



# Long-Lived stau Signature in the LHC

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Based on <sup>ATLAS collaboration</sup>  
Asai, Azuma, Endo, Hamaguchi, and Iwamoto.  
*Stau Kinks at the LHC.* Theory group (Phenomenologists)  
JHEP 1112 (2011) 077. [arXiv: 1103.1881] (hep-ph)

# Talk Plan

## 1. SUSY

- SUSY search
- The LHC experiment

## 2. Long-lived stau signature

- “stable stau” signature
- “stau kink” signature

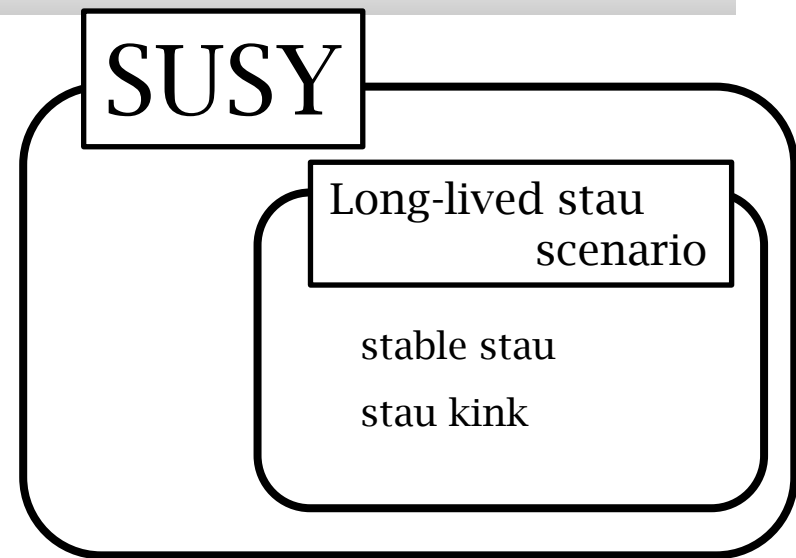
## 3. Stau Kink in detail

Based on

**Asai, Azuma, Endo, Hamaguchi, and Iwamoto.**

*Stau Kinks at the LHC.*

JHEP 1112 (2011) 077. [[arXiv: 1103.1881](https://arxiv.org/abs/1103.1881)] (hep-ph)



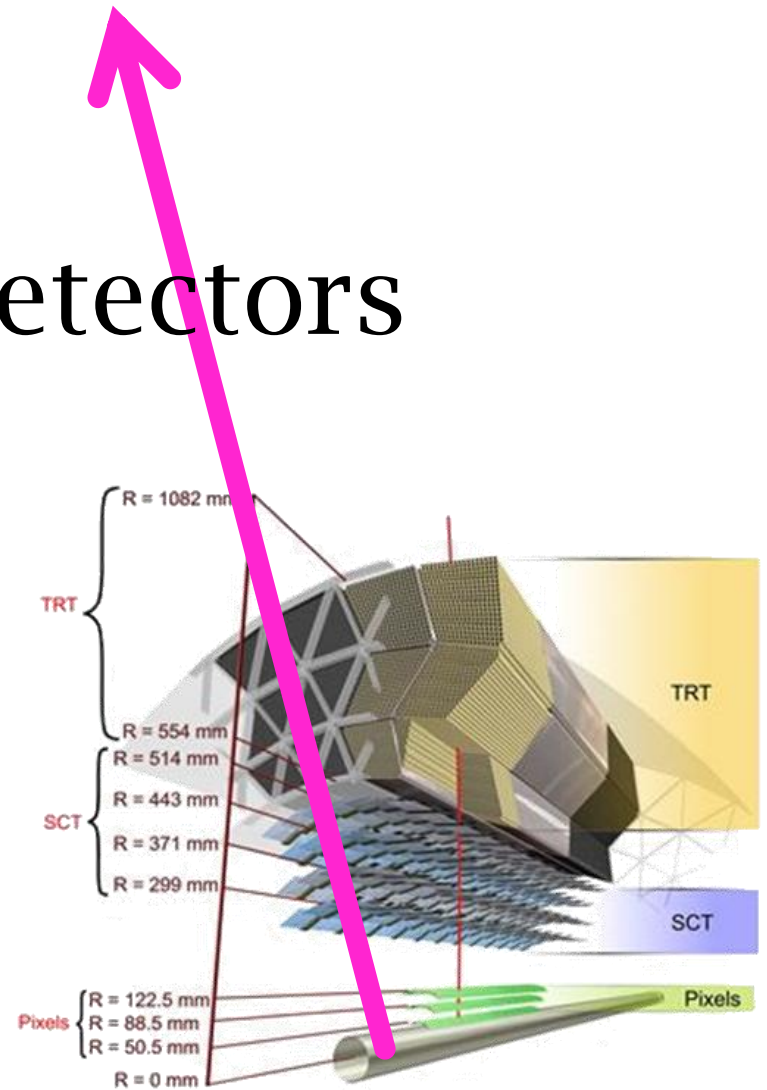
# CONCLUSION

or  
THE  
MAIN  
MESSAGE

Stau = Charged

Long-lived stau

⇒ a track in detectors



ATLAS detectors (in the LHC)

Stau = Charged

Long-lived stau

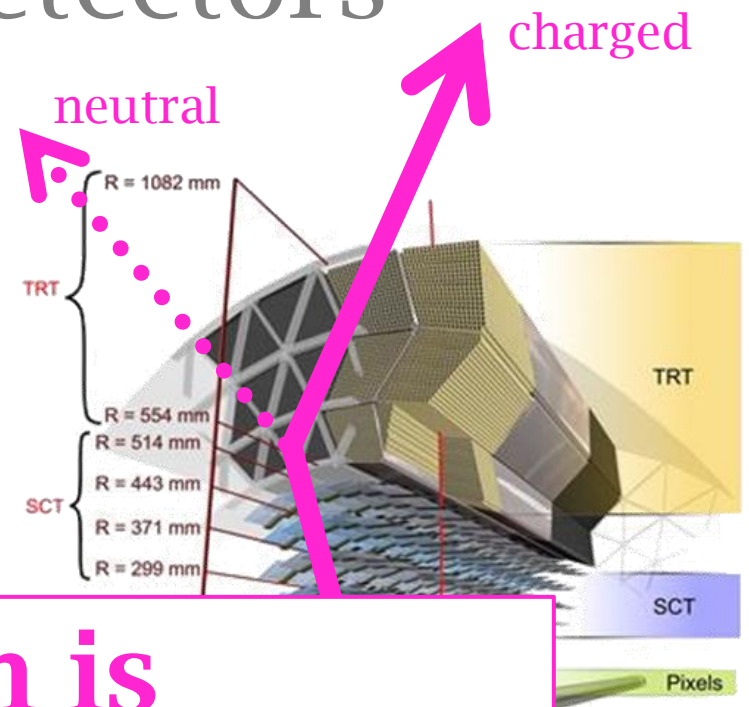
⇒ a track in detectors

When decay inside

⇒ track bends

“Stau Kink”

**Stau kink search is realistic & interesting.**



the LHC)

# 1. SUSY and its Signature

# Standard Model

 Successful!

 “Hierarchy problem”



supersymmetry

**MSSM** [Minimal Supersymmetric Standard Model]

 “Hierarchy” solved.

 GUTs & Dark Matter?

However...



# Standard Model

😊 Successful!

😡 “Hierarchy problem”



supersymmetry

**MSSM** [Minimal Supersymmetric Standard Model]

😊 “Hierarchy” solved

😊 GUTs & Dark Matter?

**Hypothetical!**

**Not discovered yet!**

However...

# Standard Model

😊 Successful!

😡 “Hierarchy problem”



supersymmetry

**MSSM** [Minimal Supersymmetric Standard Model]

😊 “Hierarchy” solved

😊 GUTs & Dark Matter?

**Hypothetical!**

**Not discovered yet!**

We want to discover SUSY.

How to discover SUSY?

What is **characteristic** in SUSY?

Important one: *R-parity*

If R-parity is conserved...

✓ Proton decay problem avoided! 😊

✓ LSP becomes **stable!**

➤ must be **neutral**.

➤ would be a **Dark Matter** candidate. 😊

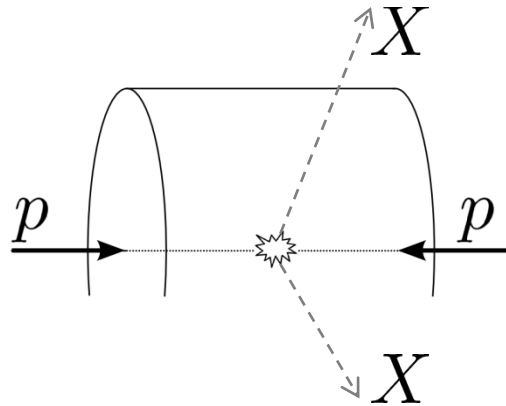
How to discover SUSY?

What is **characteristic** in SUSY?

[Case 1] Signature from **stable neutral particles**

⇒ Large missing energy  $\cancel{E}$

*main stream!*  
*Expected in many SUSY models!*



$X$  = escaping (missing) particle;  
e.g.  $\tilde{\chi}_1^0$  or  $\tilde{G}$ .

[Case 2] Signature from  
**long-lived charged particles**

expected in several models.

⇒ We will see in the next section.

Then, where can we discover SUSY?

... of course,

A satellite-style map of Europe and Asia. The text 'Bonn' is in pink at the top left. A pink dot is above the colon. The text 'LHC' is in large pink letters across the center. A pink dot is above the text '東京' (Tokyo) on the right side of the map.

Bonn

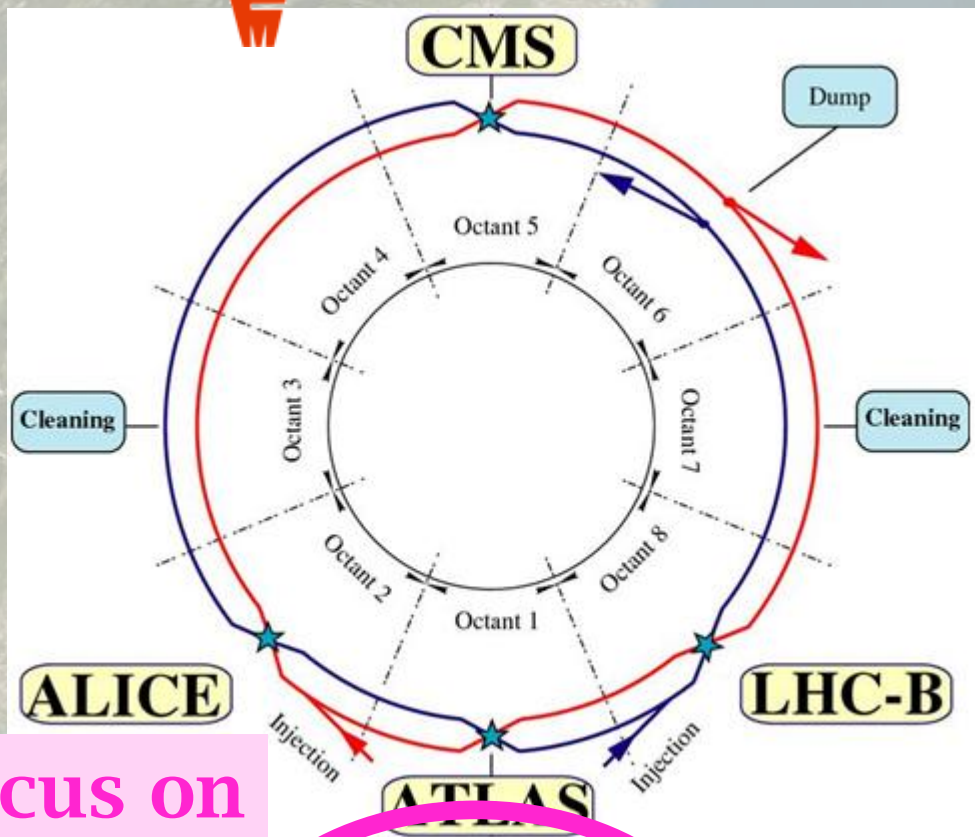
:

LHC

東京



LHC

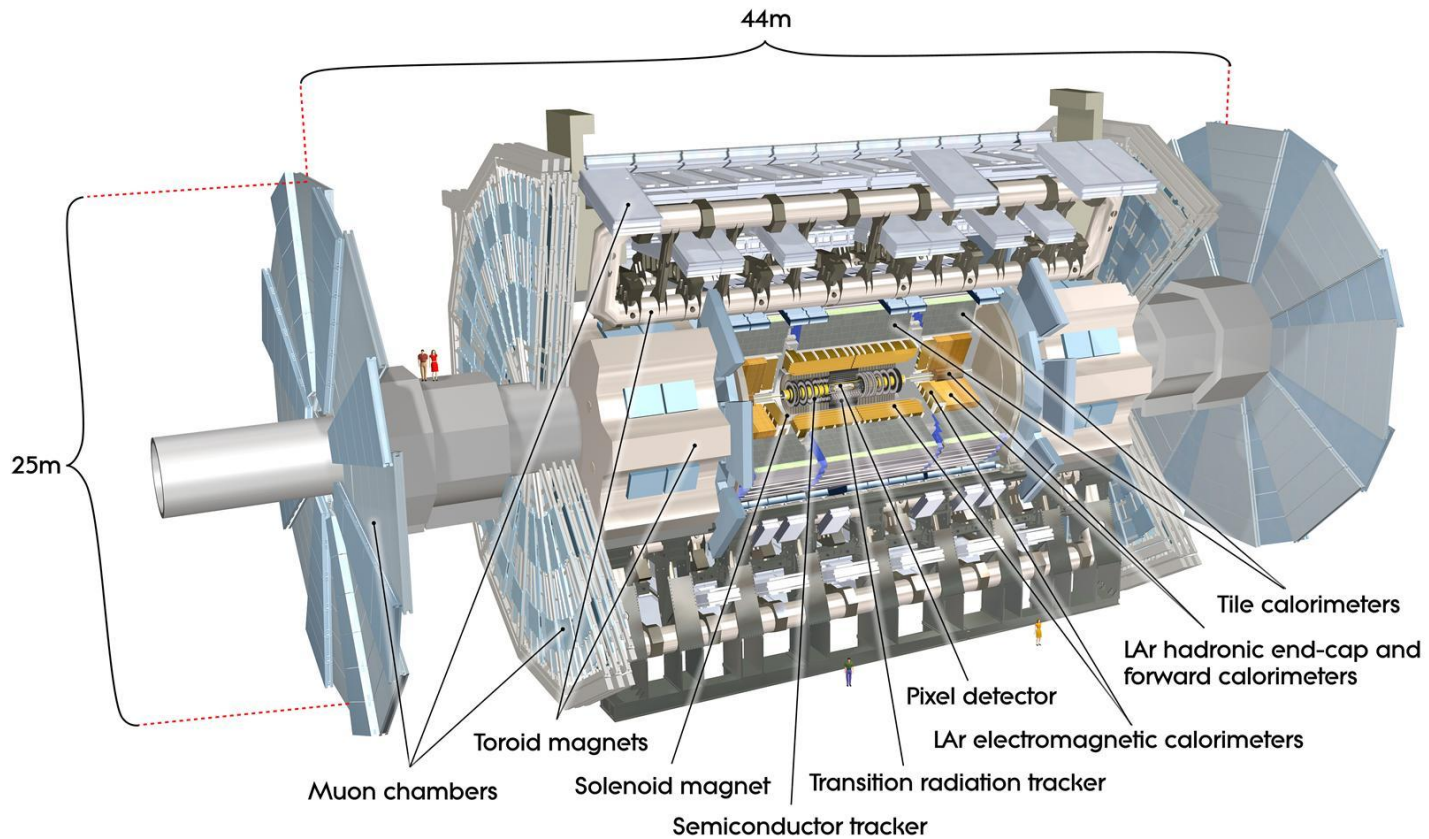


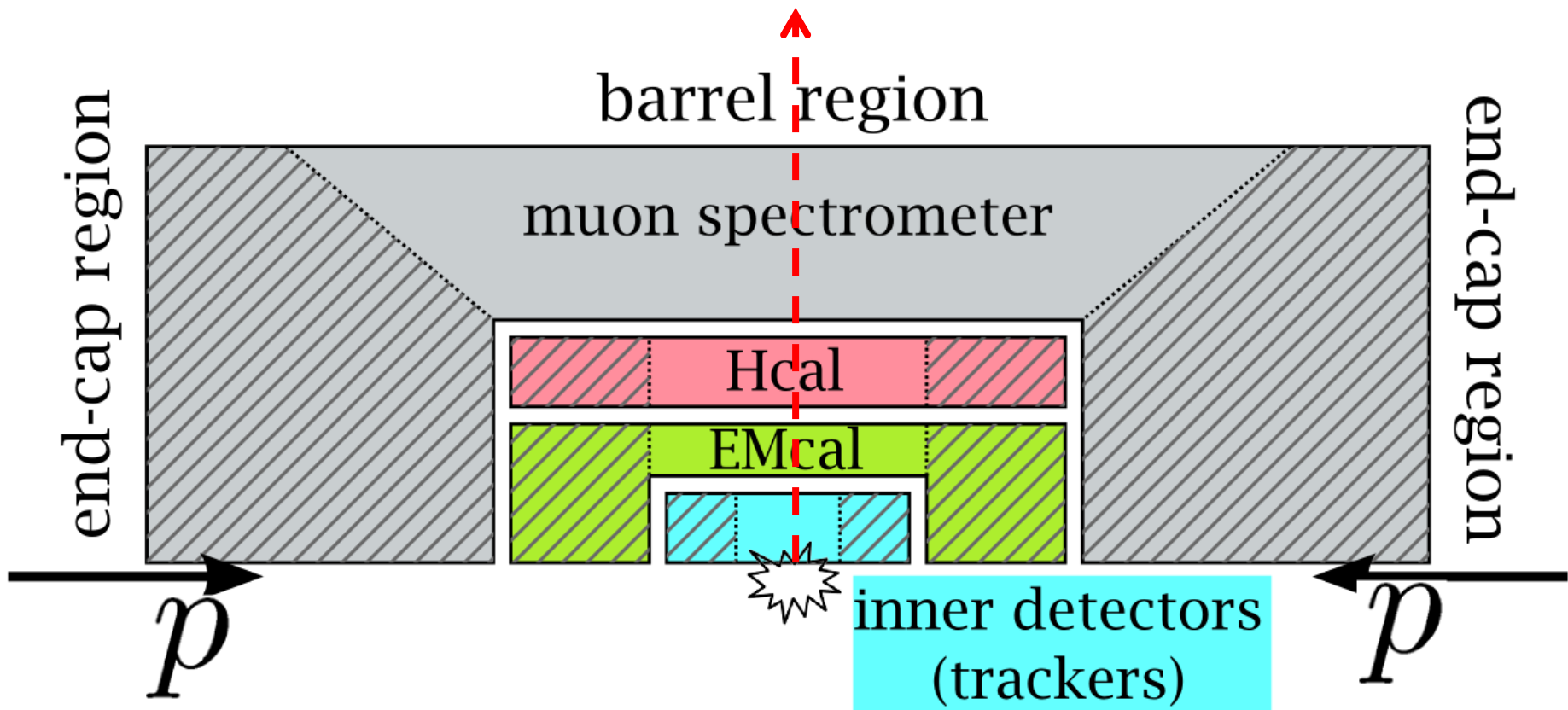
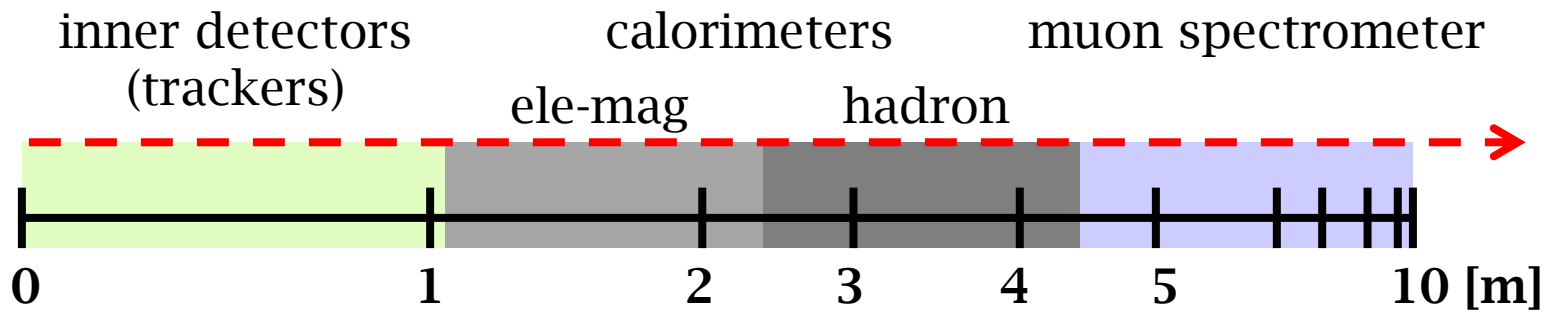
We focus on



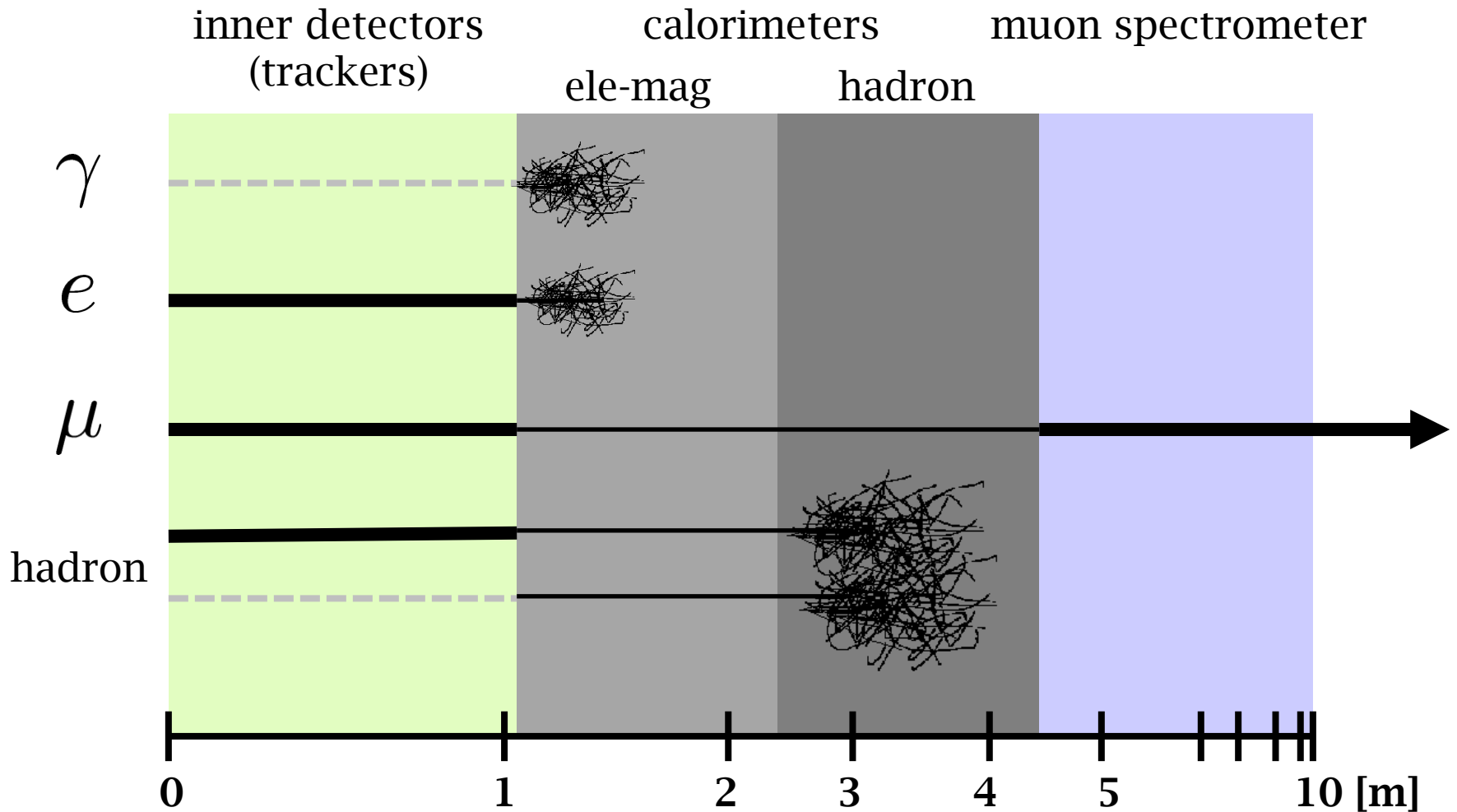


# ATLAS EXPERIMENT





[sectional (cut-away) view]



**Note**  $\tau \implies$  immediately decays into  
 leptons (35%)  $\longrightarrow e/\mu$  is observed.  
 hadrons (65%)  $\longrightarrow$  hadrons are observed.

## 2. Long-lived Stau Signature

- Scenario with Long-lived stau
- Its Signature
  1. Stable stau
  2. Stau kink

[Case 2] Signature from (We don't consider  $\tilde{e}, \tilde{\mu}$  -case for simplicity.)  
**long-lived charged particles**  $\left(\tilde{\tau}_1\right)$

expected in several models:

A)  $\tilde{G}$ -LSP,  $\tilde{\tau}_1$ -NLSP model (in GMSB framework) “weakness of gravity”

$$\tilde{\tau}_1 \rightarrow \tilde{G}\tau : c\tau \simeq 0.55 \text{ m} \left( \frac{200 \text{ GeV}}{m_{\tilde{\tau}_1}} \right)^5 \left( \frac{m_{\tilde{G}}}{1 \text{ keV}} \right)^2$$

B)  $\tilde{\tau}_1$ -LSP with tiny  $R$ -parity violation “tiny R-parity violation”

$$c\tau \sim O(1 \text{ m}) \text{ if RpV couplings } \sim 10^{-8}.$$

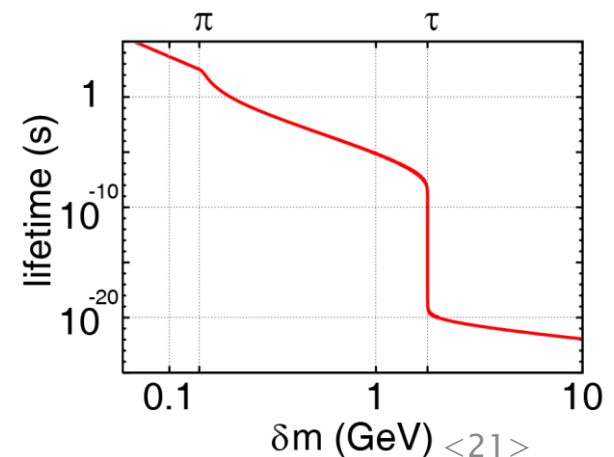
C) Coannihilation region “phase-space suppression”

$$(\tilde{\chi}_1^0\text{-LSP, } \tilde{\tau}_1\text{-NLSP, } m_{\tilde{\tau}_1} \simeq m_{\tilde{\chi}_1^0})$$

$$\tilde{\tau}_1 \rightarrow \tilde{\chi}_1^0 l \nu \bar{\nu}, \tilde{\chi}_1^0 \pi \nu$$

Right: lifetime @  $m_{\tilde{\tau}_1} = 300 \text{ GeV}$   
 $(\theta_\tau = 0.33)$

Jittoh *et al.*, PRD73.055009 [hep-ph/0512197]



[Case 2] Signature from  
**long-lived charged particles**  $(\tilde{\tau}_1)$

signature depends on where stau decays.

a) **outside** detectors

b) **inside** a detector

c) at the very **center**

[Case 2] Signature from

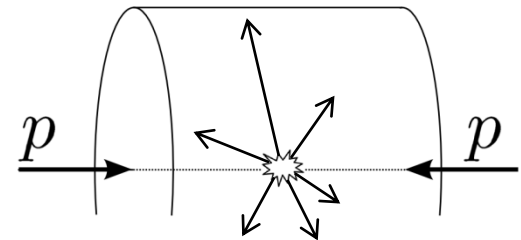
**long-lived charged particles**  $(\tilde{\tau}_1)$

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Signature depends on the decay mode.

(tau-rich or lepton-rich signature, etc...)

[Case 2] Signature from

**long-lived charged particles**  $(\tilde{\tau}_1)$

signature depends on where stau decays.

a) **outside** detectors

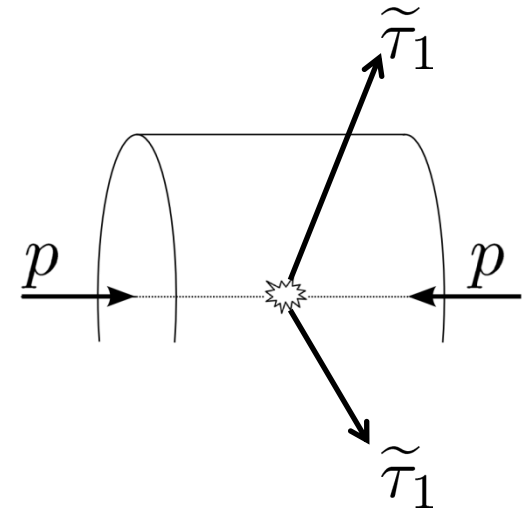
$\Rightarrow$  heavy “ $\mu$ -like” track

b) **inside** a detector

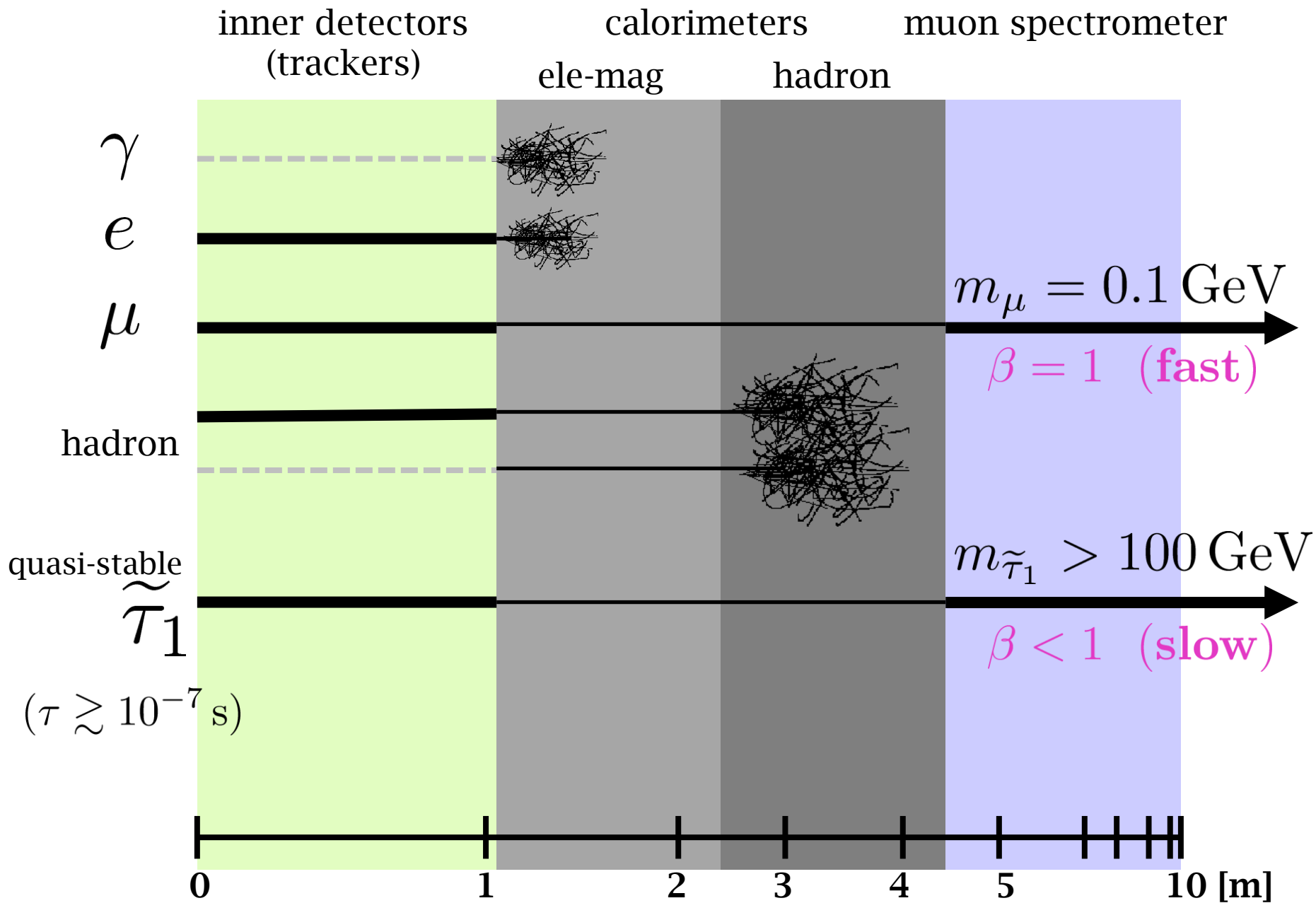
c) at the very **center**

Signature depends on the decay mode.

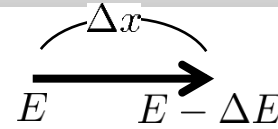
(tau-rich or lepton-rich signature, etc...)







# How to measure velocity $\beta$ ?



1. Energy Deposit  $\frac{dE}{dx}$

➤ Bethe-Bloch formula  $\left| \frac{dE}{dx} \right| = \frac{N_A Z}{A} \frac{4\pi Q^2 \alpha^2}{m_e} \left( \frac{1}{\beta^2} \log \frac{2m_e \beta^2}{I(1 - \beta^2)} - 1 \right)$

➤ A function of velocity  $\beta$  !

RED: material dependent  
BLUE: constants

➤ measured at an inner detector

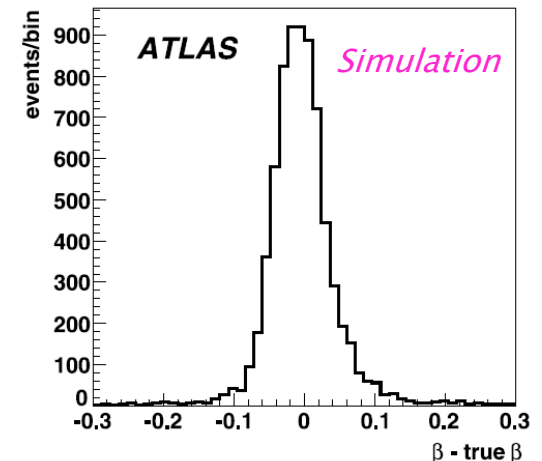
2. Time of flight (TOF)

➤ at Tracker and MS

➤ 1ns resolution

➔  $\beta \in [0.6, 0.9]$  is distinguishable from  $\beta = 1$ .

$$\Delta\beta \sim 0.05$$



ATLAS coll., EPJ C62 (2009) 281

# Current Bounds on *STABLE* $\tilde{\tau}_1$

## ◎ LEP

- DELPHI [PLB478.65; hep-ex/0103038]

$$\tilde{\tau}_{R/L} > 88/87.5 \text{ GeV} \quad (\text{direct production})$$

- OPAL [PLB572.8; hep-ex/0305031]

$$\tilde{\tau}_{R/L} > 98.0/98.5 \text{ GeV} \quad (\text{direct production})$$

## ◎ LHC

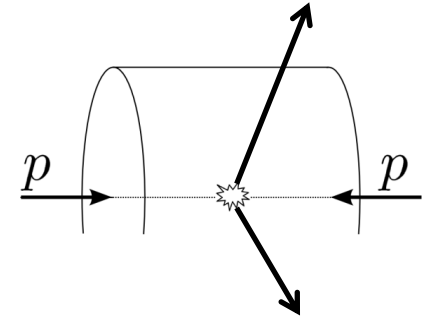
- ATLAS (37pb<sup>-1</sup>) [PLB703.428; 1106.4495]

$$\tilde{\tau}_1 > 136 \text{ GeV} \quad (\text{assuming a GMSB model})$$

$$\tilde{\tau}_1 > 110 \text{ GeV} \quad (\text{EW production = generic})$$

- CMS (1.1fb<sup>-1</sup>) [CMS-PAS-EXO-11-022]

$$\tilde{\tau}_1 > 293 \text{ GeV} \quad (\text{assuming a GMSB model})$$



[Case 2] Signature from

**long-lived charged particles**  $(\tilde{\tau}_1)$

signature depends on where stau decays.

a) **outside** detectors

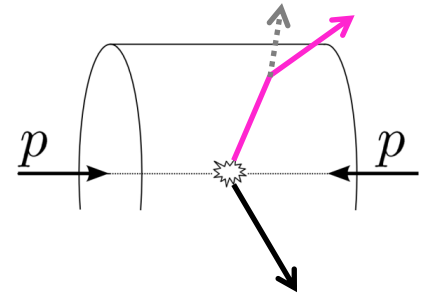
⇒ heavy “ $\mu$ -like” track

b) **inside** a detector

⇒ Kink track etc.

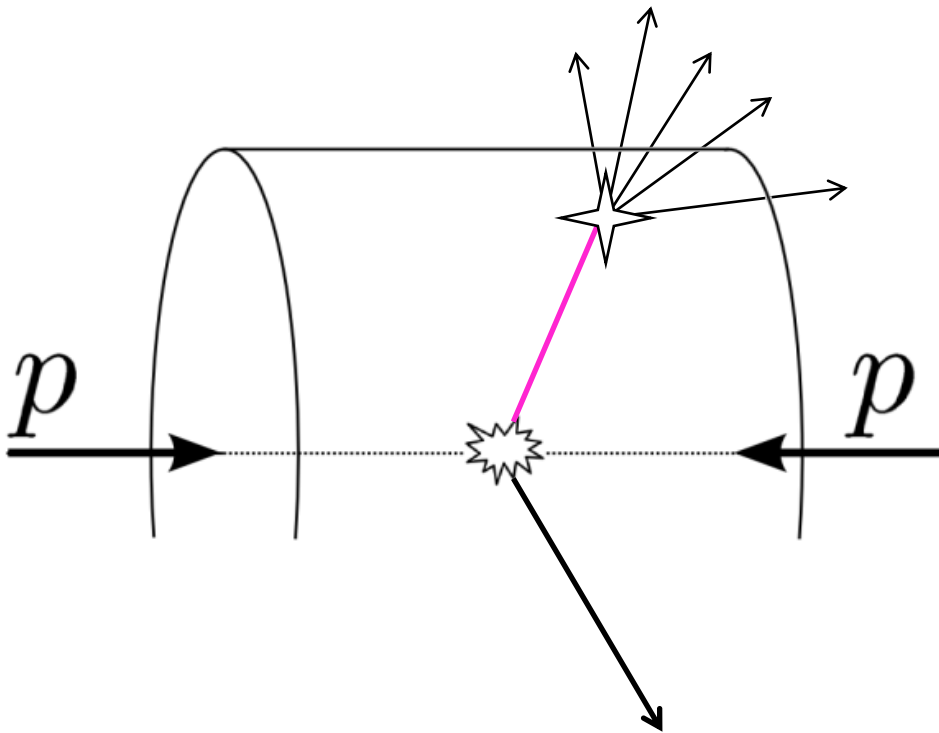
c) at the very **center**

⇒ depends on the decay mode.  
(tau-rich or lepton-rich signature, etc...)



Signature depends on...

Decay into WHAT?



**Complicated,**  
determined by underlying model...

# Long-lived stau scenarios

A)  $\tilde{G}$ -LSP,  $\tilde{\tau}_1$ -NLSP model  $\implies \tilde{\tau}_1 \rightarrow \tau \tilde{G}$

B)  $\tilde{\tau}_1$ -LSP with tiny RpV

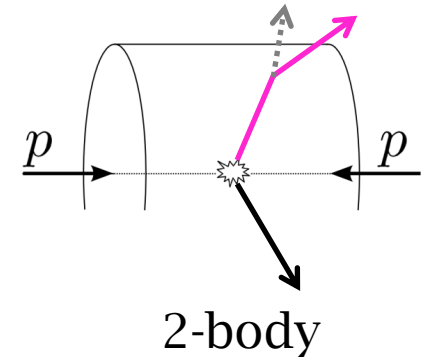
$$W = \frac{1}{2} \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \frac{1}{2} \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k$$

$\lambda_{i3k}, \lambda_{ij3} \implies \tilde{\tau}_1 \rightarrow e\nu, \mu\nu, \tau\nu$

$\lambda_{121}, \lambda_{122} \implies$  4-body decay

$\lambda' \implies$  hadron or 4-body

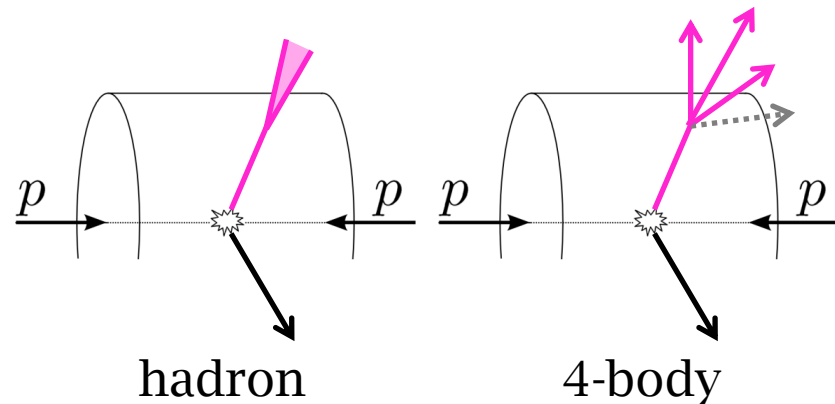
$\lambda'' \implies$  4-body decay



C) Coannihilation region

$(\tilde{\chi}_1^0$ -LSP,  $\tilde{\tau}_1$ -NLSP,  $m_{\tilde{\tau}_1} \simeq m_{\tilde{\chi}_1^0}$ )

$\tilde{\tau}_1 \rightarrow \tilde{\chi}_1^0 \nu \pi, \tilde{\chi}_1^0 l \nu \bar{\nu}$



# Long-lived stau scenarios

$\tilde{\tau}_1$ - $\tau$ -kink

A)  $\tilde{G}$ -LSP,  $\tilde{\tau}_1$ -NLSP model  $\Rightarrow \tilde{\tau}_1 \rightarrow \tau \tilde{G}$

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$$W = \frac{1}{2} \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \frac{1}{2} \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k$$

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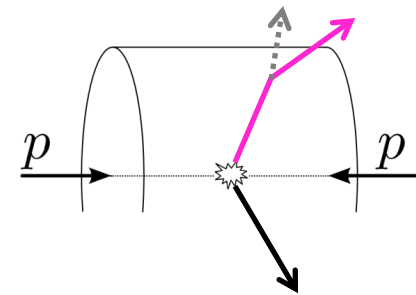
$\tilde{\tau}_1$ -( $e, \mu, \tau$ )-kink

C) Coannihilation region

( $\tilde{\chi}_1^0$ -LSP,  $\tilde{\tau}_1$ -NLSP,  $m_{\tilde{\tau}_1} \simeq m_{\tilde{\chi}_1^0}$ )

$\tilde{\tau}_1 \rightarrow \tilde{\chi}_1^0 \nu \pi, \tilde{\chi}_1^0 l \nu \bar{\nu}$

kink, but “soft”...



2-body (kink track)

Kink track

=

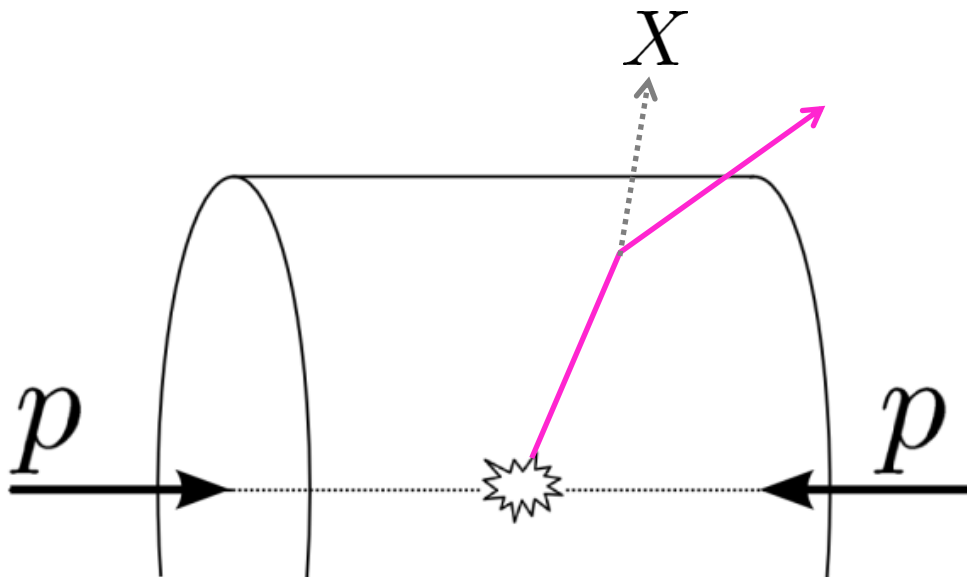
stau track

+

daughter track

by a tracker

by a tracker

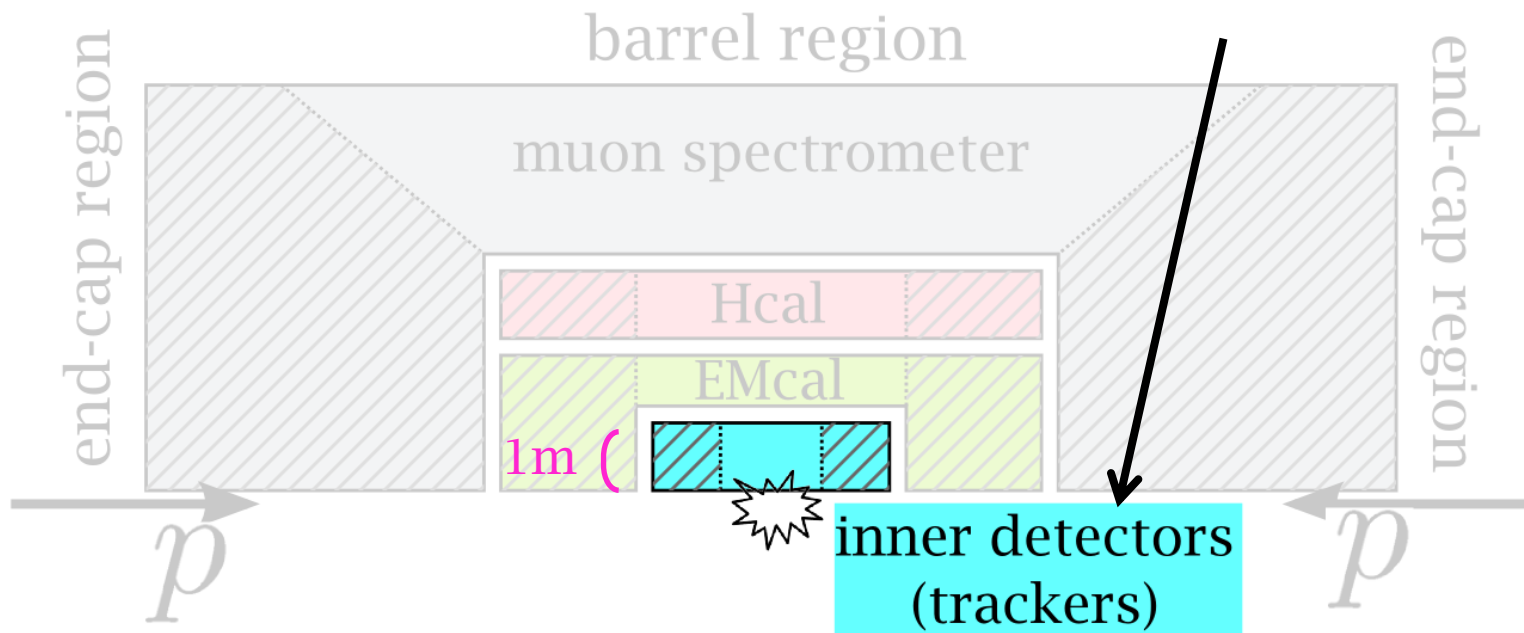




Kink track id. = id. of stau track + id. of daughter track  
by a tracker by a tracker

↓

We have to do “two” id. in

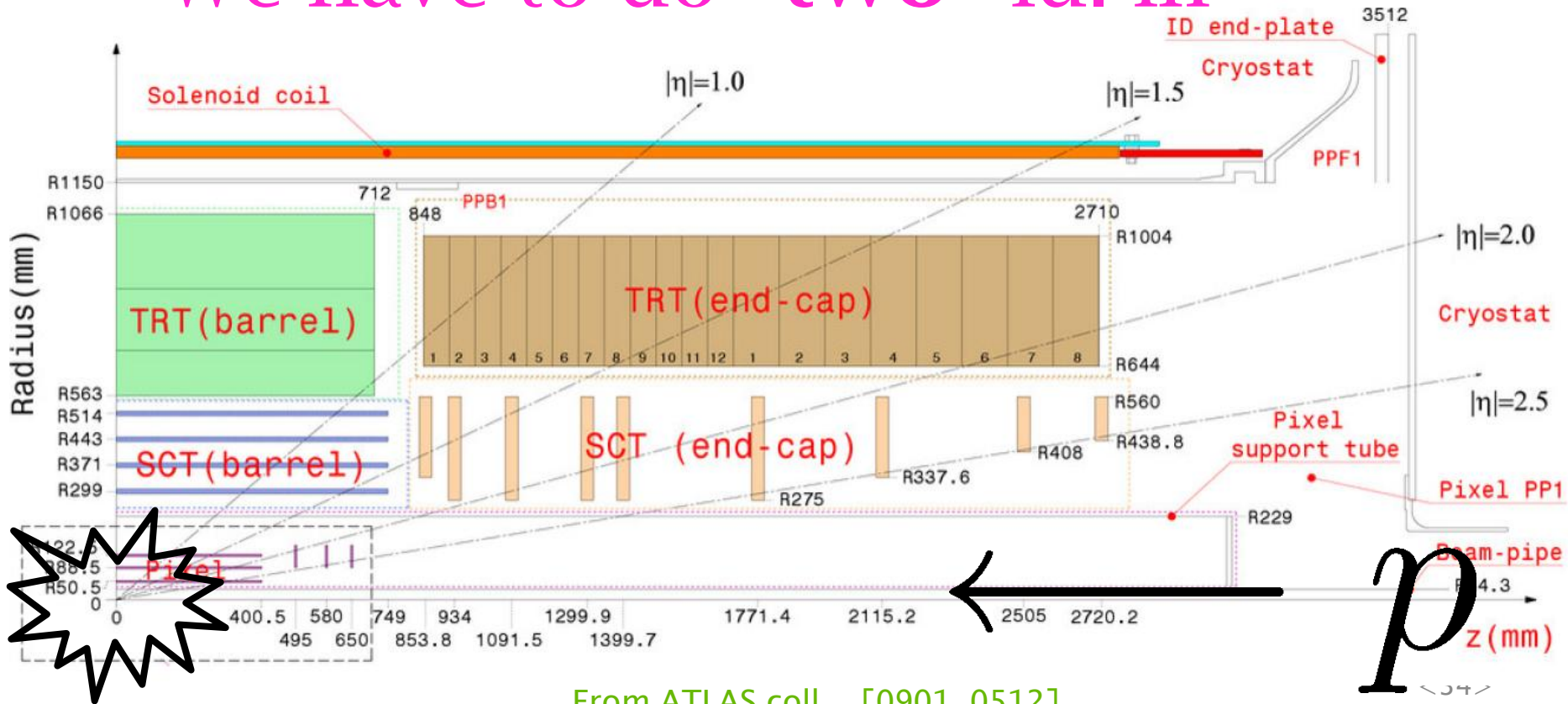


Kink track id. = id. of stau track + id. of daughter track

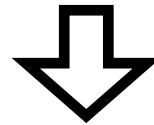
by a tracker

by a tracker

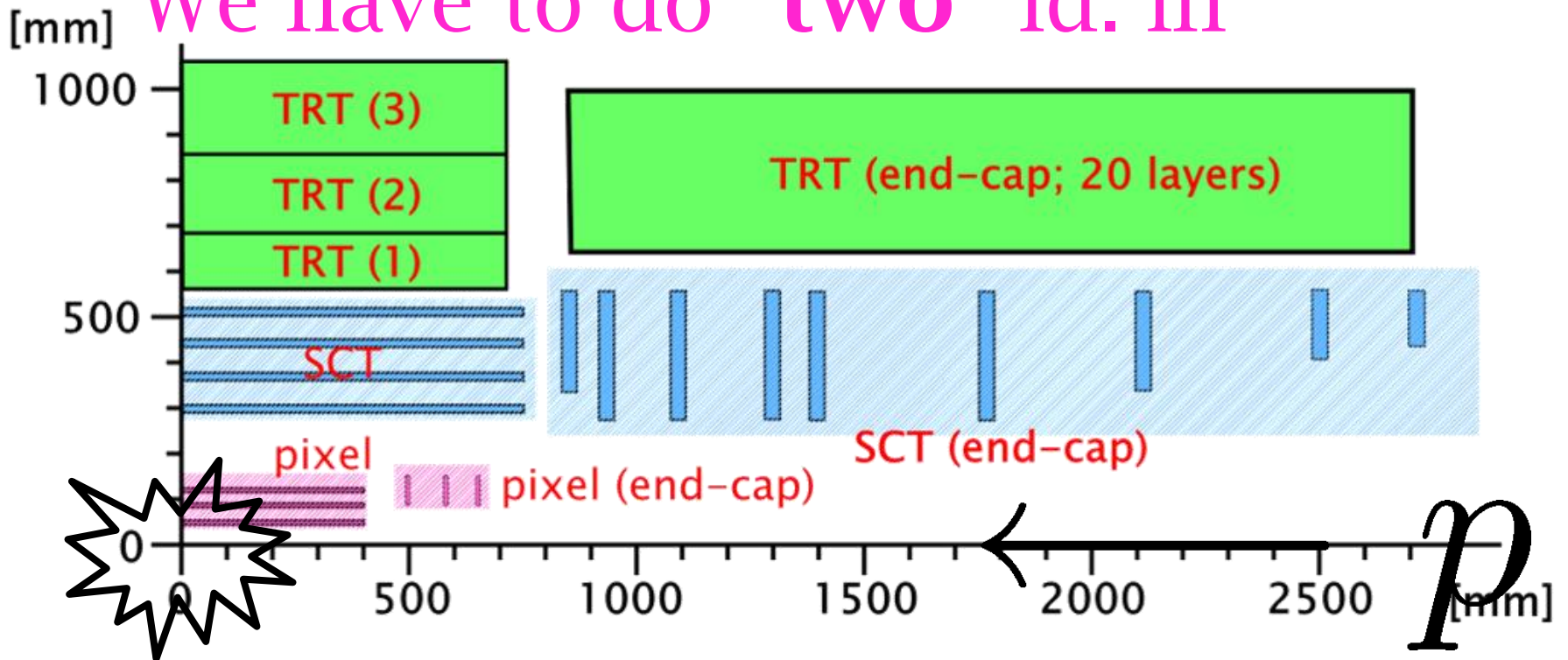
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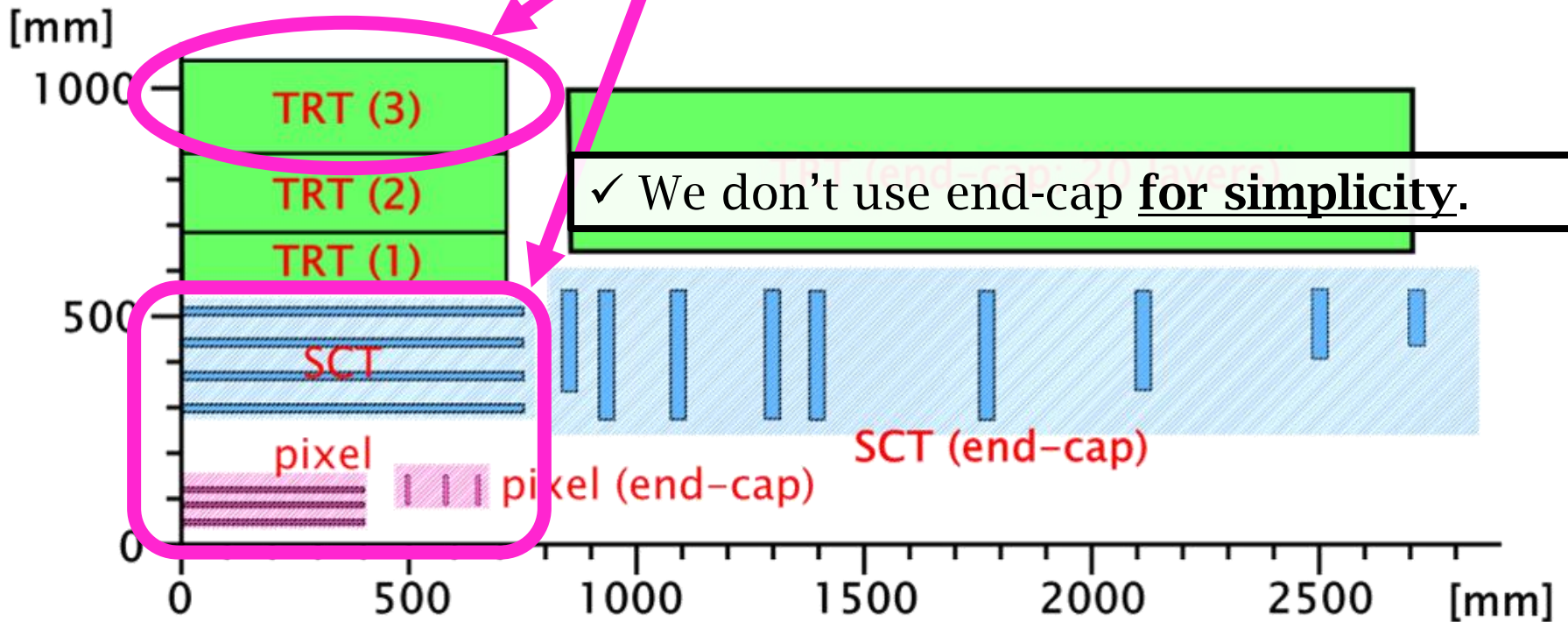
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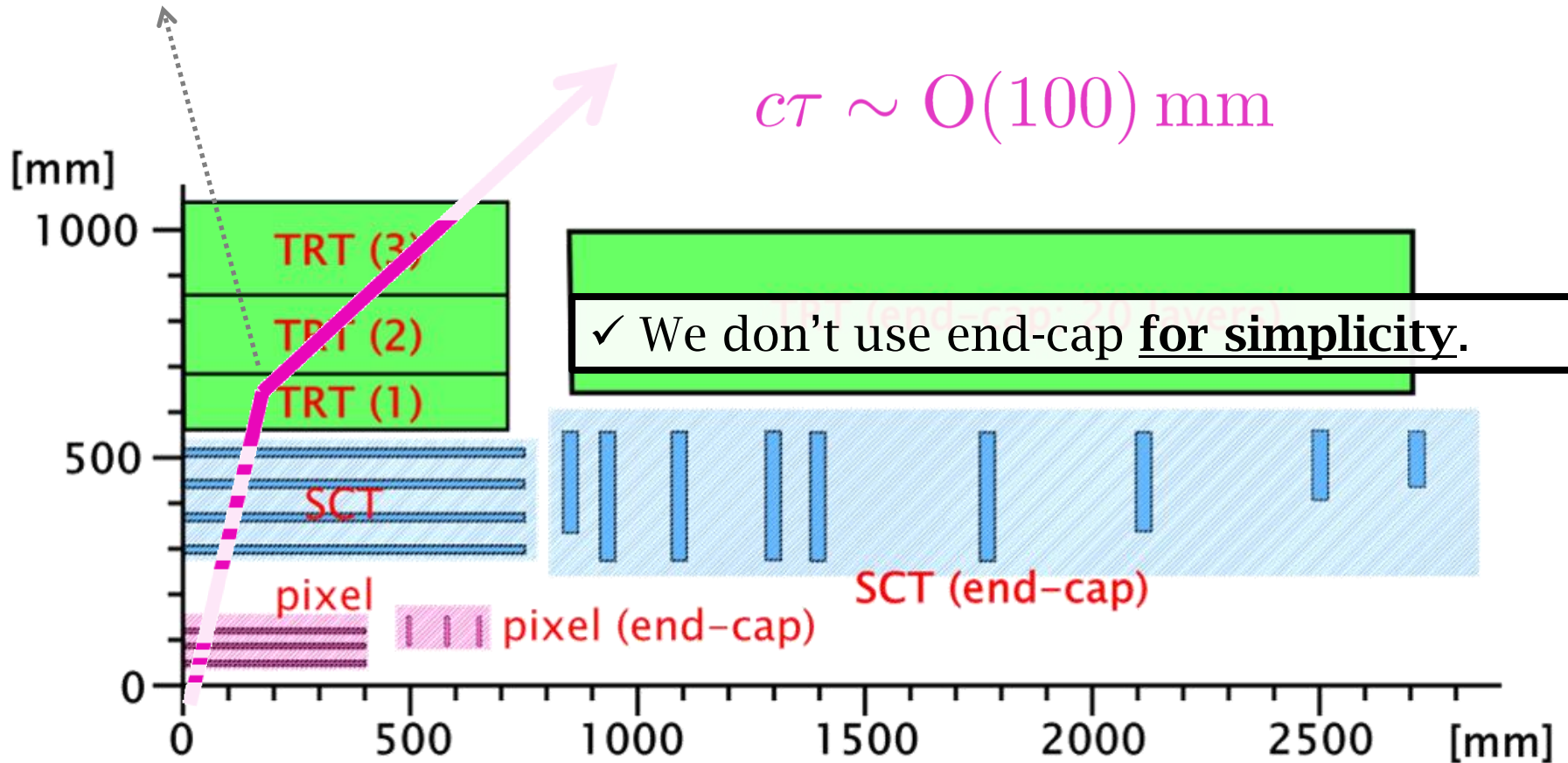
We have to do “two” id. in



Kink track id. = id. of stau track + id. of daughter track  
by a tracker by a tracker



Kink at **TRT 1<sup>st</sup> or 2<sup>nd</sup> module**  
can be observed.



# With this method, we can observe kinks.

➤ Sweet range  $C\mathcal{T}$  (of stau)  $\sim O(0.1 - 10)$  m

➤ 300GeV stau can be observed.

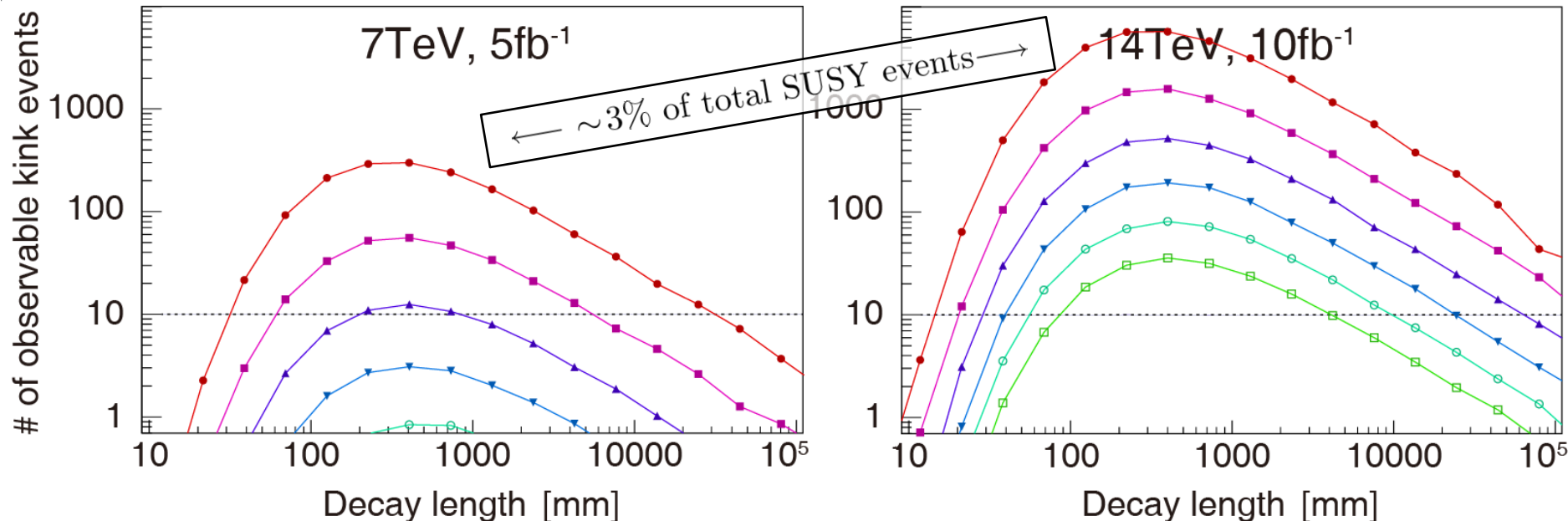
✓ Some CMSSM-model is assumed.

✓ Efficiencies are considered.

✓ Background events are fairly suppressed.

Let's see in detail.

●	$\tilde{\tau}, \tilde{g} = 103, 715$ GeV
■	140, 932
▲	176, 1145
▼	212, 1355
○	248, 1562
□	283, 1768



# 3. Stau Kinks in detail

# Stau kinks *in detail*

## ⊙ Technical topics (**experiment**)

- ◆ trigger and efficiency
- ◆ track reconstructions and efficiency
- ◆ background events
- ◆ Monte Carlo simulation

etc...

## ⊙ Physical topics (**phenomenology**)

- We can **discriminate** the models!!



# Long-lived stau scenarios

$\tilde{\tau}_1$ - $\tau$ -kink

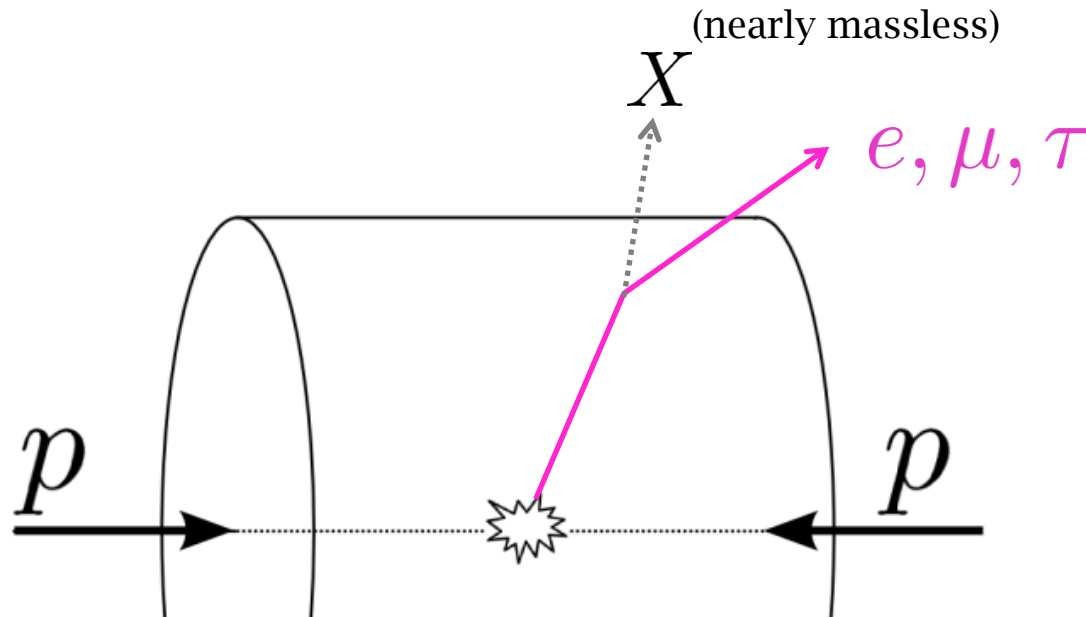
A)  $\tilde{G}$ -LSP,  $\tilde{\tau}_1$ -NLSP model  $\Rightarrow \tilde{\tau}_1 \rightarrow \tau \tilde{G}$

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$$W = \frac{1}{2} \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \frac{1}{2} \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k$$

$\lambda_{i3k}, \lambda_{ij3} \Rightarrow \tilde{\tau}_1 \rightarrow e\nu, \mu\nu, \tau\nu$

$\tilde{\tau}_1$ -( $e, \mu, \tau$ )-kink



$$W = \frac{1}{2}\lambda_{ijk}L_iL_j\bar{E}_k + \lambda'_{ijk}L_iQ_j\bar{D}_k + \frac{1}{2}\lambda''_{ijk}\bar{U}_i\bar{D}_j\bar{D}_k$$

**Kink type**  $\longrightarrow$  **daughter signature**

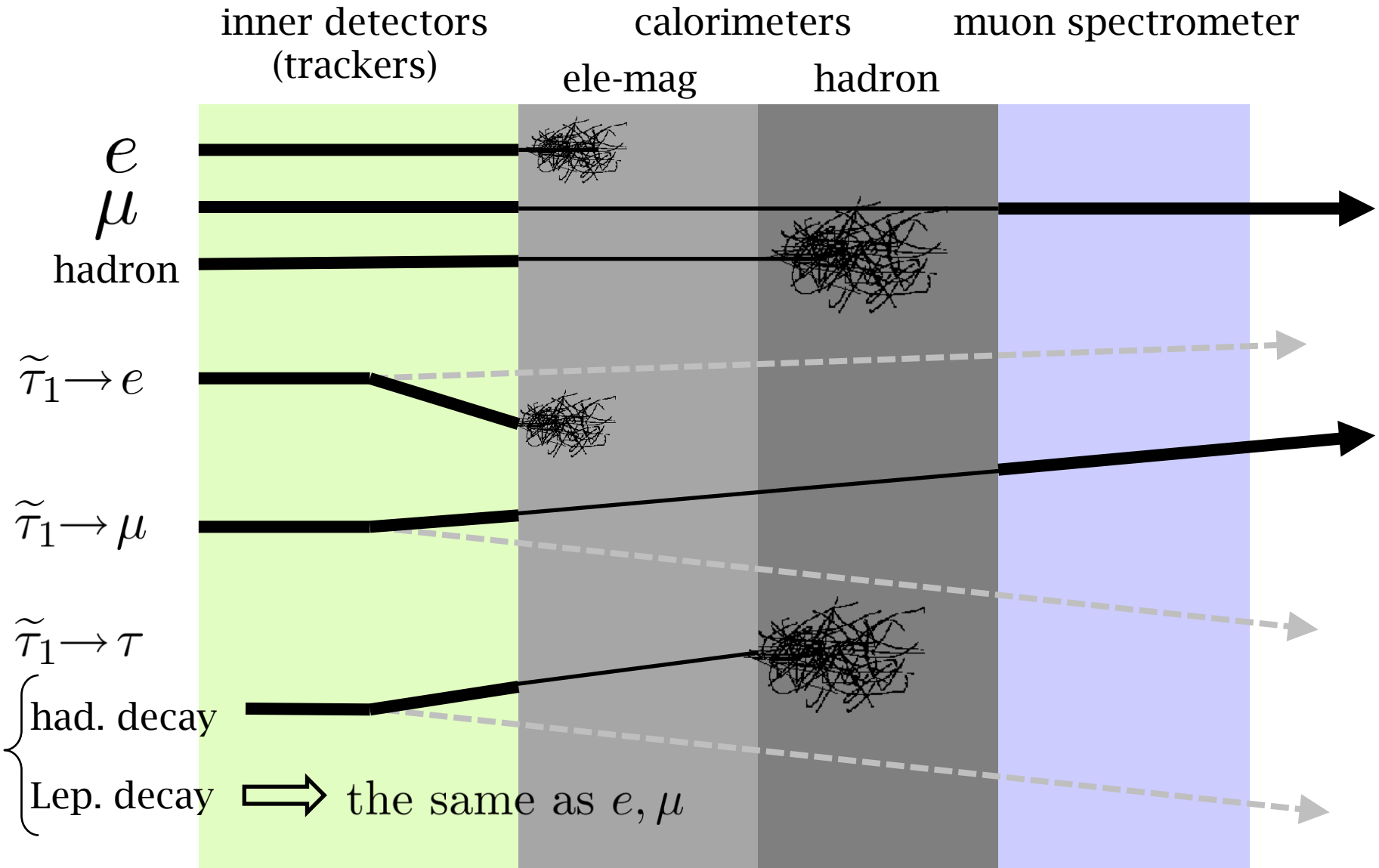
Models	$e$ : $\mu$ : $\tau$	$e$	$\mu$	$\tau$ -jet
Gravitino LSP	0 : 0 : 1	18%	17%	65%
$\lambda_{123}$	1 : 1 : 0	50%	50%	—
$\lambda_{i31}$	1 : 0 : 0	100%	—	—
$\lambda_{i32}$	0 : 1 : 0	—	100%	—
$\lambda_{133}$	$\sin^2\theta$ : 0 : 1	* 59%	* 9%	* 32%
$\lambda_{233}$	0 : $\sin^2\theta$ : 1	* 9%	* 59%	* 32%

\*depending on stau mixing angle  $\theta$ ; values are for  $\theta = 1$ .

## Daughter lepton discrimination

$\Rightarrow$  Ratio of the daughter leptons  
= Underlying models

$$\left[ \begin{array}{l} \underline{\lambda_{123}} \tilde{\tau} \rightarrow L_1L_2 \quad \rightsquigarrow e : \mu = 1 : 1 \\ \underline{\lambda_{i3k}} \tilde{\tau} \rightarrow L_i\bar{E}_k \quad \rightsquigarrow l_k + \nu_i \\ \underline{\lambda_{i33}} \tilde{\tau} \rightarrow L_iL_3, L_i\bar{E}_3 \quad \rightsquigarrow l_i : \tau = \sin^2\theta : 1 \end{array} \right]$$



**Daughter lepton can be distinguished.**

[m]

# (phenomenological) Conclusion

## ⊙ Stau (slepton) in-flight-decay

⇒ observable as kink events.

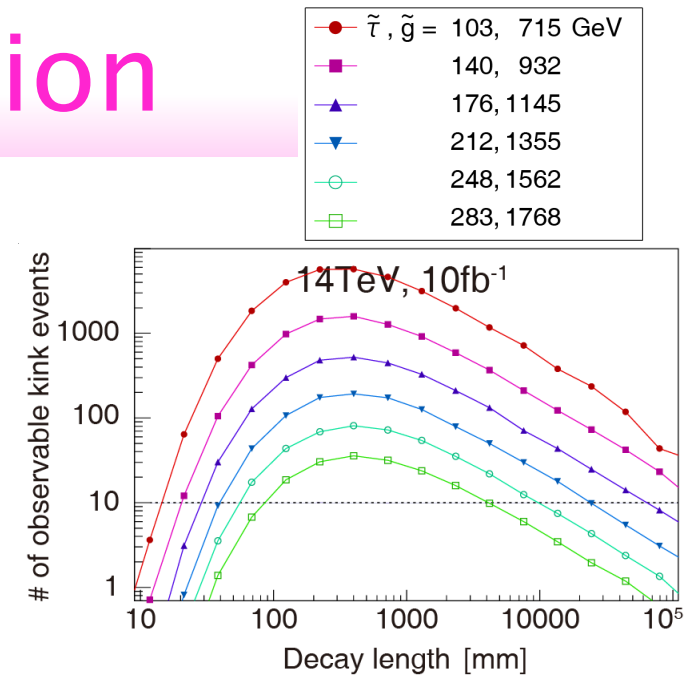
- decay length:  $c\tau \sim O(0.1 - 100)\text{m}$
- Stau mass:  $m \lesssim 300\text{ GeV}$

(Much more luminosity allows us to go further.)

## ⊙ This decay length corresponds to

- gravitino model:  $m_{\tilde{G}} \sim 0.1 - 10\text{ keV}$
- R-parity violation:  $\lambda \sim O(10^{-8} - 10^{-9})$

## ⊙ Model discrimination is possible.



# Stau kinks *in detail*

- ⊙ Technical topics (**experiment**)

- ◆ trigger and efficiency
- ◆ track reconstructions and efficiency
- ◆ background events
- ◆ Monte Carlo simulation

etc...

- ⊙ Physical topics (**phenomenology**)

- We can **discriminate** the models!!

# Monte Carlo Simulation

## Method

mass spectrum: SUSY-HIT  
event generation: Pythia6  
fast detector sim.: PGS4

- ⊙ Benchmark Point: CMSSM model

$M_0$	$M_{1/2}$	$\tan \beta$	$A_0$	$\text{sgn } \mu$
0 GeV	varied	13	0 GeV	+

$M_{1/2}$	$\tilde{\tau}$	$\tilde{g}$
300	103,	715
400	140,	932
500	176,	1145
600	212,	1355
700	248,	1562
800	283,	1768

[GeV]

- ⊙ PGS4-based fast detector simulation

# Monte Carlo Simulation

## Event selection

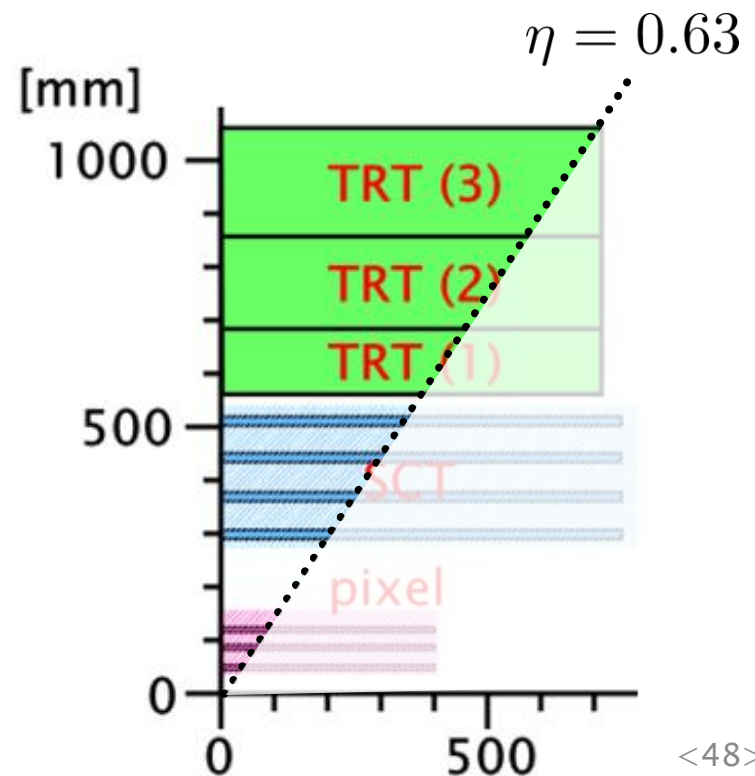
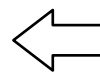
- triggering issue
  - 1 jet with  $P_T > 120$  GeV.
  - $\cancel{E}_T > 100$  GeV.
- $\tilde{\tau}_1$  must be
  - $|\eta| < 0.63$ .
  - $P_T > 100$  GeV.
  - decay in TRT 1st or 2nd module.
- The kink must be
  - azimuthal opening angle  $0.1 < \Delta\phi < \pi/2$ .
- daughter particle must be
  - not into end-cap; stay in barrel region.
  - $P_T > 10$  GeV (efficiency 0.6) or  $> 20$  GeV (0.7).

# Monte Carlo Simulation

## Event selection

- triggering issue
  - 1 jet with  $P_T > 120$  GeV.
  - $\cancel{E}_T > 100$  GeV.
- $\tilde{\tau}_1$  must be
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- The kink must be
  - azimuthal opening angle  $0.1 < \Delta\phi$
- daughter particle must be
  - not into end-cap; stay in barrel reg
  - $P_T > 10$  GeV (efficiency 0.6) or  $> 100$  GeV

← Trigger: 1jet(70) + MET(40) is “stable” (90% eff.) above this point.

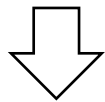




# Monte Carlo Simulation

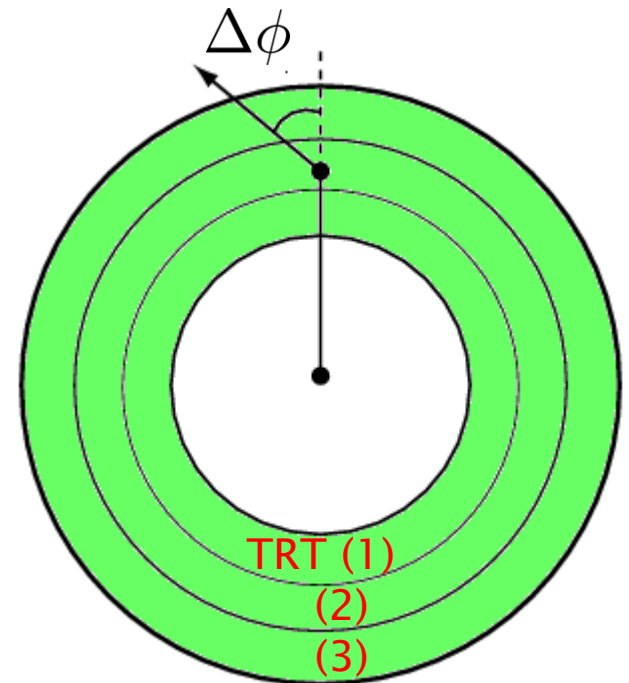
TRT =  $r-\phi$  information

(know nothing about on z-direction.)



“azimuthal opening angle” can be measured.

- $|\eta| < 0.63$ .
- $P_T > 100$  GeV.
- decay in TRT 1st or 2nd module.
- The kink must be
  - azimuthal opening angle  $0.1 < \Delta\phi < \pi/2$ .
- daughter particle must be
  - not into end-cap; stay in barrel region.
  - $P_T > 10$  GeV (efficiency 0.6) or  $> 20$  GeV (0.7).

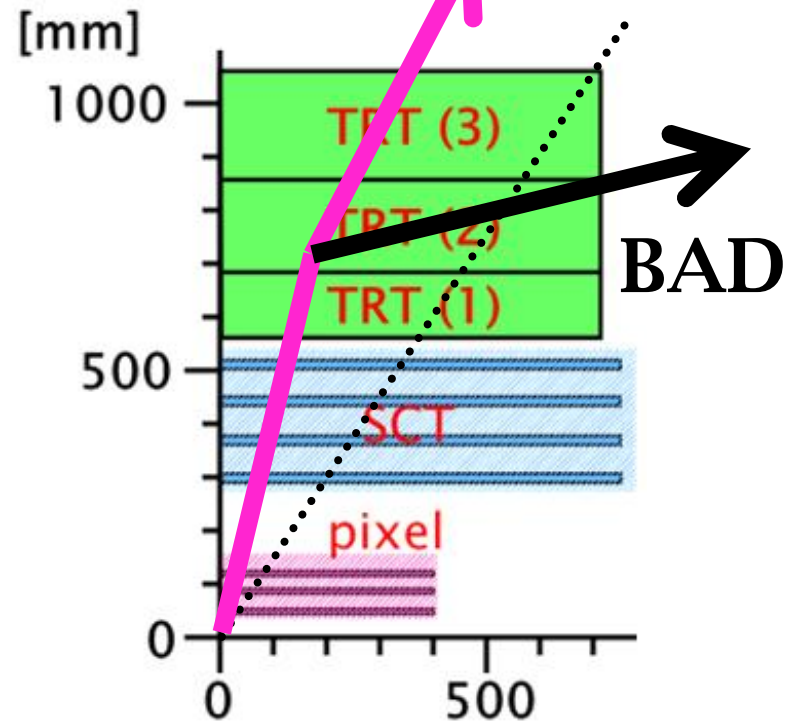


$r-\phi$  projected plane view

# Monte Carlo Simulation **GOOD**

## Event selection

- triggering issue
  - 1 jet with  $P_T > 120$  GeV.
  - $\cancel{E}_T > 100$  GeV.
- $\tilde{\tau}_1$  must be
  - $|\eta| < 0.63$ .
  - $P_T > 100$  GeV.
  - decay in TRT 1st or 2nd module.
- The kink must be
  - azimuthal opening angle  $0.1 < \Delta\phi < \pi/2$ .
- daughter particle must be
  - not into end-cap; stay in barrel region.
  - $P_T > 10$  GeV (efficiency 0.6) or  $> 20$  GeV (0.7).



Daughter must go through TRT (3).

**in order to the daughter reconstruction.**

# Monte Carlo Simulation

## Event selection

100%

- triggering issue
    - 1 jet with  $P_T > 120$  GeV.
    - $\cancel{E}_T > 100$  GeV.

---
- $\tilde{\tau}_1$  must be
  - $|\eta| < 0.63$ . ..... ~ 85%
  - $P_T > 100$  GeV. ..... ~ 35%
  - decay in TRT 1st or 2nd module. ..... ~ 33%

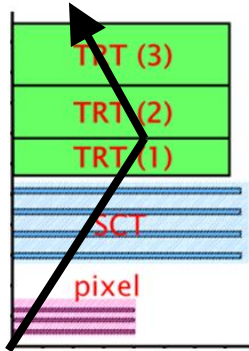
---
- The kink must be
  - azimuthal opening angle  $0.1 < \Delta\phi < \pi/2$ . ..... ~ 4%

---
- daughter particle must be
  - not into end-cap; stay in barrel region.
  - $P_T > 10$  GeV (efficiency 0.6) or  $> 20$  GeV (0.7). ..... ~ 3%

---
- ..... ~ 2%

# Possible Background Events

- Stable charged hadrons: Hit to detector material



⇒ Few hadrons have  $P_T > 100$  GeV.  
Few hadrons interact with material.

suppressed.

- In-flight-decay of hadrons

- hadron decay ... small  $\Delta m$
- small  $\Delta m +$  large  $P_T \implies$  small kink angle

suppressed.

- “false” tracks from noise

⇒ We require two tracks (mother & daughter)

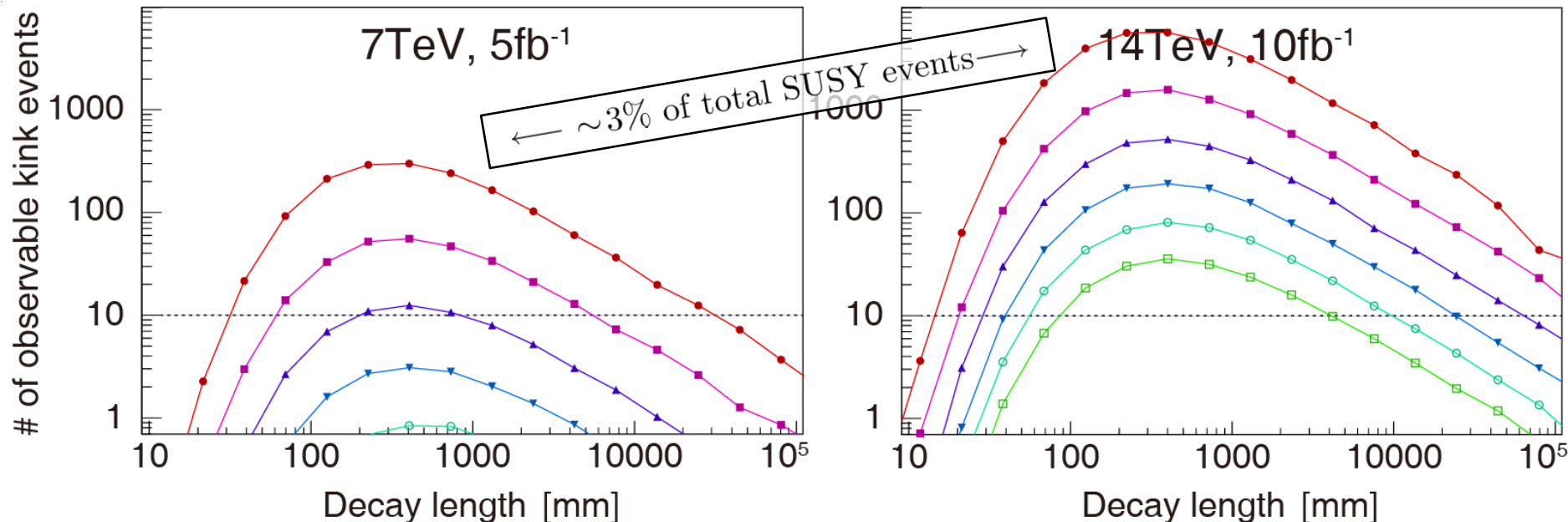
ignorable.

**background events are ignorable!**

# Numerical Results (again)

- Sweet range  $CT$  (of stau)  $\sim O(0.1 - 10)$  m
- **300GeV** stau can be observed.
- ✓ Some CMSSM-model is assumed.
- ✓ Efficiencies are considered.
- ✓ Background events are fairly suppressed.

●	$\tilde{\tau}, \tilde{g} = 103, 715$ GeV
■	140, 932
▲	176, 1145
▼	212, 1355
○	248, 1562
□	283, 1768



# Conclusion (again)

- Stau (slepton) in-flight-decay

⇒ observable as kink events.

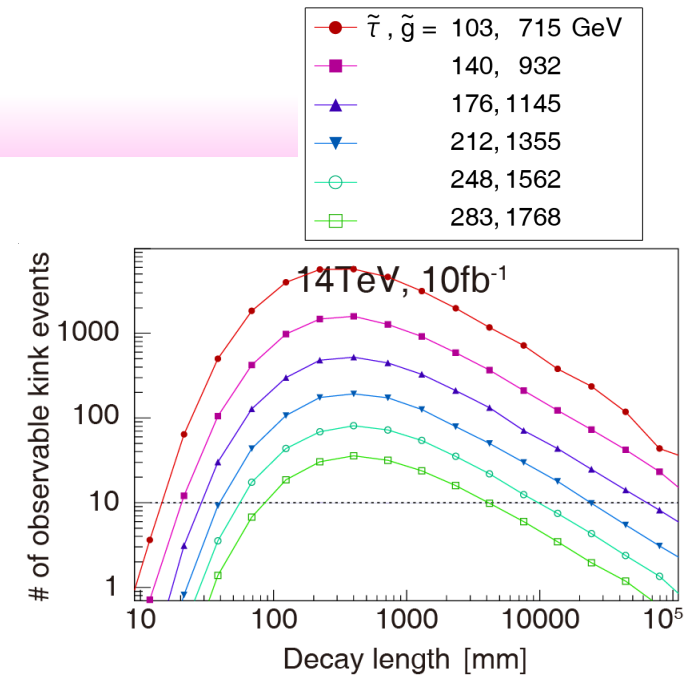
- decay length:  $c\tau \sim O(0.1 - 100)\text{m}$
- Stau mass:  $m \lesssim 300\text{ GeV}$

(Much more luminosity allows us to go further.)

- This decay length corresponds to

- gravitino model:  $m_{\tilde{G}} \sim 0.1 - 10\text{ keV}$
- R-parity violation:  $\lambda \sim O(10^{-8} - 10^{-9})$

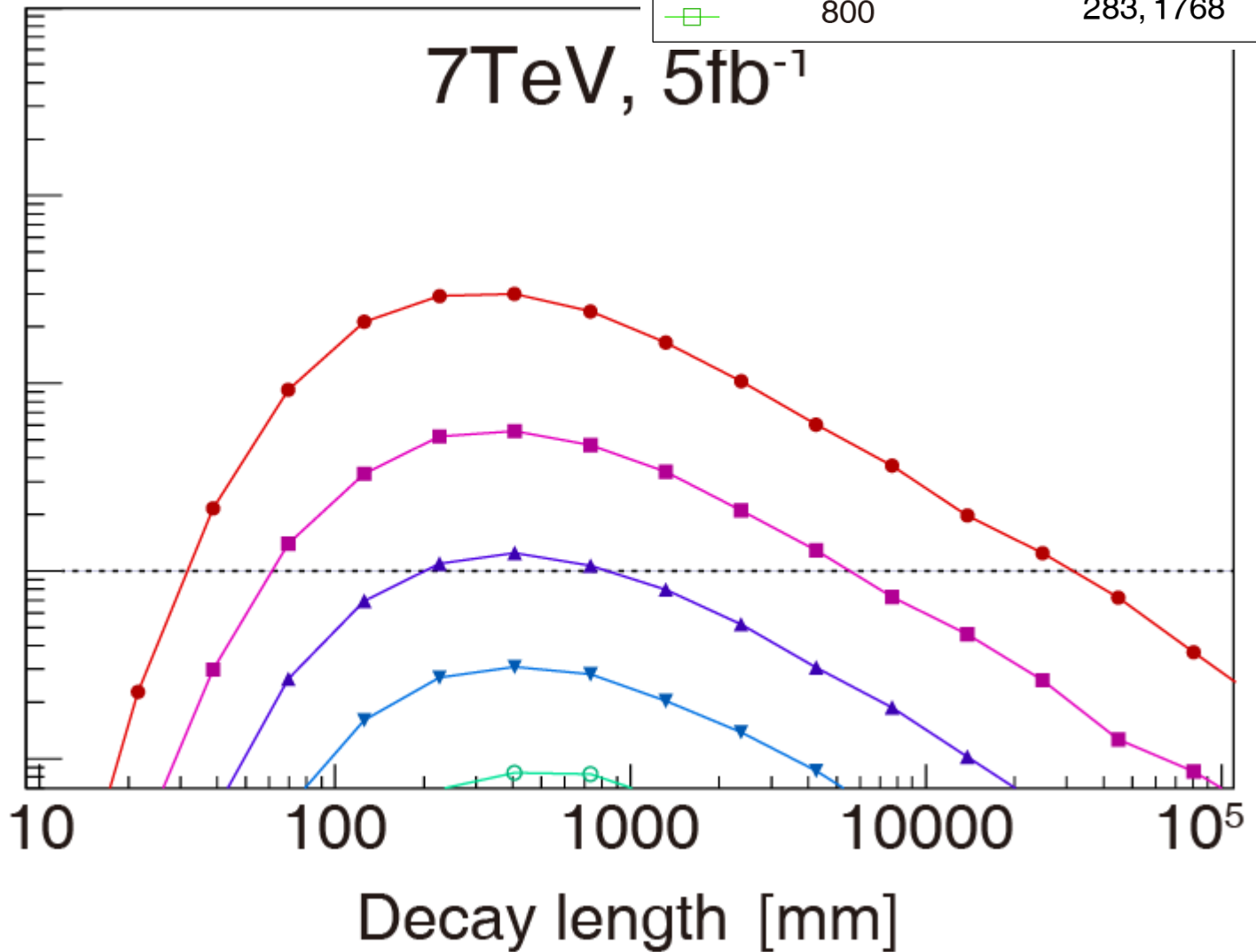
- Model discrimination is possible.



# of observable kink events

7TeV, 5fb<sup>-1</sup>

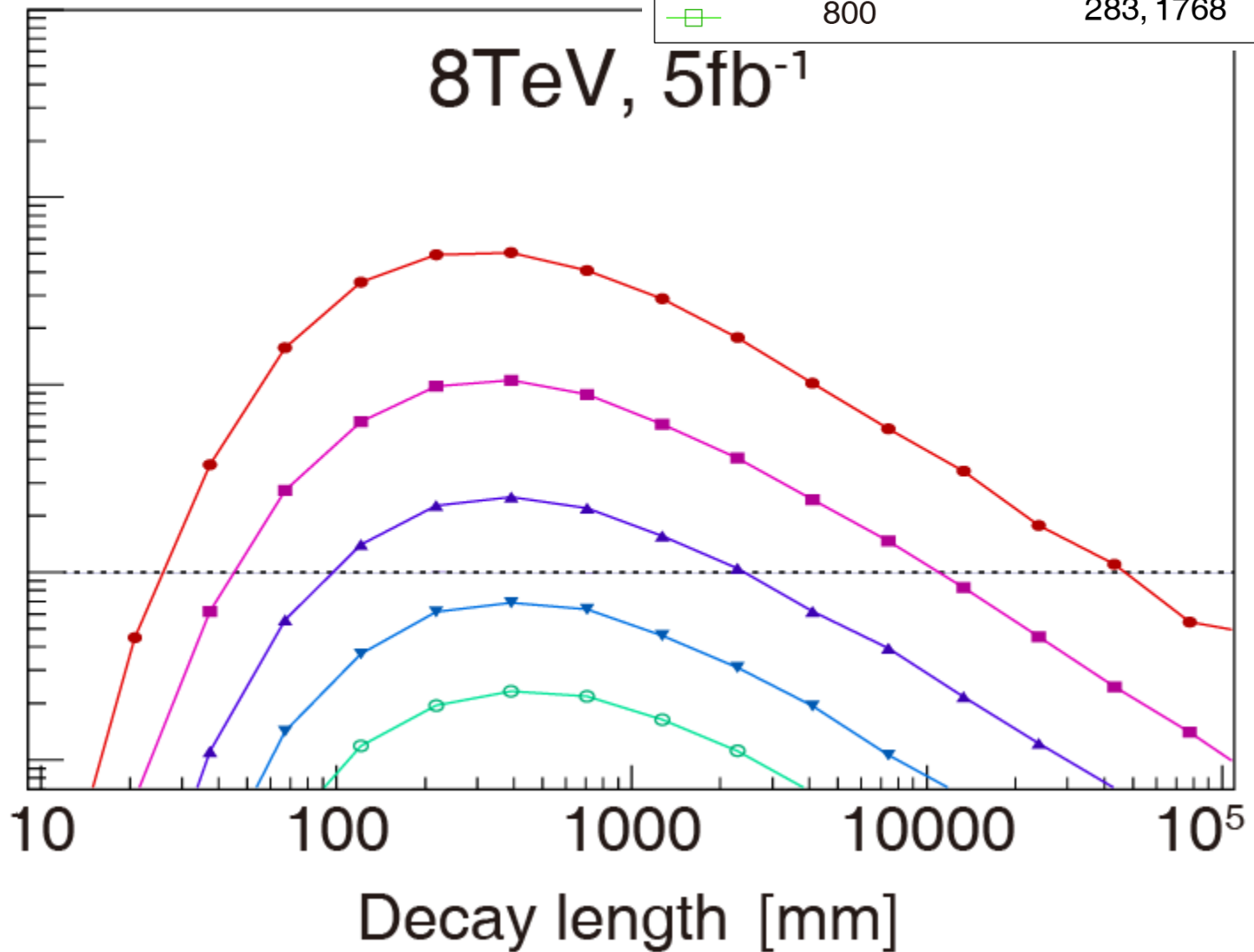
●	$M_{1/2} = 300\text{GeV}$	$\tilde{\tau}, \tilde{g} = 103, 715\text{ GeV}$
■	400	140, 932
▲	500	176, 1145
▼	600	212, 1355
○	700	248, 1562
□	800	283, 1768



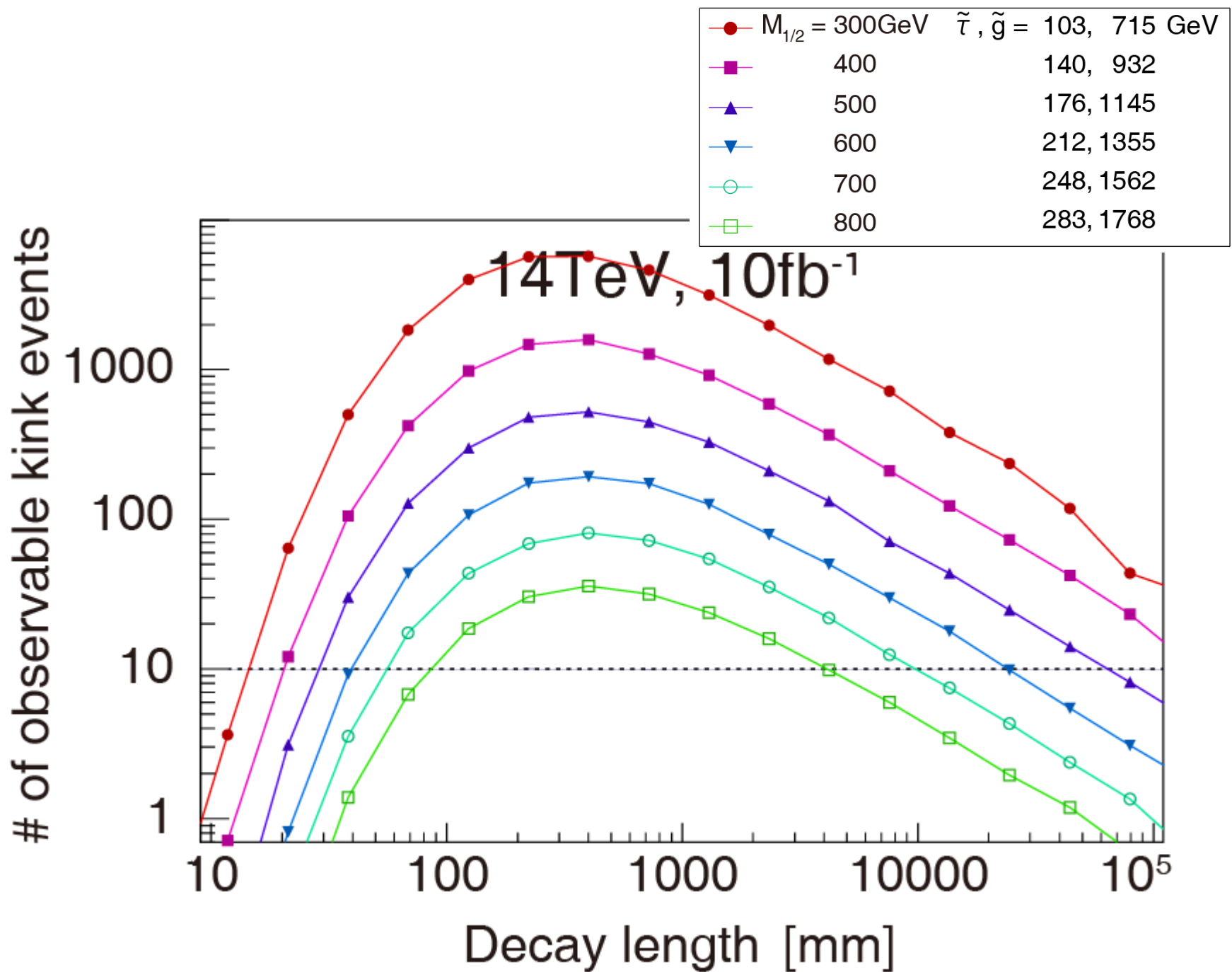
# of observable kink events

8TeV, 5fb<sup>-1</sup>

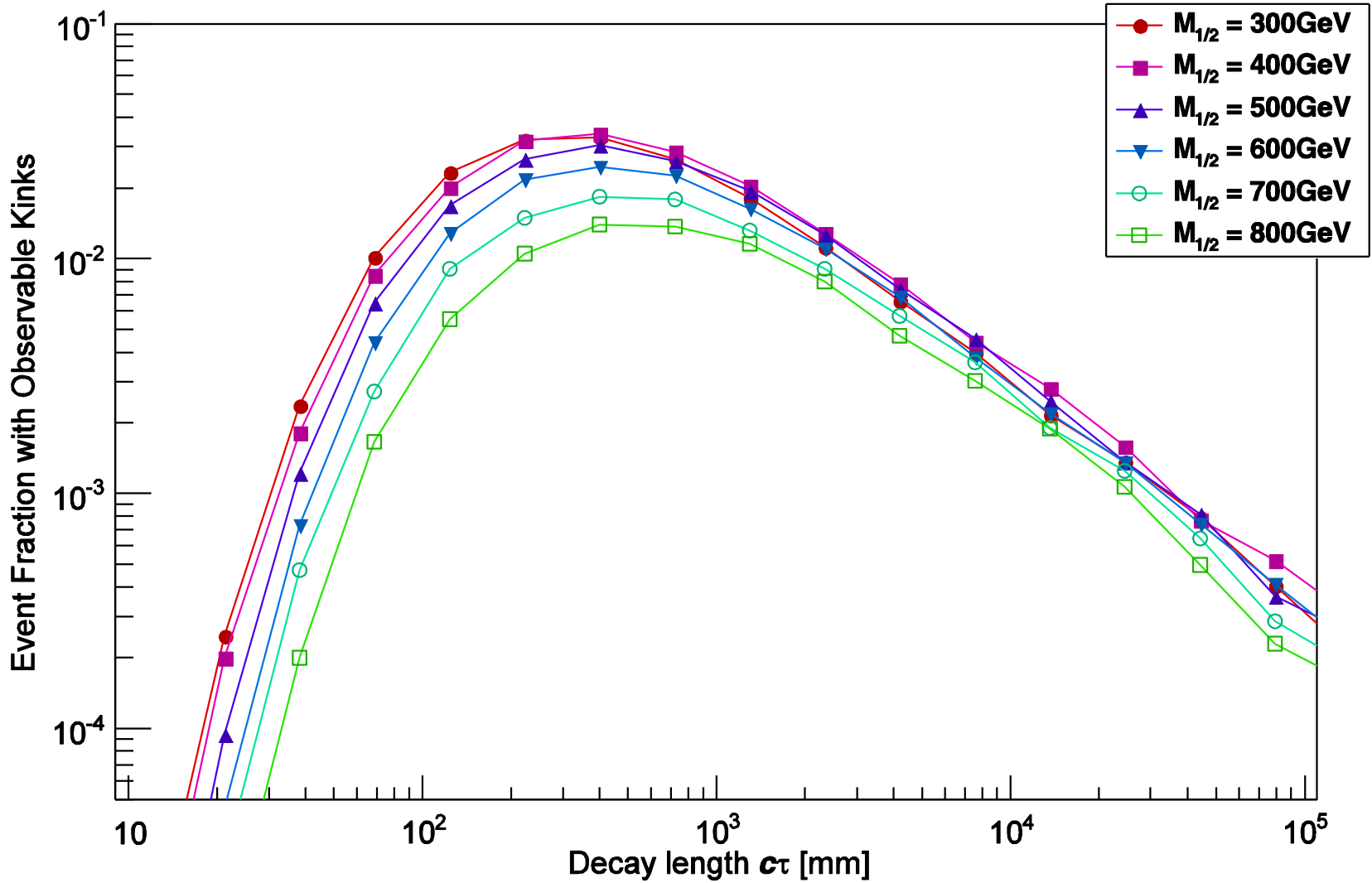
●	M <sub>1/2</sub> = 300GeV	$\tilde{\tau}, \tilde{g} = 103, 715$ GeV
■	400	140, 932
▲	500	176, 1145
▼	600	212, 1355
○	700	248, 1562
□	800	283, 1768



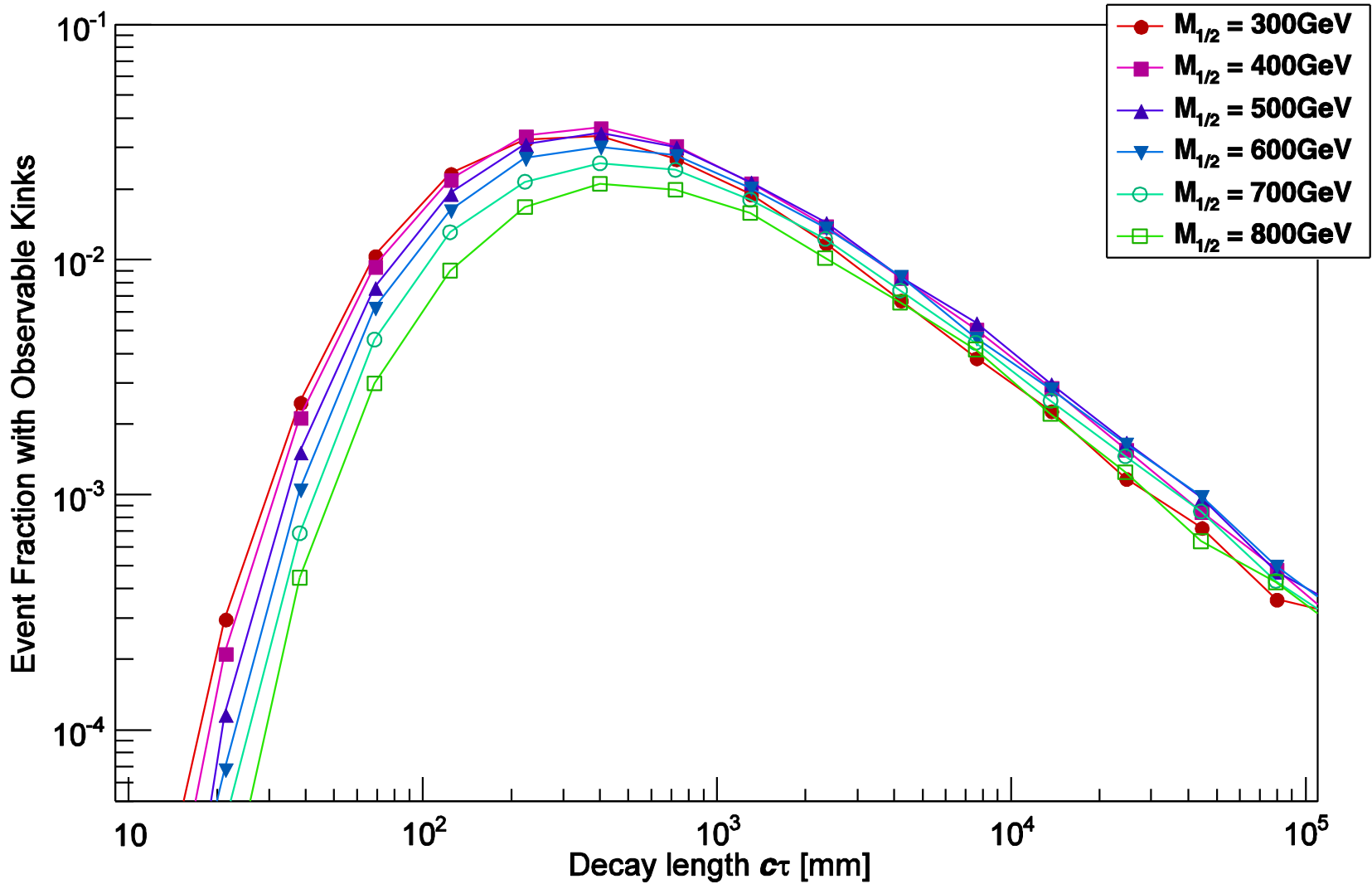




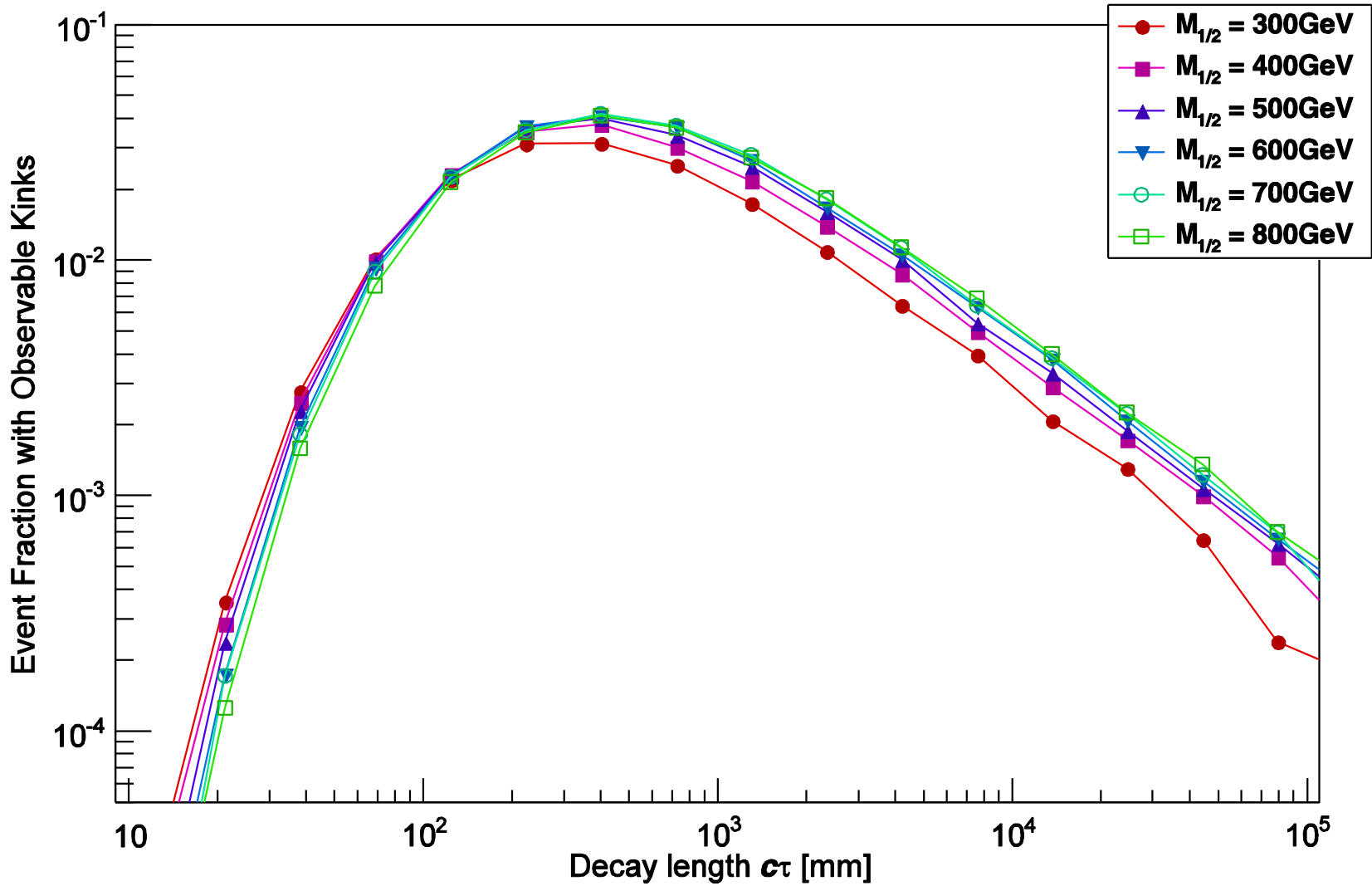
7TeV, 5fb<sup>-1</sup>



8TeV, 5fb<sup>-1</sup>



# 14TeV, 10fb<sup>-1</sup>



# Monte Carlo System

