



SUSY with light **electroweakino**

—A self introduction

[Sho IWAMOTO](#) (岩本 祥)

17 Dec. 2014

[Joint HEP Seminar](#) @ Tel Aviv University

References:

- * M. Endo, K. Hamaguchi, S. I., and T. Yoshinaga [[1303.4256](#)]
- * S. I., T. T. Yanagida, N. Yokozaki [[1407.4226](#)]

◎ **SUSY** × **LHC**

× ILC

× astrophys (AMS-02)

× cosmology (baryogenesis)

◎ Tevatron top forward-backward anomaly

◎ CMS W' searches at LHC

- ◎ **SUSY** × **LHC**
 - × ILC
 - × astrophys (**AMS-02**)
 - × cosmology (baryogenesis)
- ◎ **Tevatron** top forward-backward anomaly
- ◎ **CMS** W' searches at LHC

Experiments × Phenomenology

0. Introduction

- Why I like SUSY?
- A nightmare
- Muon $g-2$ anomaly

1. Muon $g-2$ v.s. LHC

2. CP-safe Gravity Mediation

MSSM (with R -parity)

- ✓ Fermion \longleftrightarrow Boson (“unification”)
- ✓ Radiative EWSB [Electroweak symmetry breaking]
... $m_h \sim 125 \text{ GeV}$
- ✓ Dark matter
- ✓ $(g - 2)_\mu$ anomaly
- ✓ Hierarchy problem

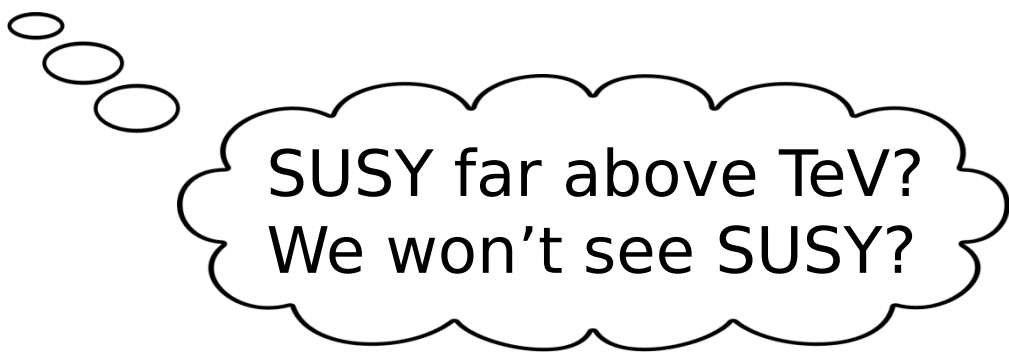
MSSM (with R -parity)

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$$m_h \sim 125 \text{ GeV}$$

$$\Rightarrow \Delta(m_h)^{\text{loop}} : \text{large}$$

$$\Rightarrow m_{\tilde{t}} = \mathcal{O}(1-10) \text{ TeV in MSSM}$$



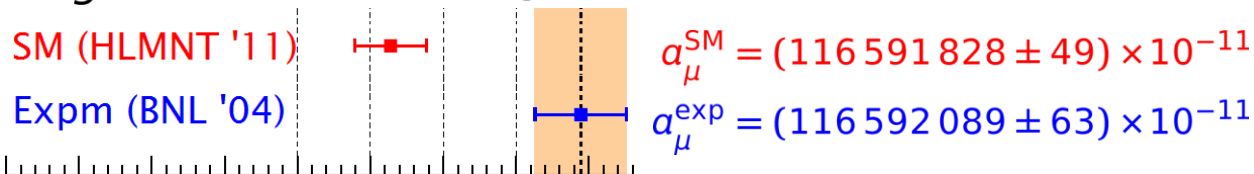
$$m_h^2(\text{MSSM}) \approx \underbrace{m_Z^2}_{\text{tree}} + \underbrace{\frac{3g_W^2 m_t^4}{8\pi^2 m_W^2} \left(\ln \frac{m_{\tilde{t}}^2}{m_t^2} + \frac{12\alpha^2 - \alpha^4}{12} \right)}_{\Delta(m_h^2)^{\text{loop}}}$$

where $\alpha := A_t/m_{\tilde{t}}$.
(stop mixing parameter)

Introduction: Muon g-2 anomaly, solved by SUSY

$$\left(a_\mu := \frac{g_\mu - 2}{2} \right)$$

Muon $g-2$ [anomalous magnetic moment]



3.3 σ

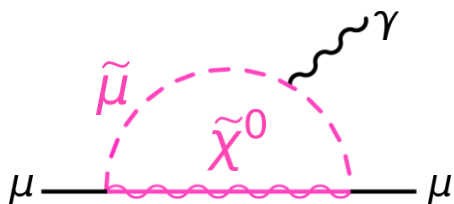
Hagiwara, Liao, Martin, Nomura, Teubner [[1105.3149](#)]

⇒ Loop contrib. from

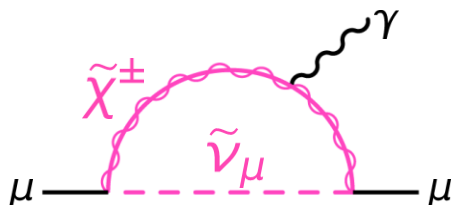
EWKino with $m = O(100)$ GeV

$(\tilde{\mu}, \tilde{\nu}_\mu, \tilde{\chi}^0, \tilde{\chi}^\pm)$

(and large $\tan \beta$, positive μ).



$$a_\mu^{\text{SUSY}}(\tilde{\chi}^0, \tilde{\mu}) \approx \frac{g_Y^2}{(4\pi)^2} \frac{m_\mu^2}{m_{\text{soft}}^2} \text{sgn}(\mu) \tan \beta + \dots,$$



$$a_\mu^{\text{SUSY}}(\tilde{\chi}^\pm, \tilde{\nu}_\mu) \approx \frac{g_2^2}{(4\pi)^2} \frac{m_\mu^2}{m_{\text{soft}}^2} \text{sgn}(\mu) \tan \beta.$$

Lopez, Nanopoulos, Wang [[ph/9308336](#)]

Chattopadhyay, Nath [[ph/9507386](#)]

Moroi [[ph/9512396](#)]

$W \ni \mu H_u H_d$ (Higgsino mass term), $\tan \beta = \langle H_u \rangle / \langle H_d \rangle$,
 m_{soft} : SUSY-particle mass-scale, g_i : Gauge couplings.

⊙ Higgs mass

$$\Rightarrow \tilde{t} = O(1-10) \text{ TeV}$$

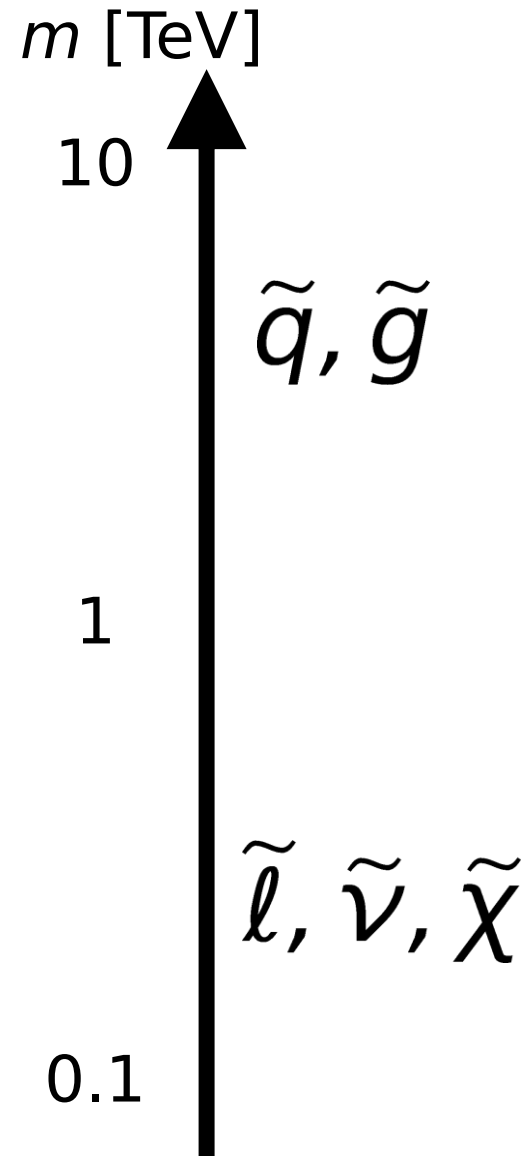
➤ Together with 8TeV LHC

$$\Rightarrow \text{colored} = O(1-10) \text{ TeV}$$

⊙ $(g - 2)_\mu$ anomaly

$$\Rightarrow \text{EWKino} = O(100) \text{ GeV}$$

(non-colored)



if EWKino = O(100)GeV,

SUSY → $(g - 2)_\mu$

$\tilde{l}, \tilde{\nu}, \tilde{\chi}^0, \tilde{\chi}^\pm$

$(g - 2)_\mu$ anomaly

O(100)GeV EWKino

Discovery @ LHC!

if $EWKino = O(100)GeV,$

SUSY \rightarrow $(g - 2)_\mu$

$\tilde{l}, \tilde{\nu}, \tilde{\chi}^0, \tilde{\chi}^\pm$

$(g - 2)_\mu$ anomaly

$O(100)GeV$ EWKino

Discovery @ LHC!

Which channel?

Reach?

~~SUSY Model?~~

if EWKino = $O(100)\text{GeV}$,

SUSY \rightarrow $(g - 2)_\mu$

$\tilde{l}, \tilde{\nu}, \tilde{\chi}^0, \tilde{\chi}^\pm$

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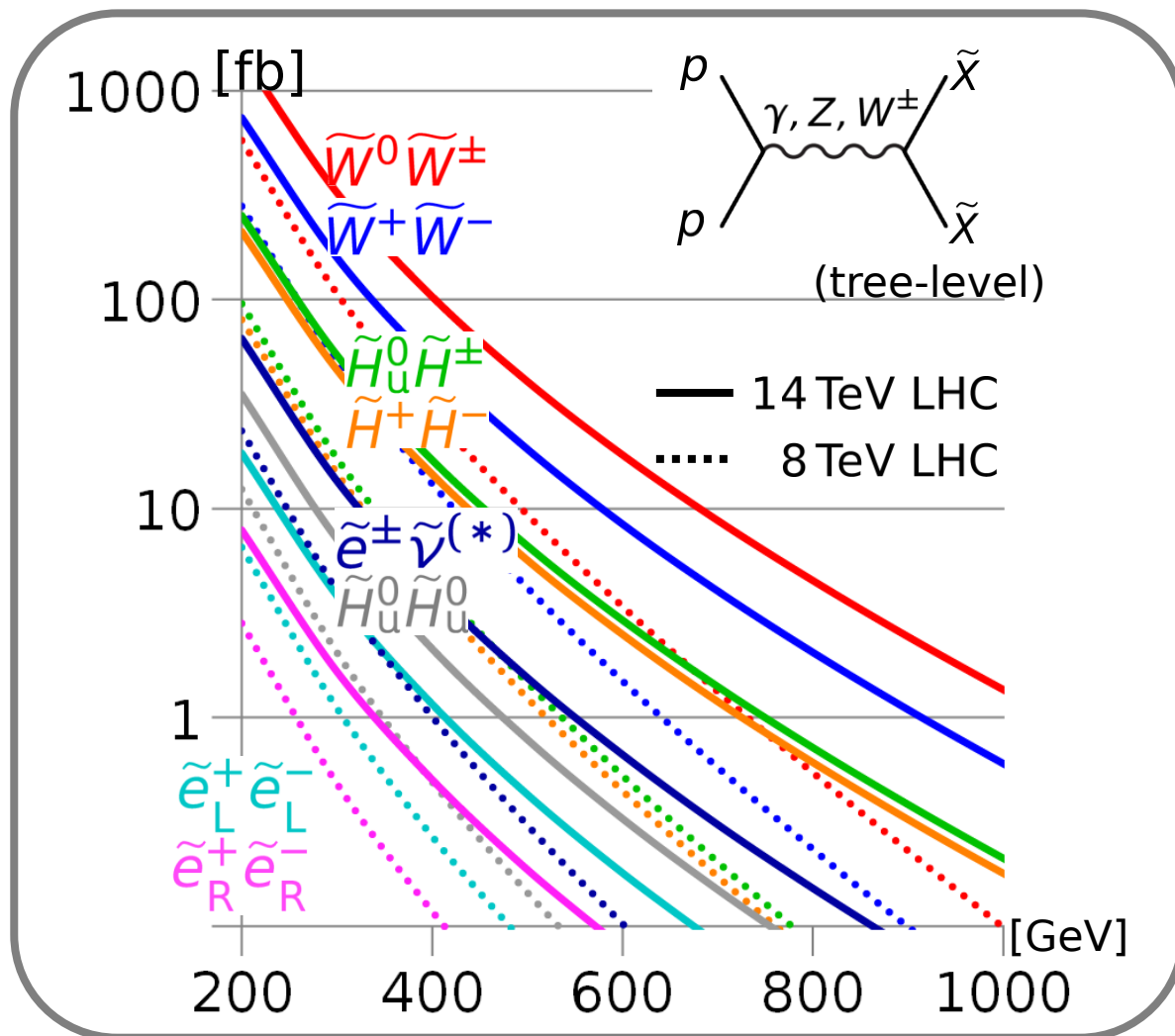
- Why I like SUSY?
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Production

$pp \rightarrow \tilde{\chi}^0 \tilde{\chi}^+, \tilde{\chi}^+ \tilde{\chi}^-$ (Wino-like)



m [TeV]

10

\tilde{q}, \tilde{g}

1

$\tilde{l}, \tilde{\nu}, \tilde{\chi}$

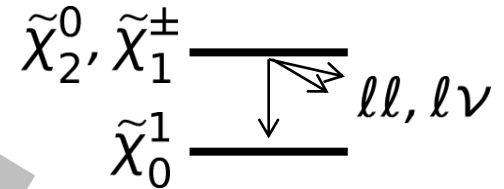
0.1

[GeV]

Production

$pp \rightarrow \tilde{\chi}^0 \tilde{\chi}^+, \tilde{\chi}^+ \tilde{\chi}^-$ (Wino-like)

Sensitive decay channels



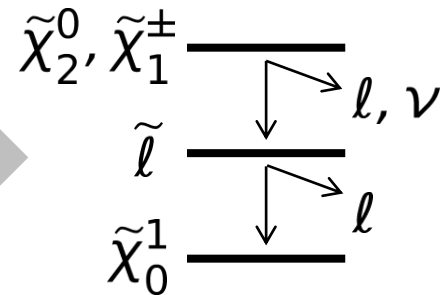
Non Z-like leptons!

LSP < \tilde{l} < $\tilde{\chi}_2^0, \tilde{\chi}_1^+$?

NO

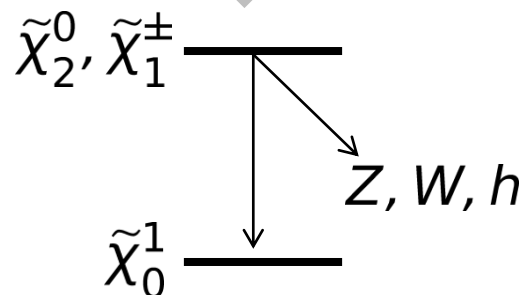
if degenerated

YES



$\tilde{\chi}^0 \tilde{\chi}^+ \rightarrow 3l + \cancel{E}_T$

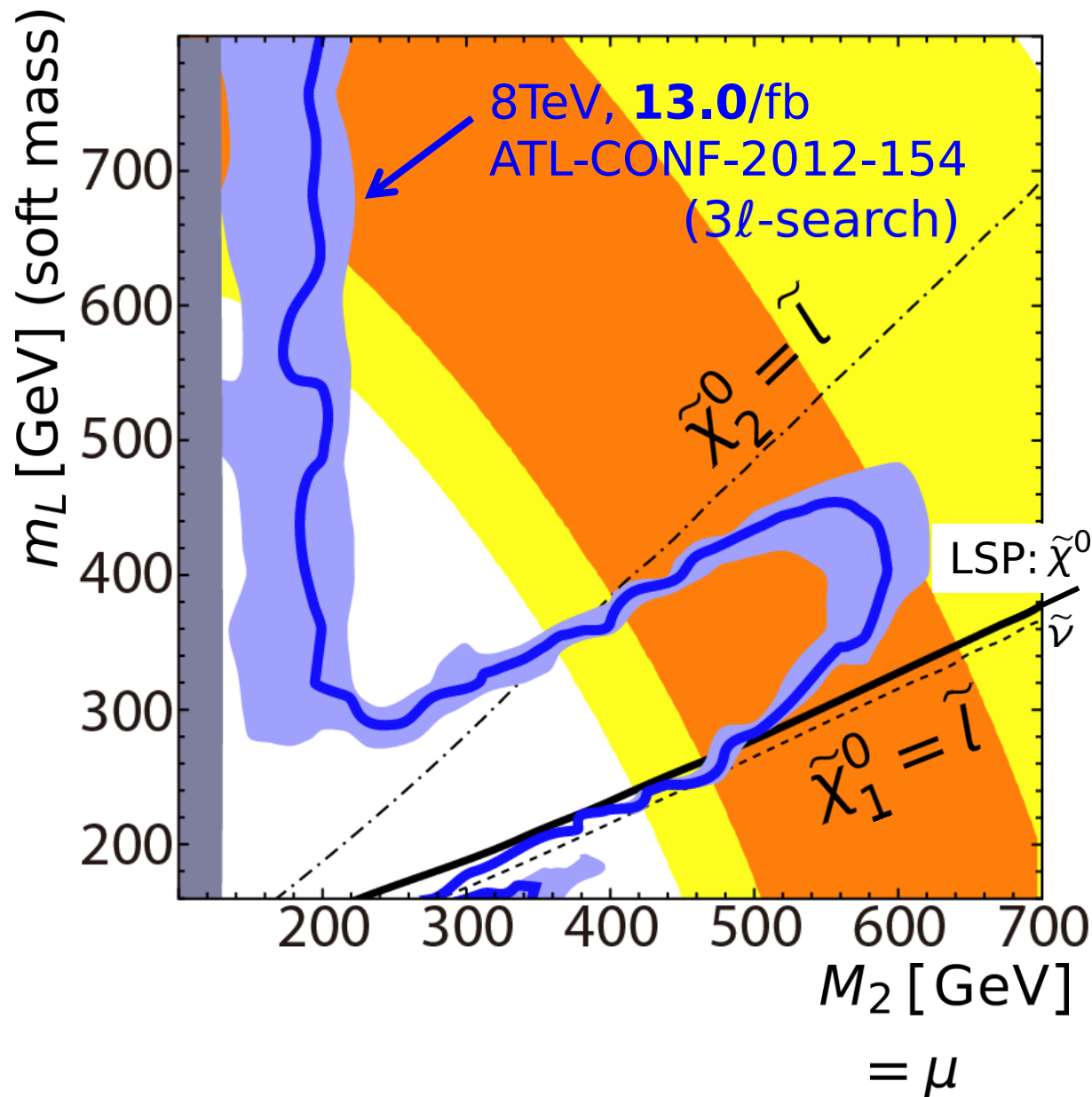
Non Z-like leptons!



$\tilde{\chi}^0 \tilde{\chi}^+ \rightarrow ZW(hW) + \cancel{E}_T$

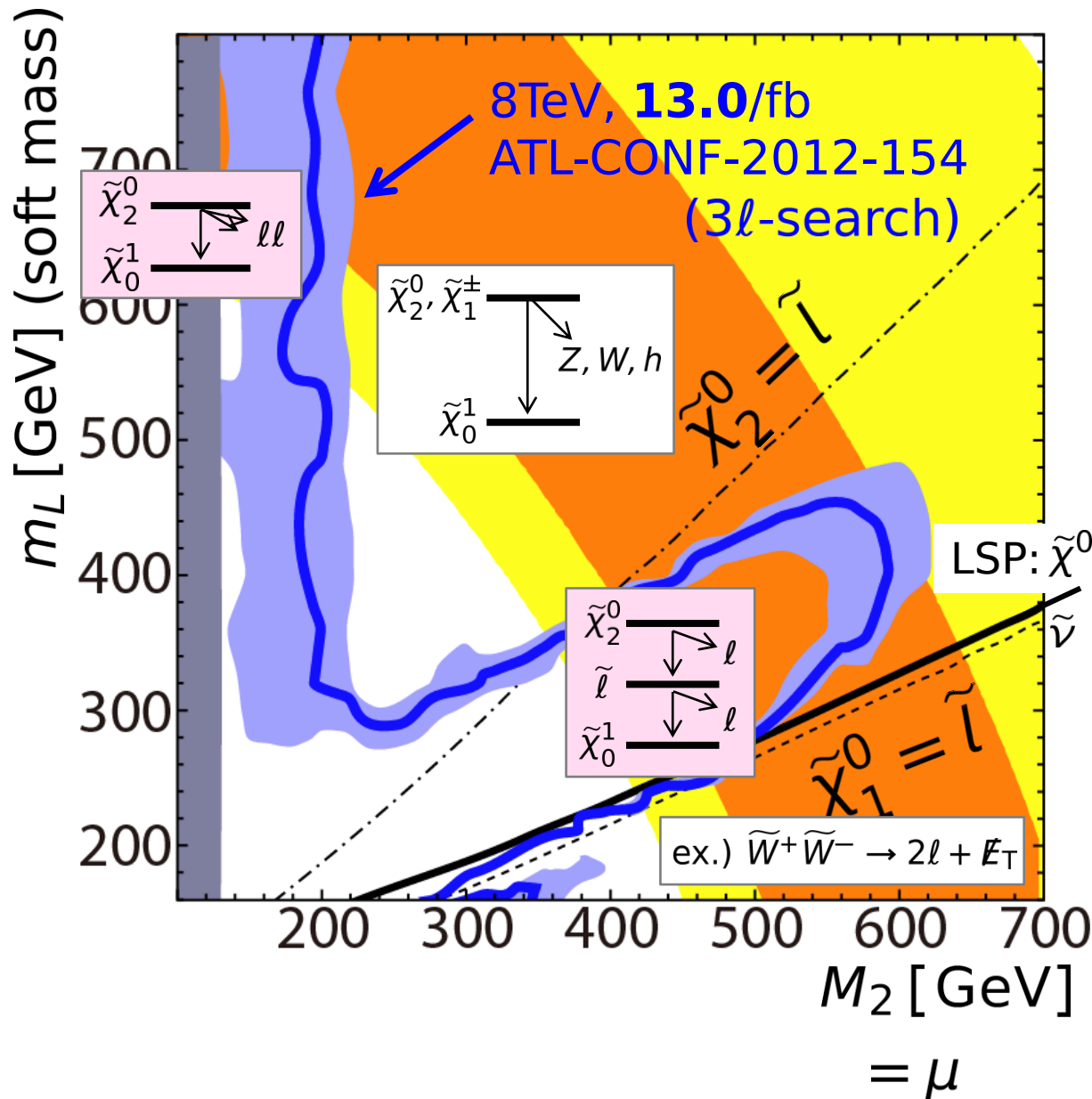
$\rightarrow 3l$

but large SM BKG...



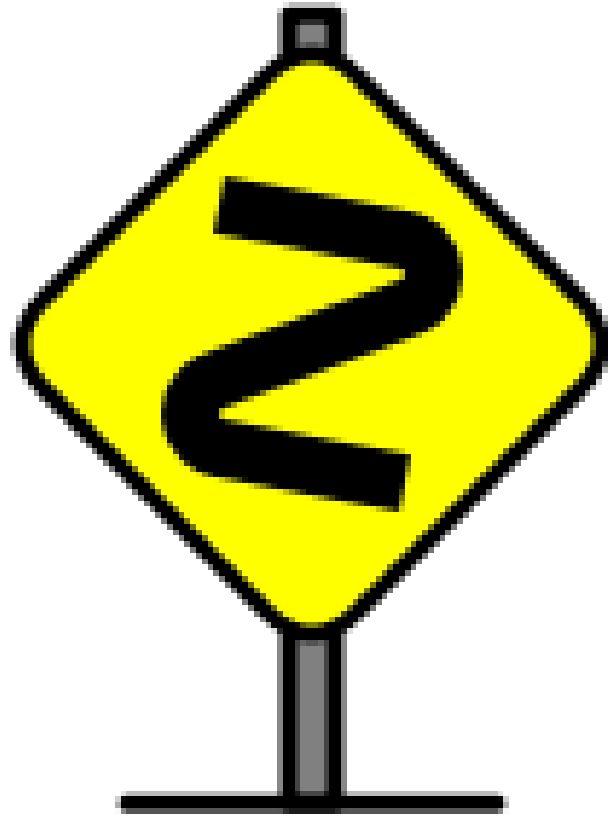
g - 2	1 σ	2 σ
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- $\tan \beta = 40$
- $M_1 = M_2/2$
- $\tilde{l}_R \gg 1 \text{ TeV}$
(as well as \tilde{q} & \tilde{g})
- $\tilde{\tau}_{L,R} \gg 1 \text{ TeV}$
- Long-lived LSP
(R-parity)
- Ω_{DM} not cared



g - 2 1σ 2σ

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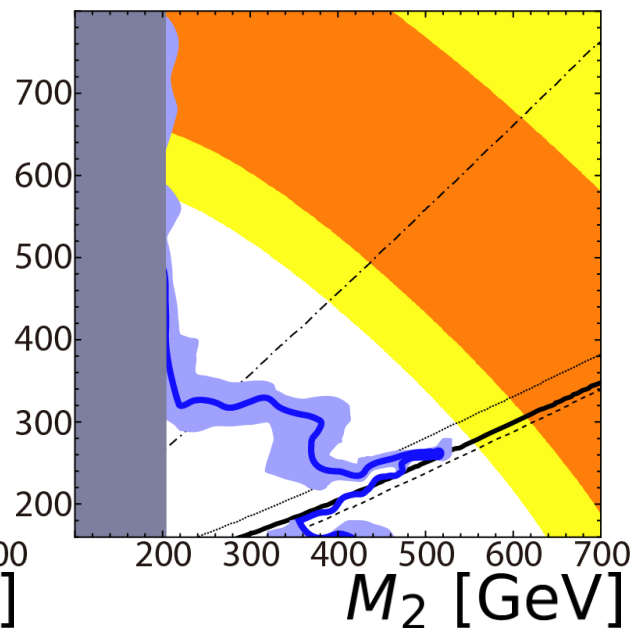
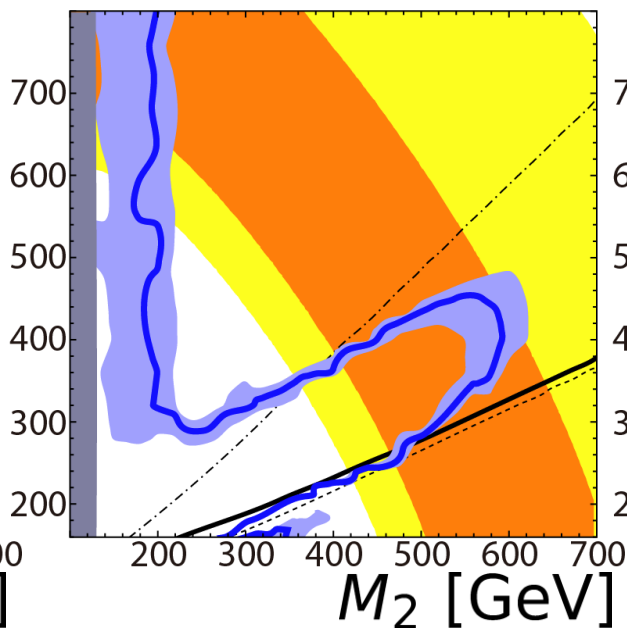
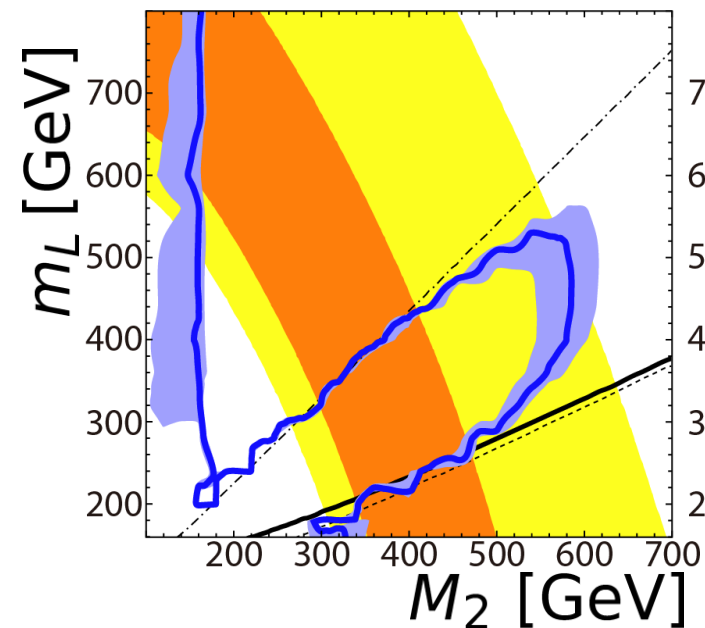




$$\mu = 2M_2$$

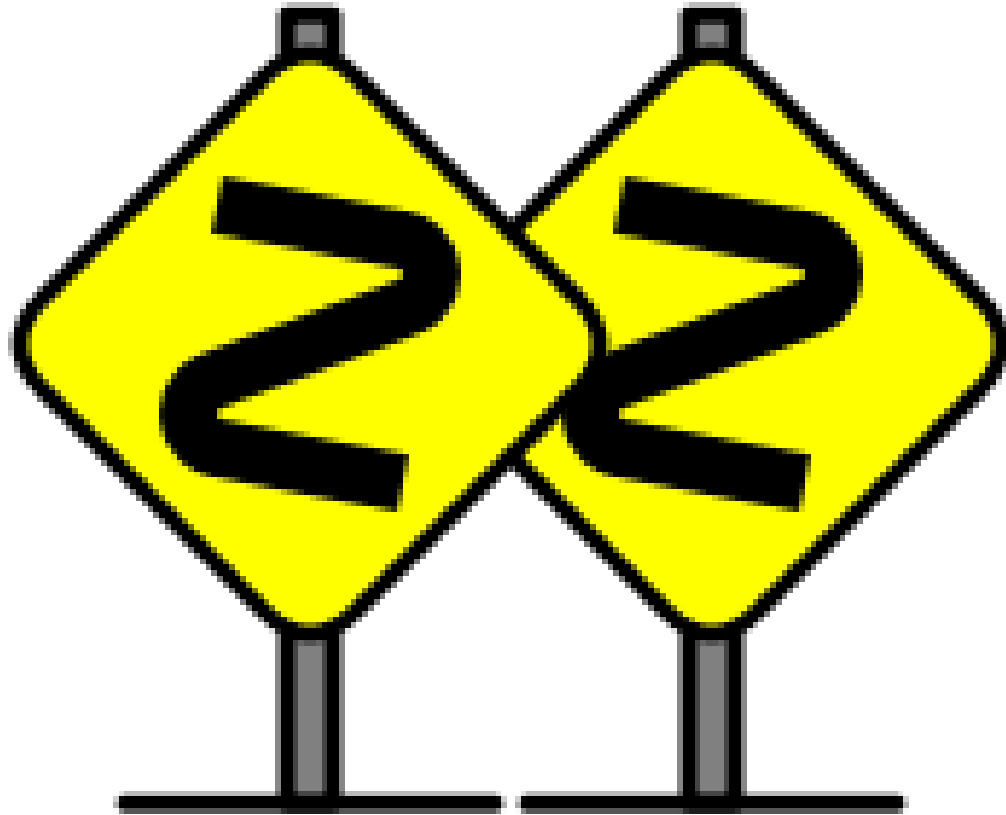
$$\mu = M_2$$

$$\mu = 0.5M_2$$



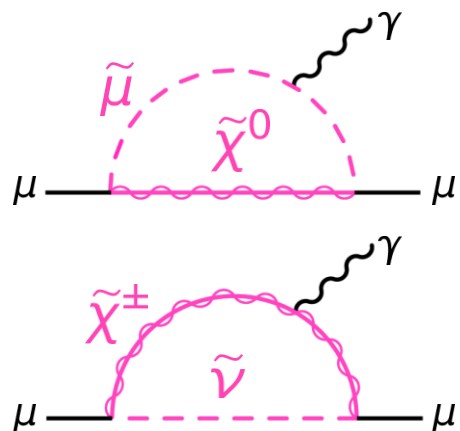
\tilde{H} is irrelevant
↓
 \tilde{W} simply to 3ℓ
via $\tilde{\ell}$

\tilde{W} can decay
via \tilde{H}
↓
Less/softer ℓ





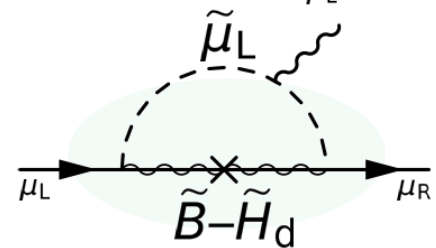
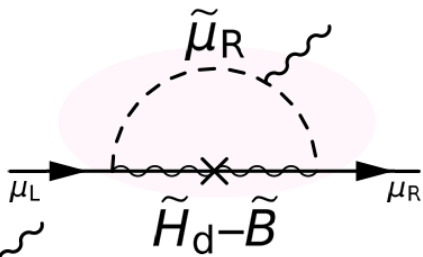
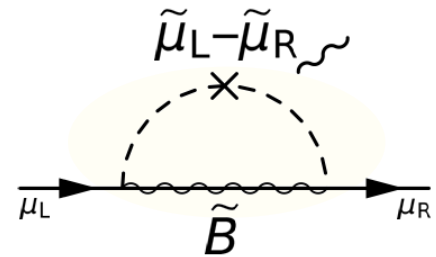
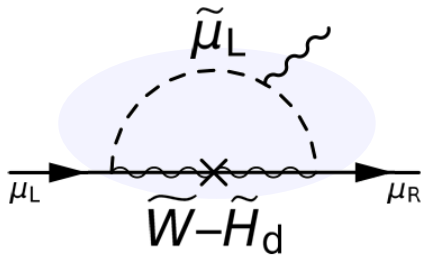
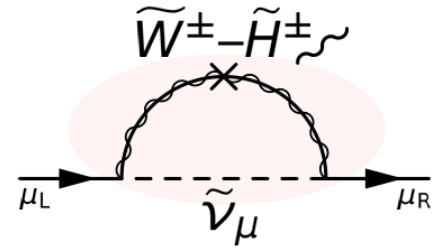
SUSY contribution to g-2 : mass basis



$$a_{\mu}^{\text{SUSY}}(\tilde{\chi}^0, \tilde{\mu}) \approx \frac{g_Y^2}{(4\pi)^2} \frac{m_{\mu}^2}{m_{\text{soft}}^2} \text{sgn}(\mu) \tan \beta + \dots,$$

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$$\frac{g_2^2 m_\mu^2}{8\pi^2} \frac{M_2 \mu \tan \beta}{m_{\tilde{\nu}_\mu}^4} \cdot F_a \left(\frac{M_2}{m_{\tilde{\nu}_\mu}}, \frac{\mu}{m_{\tilde{\nu}_\mu}} \right)$$

$$-\frac{g_2^2 m_\mu^2}{16\pi^2} \frac{M_2 \mu \tan \beta}{m_{\tilde{\mu}_L}^4} \cdot F_b \left(\frac{M_2}{m_{\tilde{\mu}_L}}, \frac{\mu}{m_{\tilde{\mu}_L}} \right)$$

$$\frac{g_Y^2 m_\mu^2}{8\pi^2} \frac{\mu \tan \beta}{M_1^3} \cdot F_b \left(\frac{m_{\tilde{\mu}_L}}{M_1}, \frac{m_{\tilde{\mu}_R}}{M_1} \right)$$

$$-\frac{g_Y^2 m_\mu^2}{8\pi^2} \frac{M_1 \cdot \mu \tan \beta}{m_{\tilde{\mu}_R}^4} \cdot F_b \left(\frac{M_1}{m_{\tilde{\mu}_R}}, \frac{\mu}{m_{\tilde{\mu}_R}} \right)$$

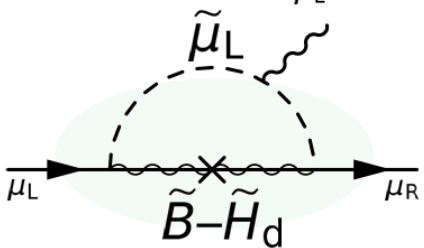
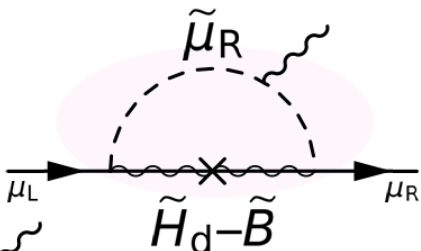
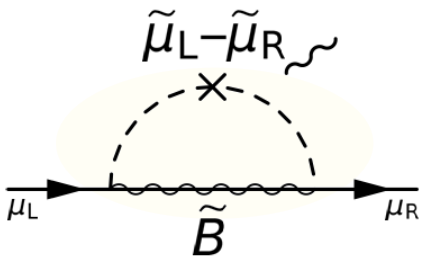
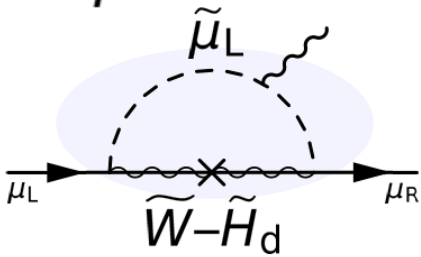
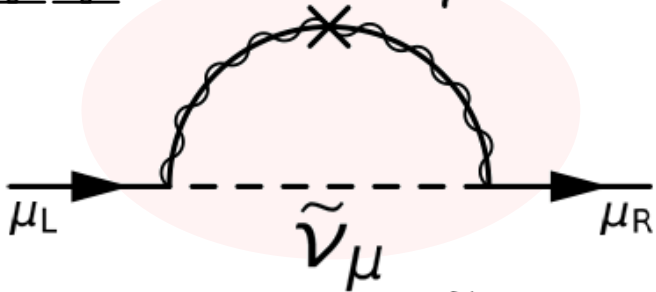
$$\frac{g_Y^2 m_\mu^2}{16\pi^2} \frac{M_1 \mu \tan \beta}{m_{\tilde{\mu}_L}^4} \cdot F_b \left(\frac{M_1}{m_{\tilde{\mu}_L}}, \frac{\mu}{m_{\tilde{\mu}_L}} \right)$$

* F_a, F_b are loop functions (>0).

SUSY contribution to g-2 : gauge basis



$\widetilde{W}^\pm - \widetilde{H}^\pm$



$$\frac{g_2^2 m_\mu^2}{8\pi^2} \frac{M_2 \mu \tan \beta}{m_{\widetilde{\nu}_\mu}^4} \cdot F_a \left(\frac{M_2}{m_{\widetilde{\nu}_\mu}}, \frac{\mu}{m_{\widetilde{\nu}_\mu}} \right)$$

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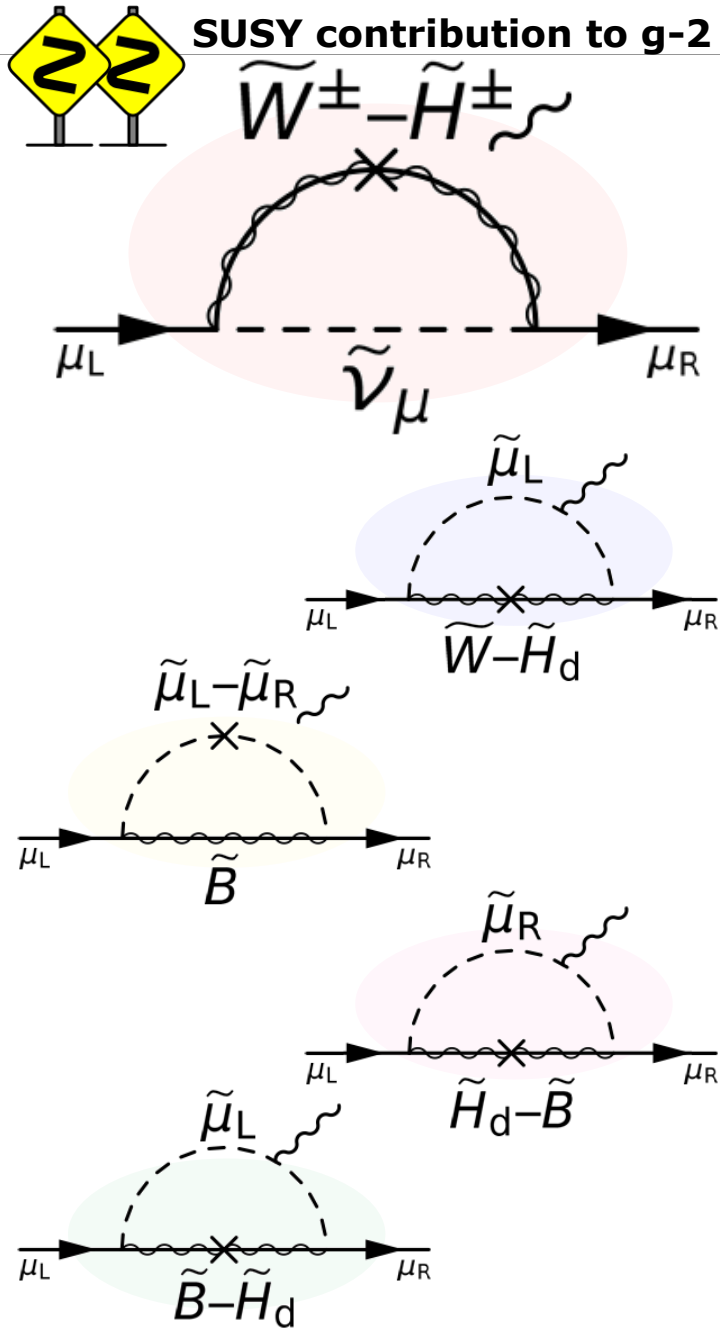
$$\frac{g_Y^2 m_\mu^2}{8\pi^2} \frac{\mu \tan \beta}{M_1^3} \cdot F_b \left(\frac{m_{\widetilde{\mu}_L}}{M_1}, \frac{m_{\widetilde{\mu}_R}}{M_1} \right)$$

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SUSY contribution to g-2 : gauge basis



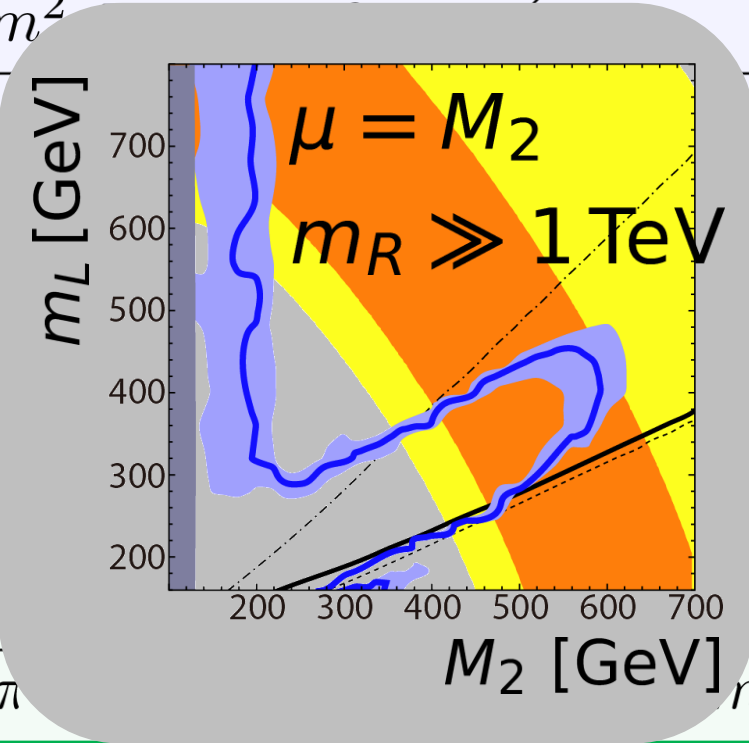
$$\frac{g_2^2 m_\mu^2}{8\pi^2} \frac{M_2 \mu \tan \beta}{m_{\tilde{\nu}_\mu}^4} \cdot F_a \left(\frac{M_2}{m_{\tilde{\nu}_\mu}}, \frac{\mu}{m_{\tilde{\nu}_\mu}} \right)$$

$$\frac{g_2^2 m_\mu^2}{16\pi^2} \frac{\mu}{m_{\tilde{\mu}_L}}$$

$$\frac{g_Y^2 \mu^2}{8\pi^2} \frac{\mu}{m_{\tilde{\mu}_R}}$$

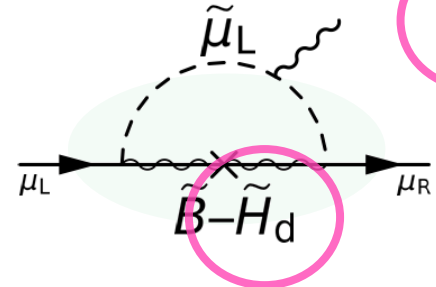
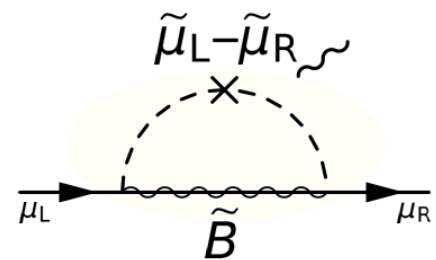
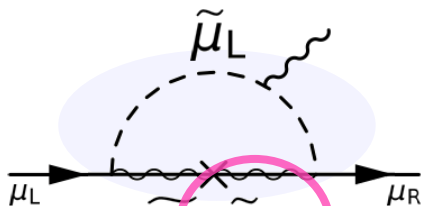
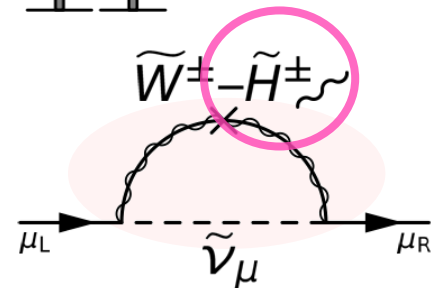
$$\frac{g_Y^2 \mu^2}{8\pi^2} \frac{\mu}{m_{\tilde{\mu}_R}}$$

$$\frac{g_Y^2 \mu^2}{16\pi^2} \frac{\mu}{m_{\tilde{\mu}_L}}$$



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SUSY contribution to g-2 : gauge basis



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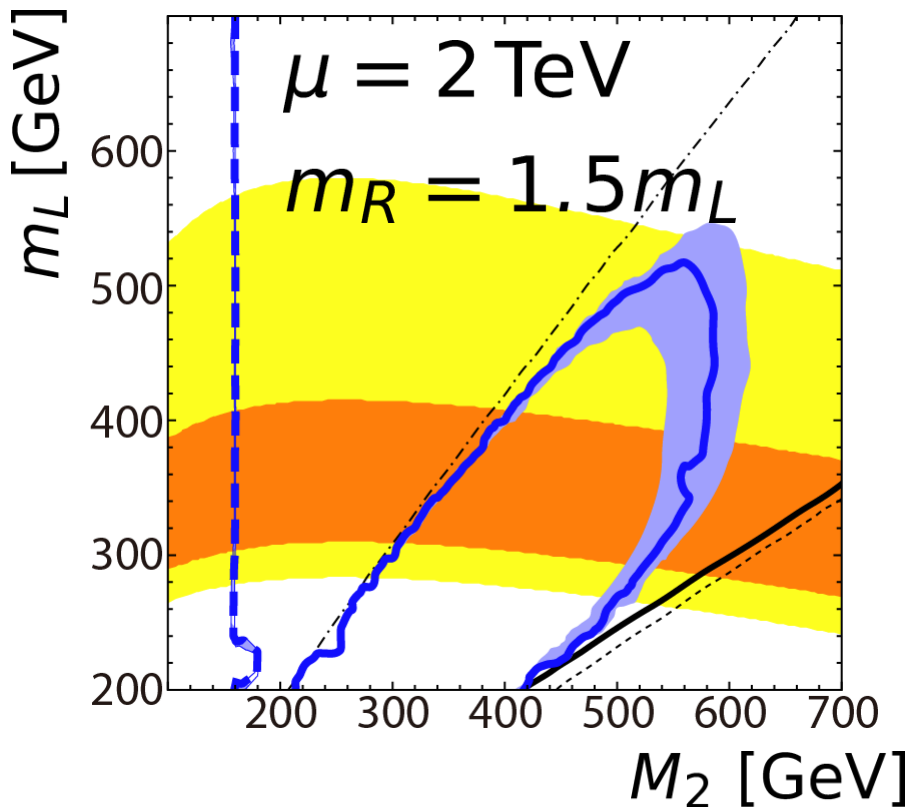
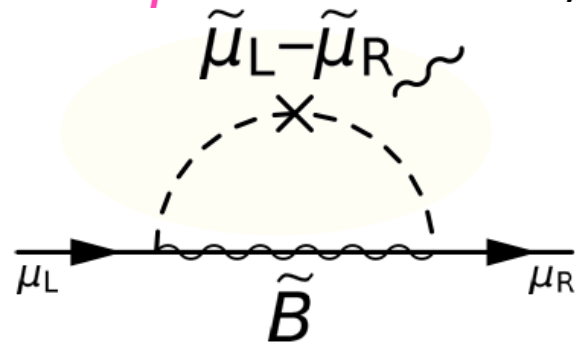
$$\frac{g_Y^2 m_\mu^2}{16\pi^2} \frac{M_1 \mu \tan \beta}{m_{\tilde{\mu}_L}^4} \cdot F_b \left(\frac{M_1}{m_{\tilde{\mu}_L}}, \frac{\mu}{m_{\tilde{\mu}_L}} \right)$$

$$M_{\tilde{\mu}}^2 = \begin{pmatrix} m(l_L)^2 & m_\mu (A_\mu^* - \mu \tan \beta) \\ m_\mu (A_\mu^* - \mu \tan \beta) & m(l_R)^2 \end{pmatrix}.$$



For $\mu \gg 100 \text{ GeV}$,

$$\frac{g_Y^2 m_\mu^2}{8\pi^2} \frac{\mu \tan \beta}{M_1^3} \cdot F_b \left(\frac{m_{\tilde{\mu}_L}}{M_1}, \frac{m_{\tilde{\mu}_R}}{M_1} \right)$$



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SUSY \rightarrow $(g - 2)_\mu$

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$(g - 2)_\mu$ anomaly

$O(100)GeV$ EWKino

Discovery @ LHC!

Which channel?

Reach?

~~SUSY Model?~~

if EWKino = O(100)GeV,

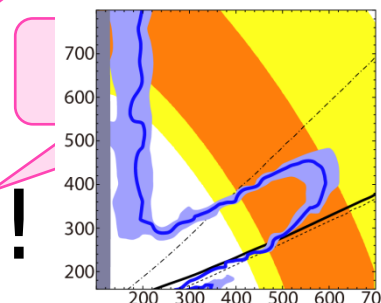
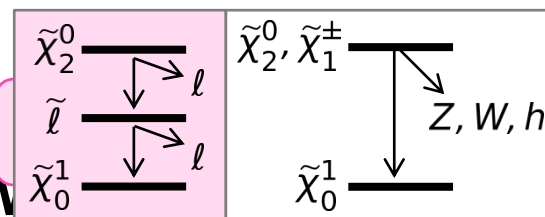
SUSY \rightarrow $(g - 2)_\mu$

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$(g - 2)_\mu$ anomaly

O(100)GeV EW

Discovery @ LHC!



SUSY Model?

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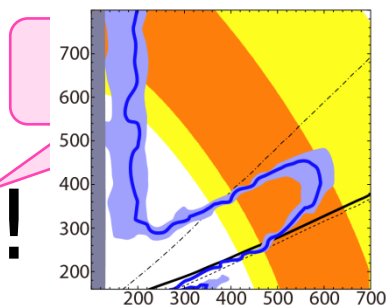
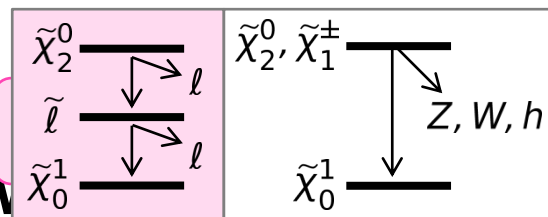
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$(g - 2)_\mu$ anomaly

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Discovery @ LHC!



SUSY Model?

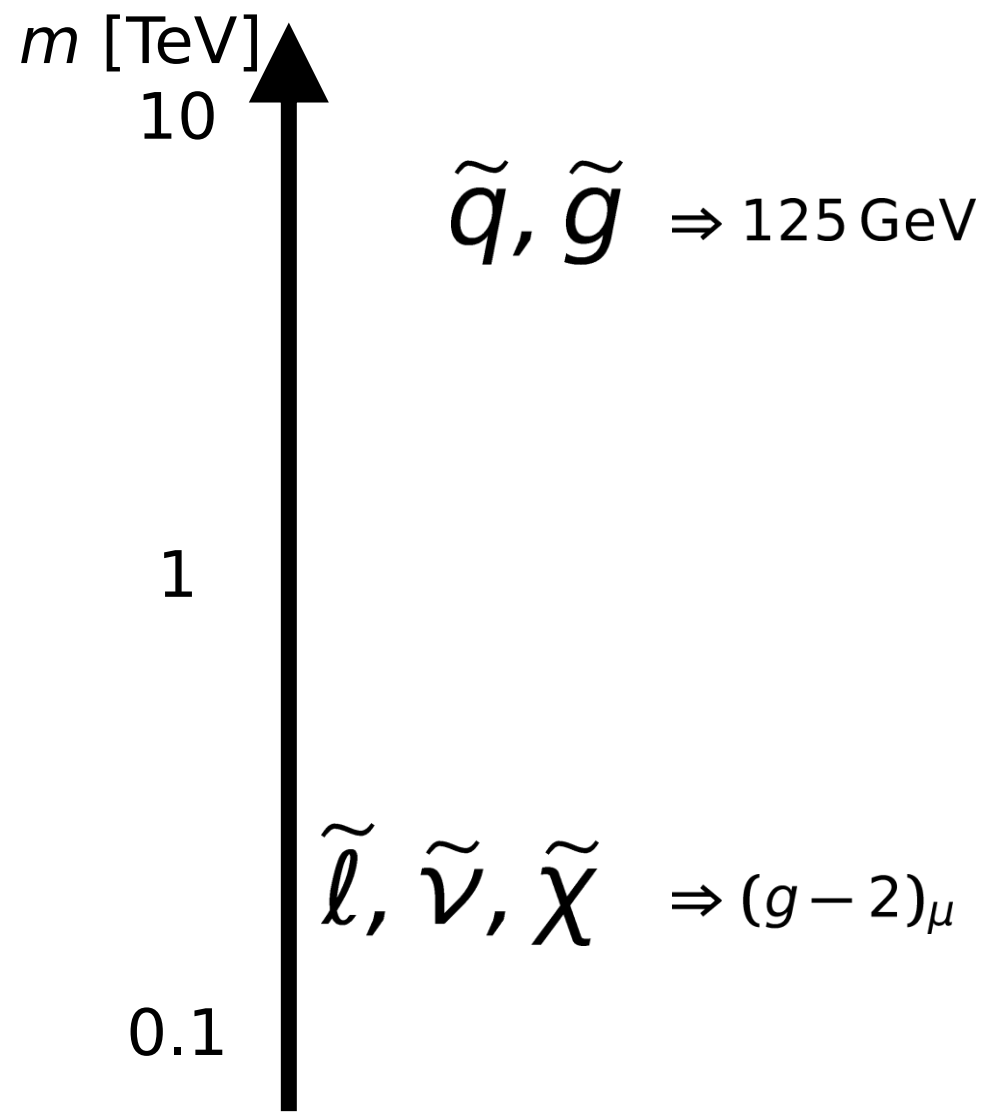
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◎ SUSY model?



⊙ SUSY model? → ~~CMSSM~~, ~~mGMSB~~

See also:

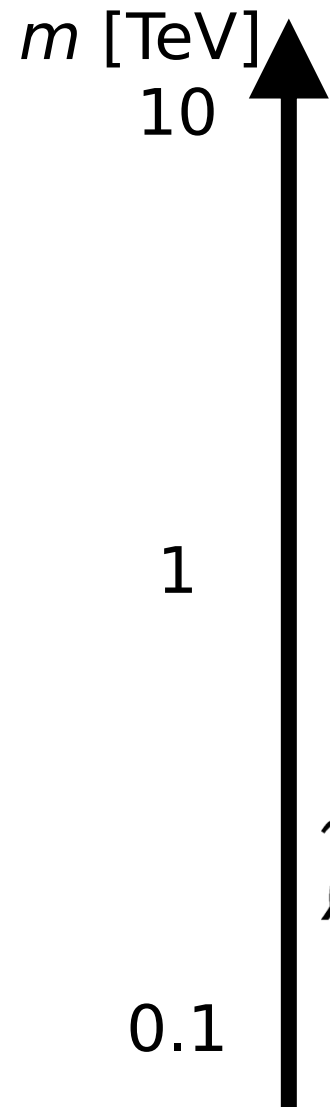
Endo, Hamaguchi, SI, Yokozaki

[[1108.3071](#)]

EHSI + Nakayama [[1112.6412](#)]

[constrained MSSM]

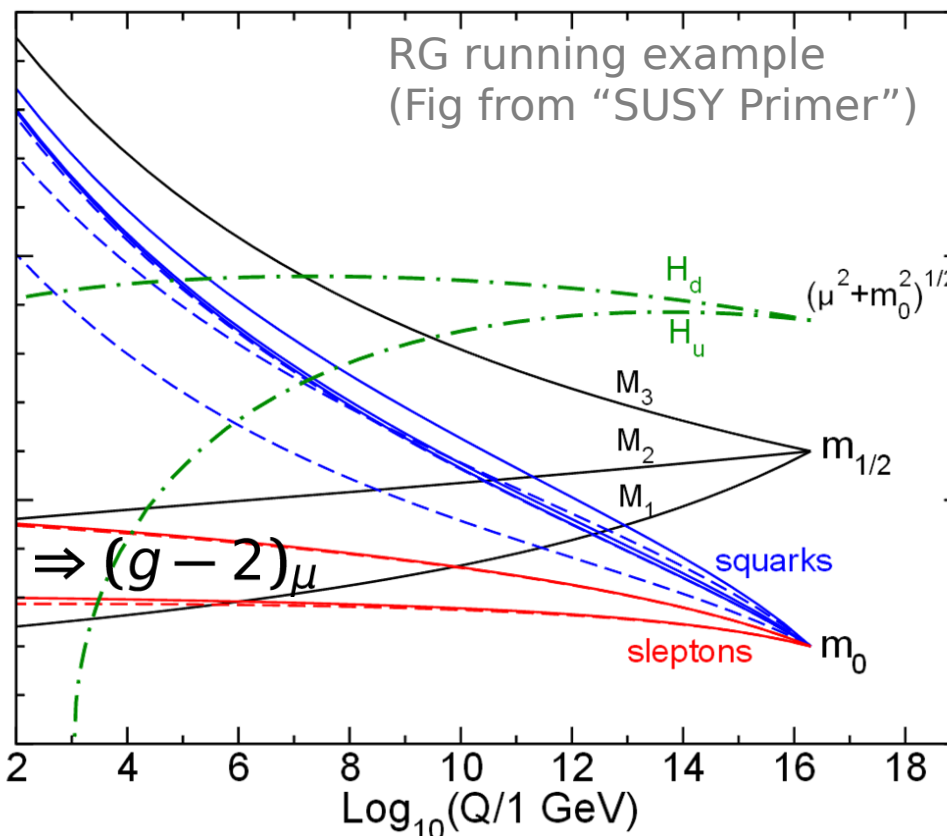
[minimal gauge-mediated SUSY breaking]



~~\tilde{q}, \tilde{g}~~ ⇒ ~~125 GeV~~

\tilde{q}, \tilde{g}

$\tilde{l}, \tilde{\nu}, \tilde{\chi}$



⊙ SUSY model? → ~~CMSSM~~ = $(m_0, M_{1/2}, A_0, \tan \beta, \text{sgn } \mu)$

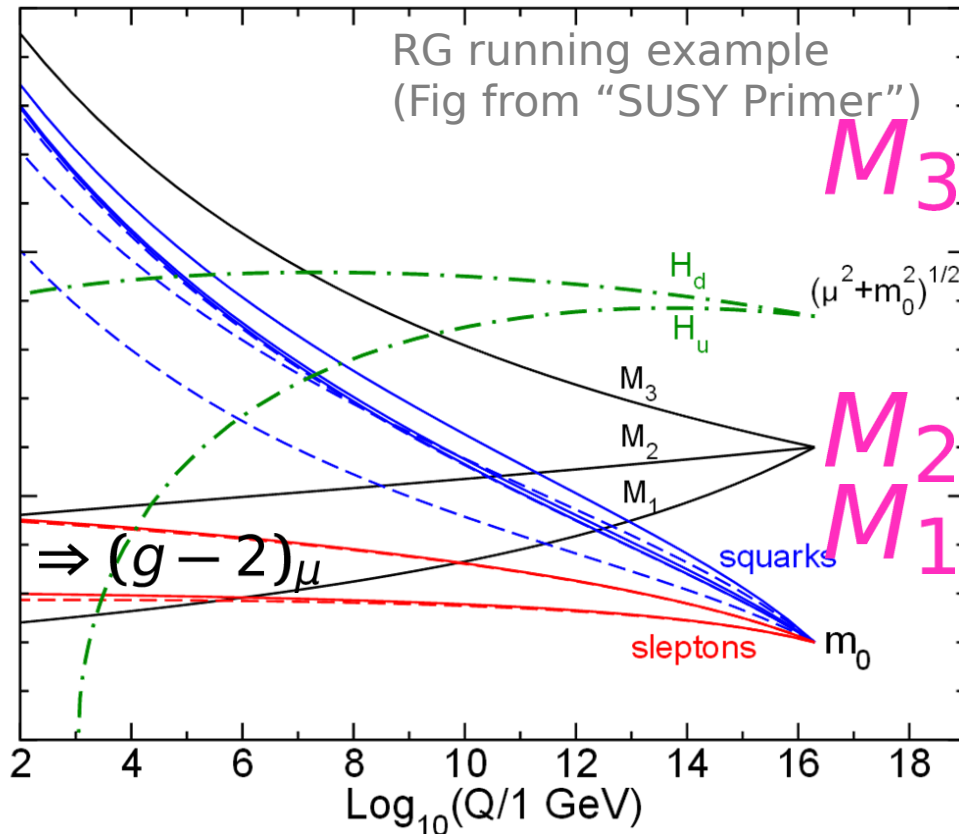
EXTEND

“NUGM” (M_1, M_2, M_3) for $\tilde{B}, \tilde{W}, \tilde{g}$
 (non-universal gaugino mass)

$\tilde{q}, \tilde{g} \Rightarrow 125 \text{ GeV}$

\tilde{q}, \tilde{g}

$\tilde{l}, \tilde{\nu}, \tilde{\chi}$



⊙ SUSY model? → ~~CMSSM~~ = $(m_0, M_{1/2}, A_0, \tan \beta, \text{sgn } \mu)$

EXTEND

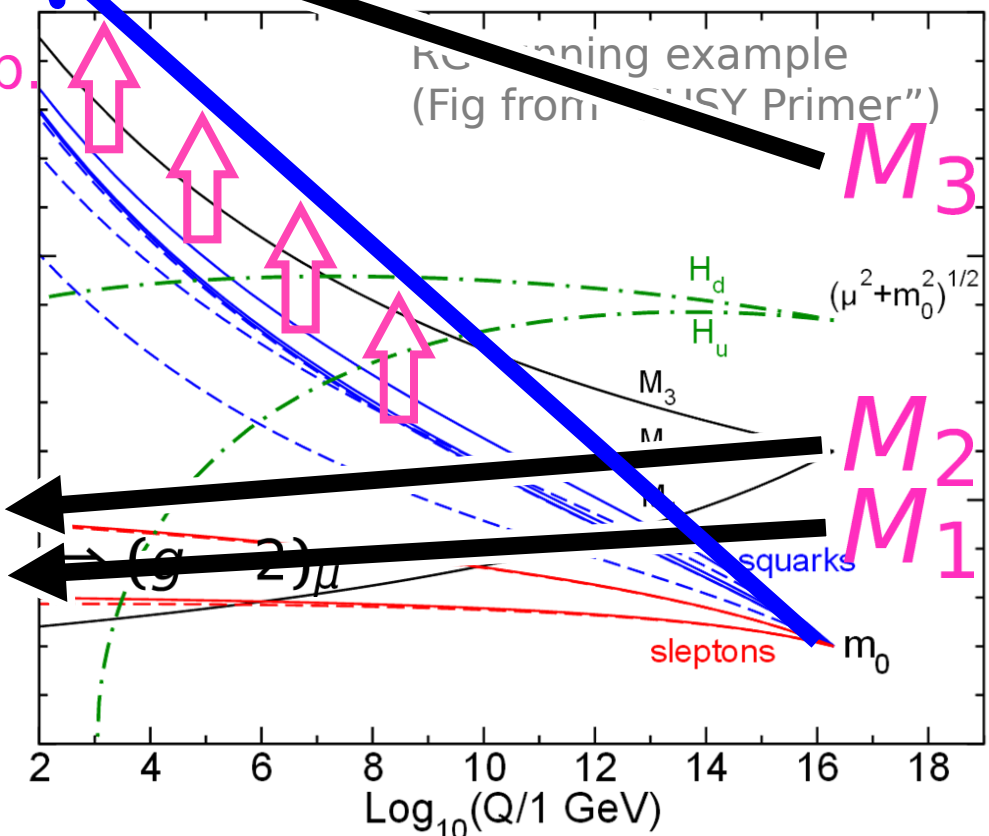
“NUGM” (M_1, M_2, M_3) for $\tilde{B}, \tilde{W}, \tilde{g}$
(non-universal gaugino mass)

m [TeV]
10
1
0.1

\tilde{q}, \tilde{g} ⇒ 125 GeV

Large M_3 contrib.

$\tilde{l}, \tilde{\nu}, \tilde{\chi}$



◎ SUSY model? → ~~CMSSM~~ = $(m_0, M_{1/2}, A_0, \tan \beta, \text{sgn } \mu)$
↓ EXTEND
“NUGM” (M_1, M_2, M_3) for $\tilde{B}, \tilde{W}, \tilde{g}$
(non-universal gaugino mass)

- Just a “parameter set”.
- Flavor & CP problems.

“CP-safe Gravity Mediation”

- Models for NUGM.
- CP-problems partially solved.

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partially

②

“CP-safe Gravity Mediation”

- Models for NUGM.
- CP-problems partially solved.

◎ CP problem?

- MSSM ... \mathbb{C} -parameters $(A_{u,d,e}, M_{1,2,3}, \mu, B)$
- Unaligned \rightarrow CP-violation

EDM $\propto \text{Im} \left[\begin{array}{c} \tilde{l}/\tilde{\nu} \\ \tilde{\chi}^0/\tilde{\chi}^\pm \end{array} \right] < 8.7 \times 10^{-29} \text{ e cm}$
[electric dipole moment] (electron) ACME collab. [[1310.7534](#)]

$m = O(100) \text{ GeV} \implies \Delta(\text{phases}) \lesssim 10^{-3} - 10^{-4}.$

$g - 2 \propto \text{Re} \left[\begin{array}{c} \tilde{l}/\tilde{\nu} \\ \tilde{\chi}^0/\tilde{\chi}^\pm \end{array} \right] \implies m = O(100) \text{ GeV}$

⊙ CP problem?

- MSSM ... \mathbb{C} -parameters ($A_{u,d,e}, M_{1,2,3}, \mu, B$)
- Unaligned \rightarrow CP-violation
- Generic problem in $O(100)\text{GeV SUSY}$.

$$\text{EDM} \propto \text{Im} \left[\text{Diagram} \right] < 8.7 \times 10^{-29} \text{ e cm} \quad (\text{electron})$$

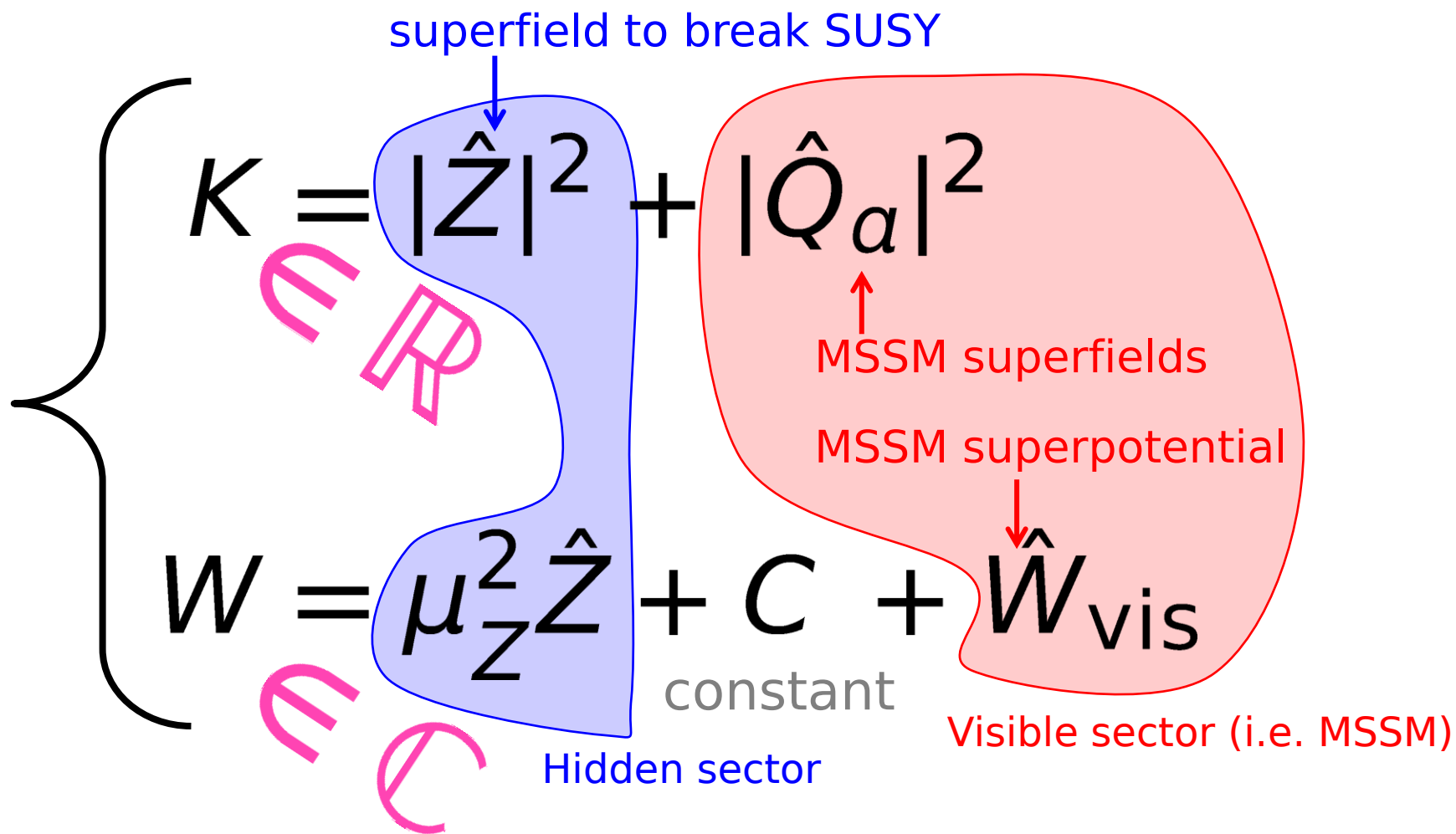
[electric dipole moment] ACME collab. [[1310.7534](#)]

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$$g - 2 \propto \text{Re} \left[\text{Diagram} \right] \implies m = O(100) \text{ GeV}$$

$$\mathcal{L} = \int d^2\theta d^2\bar{\theta} K(\Phi, \bar{\Phi}) + \left(\int d^2\theta W(\Phi) + \text{h.c.} \right) + \dots$$

Ex.) Polonyi model



$$\mathcal{L} = \int d^2\theta d^2\bar{\theta} K(\Phi, \bar{\Phi}) + \left(\int d^2\theta W(\Phi) + \text{h.c.} \right) + \dots$$

Ex.) Polonyi model

superfield to break SUSY (... \mathbb{C})

$$K = |\hat{Z}|^2 + |\hat{Q}_a|^2$$

$$W = \mu_Z^2 \hat{Z} + C + \hat{W}_{\text{vis}}$$

$\in \mathbb{C}$

→ unaligned phases

→ CP-viol.



Polonyi model: analysis

$$K = |Z|^2 + |Q_a|^2$$

$$W = \mu_Z^2 Z + C + W_{\text{vis}}$$

$$V = K_i^j F_j F^{i*} - 3e^{K/M_P^2} |W/M_P|^2 + (D\text{-term})$$

$$\langle V \rangle = 0, \quad \left\langle \frac{\partial V}{\partial Z} \right\rangle = 0 \quad \Rightarrow \quad \langle Z \rangle = (\sqrt{3} - 1) e^{i \arg(C/\mu_Z^2)} M_P \equiv z$$

↑
[vanishing cosmo. const.]

$$\langle F_Z \rangle = -e^{K/2M_P^2} \left(\frac{\partial W^*}{\partial Z^*} + \frac{\partial K}{\partial Z^*} \frac{W^*}{M_P^2} \right)$$

$$= -\sqrt{3} e^{\sqrt{3}-2} \mu_Z^{2*}$$

$$\Rightarrow \langle Z \rangle \neq 0, \quad \langle F_Z \rangle \neq 0 \quad (\text{SUSY})$$

$$V \ni -e^{K/2M_P^2} \frac{\langle F_Z \rangle}{M_P^2} \left\langle \frac{\partial K}{\partial Z} \right\rangle W_{\text{vis}} + e^{K/M_P^2} \frac{\langle W^* \rangle}{M_P^2} \left[\frac{\partial W}{\partial Q_a} Q_a - 3W_{\text{vis}} \right]$$

$$\Rightarrow B_\mu = (2/\sqrt{3} - 1) e^{-i \arg C} \frac{|\langle F_Z \rangle|}{M_P} \mu, \quad A_{ijk} = (\sqrt{3} - 1) e^{-i \arg C} \frac{|\langle F_Z \rangle|}{M_P}$$

$$\mathcal{L} \supset \int d^2\theta \left(\frac{1}{4g_i^2} + k \frac{Z}{M_P} \right) W^{i\alpha} W_\alpha^i \Rightarrow M_i = 2kg_i^2 \frac{\langle F_Z \rangle}{M_P}$$

(set $k \in \mathbb{R}$ by rotation of Z)

◎ SUSY model? → ~~CMSSM~~ = $(m_0, M_{1/2}, A_0, \tan \beta, \text{sgn } \mu)$
↓ EXTEND
“NUGM” (M_1, M_2, M_3) for $\tilde{B}, \tilde{W}, \tilde{g}$
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partially

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“CP-safe Gravity Mediation”

- Models for NUGM.
- CP-problems partially solved.



Feature: ➤ Shift symmetry on Z ($Z \rightarrow Z + i\mathcal{R}$)
 ($\mathcal{R} \in \mathbb{R}$)
 ➤ ~~SUSY~~ in Kähler K
 ($\Rightarrow \langle Z \rangle \in \mathbb{R}$)
 Izawa, Kugo, Yanagida [1008.4641]

$$\left\{ \begin{array}{l}
 K = \underline{s(x)} + |Q_a|^2 \\
 \text{complicated function} \\
 W = \underline{C} + W_{\text{vis}} \\
 \dots \text{only one } \mathbb{C} \rightarrow \text{no misalignment} \\
 \text{where } x \equiv Z + Z^* \\
 \text{[invariant]}
 \end{array} \right.$$

Note for professionals: $R(H_u H_d) = 2;$
 $m_{3/2} = \mathcal{O}(100) \text{ GeV}.$



Feature: ➤ Shift symmetry on Z ($Z \rightarrow Z + i\mathbb{R}$)
 ➤ ~~SUSY~~ in Kähler K ($\mathbb{R} \in \mathbb{R}$)
 (Izawa, Kugo, Yanagida [1008.4641]) $(\Rightarrow \langle Z \rangle \in \mathbb{R})$

$$K = \underline{s(x)} + |Q_a|^2$$

complicated function but $\mathbb{R} (\because K \in \mathbb{R})$

Example:

$$s(x) = -3 \log \frac{-f(x)}{3}; \quad f(x) = -3 + c_1 x + c_2 x^2 + c_3 x^3 + c_4 x^4$$

$$\text{with } c_2 = 3c_3^2/8c_4, \quad 768c_4^3 + 64c_1c_3c_4^2 - 3c_3^4 > 0.$$

$$\Rightarrow \langle x \rangle = -c_3/4c_4.$$

~~SUSY~~

where $x \equiv Z + Z^*$
 [invariant]

Note for professionals: $R(H_u H_d) = 2;$
 $m_{3/2} = \mathcal{O}(100) \text{ GeV}.$



CP-safe gravity mediation: Analysis

$$V = K_i^j F_j F^{i*} - 3e^{K/M_P^2} |W/M_P|^2 + (D\text{-term})$$

$$K = s(x) + |Q_a|^2$$

$$W = C + W_{\text{vis}}$$

with $x = Z + Z^*$

$$\langle V \rangle \underset{\uparrow}{=} 0, \quad \left\langle \frac{\partial V}{\partial Z} \right\rangle = 0 \quad \longleftrightarrow \quad \frac{s'(x)^2}{M_P^2 s''(x)} = 3, \quad \frac{\partial}{\partial x} \frac{s'(x)^2}{s''(x)} = 0$$

[vanishing cosmo. const.]

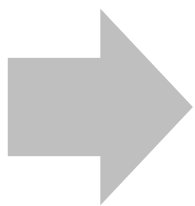
(constraints on the function s)

$$\implies \langle F_Z \rangle = -3e^{s/2M_P^2} C^* / s' \quad (\text{SUSY}),$$

$$\langle Z \rangle \neq 0$$

$$V \ni -e^{K/2M_P^2} \frac{\langle F_Z \rangle}{M_P^2} \left\langle \frac{\partial K}{\partial Z} \right\rangle W_{\text{vis}} + e^{K/M_P^2} \frac{\langle W^* \rangle}{M_P^2} \left[\frac{\partial W}{\partial Q_a} Q_a - 3W_{\text{vis}} \right]$$

$$= 3e^{K/M_P^2} \frac{C^*}{M_P^2} W_{\text{vis}} + e^{K/M_P^2} \frac{C^*}{M_P^2} \left[\frac{\partial W_{\text{vis}}}{\partial Q_a} Q_a - 3W_{\text{vis}} \right]$$



$$B_\mu / \mu = 2m_{3/2} e^{-i \arg C}$$

$$A_0 = 3m_{3/2} e^{-i \arg C}$$

$$m_0^2 = m_{3/2}^2 \quad \text{where } m_{3/2}^2 = |C|^2 e^{s/M_P^2} / M_P^4$$



Gaugino mass M_i \longleftarrow $\mathcal{L} \ni \int d^2\theta \left(\frac{1}{4g_i^2} + k_i \frac{Z}{M_P} \right) W^\alpha W_\alpha$

From anomaly of (broken) shift sym. ($\implies k_i \in \mathbb{R}$)

$$M_i = k_i g_i^2 \frac{\langle F_Z \rangle}{M_P} = k_i g_i^2 m_{3/2} \frac{3M_P}{-s'(x)} e^{-i \arg C}$$

$M_1, M_2 \ll M_3$ from "above-GUT" structure.

Ex.) $SU(5) \times U(3) \rightarrow SU(3) \times SU(2) \times U(1) \times SU(3) \times U(1)$
 $\rightarrow SU(3)_c \times SU(2)_L \times U(1)_Y$
 $k_3 = a_5 + a_3 \quad k_2 = a_5 \quad k_1 = a_5 + a_1$

$$\left\{ \begin{array}{l} B_\mu / \mu = 2m_{3/2} e^{-i \arg C} \\ A_0 = 3m_{3/2} e^{-i \arg C} \\ m_0^2 = m_{3/2}^2 \end{array} \right. \quad \text{where } m_{3/2}^2 = |C|^2 e^{s/M_P^2} / M_P^4$$

$$\begin{array}{l} K = s(x) + |Q_a|^2 \\ W = C + W_{\text{vis}} \end{array}$$



“CP-safe Gravity Mediation”

- Models for NUGM.
- CP-problems partially solved.

$$M_i = k_i g_i^2 \frac{\langle F_Z \rangle}{M_P} = k_i g_i^2 m_{3/2} \frac{3M_P}{-s'(x)} e^{-i \arg C}$$

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\downarrow \swarrow \searrow \swarrow
 $\rightarrow SU(3)_c \times SU(2)_L \times U(1)_Y$
 $k_3 = a_5 + a_3$ $k_2 = a_5$ $k_1 = a_5 + a_1$

$$\left\{ \begin{array}{l} B_\mu / \mu = 2m_{3/2} e^{-i \arg C} \\ A_0 = 3m_{3/2} e^{-i \arg C} \\ m_0^2 = m_{3/2}^2 \end{array} \right. \quad \text{where } m_{3/2}^2 = |C|^2 e^{s/M_P^2} / M_P^4$$

$$\begin{array}{l} K = s(x) + |Q_a|^2 \\ W = C + W_{\text{vis}} \end{array}$$



“CP-safe Gravity Mediation”

- Models for NUGM.
- CP-problems **partially** solved.

$$M_i = k_i g_i^2 \frac{\langle F_Z \rangle}{M_P} = k_i g_i^2 m_{3/2} \frac{3M_P}{-s'(x)} e^{-i \arg C}$$

$M_1, M_2, M_3, B_\mu, A_u, A_d, A_e$ are aligned.

CP-violation from sfermion $(m_0^2)_{ij}$ still remains.

→ should be solved together with flavor problem.

(OR: $m_0^2 = 0$?)

$$\left\{ \begin{array}{l} B_\mu / \mu = 2m_{3/2} e^{-i \arg C} \\ A_0 = 3m_{3/2} e^{-i \arg C} \\ m_0^2 = m_{3/2}^2 \end{array} \right. \quad \text{where } m_{3/2}^2 = |C|^2 e^{s/M_P^2} / M_P^4$$

$$\begin{array}{l} K = s(x) + |Q_a|^2 \\ W = C + W_{\text{vis}} \end{array}$$

Summary

if EWKino = O(100)GeV,

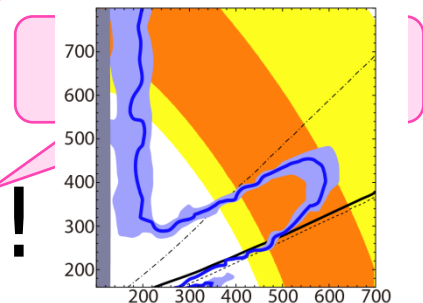
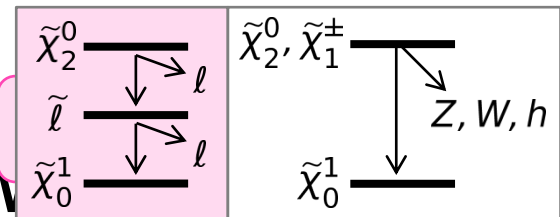
SUSY \rightarrow $(g - 2)_\mu$

$\tilde{l}, \tilde{\nu}, \tilde{\chi}^0, \tilde{\chi}^\pm$

$(g - 2)_\mu$ anomaly

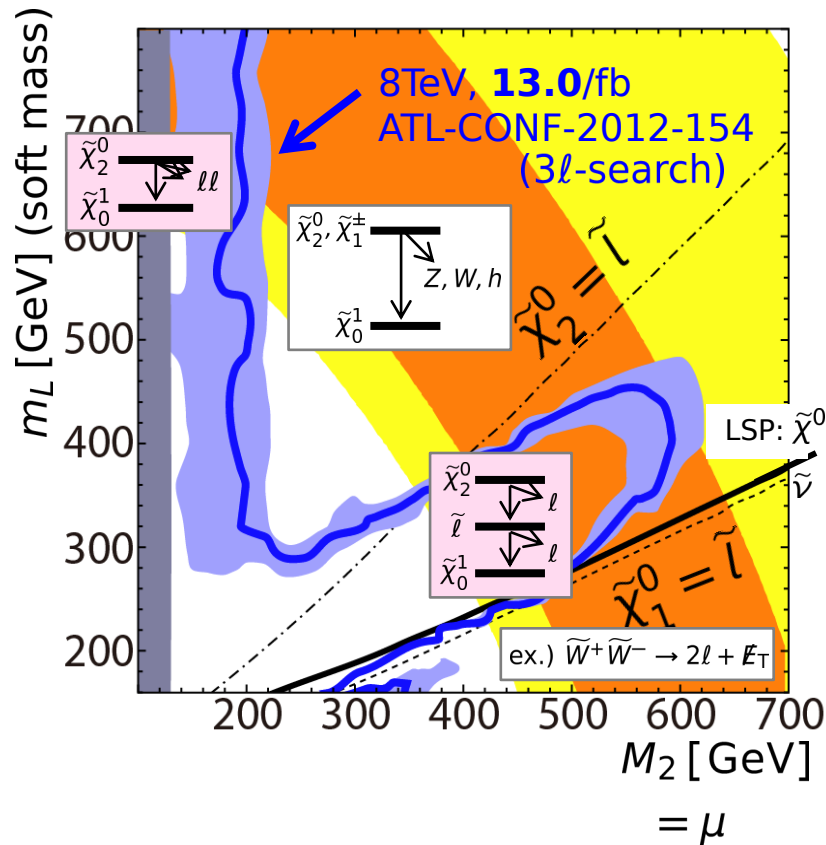
O(100)GeV EW

Discovery @ LHC!



“CP-safe GM” \leftarrow

SUSY Model?



“CP-safe Gravity Mediation”
 = Shift sym. of Z & ~~SUSY~~ in K

- Models for **NUGM**.
- **CP**-problems partially solved.

$$K = s(x) + |Q_a|^2$$

$$W = C + W_{\text{vis}}$$

where $x \equiv Z + Z^*$

SUSY from the viewpoint of $(g - 2)_\mu$