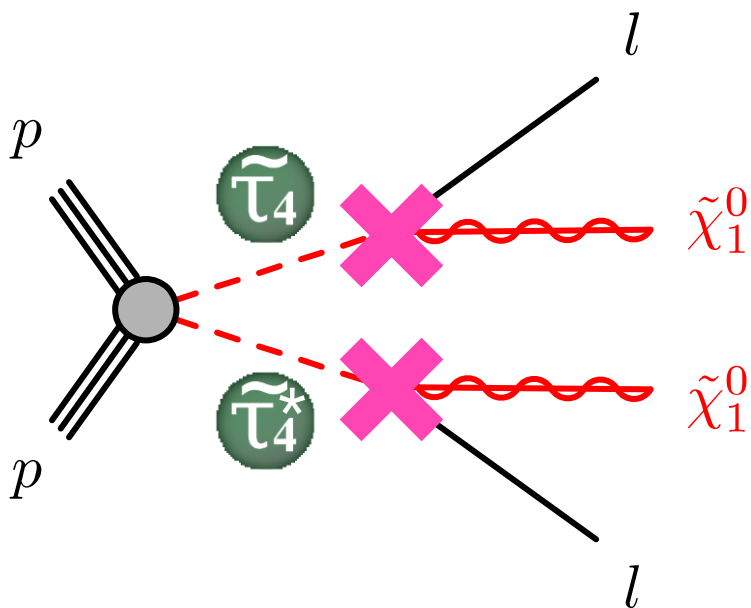
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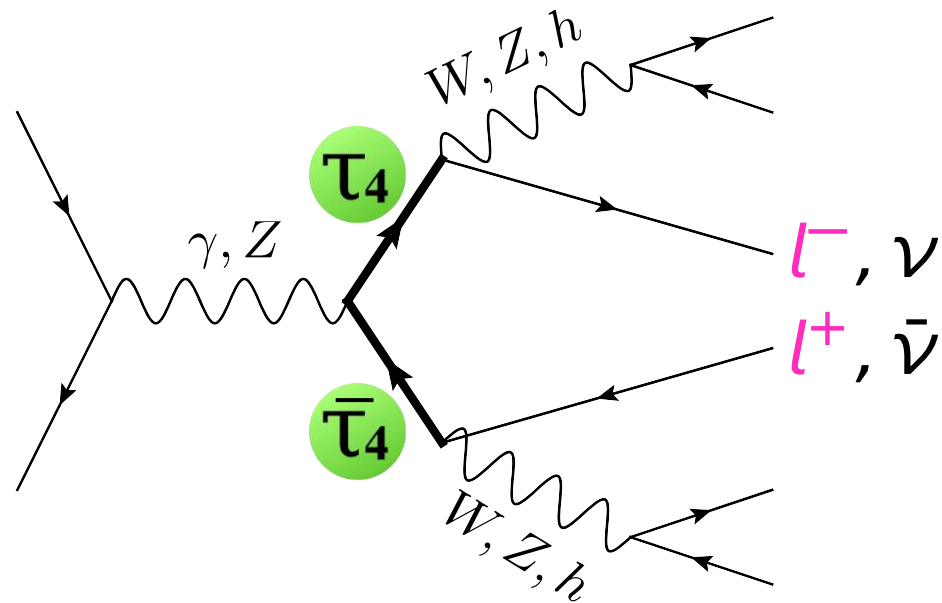
Summary

$\tilde{\tau}_4$  (4G slepton) search



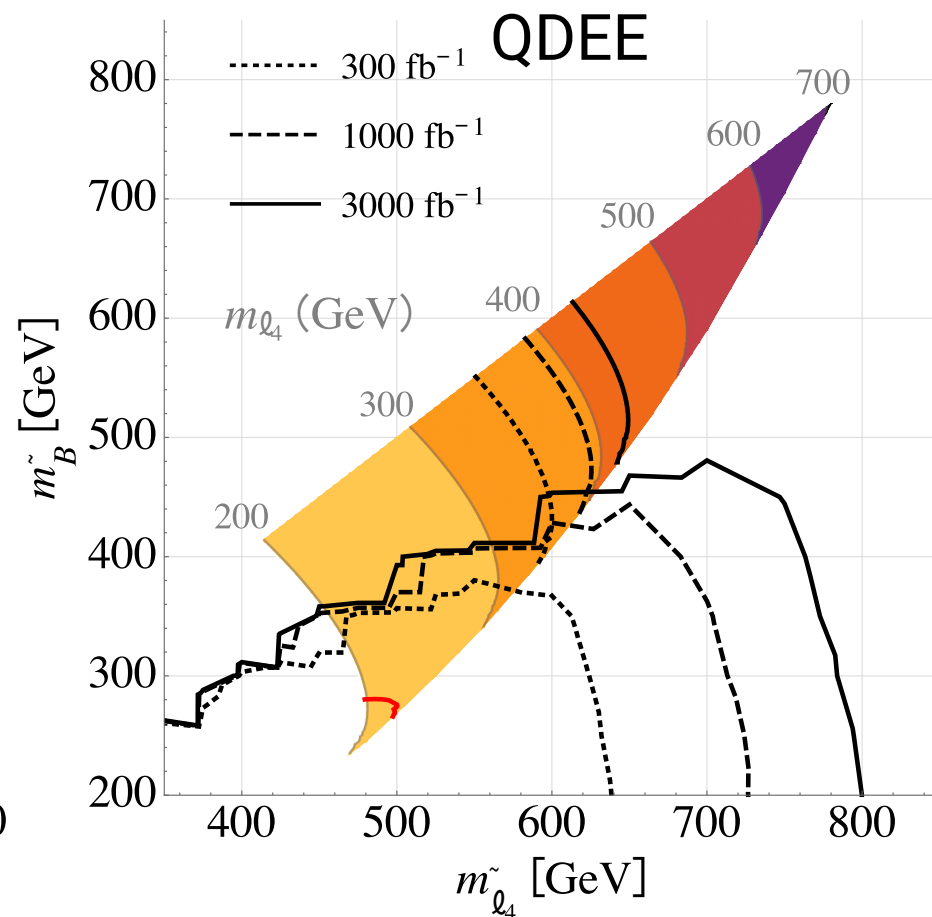
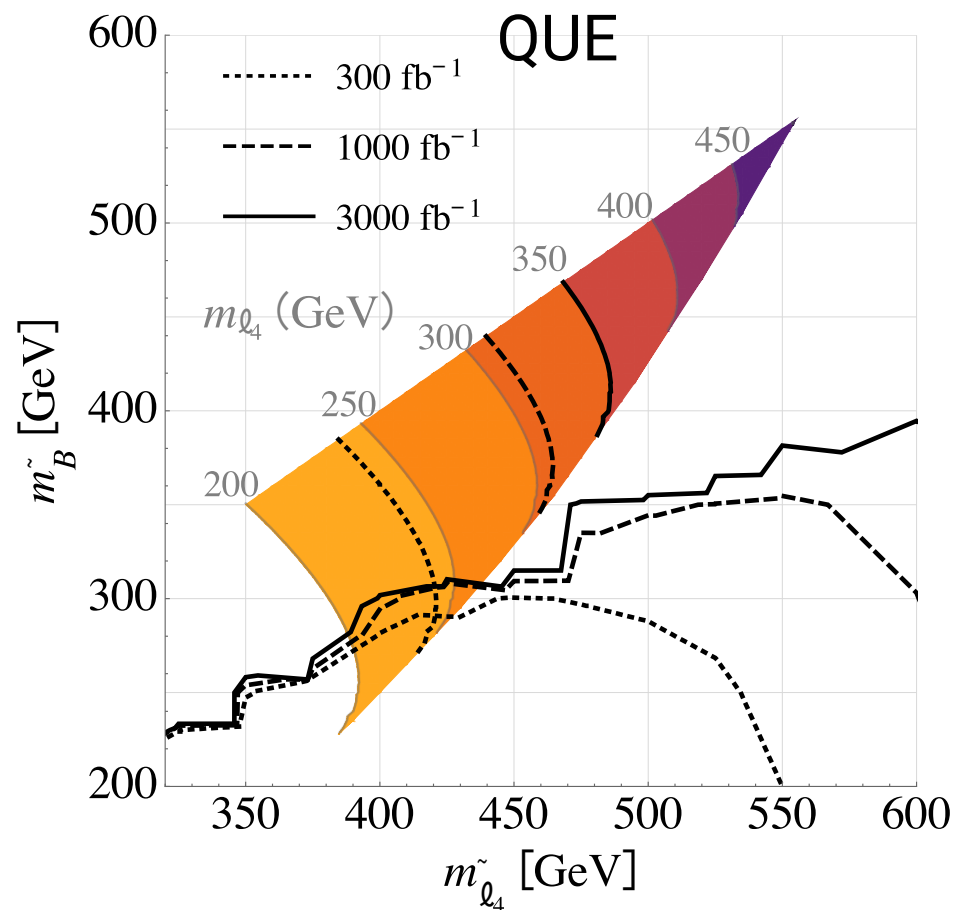
→ searches for  
2-lepton + Missing  $E_T$   
(same as MSSM slepton searches)

$\tau_4$  (4G lepton) search



→ searches for  
multi-lepton final state

■ if 4G lepton decays to electron or muon

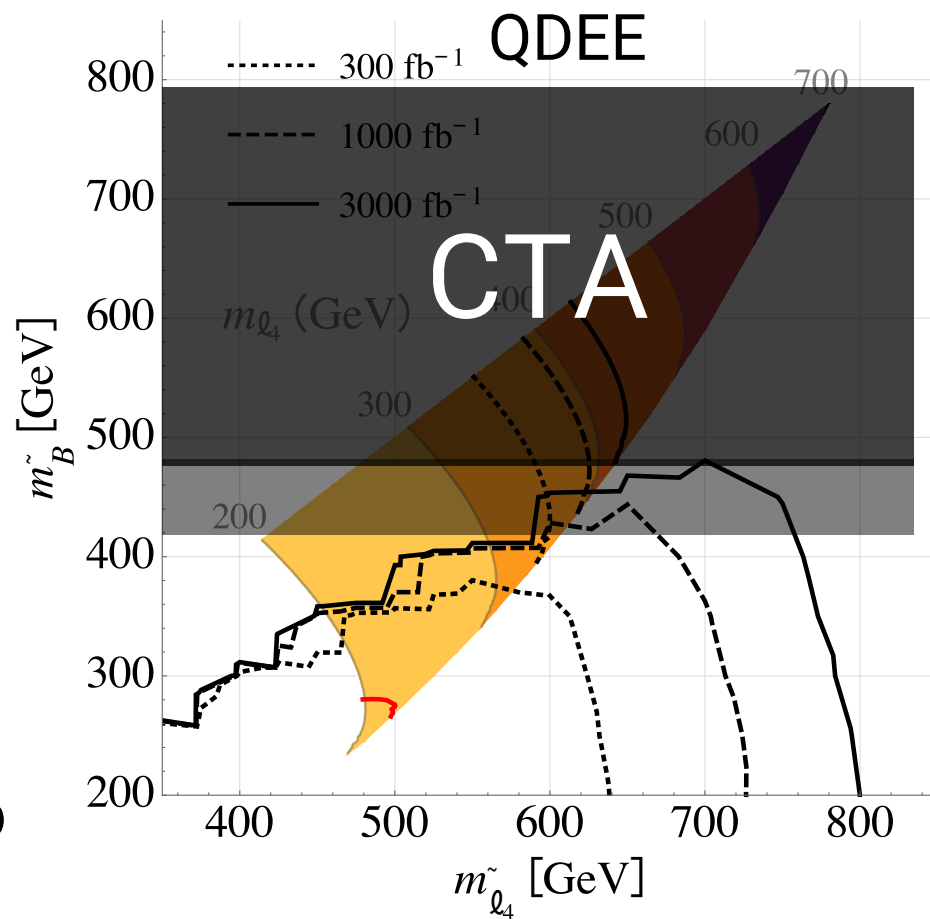
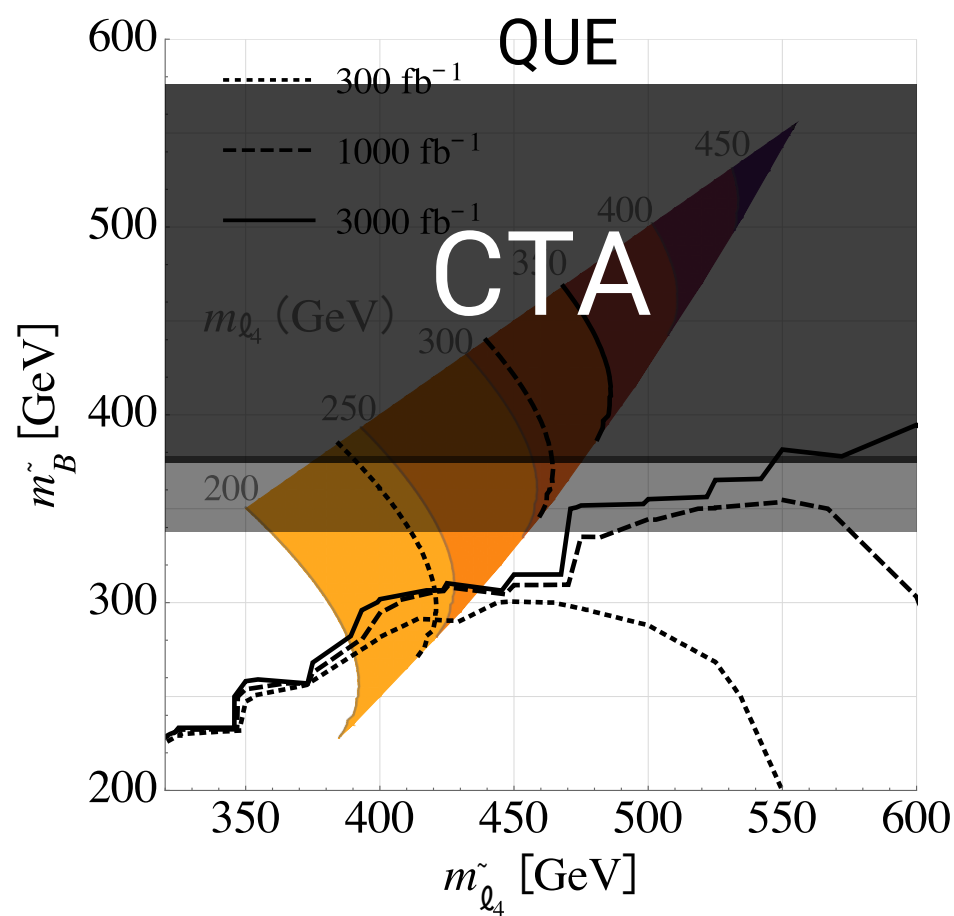


■ if 4G lepton decays to tau-lepton

LHC insensitive ... ( $\tilde{\nu} \cdot \omega \cdot \tilde{\nu}$ )



■ if 4G lepton decays to electron or muon



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Model: **MSSM4G** 

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Summary

## Summary

- Pure- $\tilde{B}$  DM **may cause** DM overabundance
- MSSM4G **is a solution to this problem.**
- MSSM4G will be **ALL explored soon.**

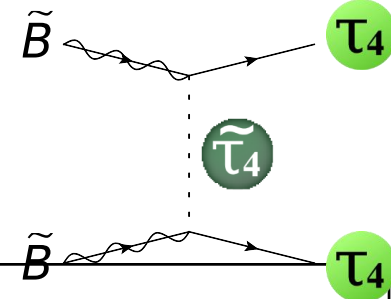
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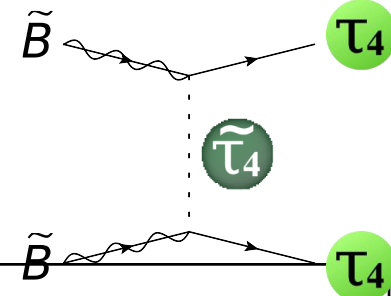
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 $\rightarrow$  correct relic density even w.  $m_{\tilde{B}} \gtrsim 100$  GeV.

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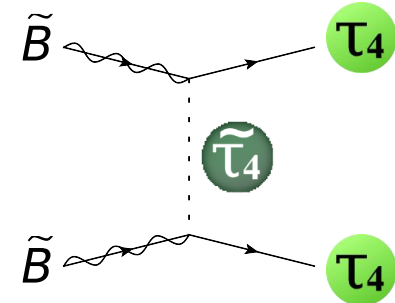
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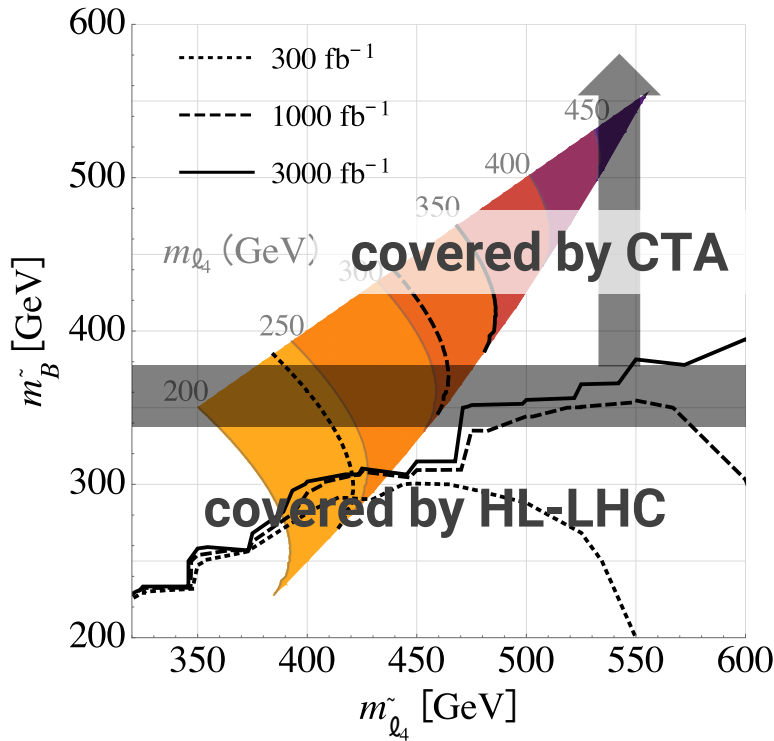


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tau-lepton

